# Extra Dimensions

For explanation of terms used and discussion of significant model dependence of following limits, see the "Extra Dimensions" review. Footnotes describe originally quoted limit.  $\delta$  indicates the number of extra dimensions.

Limits not encoded here are summarized in the "Extra Dimensions" review, where the latest unpublished results are also described.

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## Limits on R from Deviations in Gravitational Force Law

This section includes limits on the size of extra dimensions from deviations in the Newtonian  $(1/r^2)$  gravitational force law at short distances. Deviations are parametrized by a gravitational potential of the form  $V = -(G \ m \ m'/r) \ [1 + \alpha \ \exp(-r/R)]$ . For  $\delta$ toroidal extra dimensions of equal size,  $\alpha = 8\delta/3$ . Quoted bounds are for  $\delta = 2$  unless otherwise noted.

$VALUE~(\mu m)$	CL%	DOCUMENT ID		COMMENT
< 30	95	$^{ m 1}$ KAPNER	07	Torsion pendulum
• • • We do not use the	e following	g data for averages	s, fits,	limits, etc. • • •
		<sup>2</sup> XU	13	Nuclei properties
		<sup>3</sup> BEZERRA	11	Torsion oscillator
		<sup>4</sup> SUSHKOV	11	Torsion pendulum
		<sup>5</sup> BEZERRA	10	Microcantilever
		<sup>6</sup> MASUDA	09	Torsion pendulum
		<sup>7</sup> GERACI	80	Microcantilever
		<sup>8</sup> TRENKEL	80	Newton's constant
		<sup>9</sup> DECCA	07A	Torsion oscillator
< 47	95	<sup>10</sup> TU	07	Torsion pendulum
		<sup>11</sup> SMULLIN	05	Microcantilever
<130	95	<sup>12</sup> HOYLE	04	Torsion pendulum
		<sup>13</sup> CHIAVERINI	03	Microcantilever
≤ 200	95	<sup>14</sup> LONG	03	Microcantilever
<190	95	<sup>15</sup> HOYLE	01	Torsion pendulum
		<sup>16</sup> HOSKINS	85	Torsion pendulum

- $^1$  KAPNER 07 search for new forces, probing a range of  $\alpha \simeq 10^{-3} \text{--} 10^5$  and length scales  $R \simeq 10 \text{--} 1000~\mu\text{m}$ . For  $\delta = 1$  the bound on R is 44  $\mu\text{m}$ . For  $\delta = 2$ , the bound is expressed in terms of  $M_*$ , here translated to a bound on the radius. See their Fig. 6 for details on the bound.
- $^2$  XU 13 obtain constraints on non-Newtonian forces with strengths  $|\alpha| \simeq 10^{34} 10^{36}$  and length scales  $R \simeq 1$ –10 fm. See their Fig. 4 for more details. These constraints do not place limits on the size of extra flat dimensions.
- $^3$  BEZERRA 11 obtain constraints on non-Newtonian forces with strengths  $10^{11}\lesssim |\alpha|\lesssim 10^{18}$  and length scales R=30–1260 nm. See their Fig. 2 for more details. These constraints do not place limits on the size of extra flat dimensions.
- $^4$  SUSHKOV 11 obtain improved limits on non-Newtonian forces with strengths  $10^7 \lesssim |\alpha| \lesssim 10^{11}$  and length scales 0.4  $\mu \rm m < R < 4~\mu m$  (95% CL). See their Fig. 2. These bounds do not place limits on the size of extra flat dimensions. However, a model dependent bound of  $M_* > 70$  TeV is obtained assuming gauge bosons that couple to baryon number also propagate in  $(4+\delta)$  dimensions.
- $^5$  BEZERRA 10 obtain improved constraints on non-Newtonian forces with strengths  $10^{19}\lesssim |\alpha|\lesssim 10^{29}$  and length scales R=1.6–14 nm (95% CL). See their Fig. 1. This bound does not place limits on the size of extra flat dimensions.
- $^6$  MASUDA 09 obtain improved constraints on non-Newtonian forces with strengths  $10^9 \lesssim |\alpha| \lesssim 10^{11}$  and length scales R=1.0–2.9  $\mu m$  (95% CL). See their Fig. 3. This bound does not place limits on the size of extra flat dimensions.
- $^7$  GERACI 08 obtain improved constraints on non-Newtonian forces with strengths  $|\alpha|>14,000$  and length scales R=5–15  $\mu\mathrm{m}.$  See their Fig. 9. This bound does not place limits on the size of extra flat dimensions.
- 8 TRENKEL 08 uses two independent measurements of Newton's constant G to constrain new forces with strength  $|\alpha| \simeq 10^{-4}$  and length scales R=0.02-1 m. See their Fig. 1. This bound does not place limits on the size of extra flat dimensions.
- <sup>9</sup> DECCA 07A search for new forces and obtain bounds in the region with strengths  $|\alpha| \simeq 10^{13}$ – $10^{18}$  and length scales R=20–86 nm. See their Fig. 6. This bound does not place limits on the size of extra flat dimensions.
- $^{10}$  TU 07 search for new forces probing a range of  $|\alpha| \simeq 10^{-1}$ – $10^5$  and length scales  $R \simeq 20$ – $1000~\mu m$ . For  $\delta = 1$  the bound on R is 53  $\mu m$ . See their Fig. 3 for details on the bound
- $^{11}$  SMULLIN 05 search for new forces, and obtain bounds in the region with strengths  $\alpha \simeq 10^3 \text{--}10^8$  and length scales  $R=6\text{--}20~\mu\text{m}$ . See their Figs. 1 and 16 for details on the bound. This work does not place limits on the size of extra flat dimensions.
- <sup>12</sup> HOYLE 04 search for new forces, probing  $\alpha$  down to  $10^{-2}$  and distances down to  $10\mu$ m. Quoted bound on R is for  $\delta=2$ . For  $\delta=1$ , bound goes to 160  $\mu$ m. See their Fig. 34 for details on the bound.
- <sup>13</sup> CHIAVERINI 03 search for new forces, probing  $\alpha$  above  $10^4$  and  $\lambda$  down to  $3\mu$ m, finding no signal. See their Fig. 4 for details on the bound. This bound does not place limits on the size of extra flat dimensions.
- $^{14}$  LONG 03 search for new forces, probing  $\alpha$  down to 3, and distances down to about  $10\mu\mathrm{m}$ . See their Fig. 4 for details on the bound.
- <sup>15</sup> HOYLE 01 search for new forces, probing  $\alpha$  down to  $10^{-2}$  and distances down to  $20\mu m$ . See their Fig. 4 for details on the bound. The quoted bound is for  $\alpha \geq 3$ .
- 16 HOSKINS 85 search for new forces, probing distances down to 4 mm. See their Fig. 13 for details on the bound. This bound does not place limits on the size of extra flat dimensions.

## Limits on R from On-Shell Production of Gravitons: $\delta = 2$

This section includes limits on on-shell production of gravitons in collider and astrophysical processes. Bounds quoted are on R, the assumed common radius of the flat extra dimensions, for  $\delta=2$  extra dimensions. Studies often quote bounds in terms of derived parameter; experiments are actually sensitive to the masses of the KK gravitons:  $m_{\vec{n}}=|\vec{n}|/R$ . See the Review on "Extra Dimensions" for details. Bounds are given in  $\mu m$  for  $\delta=2$ .

VALU	$JE\left( \mum ight)$	CL%		DOCUMENT ID		TECN	COMMENT
<	10.9	95	1	AABOUD	<b>16</b> D	ATLS	$pp \rightarrow jG$
<	0.00016	95	2	HANNESTAD	03		Neutron star heating
• •	• We do not use the	following	g d	ata for averages,	fits,	limits, e	tc. • • •
<	90	95		AABOUD		ATLS	$pp \rightarrow \gamma G$
			4	KHACHATRY	.16N	CMS	$pp \rightarrow \gamma G$
<	17.2	95		AAD	<b>15</b> BH	ATLS	$pp \rightarrow jG$
			_	AAD		ATLS	$pp \rightarrow \gamma G$
<	15	95		KHACHATRY	.15AL	CMS	$pp \rightarrow jG$
<	25	95		AAD	<b>13</b> AD	ATLS	$pp \rightarrow jG$
< 1	127	95		AAD	<b>13</b> C	ATLS	$pp \rightarrow \gamma G$
<	34.4	95		AAD	<b>13</b> D	ATLS	$pp \rightarrow jj$
<	0.0087	95		AJELLO	12	FLAT	Neutron star $\gamma$ sources
<	23	95		CHATRCHYAN	<b>12</b> AP	CMS	$pp \rightarrow jG$
<	92	95		AAD		ATLS	$pp \rightarrow jG$
<	72	95	14	CHATRCHYAN	<b>11</b> U	CMS	$pp \rightarrow jG$
< 2	245	95	15	AALTONEN		CDF	$p\overline{p} \rightarrow \gamma G, jG$
< 6	515	95	16	ABAZOV	<b>08</b> S	D0	$p\overline{p} \rightarrow \gamma G$
<	0.916	95	17	DAS	80		Supernova cooling
< 3	350	95	18	ABULENCIA,A	06	CDF	$p\overline{p} \rightarrow jG$
< 2	270	95		ABDALLAH	<b>05</b> B	DLPH	$e^+e^-  ightarrow \gamma G$
< 2	210	95	20	ACHARD	04E	L3	$e^+e^- \rightarrow \gamma G$
< 4	180	95		ACOSTA	04C	CDF	$\overline{p}p \rightarrow jG$
<	0.00038	95		CASSE	04		Neutron star $\gamma$ sources
< 6	510	95	23	ABAZOV	03	D0	$\overline{p}p \rightarrow jG$
<	0.96	95	24	HANNESTAD	03		Supernova cooling
<	0.096	95	25		03		Diffuse $\gamma$ background
<	0.051	95	26		03		Neutron star $\gamma$ sources
< 3	300	95	27	HEISTER	<b>03</b> C	ALEP	$e^+e^-  ightarrow \gamma G$
				FAIRBAIRN	01		Cosmology
<	0.66	95	29	HANHART	01		Supernova cooling
•	-	-	30	CASSISI	00		Red giants
<13	300	95		ACCIARRI		L3	$e^+e^- \rightarrow ZG$
							· - ·

<sup>&</sup>lt;sup>1</sup> AABOUD 16D search for  $pp \to jG$ , using 3.2 fb<sup>-1</sup> of data at  $\sqrt{s} = 13$  TeV to place lower limits on  $M_D$  for two to six extra dimensions (see their Table X), from which this bound on R is derived.

<sup>&</sup>lt;sup>2</sup> HANNESTAD 03 obtain a limit on R from the heating of old neutron stars by the surrounding cloud of trapped KK gravitons. Limits for all  $\delta \leq 7$  are given in their Tables V and VI. These limits supersede those in HANNESTAD 02.

<sup>&</sup>lt;sup>3</sup> AABOUD 16F search for  $pp \to \gamma G$ , using 3.2 fb<sup>-1</sup> of data at  $\sqrt{s} = 13$  TeV to place limits on  $M_D$  for two to six extra dimensions (see their Figure 9), from which this bound on R is derived.

<sup>&</sup>lt;sup>4</sup> KHACHATRYAN 16N search for  $pp \rightarrow \gamma G$ , using 19.6 fb<sup>-1</sup> of data at  $\sqrt{s} = 8$  TeV to place limits on  $M_D$  for three to six extra dimensions (see their Table 5).

- <sup>5</sup> AAD 15BH search for  $pp \to jG$ , using 20.3 fb<sup>-1</sup> of data at  $\sqrt{s} = 8$  TeV to place bounds on  $M_D$  for two to six extra dimensions, from which this bound on R is derived. See their Figure 9 for bounds on all  $\delta \leq 6$ .
- <sup>6</sup> AAD 15CS search for  $pp \to \gamma G$ , using 20.3 fb<sup>-1</sup> of data at  $\sqrt{s} = 8$  TeV to place lower limits on  $M_D$  for two to six extra dimensions (see their Fig. 18).
- <sup>7</sup> KHACHATRYAN 15AL search for  $pp \to jG$ , using 19.7 fb<sup>-1</sup> of data at  $\sqrt{s} = 8$  TeV to place bounds on  $M_D$  for two to six extra dimensions (see their Table 7), from which this bound on R is derived.
- <sup>8</sup> AAD 13AD search for  $pp \to jG$ , using 4.7 fb<sup>-1</sup> of data at  $\sqrt{s}=7$  TeV to place bounds on  $M_D$  for two to six extra dimensions, from which this bound on R is derived. See their Table 8 for bounds on all  $\delta \leq 6$ .
- <sup>9</sup> AAD 13C search for  $pp \to \gamma G$ , using 4.6 fb<sup>-1</sup> of data at  $\sqrt{s} = 7$  TeV to place bounds on  $M_D$  for two to six extra dimensions, from which this bound on R is derived.
- $^{10}$  AAD 13D search for the dijet decay of quantum black holes in 4.8 fb $^{-1}$  of data produced in pp collisions at  $\sqrt{s}=7$  TeV to place bounds on  $M_D$  for two to seven extra dimensions, from which these bounds on R are derived. Limits on  $M_D$  for all  $\delta \leq 7$  are given in their Table 3.
- $^{11}$  AJELLO 12 obtain a limit on R from the gamma-ray emission of point  $\gamma$  sources that arise from the photon decay of KK gravitons which are gravitationally bound around neutron stars. Limits for all  $\delta \leq 7$  are given in their Table 7.
- <sup>12</sup> CHATRCHYAN 12AP search for  $pp \to jG$ , using 5.0 fb<sup>-1</sup> of data at  $\sqrt{s}=7$  TeV to place bounds on  $M_D$  for two to six extra dimensions, from which this bound on R is derived. See their Table 7 for bounds on all  $\delta \leq 6$ .
- <sup>13</sup> AAD 11S search for  $pp \to jG$ , using 33 pb<sup>-1</sup> of data at  $\sqrt{s}=7$  TeV, to place bounds on  $M_D$  for two to four extra dimensions, from which these bounds on R are derived. See their Table 3 for bounds on all  $\delta \leq 4$ .
- <sup>14</sup> CHATRCHYAN 110 search for  $pp \to jG$ , using 36 pb<sup>-1</sup> of data at  $\sqrt{s}=7$  TeV, to place bounds on  $M_D$  for two to six extra dimensions, from which these bounds on R are derived. See their Table 3 for bounds on all  $\delta \leq 6$ .
- $^{15}$  AALTONEN 08AC search for  $p\overline{p}\to \gamma\,G$  and  $p\overline{p}\to j\,G$  at  $\sqrt{s}=1.96$  TeV with 2.0 fb $^{-1}$  and 1.1 fb $^{-1}$  respectively, in order to place bounds on the fundamental scale and size of the extra dimensions. See their Table III for limits on all  $\delta\leq 6$ .
- <sup>16</sup> ABAZOV 08S search for  $p\overline{p} \to \gamma G$ , using 1 fb<sup>-1</sup> of data at  $\sqrt{s} = 1.96$  TeV to place bounds on  $M_D$  for two to eight extra dimensions, from which these bounds on R are derived. See their paper for intermediate values of  $\delta$ .
- $^{17}$  DAS 08 obtain a limit on R from Kaluza-Klein graviton cooling of SN1987A due to plasmon-plasmon annihilation.
- <sup>18</sup> ABULENCIA,A 06 search for  $p\overline{p}\to jG$  using 368 pb<sup>-1</sup> of data at  $\sqrt{s}=1.96$  TeV. See their Table II for bounds for all  $\delta\le 6$ .
- $^{19}$  ABDALLAH 05B search for e $^+$ e $^- \to \gamma \, G$  at  $\sqrt{s}=$  180–209 GeV to place bounds on the size of extra dimensions and the fundamental scale. Limits for all  $\delta \leq 6$  are given in their Table 6. These limits supersede those in ABREU 00Z.
- $^{20}$  ACHARD 04E search for  $e^+\,e^- \to \gamma\,G$  at  $\sqrt{s}=189$ –209 GeV to place bounds on the size of extra dimensions and the fundamental scale. See their Table 8 for limits with  $\delta\,\leq\,8$ . These limits supersede those in ACCIARRI 99R.
- <sup>21</sup> ACOSTA 04C search for  $\overline{p}p \rightarrow jG$  at  $\sqrt{s}=1.8$  TeV to place bounds on the size of extra dimensions and the fundamental scale. See their paper for bounds on  $\delta=4$ , 6.
- $^{22}$  CASSE 04 obtain a limit on R from the gamma-ray emission of point  $\gamma$  sources that arises from the photon decay of gravitons around newly born neutron stars, applying the technique of HANNESTAD 03 to neutron stars in the galactic bulge. Limits for all  $\delta \leq 7$  are given in their Table I.
- <sup>23</sup> ABAZOV 03 search for  $p\overline{p} \to jG$  at  $\sqrt{s}{=}1.8$  TeV to place bounds on  $M_D$  for 2 to 7 extra dimensions, from which these bounds on R are derived. See their paper for bounds on intermediate values of  $\delta$ . We quote results without the approximate NLO scaling introduced in the paper.

<sup>26</sup> HANNESTAD 03 obtain a limit on R from gravitons emitted in two recent supernovae and which subsequently decay, creating point  $\gamma$  sources. Limits for all  $\delta \leq 7$  are given in their Tables V and VI. These limits are corrected in the published erratum.

<sup>27</sup> HEISTER 03C use the process  $e^+e^- \rightarrow \gamma G$  at  $\sqrt{s}=189$ –209 GeV to place bounds on the size of extra dimensions and the scale of gravity. See their Table 4 for limits with  $\delta \leq 6$  for derived limits on  $M_D$ .

<sup>28</sup> FAIRBAIRN 01 obtains bounds on R from over production of KK gravitons in the early universe. Bounds are quoted in paper in terms of fundamental scale of gravity. Bounds depend strongly on temperature of QCD phase transition and range from R< 0.13  $\mu$ m to 0.001  $\mu$ m for  $\delta$ =2; bounds for  $\delta$ =3,4 can be derived from Table 1 in the paper.

<sup>29</sup> HANHART 01 obtain bounds on *R* from limits on graviton cooling of supernova SN 1987a using numerical simulations of proto-neutron star neutrino emission.

 $^{30}$  CASSISI 00 obtain rough bounds on  $M_D$  (and thus R) from red giant cooling for  $\delta{=}2,3.$  See their paper for details.

<sup>31</sup> ACCIARRI 99S search for  $e^+e^- \rightarrow ZG$  at  $\sqrt{s}$ =189 GeV. Limits on the gravity scale are found in their Table 2, for  $\delta \leq 4$ .

# Mass Limits on M<sub>TT</sub>

This section includes limits on the cut-off mass scale,  $M_{TT}$ , of dimension-8 operators from KK graviton exchange in models of large extra dimensions. Ambiguities in the UV-divergent summation are absorbed into the parameter  $\lambda$ , which is taken to be  $\lambda=\pm 1$  in the following analyses. Bounds for  $\lambda=-1$  are shown in parenthesis after the bound for  $\lambda=+1$ , if appropriate. Different papers use slightly different definitions of the mass scale. The definition used here is related to another popular convention by  $M_{TT}^4=(2/\pi)~\Lambda_T^4$ , as discussed in the above Review on "Extra Dimensions."

VALUE (TeV)	· <u></u>	CL%	DOCUMENT ID		TECN	COMMENT
> 6.3		95		. <b>15</b> J		$pp \rightarrow \text{dijet, ang. distrib.}$
>20.6	(> 15.7)	95	<sup>2</sup> GIUDICE	03	RVUE	Dim-6 operators
• • • We d	o not use t	he follow	ing data for averag	es, fit	s, limits	, etc. • • •
> 3.7		95	<sup>3</sup> KHACHATRY	.15AE	CMS	$pp  ightarrow e^+e^-$ , $\mu^+\mu^-$
> 3.8		95		<b>14</b> BE	ATLS	$pp \rightarrow e^+e^-, \mu^+\mu^-$
> 2.94	(>2.52)	95		<b>13</b> AS	ATLS	$pp \rightarrow \gamma \gamma$
> 3.2		95	_	13E	ATLS	$pp \rightarrow e^+e^-, \mu^+\mu^-, \gamma\gamma$
> 2.66	(>2.27)	95		12Y	ATLS	$pp \rightarrow \gamma \gamma$
				12	RVUE	Electroweak
> 2.86		95	<sup>9</sup> CHATRCHYAN			$pp \rightarrow e^+e^-, \mu^+\mu^-$
> 2.84	(>2.41)	95	<sup>10</sup> CHATRCHYAN	<b>12</b> R	CMS	$pp \rightarrow \gamma\gamma$
> 0.90	(>0.92)	95			H1	$e^{\pm} p \rightarrow e^{\pm} X$
> 1.74	(>1.71)	95	<sup>12</sup> CHATRCHYAN	11A	CMS	$pp \rightarrow \gamma \gamma$
> 1.48		95		09AE	D0	$p\overline{p}  o  ext{dijet}$ , ang. distrib.
> 1.45		95		<b>09</b> D	D0	$p\overline{p}  ightarrow e^+e^-$ , $\gamma\gamma$
> 1.1	(> 1.0)	95		07A	ALEP	$e^+e^-  ightarrow e^+e^-$
> 0.898	(> 0.998)	95	<sup>16</sup> ABDALLAH	<b>06</b> C	DLPH	$e^+e^- \rightarrow \ell^+\ell^-$
> 0.853	(> 0.939)	95	<sup>17</sup> GERDES	06		$p\overline{p} \rightarrow e^+e^-, \gamma\gamma$
> 0.96	(>0.93)	95	<sup>18</sup> ABAZOV	05∨	D0	$p\overline{p} \rightarrow \mu^{+}\mu^{-}$
> 0.78	(> 0.79)	95	<sup>19</sup> CHEKANOV	<b>04</b> B	ZEUS	$e^{\pm}p \rightarrow e^{\pm}X$

<sup>&</sup>lt;sup>24</sup> HANNESTAD 03 obtain a limit on R from graviton cooling of supernova SN1987a. Limits for all  $\delta \leq 7$  are given in their Tables V and VI.

<sup>&</sup>lt;sup>25</sup> HANNESTAD 03 obtain a limit on R from gravitons emitted in supernovae and which subsequently decay, contaminating the diffuse cosmic  $\gamma$  background. Limits for all  $\delta \leq 7$  are given in their Tables V and VI. These limits supersede those in HANNESTAD 02.

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<sup>20</sup> ABBIENDI
                                                               03D OPAL e^+e^- \rightarrow \gamma \gamma
> 0.805
               (>0.956)95
                                       <sup>21</sup> ACHARD
               (> 0.7)
                                                               03D
                                                                    L3
> 0.7
                             95
                                       <sup>22</sup> ADLOFF
> 0.82
               (>0.78)
                             95
                                                               03
                                                                     H1
                                       <sup>23</sup> GIUDICE
                                                               03
                                                                     RVUE
> 1.28
               (>1.25)
                             95
                                       <sup>24</sup> HEISTER
                                                               03C ALEP
> 0.80
               (>0.85)
                             95
                                       <sup>25</sup> ACHARD
                                                               02D L3
> 0.84
               (> 0.99)
                                       <sup>26</sup> ABBOTT
                                                               01
                                                                     D0
                                                                                p\overline{p} \rightarrow e^+e^-, \gamma\gamma
> 1.2
               (>1.1)
                                                               OOR OPAL e^+e^- 
ightarrow \mu^+\mu^-
                                       <sup>27</sup> ABBIENDI
> 0.60
               (> 0.63)
                                                               OOR OPAL e^+e^- \rightarrow \tau^+\tau^-
                                       <sup>27</sup> ABBIENDI
> 0.63
               (> 0.50)
                             95
                                       <sup>27</sup> ABBIENDI
                                                               00R OPAL e^+e^- \to \mu^+\mu^-, \tau^+\tau^-
> 0.68
               (> 0.61)
                                       <sup>28</sup> ABREU
                                                               00A DLPH e^+e^- \rightarrow \gamma \gamma
                                       <sup>29</sup> ABREU
                                                               00S DLPH e^+e^- \to \mu^+\mu^-, \tau^+\tau^-
> 0.680
               (>0.542) 95
                                       <sup>30</sup> CHANG
                                                               00B RVUE Electroweak
> 15-28
                             99.7
                                       <sup>31</sup> CHEUNG
                                                                     RVUE e^+e^- \rightarrow \gamma \gamma
> 0.98
                             95
                                                               00
                                       <sup>32</sup> GRAESSER
> 0.29-0.38
                             95
                                                                     RVUE (g-2)_{II}
                                       <sup>33</sup> HAN
                                                                     RVUE Electroweak
> 0.50-1.1
                             95
                                                               00
                                       <sup>34</sup> MATHEWS
                             95
                                                                     RVUE \overline{p}p \rightarrow jj
> 2.0
               (> 2.0)
                                       35 MELE
                                                                     RVUE e^+e^- \rightarrow VV
                                                               00
> 1.0
               (>1.1)
                                       <sup>36</sup> ABBIENDI
                                                               99P OPAL
                                       <sup>37</sup> ACCIARRI
                                                               99M L3
                                       <sup>38</sup> ACCIARRI
                                                               99s
                                                                                e^+e^- \rightarrow e^+e^-
                                       <sup>39</sup> BOURILKOV
> 1.412
               (>1.077)95
                                                               99
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- <sup>1</sup> KHACHATRYAN 15J use dijet angular distributions in 19.7 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=8$  TeV to place a lower bound on  $\Lambda_T$ , here converted to  $M_{TT}$ .
- $^2$  GIUDICE 03 place bounds on  $\Lambda_6$ , the coefficient of the gravitationally-induced dimension- 6 operator  $(2\pi\lambda/\Lambda_6^2)(\sum\overline{f}\gamma_\mu\gamma^5f)(\sum\overline{f}\gamma^\mu\gamma^5f)$ , using data from a variety of experiments. Results are quoted for  $\lambda=\pm 1$  and are independent of  $\delta$ .
- <sup>3</sup> KHACHATRYAN 15AE use 20.6 (19.7) fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=8$  TeV in the dimuon (dielectron) channel to place a lower limit on  $\Lambda_T$ , here converted to  $M_{TT}$ .
- $^4$  AAD 14BE use 20 fb $^{-1}$  of data from  $p\,p$  collisions at  $\sqrt{s}=8$  TeV in the dilepton channel to place lower limits on  $M_{TT}$  (equivalent to their  $M_{\mbox{\scriptsize $S$}}).$
- <sup>5</sup> AAD 13AS use 4.9 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=7$  TeV to place lower limits on  $M_{TT}$  (equivalent to their  $M_{S}$ ).
- <sup>6</sup> AAD 13E use 4.9 and 5.0 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=7$  TeV in the dielectron and dimuon channels, respectively, to place lower limits on  $M_{TT}$  (equivalent to their  $M_S$ ). The dielectron and dimuon channels are combined with previous results in the diphoton channel to set the best limit. Bounds on individual channels and different priors can be found in their Table VIII.
- <sup>7</sup> AAD 12Y use 2.12 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=7$  TeV to place lower limits on  $M_{TT}$  (equivalent to their  $M_{S}$ ).
- <sup>8</sup> BAAK 12 use electroweak precision observables to place bounds on the ratio  $\Lambda_T/M_D$  as a function of  $M_D$ . See their Fig. 22 for constraints with a Higgs mass of 120 GeV.
- <sup>9</sup> CHATRCHYAN 12J use approximately 2 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=7$  TeV in the dielectron and dimuon channels to place lower limits on  $\Lambda_T$ , here converted to  $M_{TT}$ .
- $^{10}$  CHATRCHYAN 12R use 2.2 fb $^{-1}$  of data from pp collisions at  $\sqrt{s}=7$  TeV to place lower limits on  $M_{TT}$  (equivalent to their  $M_S$ ).
- $^{11}$  AARON 11C search for deviations in the differential cross section of  $e^{\pm}p \rightarrow e^{\pm}X$  in 446 pb $^{-1}$  of data taken at  $\sqrt{s}=$  301 and 319 GeV to place a bound on  $M_{TT}$ .
- $^{12}$  CHATRCHYAN  $^{11}$ A use 36 pb $^{-1}$  of data from pp collisions at  $\sqrt{s}=7$  TeV to place lower limits on  $\Lambda_T$  , here converted to  $M_{TT}$  .

- $^{13}$  ABAZOV 09AE use dijet angular distributions in 0.7 fb $^{-1}$  of data from  $p\overline{p}$  collisions at  $\sqrt{s}=1.96$  TeV to place lower bounds on  $\Lambda_T$  (equivalent to their  $M_S$ ), here converted to  $M_{TT}$ .
- <sup>14</sup> ABAZOV 09D use 1.05 fb<sup>-1</sup> of data from  $p\bar{p}$  collisions at  $\sqrt{s}=1.96$  TeV to place lower bounds on  $\Lambda_T$  (equivalent to their  $M_s$ ), here converted to  $M_{TT}$ .
- $^{15}$  SCHAEL 07A use  $e^+\,e^-$  collisions at  $\sqrt{s}=$  189–209 GeV to place lower limits on  $\Lambda_T$  , here converted to limits on  $M_{TT}$  .
- $^{16}$  ABDALLAH 06C use  $e^+\,e^-$  collisions at  $\sqrt{s}\sim 130$ –207 GeV to place lower limits on  $M_{TT}$ , which is equivalent to their definition of  $M_{s}$ . Bound shown includes all possible final state leptons,  $\ell=e,\,\mu,\,\tau.$  Bounds on individual leptonic final states can be found in their Table 31.
- $^{17}$  GERDES 06 use 100 to 110 pb $^{-1}$  of data from  $p\overline{p}$  collisions at  $\sqrt{s}=1.8$  TeV, as recorded by the CDF Collaboration during Run I of the Tevatron. Bound shown includes a K-factor of 1.3. Bounds on individual  $e^+e^-$  and  $\gamma\gamma$  final states are found in their Table I.
- <sup>18</sup> ABAZOV 05V use 246 pb<sup>-1</sup> of data from  $p\overline{p}$  collisions at  $\sqrt{s} = 1.96$  TeV to search for deviations in the differential cross section to  $\mu^+\mu^-$  from graviton exchange.
- <sup>19</sup> CHEKANOV 04B search for deviations in the differential cross section of  $e^{\pm}p \rightarrow e^{\pm}X$  with 130  $pb^{-1}$  of combined data and  $Q^2$  values up to 40,000 GeV<sup>2</sup> to place a bound on  $M_{TT}$ .
- <sup>20</sup> ABBIENDI 03D use  $e^+e^-$  collisions at  $\sqrt{s}$ =181–209 GeV to place bounds on the ultraviolet scale  $M_{TT}$ , which is equivalent to their definition of  $M_s$ .
- <sup>21</sup> ACHARD 03D look for deviations in the cross section for  $e^+e^- \rightarrow ZZ$  from  $\sqrt{s}=200$ –209 GeV to place a bound on  $M_{TT}$ .
- <sup>22</sup> ADLOFF 03 search for deviations in the differential cross section of  $e^{\pm}p \rightarrow e^{\pm}X$  at  $\sqrt{s}$ =301 and 319 GeV to place bounds on  $M_{TT}$ .
- $^{23}$  GIUDICE 03 review existing experimental bounds on  $M_{TT}$  and derive a combined limit.
- <sup>24</sup> HEISTER 03C use  $e^+e^-$  collisions at  $\sqrt{s}=$  189–209 GeV to place bounds on the scale of dim-8 gravitational interactions. Their  $M_s^\pm$  is equivalent to our  $M_{TT}$  with  $\lambda=\pm 1$ .
- $^{25}$  ACHARD 02 search for s-channel graviton exchange effects in  $e^+e^-\to\gamma\gamma$  at  $E_{\rm cm}=192$ –209 GeV.
- <sup>26</sup> ABBOTT 01 search for variations in differential cross sections to  $e^+e^-$  and  $\gamma\gamma$  final states at the Tevatron.
- <sup>27</sup> ABBIENDI 00R uses  $e^+e^-$  collisions at  $\sqrt{s}$ = 189 GeV.
- <sup>28</sup> ABREU 00A search for s-channel graviton exchange effects in  $e^+e^- \rightarrow \gamma\gamma$  at  $E_{\rm cm}=$  189–202 GeV.
- ABREU 00S uses  $e^+e^-$  collisions at  $\sqrt{s}$ =183 and 189 GeV. Bounds on  $\mu$  and  $\tau$  individual final states given in paper.
- $^{30}$  CHANG 00B derive  $3\sigma$  limit on  $M_{TT}$  of (28,19,15) TeV for  $\delta$ =(2,4,6) respectively assuming the presence of a torsional coupling in the gravitational action. Highly model dependent.
- <sup>31</sup> CHEUNG 00 obtains limits from anomalous diphoton production at OPAL due to graviton exchange. Original limit for  $\delta$ =4. However, unknown UV theory renders  $\delta$  dependence unreliable. Original paper works in HLZ convention.
- $^{32}$  GRAESSER 00 obtains a bound from graviton contributions to g-2 of the muon through loops of 0.29 TeV for  $\delta=2$  and 0.38 TeV for  $\delta=4,6$ . Limits scale as  $\lambda^{1/2}$ . However calculational scheme not well-defined without specification of high-scale theory. See the "Extra Dimensions Review."
- 33 HAN 00 calculates corrections to gauge boson self-energies from KK graviton loops and constrain them using S and T. Bounds on  $M_{TT}$  range from 0.5 TeV ( $\delta$ =6) to 1.1 TeV ( $\delta$ =2); see text. Limits have strong dependence,  $\lambda^{\delta+2}$ , on unknown  $\lambda$  coefficient.
- <sup>34</sup> MATHEWS 00 search for evidence of graviton exchange in CDF and DØ dijet production data. See their Table 2 for slightly stronger  $\delta$ -dependent bounds. Limits expressed in terms of  $\widetilde{M}_S^4 = M_{TT}^4/8$ .

<sup>35</sup> MELE 00 obtains bound from KK graviton contributions to  $e^+e^- \rightarrow VV$  ( $V=\gamma,W,Z$ ) at LEP. Authors use Hewett conventions.

 $^{36}$  ABBIENDI 99P search for s-channel graviton exchange effects in  $e^+e^- \to \gamma\gamma$  at  $E_{\rm cm}{=}189$  GeV. The limits  $G_+>660$  GeV and  $G_->634$  GeV are obtained from combined  $E_{\rm cm}{=}183$  and 189 GeV data, where  $G_\pm$  is a scale related to the fundamental gravity scale.

<sup>37</sup> ACCIARRI 99M search for the reaction  $e^+e^- \to \gamma G$  and s-channel graviton exchange effects in  $e^+e^- \to \gamma \gamma$ ,  $W^+W^-$ , ZZ,  $e^+e^-$ ,  $\mu^+\mu^-$ ,  $\tau^+\tau^-$ ,  $q\overline{q}$  at  $E_{\rm cm}=$ 183 GeV. Limits on the gravity scale are listed in their Tables 1 and 2.

38 ACCIARRI 99S search for the reaction  $e^+e^- \rightarrow ZG$  and s-channel graviton exchange effects in  $e^+e^- \rightarrow \gamma\gamma$ ,  $W^+W^-$ , ZZ,  $e^+e^-$ ,  $\mu^+\mu^-$ ,  $\tau^+\tau^-$ ,  $q\overline{q}$  at  $E_{\rm cm}=$ 189 GeV. Limits on the gravity scale are listed in their Tables 1 and 2.

<sup>39</sup> BOURILKOV 99 performs global analysis of LEP data on  $e^+e^-$  collisions at  $\sqrt{s}$ =183 and 189 GeV. Bound is on  $\Lambda_T$ .

# Limits on $1/R = M_c$

This section includes limits on  $1/R=M_{\rm C}$ , the compactification scale in models with one TeV-sized extra dimension, due to exchange of Standard Model KK excitations. Bounds assume fermions are not in the bulk, unless stated otherwise. See the "Extra Dimensions" review for discussion of model dependence.

VALUE (TeV)	CL%	DOCUMENT ID		TECN	COMMENT
>4.16	95	<sup>1</sup> AAD	<b>12</b> CC	ATLS	$pp  o \ell \overline{\ell}$
>6.1		^	04	RVUE	Electroweak
• • • We do not	use the f	following data for av	erage	s, fits, li	mits, etc. • • •
>3.8	95	<sup>3</sup> ACCOMANDO	15	RVUE	Electroweak
>3.40	95	<sup>4</sup> KHACHATRY			
		<sup>5</sup> CHATRCHYAN	l 13AQ	CMS	$pp \rightarrow \ell X$
>1.38	95				$pp \rightarrow \gamma\gamma$ , $\delta$ =6, $M_D$ =5 TeV
>0.715	95	<sup>7</sup> EDELHAUSER	13	RVUE	$pp \rightarrow \ell \overline{\ell} + X$
>1.40	95				$pp \rightarrow \gamma\gamma$ , $\delta$ =6, $M_D$ =5 TeV
>1.23	95	<sup>9</sup> AAD	12X	ATLS	$pp \rightarrow \gamma\gamma$ , $\delta$ =6, $M_D$ =5 TeV
>0.26	95	<sup>10</sup> ABAZOV	12M	D0	$ ho  \overline{ ho}  ightarrow  ho  \mu  \mu$
>0.75	95		12	RVUE	Electroweak
		<sup>12</sup> FLACKE	12	RVUE	
>0.43	95	<sup>13</sup> NISHIWAKI			
>0.729	95				$pp \rightarrow \gamma\gamma$ , $\delta$ =6, $M_D$ =5 TeV
>0.961	95	<sup>15</sup> AAD			$pp \rightarrow \gamma\gamma$ , $\delta$ =6, $M_D$ =5 TeV
>0.477	95	<sup>16</sup> ABAZOV			$p\overline{p} \rightarrow \gamma\gamma$ , $\delta$ =6, $M_D$ =5 TeV
>1.59	95				
>0.6	95		07	RVUE	$\overline{B} \rightarrow X_{S} \gamma$
>0.6	90	<sup>19</sup> GOGOLADZE	06	RVUE	Electroweak
>3.3	95	<sup>20</sup> CORNET	00	RVUE	Electroweak
> 3.3–3.8	95	<sup>21</sup> RIZZO	00	RVUE	Electroweak

 $<sup>^1</sup>$  AAD 12CC use 4.9 and 5.0 fb $^{-1}$  of data from pp collisions at  $\sqrt{s}=7$  TeV in the dielectron and dimuon channels, respectively, to place a lower bound on the mass of the lightest KK  $Z/\gamma$  boson (equivalent to  $1/R=M_{\rm C}$ ). The limit quoted here assumes a flat prior corresponding to when the pure  $Z/\gamma$  KK cross section term dominates. See their Section 15 for more details.

<sup>&</sup>lt;sup>2</sup>BARBIERI 04 use electroweak precision observables to place a lower bound on the compactification scale 1/R. Both the gauge bosons and the Higgs boson are assumed to propagate in the bulk.

- <sup>3</sup> ACCOMANDO 15 use electroweak precision observables to place a lower bound on the compactification scale 1/R. See their Fig. 2 for the bound as a function of  $\sin\beta$ , which parametrizes the VEV contribution from brane and bulk Higgs fields. The quoted value is for the minimum bound which occurs at  $\sin\beta = 0.45$ .
- <sup>4</sup> KHACHATRYAN 15T use 19.7 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=8$  TeV to place a lower bound on the compactification scale 1/R.
- <sup>5</sup> CHATRCHYAN 13AQ use 5.0 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=7$  TeV and a further 3.7 fb<sup>-1</sup> of data at  $\sqrt{s}=8$  TeV to place a lower bound on the compactification scale 1/R, in models with universal extra dimensions and Standard Model fields propagating in the bulk. See their Fig. 5 for the bound as a function of the universal bulk fermion mass parameter  $\mu$ .
- <sup>6</sup> CHATRCHYAN 13W use diphoton events with large missing transverse momentum in 4.93 fb<sup>-1</sup> of data produced from pp collisions at  $\sqrt{s}=7$  TeV to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale  $\Lambda$ , for the radiative corrections to the Kaluza-Klein masses, satisfies  $\Lambda/M_C=20$ . The model parameters are chosen such that the decay  $\gamma^*\to G\gamma$  occurs with an appreciable branching fraction.
- $^7$  EDELHAUSER 13 use 19.6 and 20.6 fb $^{-1}$  of data from pp collisions at  $\sqrt{s}=8$  TeV analyzed by the CMS Collaboration in the dielectron and dimuon channels, respectively, to place a lower bound on the mass of the second lightest Kaluza-Klein  $Z/\gamma$  boson (converted to a limit on  $1/R=M_{\rm C}$ ). The bound assumes Standard Model fields propagating in the bulk and that the cutoff scale  $\Lambda$ , for the radiative corrections to the Kaluza-Klein masses, satisfies  $\Lambda/M_{\rm C}=20$ .
- <sup>8</sup> AAD 12CP use diphoton events with large missing transverse momentum in 4.8 fb<sup>-1</sup> of data produced from pp collisions at  $\sqrt{s}=7$  TeV to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale  $\Lambda$ , for the radiative corrections to the Kaluza-Klein masses, satisfies  $\Lambda/M_c=20$ . The model parameters are chosen such that the decay  $\gamma^*\to G\gamma$  occurs with an appreciable branching fraction.
- $^9$  AAD 12X use diphoton events with large missing transverse momentum in 1.07 fb $^{-1}$  of data produced from pp collisions at  $\sqrt{s}=7$  TeV to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale  $\varLambda$ , for the radiative corrections to the Kaluza-Klein masses, satisfies  $\varLambda/M_C=20$ . The model parameters are chosen such that the decay  $\gamma^*\to G\gamma$  occurs with an appreciable branching fraction.
- $^{10}$  ABAZOV 12M use same-sign dimuon events in 7.3 fb $^{-1}$  of data from  $p\overline{p}$  collisions at  $\sqrt{s}=1.96$  TeV to place a lower bound on the compactification scale 1/R, in models with universal extra dimensions where all Standard Model fields propagate in the bulk.
- $^{11}$  BAAK 12 use electroweak precision observables to place a lower bound on the compactification scale 1/R, in models with universal extra dimensions and Standard Model fields propagating in the bulk. Bound assumes a 125 GeV Higgs mass. See their Fig. 25 for the bound as a function of the Higgs mass.
- $^{12}$  FLACKE 12 use electroweak precision observables to place a lower bound on the compactification scale 1/R, in models with universal extra dimensions and Standard Model fields propagating in the bulk. See their Fig. 1 for the bound as a function of the universal bulk fermion mass parameter  $\mu$ .
- $^{13}\,\text{NISHIWAKI}$  12 use up to 2 fb $^{-1}$  of data from the ATLAS and CMS experiments that constrains the production cross section of a Higgs-like particle to place a lower bound on the compactification scale 1/R in universal extra dimension models. The quoted bound assumes Standard Model fields propagating in the bulk and a 125 GeV Higgs mass. See their Fig. 1 for the bound as a function of the Higgs mass.
- <sup>14</sup> AAD 11F use diphoton events with large missing transverse energy in 3.1 pb<sup>-1</sup> of data produced from pp collisions at  $\sqrt{s}=7$  TeV to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale  $\Lambda$ , for the radiative corrections to the Kaluza-Klein masses,

- satisfies  $\Lambda/\mathrm{M}_c=$  20. The model parameters are chosen such that the decay  $\gamma^* \to ~G \gamma$ occurs with an appreciable branching fraction.
- $^{15}$  AAD  $^{11}$ X use diphoton events with large missing transverse energy in  $^{36}$  pb $^{-1}$  of data produced from pp collisions at  $\sqrt{s}=7$  TeV to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale  $\Lambda$ , for the radiative corrections to the Kaluza-Klein masses, satisfies  $\Lambda/M_C=$  20. The model parameters are chosen such that the decay  $\gamma^* \to G \gamma$ occurs with an appreciable branching fraction.
- $^{16}\,\mathrm{ABAZOV}$  10P use diphoton events with large missing transverse energy in 6.3 fb $^{-1}$  of data produced from  $p\bar{p}$  collisions at  $\sqrt{s}=1.96$  TeV to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale  $\Lambda$ , for the radiative corrections to the Kaluza-Klein masses, satisfies  $\Lambda/M_c=20$ . The model parameters are chosen such that the decay  $\gamma^* \rightarrow G \gamma$  occurs with an appreciable branching fraction.
- $^{17}$  ABAZOV 09AE use dijet angular distributions in 0.7 fb $^{-1}$  of data from  $p\overline{p}$  collisions at  $\sqrt{s}=1.96$  TeV to place a lower bound on the compactification scale.
- $^{18}$  HAISCH 07 use inclusive  $\overline{B}$ -meson decays to place a Higgs mass independent bound on the compactification scale 1/R in the minimal universal extra dimension model.
- $^{
  m 19}\,{\sf GOGOLADZE}$  06 use electroweak precision observables to place a lower bound on the compactification scale in models with universal extra dimensions. Bound assumes a 115 GeV Higgs mass. See their Fig. 3 for the bound as a function of the Higgs mass.
- $^{20}$  CORNET 00 translates a bound on the coefficient of the 4-fermion operator  $(\overline{\ell}\gamma_{\mu}\tau^a\ell)(\overline{\ell}\gamma^{\mu}\tau^a\ell)$  derived by Hagiwara and Matsumoto into a limit on the mass scale of KK W bosons.
- 21 RIZZO 00 obtains limits from global electroweak fits in models with a Higgs in the bulk (3.8 TeV) or on the standard brane (3.3 TeV).

# Limits on Kaluza-Klein Gravitons in Warped Extra Dimensions

This section places limits on the mass of the first Kaluza-Klein (KK) excitation of the graviton in the warped extra dimension model of Randall and Sundrum. Bounds in parenthesis assume Standard Model fields propagate in the bulk. Experimental bounds depend strongly on the warp parameter, k. See the "Extra Dimensions" review for a full discussion.

Here we list limits for the value of the warp parameter  $k/\overline{M}_P = 0.1$ .

VALUE (TeV)	CL%	DOCUMENT ID TECN COMMENT
>3.3	95	$^{1}$ KHACHATRY16M CMS $pp  ightarrow  G  ightarrow  \gamma \gamma$
ullet $ullet$ We do not	use the f	ollowing data for averages, fits, limits, etc. • • •
		<sup>2</sup> AABOUD 16AE ATLS $pp \rightarrow G \rightarrow WW,ZZ$
		$^3$ AABOUD 16H ATLS $pp  o G  o \gamma \gamma$
		$^4$ AABOUD 161 ATLS $pp \rightarrow G \rightarrow hh$
		$^{5}$ AAD 16R ATLS $pp \rightarrow G \rightarrow WW,ZZ$
		$^6$ KHACHATRY16BQ CMS $pp  o G  o hh$
>2.66	95	$^{7}$ AAD 15AD ATLS $pp \rightarrow G \rightarrow \gamma \gamma$
		${}^{8}$ AAD 15AU ATLS $pp \rightarrow G \rightarrow ZZ$
		$^9$ AAD 15AZ ATLS $pp \rightarrow G \rightarrow WW$
		$^{10}$ AAD $^{15}$ BK ATLS $^{}$
		$^{11}$ AAD $^{15}$ CT ATLS $^{17}$ PP $\rightarrow$ $^{17}$ G $\rightarrow$ $^{17}$ WW, $^{17}$ Z
>2.73	95	$^{12}$ KHACHATRY15AE CMS $pp  ightarrow e^+e^-, \; \mu^+\mu^-$
		$^{13}$ KHACHATRY15R CMS $pp  ightarrow G  ightarrow hh$
>2.68	95	14 AAD 14V ATLS $pp \rightarrow G \rightarrow e^+e^-, \mu^+\mu^-$
		<sup>15</sup> KHACHATRY14A CMS $pp \rightarrow G \rightarrow WW, ZZ, WZ$
>1.23 (>0.84)	95	16 AAD 13A ATLS $pp \rightarrow G \rightarrow WW$
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95		ATLS	$pp \rightarrow G \rightarrow WW$
95			$pp \rightarrow \gamma \gamma$ , $e^+e^-$ , $\mu^+\mu^-$
95	<sup>19</sup> CHATRCHYAN 13AF	CMS	$pp  ightarrow e^+e^-$ , $\mu^+\mu^-$
	<sup>20</sup> CHATRCHYAN 13U	CMS	$pp \rightarrow G \rightarrow ZZ$
95		ATLS	$pp \rightarrow G \rightarrow ZZ$
95		ATLS	$pp \rightarrow G \rightarrow \ell \overline{\ell}$
95			$pp \rightarrow \gamma \gamma$ , $e^+e^-$ , $\mu^+\mu^-$
	<sup>24</sup> AALTONEN 12V	CDF	$p\overline{p} \rightarrow G \rightarrow ZZ$
			Electroweak
95	<sup>26</sup> CHATRCHYAN 12R	CMS	$pp \rightarrow G \rightarrow \gamma \gamma$
95			$pp \rightarrow G \rightarrow \ell \overline{\ell}$
			$p\overline{p} \rightarrow G \rightarrow ZZ$
95	<sup>29</sup> AALTONEN 11R	CDF	$p\overline{p} \rightarrow G \rightarrow e^+e^-, \gamma\gamma$
95		D0	$p\overline{p} \to G \to WW$
95		CMS	$pp \rightarrow G \rightarrow \ell \overline{\ell}$
		CDF	$p\overline{p} \rightarrow G \rightarrow WW$
		D0	$p\overline{p} \rightarrow G \rightarrow e^+e^-, \gamma\gamma$
	<sup>34</sup> AALTONEN 08S	CDF	$p\overline{p} \rightarrow G \rightarrow ZZ$
		D0	$p\overline{p} \rightarrow G \rightarrow e^+e^-, \gamma\gamma$
	<sup>36</sup> AALTONEN 07G	CDF	$p\overline{p} \rightarrow G \rightarrow \gamma\gamma$
		CDF	$p\overline{p}  ightarrow G  ightarrow e\overline{e}$
		D0	$p\overline{p}  ightarrow G  ightarrow \ell\ell$ , $\gamma\gamma$
	<sup>39</sup> ABULENCIA 05A	CDF	$p\overline{p}  ightarrow  G  ightarrow  \ell\overline{\ell}$
	95 95 95 95 95 95 95	95	95

 $<sup>^1</sup>$  KHACHATRYAN 16M use 19.7 fb $^{-1}$  and 3.3 fb $^{-1}$  of data from pp collisions at  $\sqrt{s}=8$  TeV and 13 TeV, respectively, in the diphoton channel to place a lower limit on the mass of the lightest KK graviton. See their paper for limits with other warp parameter values  $k/\overline{M}_P=0.01$  and 0.2.

<sup>&</sup>lt;sup>2</sup> AABOUD 16AE use 3.2 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=13$  TeV to place a lower bound on the mass of the lightest KK graviton. See their Figure 8 for the limit on the KK graviton mass as a function of the cross section times branching fraction for  $k/\overline{M}_P=1$ .

 $<sup>^3</sup>$  AABOUD 16H use 3.2 fb $^{-1}$  of data from  $p\,p$  collisions at  $\sqrt{s}=13$  TeV in the diphoton channel to place a lower limit on the mass of the lightest KK graviton. See their Figure 11 for limits on the cross section times branching fraction as a function of the graviton mass with warp parameter values  $k/\overline{M}_P$  between 0.01 and 0.3.

<sup>&</sup>lt;sup>4</sup> AABOUD 16I use 3.2 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=13$  TeV to search for Higgs boson pair production in the  $b\overline{b}b\overline{b}$  final state. See their Figure 10 for limits on the cross section times branching fraction as a function of the KK graviton mass with warp parameter values  $k/\overline{M}_P=1.0$  and 2.0.

 $<sup>^5</sup>$  AAD 16R use 20.3 fb $^{-1}$  of data from pp collisions at  $\sqrt{s}=8$  TeV to place a lower bound on the mass of the lightest KK graviton. See their Figure 4 for the limit on the KK graviton mass as a function of the cross section times branching fraction.

 $<sup>^6</sup>$  KHACHATRYAN  $^{16}$ BQ use  $^{19.7}$  fb $^{-1}$  of data from pp collisions at  $\sqrt{s}=8$  TeV to search for Higgs boson pair production in the  $\gamma\gamma b\overline{b}$  final state. See their Figure 9 for limits on the cross section times branching fraction as a function of the KK graviton mass with a warp parameter value  $k/\overline{M}_P=0.2.$ 

 $<sup>^7</sup>$  AAD 15AD use 20.3 fb $^{-1}$  of data from pp collisions at  $\sqrt{s}=8$  TeV in the diphoton channel to place a lower limit on the mass of the lightest KK graviton. See their Table IV for limits with warp parameter values  $k/\overline{M}_P$  between 0.01 and 0.1.

- <sup>8</sup> AAD 15AU use 20 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=8$  TeV to search for KK gravitons in a warped extra dimension decaying to ZZ dibosons. See their Figure 2 for limits on the KK graviton mass as a function of the cross section times branching fraction.
- $^9$  AAD 15AZ use 20.3 fb $^{-1}$  of data from pp collisions at  $\sqrt{s}=8$  TeV to place a lower bound on the mass of the lightest KK graviton. See their Figure 2 for limits on the KK graviton mass as a function of the cross section times branching ratio.
- $^{10}$  AAD 15BK use 19.5 fb $^{-1}$  of data from pp collisions at  $\sqrt{s}=8$  TeV to search for Higgs boson pair production in the  $b\overline{b}b\overline{b}$  final state, and exclude masses of the lightest KK graviton. See their Table 9 for the excluded mass ranges with warp parameter values  $k/\overline{M}_P=1.0,\ 1.5,\ {\rm and}\ 2.0.$
- $^{11}$  AAD 15CT use 20.3 fb $^{-1}$  of data from pp collisions at  $\sqrt{s}=8$  TeV to place a lower bound on the mass of the lightest KK graviton. See their Figures 6b and 6c for the limit on the KK graviton mass as a function of the cross section times branching fraction.
- $^{12}$  KHACHATRYAN 15AE use 20.6 (19.7) fb $^{-1}$  of data from pp collisions at  $\sqrt{s}=8$  TeV in the dimuon (dielectron) channel to place a lower bound on the mass of the lightest KK graviton.
- <sup>13</sup> KHACHATRYAN 15R use 17.9 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=8$  TeV to search for Higgs boson pair production in the  $b\overline{b}b\overline{b}$  final state, and exclude a KK graviton with mass from 380 to 830 GeV.
- <sup>14</sup> AAD 14V use 20 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s} = 8$  TeV in the dielectron and dimuon channels to place a lower bound on the mass of the lightest KK graviton.
- $^{15}$  KHACHATRYAN 14A use 19.7 fb $^{-1}$  of data from  $p\,p$  collisions at  $\sqrt{s}=8$  TeV to search for KK gravitons in a warped extra dimension decaying to dibosons. See their Figure 9 for limits on the cross section times branching fraction as a function of the KK graviton mass.
- $^{16}$  AAD 13A use 4.7 fb $^{-1}$  of data from pp collisions at  $\sqrt{s}=7$  TeV in the  $\ell\nu\ell\nu$  channel, to place a lower bound on the mass of the lightest KK graviton.
- $^{17}$  AAD 13AO use 4.7 fb $^{-1}$  of data from pp collisions at  $\sqrt{s}=7$  TeV in the  $\ell\nu jj$  channel, to place a lower bound on the mass of the lightest KK graviton.
- $^{18}$  AAD 13AS use 4.9 fb $^{-1}$  of data from pp collisions at  $\sqrt{s}=7$  TeV in the diphoton channel to place lower limits on the mass of the lightest KK graviton. The diphoton channel is combined with previous results in the dielectron and dimuon channels to set the best limit. See their Table 2 for warp parameter values  $k/\overline{M}_P$  between 0.01 and 0.1.
- <sup>19</sup> CHATRCHYAN 13AF use 5.3 and 4.1 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=7$  TeV and 8 TeV, respectively, in the dielectron and dimuon channels, to place a lower bound on the mass of the lightest KK graviton.
- <sup>20</sup> CHATRCHYAN 13U use 5 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=7$  TeV to search for KK gravitons in a warped extra dimension decaying to ZZ dibosons. See their Figure 5 for limits on the lightest KK graviton mass as a function of  $k/\overline{M}_P$ .
- <sup>21</sup> AAD 12AD use 1.02 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=7$  TeV to search for KK gravitons in a warped extra dimension decaying to ZZ dibosons in the IIjj and IIII channels ( $\ell=e, \mu$ ). The limit is quoted for the combined IIjj+IIII channels. See their Figure 5 for limits on the cross section  $\sigma(G \to ZZ)$  as a function of the graviton mass.
- <sup>22</sup> AAD 12CC use 4.9 and 5.0 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=7$  TeV in the dielectron and dimuon channels, respectively, to place a lower bound on the mass of the lightest KK graviton. See their Figure 5 for limits on the lightest KK graviton mass as a function of  $k/\overline{M}p$ .
- <sup>23</sup> AAD 12Y use 2.12 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=7$  TeV in the diphoton channel to place lower limits on the mass of the lightest KK graviton. The diphoton channel is combined with previous results in the dielectron and dimuon channels to set the best limit. See their Table 3 for warp parameter values  $k/\overline{M}_P$  between 0.01 and 0.1.
- <sup>24</sup> AALTONEN 12V use 6 fb<sup>-1</sup> of data from  $p\overline{p}$  collisions at  $\sqrt{s}=1.96$  TeV to search for KK gravitons in a warped extra dimension decaying to ZZ dibosons in the IIjj and IIII channels ( $\ell=e, \mu$ ). It provides improved limits over the previous analysis in

- AALTONEN 11G. See their Figure 16 for limits from all channels combined on the cross section times branching ratio  $\sigma(p\overline{p}\to G^*\to ZZ)$  as a function of the graviton mass.
- $^{25}\,\text{BAAK}$  12 use electroweak precision observables to place a lower bound on the compactification scale  $k~e^{-\pi~k~R}$ , assuming Standard Model fields propagate in the bulk and the Higgs is confined to the IR brane. See their Fig. 27 for more details.
- $^{26}$  CHATRCHYAN 12R use 2.2 fb $^{-1}$  of data from pp collisions at  $\sqrt{s}=7$  TeV in the diphoton channel to place lower limits on the mass of the lightest KK graviton. See their Table III for warp parameter values  $k/\overline{M}_P$  between 0.01 and 0.1.
- $^{27}$  AAD 11AD use 1.08 and 1.21 fb $^{-1}$  of data from pp collisions at  $\sqrt{s}=7$  TeV in the dielectron and dimuon channels, respectively, to place a lower bound on the mass of the lightest graviton. For warp parameter values  $k/\overline{M}_P$  between 0.01 to 0.1 the lower limit on the mass of the lightest graviton is between 0.71 and 1.63 TeV. See their Table IV for more details.
- AALTONEN 11G use 2.5–2.9 fb $^{-1}$  of data from  $p\bar{p}$  collisions at  $\sqrt{s}=1.96$  TeV to search for KK gravitons in a warped extra dimension decaying to ZZ dibosons via the  $e\,e\,e$ ,  $e\,e\,\mu\mu$ ,  $\mu\mu\mu\mu$ ,  $e\,e\,j\,j$ , and  $\mu\mu\,j\,j$  channels. See their Fig. 20 for limits on the cross section  $\sigma(G\to ZZ)$  as a function of the graviton mass.
- <sup>29</sup> AALTONEN 11R uses 5.7 fb<sup>-1</sup> of data from  $p\overline{p}$  collisions at  $\sqrt{s}=1.96$  TeV in the dielectron channel to place a lower bound on the mass of the lightest graviton. It provides combined limits with the diphoton channel analysis of AALTONEN 11U. For warp parameter values  $k/\overline{M}_P$  between 0.01 to 0.1 the lower limit on the mass of the lightest graviton is between 612 and 1058 GeV. See their Table I for more details.
- $^{30}$  ABAZOV 11H use 5.4 fb $^{-1}$  of data from  $p\overline{p}$  collisions at  $\sqrt{s}=1.96$  TeV to place a lower bound on the mass of the lightest graviton. Their 95% C.L. exclusion limit does not include masses less than 300 GeV.
- <sup>31</sup> CHATRCHYAN 11 use 35 and 40 pb<sup>-1</sup> of data from pp collisions at  $\sqrt{s}=7$  TeV in the dielectron and dimuon channels, respectively, to place a lower bound on the mass of the lightest graviton. For a warp parameter value  $k/\overline{M}p=0.05$ , the lower limit on the mass of the lightest graviton is 0.855 TeV.
- <sup>32</sup> AALTONEN 10N use 2.9 fb<sup>-1</sup> of data from  $p\overline{p}$  collisions at  $\sqrt{s}=1.96$  TeV to place a lower bound on the mass of the lightest graviton.
- $^{33}$  ABAZOV 10F use 5.4 fb $^{-1}$  of data from  $p\overline{p}$  collisions at  $\sqrt{s}=1.96$  TeV to place a lower bound on the mass of the lightest graviton. For warp parameter values of  $k/\overline{M}_P$  between 0.01 and 0.1 the lower limit on the mass of the lightest graviton is between 560 and 1050 GeV. See their Fig. 3 for more details.
- <sup>34</sup> AALTONEN 08s use  $p\overline{p}$  collisions at  $\sqrt{s}=1.96$  TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to four electrons via two Z bosons using 1.1 fb<sup>-1</sup> of data. See their Fig. 8 for limits on  $\sigma \cdot B(G \to ZZ)$  versus the graviton mass.
- ABAZOV 08J use  $p\overline{p}$  collisions at  $\sqrt{s}=1.96$  TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to electrons and photons using 1 fb<sup>-1</sup> of data. For warp parameter values of  $k/\overline{M}_P$  between 0.01 and 0.1 the lower limit on the mass of the lightest excitation is between 300 and 900 GeV. See their Fig. 4 for more details.
- $^{36}$  AALTONEN 07G use  $p\overline{p}$  collisions at  $\sqrt{s}=1.96$  TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to photons using 1.2 fb $^{-1}$  of data. For warp parameter values of  $k/\overline{M}_P=0.1$ , 0.05, and 0.01 the bounds on the graviton mass are 850, 694, and 230 GeV, respectively. See their Fig. 3 for more details. See also AALTONEN 07H.
- 37 AALTONEN 07H use  $p\overline{p}$  collisions at  $\sqrt{s}=1.96$  TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to electrons using  $1.3~{\rm fb}^{-1}$  of data. For a warp parameter value of  $k/\overline{M}_P=0.1$  the bound on the graviton mass is 807 GeV. See their Fig. 4 for more details. A combined analysis with the diphoton data of AALTONEN 07G yields for  $k/\overline{M}_P=0.1$  a graviton mass lower bound of 889 GeV.

- <sup>38</sup> ABAZOV 05N use  $p\overline{p}$  collisions at  $\sqrt{s}=1.96$  TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to muons, electrons or photons, using 260 pb<sup>-1</sup> of data. For warp parameter values of  $k/\overline{M}_P=0.1$ , 0.05, and 0.01, the bounds on the graviton mass are 785, 650 and 250 GeV respectively. See their Fig. 3 for more details.
- <sup>39</sup> ABULENCIA 05A use  $p\overline{p}$  collisions at  $\sqrt{s}=1.96$  TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to muons or electrons, using 200 pb<sup>-1</sup> of data. For warp parameter values of  $k/\overline{M}_P=0.1$ , 0.05, and 0.01, the bounds on the graviton mass are 710, 510 and 170 GeV respectively.

## Limits on Kaluza-Klein Gluons in Warped Extra Dimensions

This section places limits on the mass of the first Kaluza-Klein (KK) excitation of the gluon in warped extra dimension models with Standard Model fields propagating in the bulk. Bounds are given for a specific benchmark model with  $\Gamma/m=15.3\%$  where  $\Gamma$  is the width and m the mass of the KK gluon. See the "Extra Dimensions" review for more discussion.

VALUE (TeV)	CL%	DOCUMENT ID		TECN	COMMENT
>2.5	95	<sup>1</sup> CHATRCHYAI	<b>13</b> BM	ICMS	$g_{KK}  ightarrow t \overline{t}$
• • • We do not use the	following	data for averages	s, fits,	limits, e	tc. • • •
>2.07	95	<sup>2</sup> AAD	13AQ		${\it g}_{KK}  ightarrow  {\it t}  {\overline {\it t}}  ightarrow  \ell j$
	0.5	<sup>3</sup> CHEN <sup>4</sup> AAD	13A	A.T.I. C	$\overline{B} \rightarrow X_{S} \gamma$
>1.5	95	TAAD	12BV	ATLS	$g_{KK}  ightarrow t \overline{t}  ightarrow \ell j$

 $<sup>^1</sup>$  CHATRCHYAN 13BM use 19.7 fb $^{-1}$  of data from pp collisions at  $\sqrt{s}=8$  TeV. Bound is for a width of approximately 15–20% of the KK gluon mass.

## REFERENCES FOR Extra Dimensions

AABOUD 16AE	JHEP 1609 173	M. Aaboud et al.	(ATLAS Collab.)
AABOUD 16D	PR D94 032005	M. Aaboud et al.	(ATLAS Collab.)
AABOUD 16F	JHEP 1606 059	M. Aaboud et al.	(ATLAS Collab.)
AABOUD 16H	JHEP 1609 001	M. Aaboud et al.	(ATLAS Collab.)
AABOUD 16I	PR D94 052002	M. Aaboud et al.	(ATLAS Collab.)
AAD 16R	PL B755 285	G. Aad <i>et al.</i>	(ATLAS Collab.)
KHACHATRY 16BQ	PR D94 052012	V. Khachatryan <i>et al.</i>	(CMS Collab.)
KHACHATRY 16M	PRL 117 051802	V. Khachatryan <i>et al.</i>	(CMS Collab.)
KHACHATRY 16N	PL B755 102	V. Khachatryan <i>et al.</i>	(CMS Collab.)
AAD 15AD	PR D92 032004	G. Aad <i>et al.</i>	(ATLAS Collab.)
	EPJ C75 69	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD 15AZ	EPJ C75 209	G. Aad <i>et al.</i>	(ATLAS Collab.)
Also	EPJ C75 370 (errat.)	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD 15BH	EPJ C75 299	G. Aad <i>et al.</i>	(ATLAS Collab.)
Also	EPJ C75 408 (errat.)	G. Aad <i>et al.</i>	(ATLAS Collab.)
	EPJ C75 412	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD 15CS	PR D91 012008	G. Aad <i>et al.</i>	(ATLAS Collab.)
Also	PR D92 059903 (errat.)	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD 15CT	JHEP 12 55	G. Aad <i>et al.</i>	(ATLAS Collab.)
ACCOMANDO 15	MPL A30 1540010	E. Accomando	(SHMP)
KHACHATRY 15AE		V. Khachatryan <i>et al.</i>	(CMS Collab.)
KHACHATRY 15AL		V. Khachatryan <i>et al.</i>	(CMS Collab.)
KHACHATRY 15J		V. Khachatryan <i>et al.</i>	(CMS Collab.)
KHACHATRY 15R	PL B749 560	V. Khachatryan <i>et al.</i>	(CMS Collab.)
KHACHATRY 15T	PR D91 092005	V. Khachatryan <i>et al.</i>	(CMS Collab.)
	EPJ C74 3134	G. Aad <i>et al.</i>	(ATLAS Collab.)
AAD 14V	PR D90 052005	G. Aad <i>et al.</i>	(ATLAS Collab.)
KHACHATRY 14A	JHEP 1408 174	V. Khachatryan <i>et al.</i>	(CMS Collab.)

<sup>&</sup>lt;sup>2</sup> AAD 13AQ use 4.7 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s} = 7$  TeV.

<sup>&</sup>lt;sup>3</sup> CHEN 13A place limits on the KK mass scale for a specific warped model with custodial symmetry and bulk fermions. See their Figures 4 and 5.

<sup>&</sup>lt;sup>4</sup>AAD 12BV use 2.05 fb<sup>-1</sup> of data from pp collisions at  $\sqrt{s} = 7$  TeV.

	13AO 13AQ 13AS 13C 13D 13E 13AF 13AQ 13BM	PR D87 072005 PRL 111 211804	(errat.)	G. Aad et al. S. Chatrchyan et al.		(CMS (CMS (CMS (CMS	Collab.) Collab.) Collab.) Collab.) Collab.) Collab.) Collab.)
CHEN EDELHAUSER XU	13A 13 13	CP C37 063102 JHEP 1308 091 JP G40 035107		J-B. Chen <i>et al.</i> L. Edelhauser, T. Flacke, J. Xu <i>et al.</i>	M. Kramer	`	(DALI)
AAD AAD AAD AAD AAD AAD AALTONEN ABAZOV AJELLO BAAK CHATRCHYAN	12AD 12BV 12CC 12CP 12X 12Y 12V 12M 12 12 12	PL B712 331 JHEP 1209 041 JHEP 1211 138 PL B718 411 PL B710 519 PL B710 538 PR D85 012008 PRL 108 131802 JCAP 1202 012 EPJ C72 2003 JHEP 1209 094		G. Aad et al. T. Aaltonen et al. V.M. Abazov et al. M. Ajello et al. M. Baak et al. S. Chatrchyan et al.		(D0) (Fermi-LAT (Gfitter (CMS)	Collab.) Group) Collab.)
CHATRCHYAN CHATRCHYAN FLACKE NISHIWAKI AAD AAD AAD AAD	12R 12 12	PL B711 15 PRL 108 111801 PR D85 126007 PL B707 506 PRL 107 272002 PRL 106 121803 PL B705 294 EPJ C71 1744		S. Chatrchyan et al. S. Chatrchyan et al. T. Flacke, C. Pasold K. Nishiwaki et al. G. Aad et al. G. Aad et al. G. Aad et al. G. Aad et al.		(CMS)	Collab.) Collab.) Collab.)
AALTONEN AALTONEN AALTONEN AARON ABAZOV BEZERRA	11G 11R 11U 11C 11H 11	PR D83 112008 PRL 107 051801 PR D83 011102 PL B705 52 PRL 107 011801 PR D83 075004		T. Aaltonen et al. T. Aaltonen et al. T. Aaltonen et al. F. D. Aaron et al. V.M. Abazov et al. V.B. Bezerra et al.		(CDF (CDF (CDF (H1 (D0	Collab.) Collab.) Collab.) Collab.) Collab.)
CHATRCHYAN CHATRCHYAN CHATRCHYAN SUSHKOV	11A	JHEP 1105 093 JHEP 1105 085 PRL 107 201804 PRL 107 171101		S. Chatrchyan <i>et al.</i> S. Chatrchyan <i>et al.</i> S. Chatychyan <i>et al.</i> A.O. Sushkov <i>et al.</i>		(CMS	Collab.) Collab.) Collab.)
AALTONEN ABAZOV ABAZOV BEZERRA	10N 10F 10P 10	PRL 104 241801 PRL 104 241802 PRL 105 221802 PR D81 055003		T. Aaltonen <i>et al.</i> V.M. Abazov <i>et al.</i> V.M. Abazov <i>et al.</i> V.B. Bezerra <i>et al.</i>		`(D0	Collab.) Collab.) Collab.)
ABAZOV ABAZOV MASUDA AALTONEN AALTONEN ABAZOV ABAZOV	09D 09 08AC 08S 08J 08S	PRL 103 191803 PRL 102 051601 PRL 102 171101 PRL 101 181602 PR D78 012008 PRL 100 091802 PRL 101 011601		V.M. Abazov et al. V.M. Abazov et al. M. Masuda, M. Sasaki T. Aaltonen et al. T. Aaltonen et al. V.M. Abazov et al. V.M. Abazov et al.	DK 6	(CDF (CDF (D0 (D0	Collab.) Collab.) (ICRR) Collab.) Collab.) Collab.) Collab.)
DAS GERACI TRENKEL	08 08 08	PR D78 063011 PR D78 022002 PR D77 122001		P.K. Das, V.H.S. Kumar, A.A. Geraci <i>et al.</i> C. Trenkel	P.K. Sures		(STAN)
AALTONEN AALTONEN DECCA HAISCH KAPNER	07G 07H 07A 07	PRL 99 171801 PRL 99 171802 EPJ C51 963 PR D76 034014 PRL 98 021101		T. Aaltonen et al. T. Aaltonen et al. R.S. Decca et al. U. Haisch, A. Weiler			Collab.) Collab.)
KAPNER SCHAEL TU	07 07A 07	PRL 98 021101 EPJ C49 411 PRL 98 201101		D.J. Kapner <i>et al.</i> S. Schael <i>et al.</i> LC. Tu <i>et al.</i>		(ALEPH	Collab.)
ABDALLAH ABULENCIA,A GERDES GOGOLADZE	06C 06 06 06	EPJ C45 589 PRL 97 171802 PR D73 112008 PR D74 093012		<ul><li>J. Abdallah <i>et al.</i></li><li>A. Abulencia <i>et al.</i></li><li>D. Gerdes <i>et al.</i></li><li>I. Gogoladze, C. Macesan</li></ul>	u	(DELPHI (CDF	Collab.) Collab.)

ABAZOV ABAZOV ABDALLAH ABULENCIA	05N 05V 05B 05A	PRL 95 091801 PRL 95 161602 EPJ C38 395 PRL 95 252001	V.M. Abazov et al. V.M. Abazov et al. J. Abdallah et al. A. Abulencia et al.	(D0 Collab.) (D0 Collab.) (DELPHI Collab.) (CDF Collab.)
SMULLIN ACHARD ACOSTA BARBIERI CASSE	05 04E 04C 04 04	PR D72 122001 PL B587 16 PRL 92 121802 NP B703 127 PRL 92 111102	S.J. Smullin <i>et al.</i> P. Achard <i>et al.</i> D. Acosta <i>et al.</i> R. Barbieri <i>et al.</i> M. Casse <i>et al.</i>	(L3 Collab.) (CDF Collab.)
CHEKANOV HOYLE ABAZOV ABBIENDI	04B 04 03 03D	PL B591 23 PR D70 042004 PRL 90 251802 EPJ C26 331	S. Chekanov <i>et al</i> . C.D. Hoyle <i>et al</i> . V.M. Abazov <i>et al</i> . G. Abbiendi <i>et al</i> .	(ZEUS Collab.) (WASH) (D0 Collab.)
ACHARD ADLOFF CHIAVERINI	03D 03 03	PL B572 133 PL B568 35 PRL 90 151101	P. Achard et al. C. Adloff et al. J. Chiaverini et al.	(OPAL Collab.) (L3 Collab.) (H1 Collab.)
GIUDICE HANNESTAD Also HEISTER	03 03 03C	NP B663 377 PR D67 125008 PR D69 029901(errat.) EPJ C28 1	G.F. Giudice, A. Strumia S. Hannestad, G.G. Raffelt S. Hannestad, G.G. Raffelt A. Heister <i>et al.</i>	(ALEPH Collab.)
LONG ACHARD ACHARD HANNESTAD	03 02 02D 02	Nature 421 922 PL B524 65 PL B531 28 PRL 88 071301	J.C. Long <i>et al.</i> P. Achard <i>et al.</i> P. Achard <i>et al.</i> S. Hannestad, G. Raffelt	(L3 Collab.) (L3 Collab.)
ABBOTT FAIRBAIRN HANHART	01 01 01	PRL 86 1156 PL B508 335 PL B509 1	B. Abbott <i>et al.</i> M. Fairbairn C. Hanhart <i>et al.</i>	(D0 Collab.)
HOYLE ABBIENDI ABREU ABREU CASSISI CHANG	01 00R 00A 00S 00Z 00 00B	PRL 86 1418 EPJ C13 553 PL B491 67 PL B485 45 EPJ C17 53 PL B481 323 PRL 85 3765	C.D. Hoyle et al. G. Abbiendi et al. P. Abreu et al. P. Abreu et al. P. Abreu et al. S. Cassisi et al. L.N. Chang et al.	(OPAL Collab.) (DELPHI Collab.) (DELPHI Collab.) (DELPHI Collab.)
CHEUNG CORNET GRAESSER HAN MATHEWS MELE	00 00 00 00 00 00	PR D61 015005 PR D61 037701 PR D61 074019 PR D62 125018 JHEP 0007 008 PR D61 117901	K. Cheung F. Cornet, M. Relano, J. Rico M.L. Graesser T. Han, D. Marfatia, RJ. Zhang P. Mathews, S. Raychaudhuri, K. Srid S. Mele, E. Sanchez	har
RIZZO ABBIENDI ACCIARRI ACCIARRI ACCIARRI BOURILKOV HOSKINS	00 99P 99M 99R 99S 99	PR D61 016007 PL B465 303 PL B464 135 PL B470 268 PL B470 281 JHEP 9908 006 PR D32 3084	T.G. Rizzo, J.D. Wells G. Abbiendi <i>et al.</i> M. Acciarri <i>et al.</i> M. Acciarri <i>et al.</i> M. Acciarri <i>et al.</i> D. Bourilkov J.K. Hoskins <i>et al.</i>	(OPAL Collab.) (L3 Collab.) (L3 Collab.) (L3 Collab.)