

Number of Neutrino Types

The neutrinos referred to in this section are those of the Standard $SU(2) \times U(1)$ Electroweak Model possibly extended to allow nonzero neutrino masses. Light neutrinos are those with $m < m_Z/2$. The limits are on the number of neutrino mass eigenstates, including ν_1 , ν_2 , and ν_3 .

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Number from e^+e^- Colliders

Number of Light ν Types

VALUE	DOCUMENT ID	TECN
2.9840 ± 0.0082	¹ LEP-SLC	06 RVUE
• • • We do not use the following data for averages, fits, limits, etc. • • •		
3.00 ± 0.05	² LEP	92 RVUE

¹ Combined fit from ALEPH, DELPHI, L3 and OPAL Experiments.

² Simultaneous fits to all measured cross section data from all four LEP experiments.

Number of Light ν Types from Direct Measurement of Invisible Z Width

In the following, the invisible Z width is obtained from studies of single-photon events from the reaction $e^+e^- \rightarrow \nu\bar{\nu}\gamma$. All are obtained from LEP runs in the E_{cm}^{ee} range 88–209 GeV.

VALUE	DOCUMENT ID	TECN	COMMENT
2.92 ± 0.05 OUR AVERAGE	Error includes scale factor of 1.2.		
2.84 ± 0.10 ± 0.14	ABDALLAH	05B DLPH	$\sqrt{s} = 180\text{--}209$ GeV
2.98 ± 0.05 ± 0.04	ACHARD	04E L3	1990–2000 LEP runs
2.86 ± 0.09	HEISTER	03C ALEP	$\sqrt{s} = 189\text{--}209$ GeV
2.69 ± 0.13 ± 0.11	ABBIENDI,G	00D OPAL	1998 LEP run
2.89 ± 0.32 ± 0.19	ABREU	97J DLPH	1993–1994 LEP runs
3.23 ± 0.16 ± 0.10	AKERS	95C OPAL	1990–1992 LEP runs
2.68 ± 0.20 ± 0.20	BUSKULIC	93L ALEP	1990–1991 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.84 ± 0.15 ± 0.14	ABREU	00Z DLPH	1997–1998 LEP runs
3.01 ± 0.08	ACCIARRI	99R L3	1991–1998 LEP runs
3.1 ± 0.6 ± 0.1	ADAM	96C DLPH	$\sqrt{s} = 130, 136$ GeV

Limits from Astrophysics and Cosmology

Number of Light ν Types

(“light” means $<$ about 1 MeV). See also OLIVE 81. For a review of limits based on Nucleosynthesis, Supernovae, and also on terrestrial experiments, see DENEGRI 90.

Also see “Big-Bang Nucleosynthesis” in this *Review*.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				

$0.9 < N_\nu < 8.2$		³ ICHIKAWA	07	COSM	
$3 < N_\nu < 7$	95	⁴ CIRELLI	06	COSM	
$2.7 < N_\nu < 4.6$	95	⁵ HANNESTAD	06	COSM	
$3.6 < N_\nu < 7.4$	95	⁴ SELJAK	06	COSM	
< 4.4		⁶ CYBURT	05	COSM	
< 3.3		⁷ BARGER	03C	COSM	
$1.4 < N_\nu < 6.8$		⁸ CROTTY	03	COSM	
$1.9 < N_\nu < 6.6$		⁸ PIERPAOLI	03	COSM	
$2 < N_\nu < 4$		LISI	99	BBN	
< 4.3		OLIVE	99	BBN	
< 4.9		COPI	97	Cosmology	
< 3.6		HATA	97B	High D/H quasar abs.	
< 4.0		OLIVE	97	BBN; high ⁴ He and ⁷ Li	
< 4.7		CARDALL	96B	COSM High D/H quasar abs.	
< 3.9		FIELDS	96	COSM BBN; high ⁴ He and ⁷ Li	
< 4.5		KERNAN	96	COSM High D/H quasar abs.	
< 3.6		OLIVE	95	BBN; ≥ 3 massless ν	
< 3.3		WALKER	91	Cosmology	
< 3.4		OLIVE	90	Cosmology	
< 4		YANG	84	Cosmology	
< 4		YANG	79	Cosmology	
< 7		STEIGMAN	77	Cosmology	
		PEEBLES	71	Cosmology	
< 16		⁹ SHVARTSMAN	69	Cosmology	
		HOYLE	64	Cosmology	

³ Constrains the number of neutrino types from recent CMB and large scale structure data. No priors on other cosmological parameters are used.

⁴ Constrains the number of neutrino types from recent CMB, large scale structure, Lyman-alpha forest, and SN1a data. The slight preference for $N_\nu > 3$ comes mostly from the Lyman-alpha forest data.

⁵ Constrains the number of neutrino types from recent CMB and large scale structure data. See also HAMANN 07.

⁶ Limit on the number of neutrino types based on ⁴He and D/H abundance assuming a baryon density fixed to the WMAP data. Limit relaxes to 4.6 if D/H is not used or to 5.8 if only D/H and the CMB are used. See also CYBURT 01 and CYBURT 03.

⁷ Limit on the number of neutrino types based on combination of WMAP data and big-bang nucleosynthesis. The limit from WMAP data alone is 8.3. See also KNELLER 01. $N_\nu \geq 3$ is assumed to compute the limit.

⁸ 95% confidence level range on the number of neutrino flavors from WMAP data combined with other CMB measurements, the 2dfGRS data, and HST data.

⁹ SHVARTSMAN 69 limit inferred from his equations.

Number Coupling with Less Than Full Weak Strength

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

< 20	¹⁰ OLIVE	81C	COSM
< 20	¹⁰ STEIGMAN	79	COSM

¹⁰ Limit varies with strength of coupling. See also WALKER 91.

REFERENCES FOR Limits on Number of Neutrino Types

HAMANN	07	JCAP 0708 021	J. Hamann <i>et al.</i>	
ICHIKAWA	07	JCAP 0705 007	K. Ichikawa, M. Kawasaki, F. Takahashi	
CIRELLI	06	JCAP 0612 013	M. Cirelli <i>et al.</i>	
HANNESTAD	06	JCAP 0611 016	S. Hannestad, G. Raffelt	
LEP-SLC	06	PRPL 427 257	ALEPH, DELPHI, L3, OPAL, SLD and working groups	
SELJAK	06	JCAP 0610 014	U. Seljak, A. Slosar, P. McDonald	
ABDALLAH	05B	EPJ C38 395	J. Abdallah <i>et al.</i>	(DELPHI Collab.)
CYBURT	05	ASP 23 313	R.H. Cyburt <i>et al.</i>	
ACHARD	04E	PL B587 16	P. Achard <i>et al.</i>	(L3 Collab.)
BARGER	03C	PL B566 8	V. Barger <i>et al.</i>	
CROTTY	03	PR D67 123005	P. Crotty, J. Lesgourgues, S. Pastor	
CYBURT	03	PL B567 227	R.H. Cyburt, B.D. Fields, K.A. Olive	
HEISTER	03C	EPJ C28 1	A. Heister <i>et al.</i>	(ALEPH Collab.)
PIERPAOLI	03	MNRAS 342 L63	E. Pierpaoli	
CYBURT	01	ASP 17 87	R.H. Cyburt, B.D. Fields, K.A. Olive	
KNELLER	01	PR D64 123506	J.P. Kneller <i>et al.</i>	
ABBIENDI,G	00D	EPJ C18 253	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABREU	00Z	EPJ C17 53	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACCIARRI	99R	PL B470 268	M. Acciarri <i>et al.</i>	(L3 Collab.)
LISI	99	PR D59 123520	E. Lisi, S. Sarkar, F.L. Villante	
OLIVE	99	ASP 11 403	K.A. Olive, D. Thomas	
ABREU	97J	ZPHY C74 577	P. Abreu <i>et al.</i>	(DELPHI Collab.)
COPI	97	PR D55 3389	C.J. Copi, D.N. Schramm, M.S. Turner	(CHIC)
HATA	97B	PR D55 540	N. Hata <i>et al.</i>	(OSU, PENN)
OLIVE	97	ASP 7 27	K.A. Olive, D. Thomas	(MINN, FLOR)
ADAM	96C	PL B380 471	W. Adam <i>et al.</i>	(DELPHI Collab.)
CARDALL	96B	APJ 472 435	C.Y. Cardall, G.M. Fuller	(UCSD)
FIELDS	96	New Ast 1 77	B.D. Fields <i>et al.</i>	(NDAM, CERN, MINN+)
KERNAN	96	PR D54 3681	P.S. Kernan, S. Sarkar	(CASE, OXFTP)
AKERS	95C	ZPHY C65 47	R. Akers <i>et al.</i>	(OPAL Collab.)
OLIVE	95	PL B354 357	K.A. Olive, G. Steigman	(MINN, OSU)
BUSKULIC	93L	PL B313 520	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
LEP	92	PL B276 247	LEP Collabs.	(LEP, ALEPH, DELPHI, L3, OPAL)
WALKER	91	APJ 376 51	T.P. Walker <i>et al.</i>	(HSCA, OSU, CHIC+)
DENEGRI	90	RMP 62 1	D. Denegri, B. Sadoulet, M. Spiro	(CERN, UCB+)
OLIVE	90	PL B236 454	K.A. Olive <i>et al.</i>	(MINN, CHIC, OSU+)
YANG	84	APJ 281 493	J. Yang <i>et al.</i>	(CHIC, BART)
OLIVE	81	APJ 246 557	K.A. Olive <i>et al.</i>	(CHIC, BART)
OLIVE	81C	NP B180 497	K.A. Olive, D.N. Schramm, G. Steigman	(EFI+)
STEIGMAN	79	PRL 43 239	G. Steigman, K.A. Olive, D.N. Schramm	(BART+)
YANG	79	APJ 227 697	J. Yang <i>et al.</i>	(CHIC, YALE, UVA)
STEIGMAN	77	PL 66B 202	G. Steigman, D.N. Schramm, J.E. Gunn	(YALE, CHIC+)
PEEBLES	71	Physical Cosmology	P.Z. Peebles	(PRIN)
		Princeton Univ. Press (1971)		
SHVARTSMAN	69	JETPL 9 184	V.F. Shvartsman	(MOSU)
		Translated from ZETFP 9 315.		
HOYLE	64	NAT 203 1108	F. Hoyle, R.J. Tayler	(CAMB)