

# a<sub>1</sub>(1260)

$$I^G(J^{PC}) = 1^-(1^{++})$$

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## a<sub>1</sub>(1260) MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>1230±40 OUR ESTIMATE</b>					
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
1203 ± 3		<sup>1</sup> GOMEZ-DUMM04	RVUE		$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu_\tau$
1331 ± 10 ± 3	37k	<sup>2</sup> ASNER	00	CLE2	10.6 $e^+ e^- \rightarrow \tau^+ \tau^-, \tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$
1255 ± 7 ± 6	5904	<sup>3</sup> ABREU	98G	DLPH	$e^+ e^-$
1207 ± 5 ± 8	5904	<sup>4</sup> ABREU	98G	DLPH	$e^+ e^-$
1196 ± 4 ± 5	5904	<sup>5,6</sup> ABREU	98G	DLPH	$e^+ e^-$
1240 ± 10		BARBERIS	98B		450 $pp \rightarrow p_f \pi^+ \pi^- \pi^0 p_s$
1262 ± 9 ± 7		<sup>3,7</sup> ACKERSTAFF	97R	OPAL	$E_{cm}^{ee} = 88-94, \tau \rightarrow 3\pi\nu$
1210 ± 7 ± 2		<sup>4,7</sup> ACKERSTAFF	97R	OPAL	$E_{cm}^{ee} = 88-94, \tau \rightarrow 3\pi\nu$
1211 ± 7 <sup>+50</sup> <sub>-0</sub>		<sup>4</sup> ALBRECHT	93C	ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1121 ± 8		<sup>8</sup> ANDO	92	SPEC	8 $\pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
1242 ± 37		<sup>9</sup> IVANOV	91	RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
1260 ± 14		<sup>10</sup> IVANOV	91	RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
1250 ± 9		<sup>11</sup> IVANOV	91	RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
1208 ± 15		ARMSTRONG	90	OMEG 0	300.0 $pp \rightarrow p p \pi^+ \pi^- \pi^0$
1220 ± 15		<sup>12</sup> ISGUR	89	RVUE	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1260 ± 25		<sup>13</sup> BOWLER	88	RVUE	
1166 ± 18 ± 11		BAND	87	MAC	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1164 ± 41 ± 23		BAND	87	MAC	$\tau^+ \rightarrow \pi^+ \pi^0 \pi^0 \nu$
1250 ± 40		<sup>12</sup> TORNQVIST	87	RVUE	
1046 ± 11		ALBRECHT	86B	ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1056 ± 20 ± 15		RUCKSTUHL	86	DLCO	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1194 ± 14 ± 10		SCHMIDKE	86	MRK2	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1255 ± 23		BELLINI	85	SPEC	40 $\pi^- A \rightarrow \pi^- \pi^+ \pi^- A$
1240 ± 80		<sup>14</sup> DANKOWY...	81	SPEC 0	8.45 $\pi^- p \rightarrow n 3\pi$
1280 ± 30		<sup>14</sup> DAUM	81B	CNTR	63,94 $\pi^- p \rightarrow p 3\pi$
1041 ± 13		<sup>15</sup> GAVILLET	77	HBC +	4.2 $K^- p \rightarrow \Sigma 3\pi$

<sup>1</sup> Using the data of BARATE 98R.

<sup>2</sup> From a fit to the  $3\pi$  mass spectrum including the  $K\bar{K}^*$ (892) threshold.

<sup>3</sup> Uses the model of KUHN 90.

<sup>4</sup> Uses the model of ISGUR 89.

<sup>5</sup> Includes the effect of a possible  $a_1'$  state.

<sup>6</sup> Uses the model of FEINDT 90.

- <sup>7</sup> Supersedes AKERS 95P.
- <sup>8</sup> Average and spread of values using 2 variants of the model of BOWLER 75.
- <sup>9</sup> Reanalysis of RUCKSTUHL 86.
- <sup>10</sup> Reanalysis of SCHMIDKE 86.
- <sup>11</sup> Reanalysis of ALBRECHT 86B.
- <sup>12</sup> From a combined reanalysis of ALBRECHT 86B, SCHMIDKE 86, and RUCKSTUHL 86.
- <sup>13</sup> From a combined reanalysis of ALBRECHT 86B and DAUM 81B.
- <sup>14</sup> Uses the model of BOWLER 75.
- <sup>15</sup> Produced in  $K^-$  backward scattering.

### $a_1(1260)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>250 to 600 OUR ESTIMATE</b>					
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
$480 \pm 20$		<sup>16</sup> GOMEZ-DUMM04	RVUE		$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu_\tau$
$460 \pm 85$	205	<sup>17</sup> DRUTSKOY	02	BELL	$B \rightarrow D^{(*)} K^- K^{*0}$
$814 \pm 36 \pm 13$	37k	<sup>18</sup> ASNER	00	CLE2	$10.6 e^+ e^- \rightarrow$ $\tau^+ \tau^-, \tau^- \rightarrow$ $\pi^- \pi^0 \pi^0 \nu_\tau$
$450 \pm 50$	22k	<sup>19</sup> AKHMETSHIN 99E	CMD2		$1.05-1.38 e^+ e^- \rightarrow$ $\pi^+ \pi^- \pi^0 \pi^0$
$570 \pm 10$		<sup>20</sup> BONDAR	99	RVUE	$e^+ e^- \rightarrow 4\pi, \tau \rightarrow$ $3\pi \nu_\tau$
$587 \pm 27 \pm 21$	5904	<sup>21</sup> ABREU	98G	DLPH	$e^+ e^-$
$478 \pm 3 \pm 15$	5904	<sup>22</sup> ABREU	98G	DLPH	$e^+ e^-$
$425 \pm 14 \pm 8$	5904	<sup>23,24</sup> ABREU	98G	DLPH	$e^+ e^-$
$400 \pm 35$		BARBERIS	98B		$450 pp \rightarrow$ $p_f \pi^+ \pi^- \pi^0 p_s$
$621 \pm 32 \pm 58$		<sup>21,25</sup> ACKERSTAFF	97R	OPAL	$E_{cm}^{ee} = 88-94, \tau \rightarrow$ $3\pi \nu$
$457 \pm 15 \pm 17$		<sup>22,25</sup> ACKERSTAFF	97R	OPAL	$E_{cm}^{ee} = 88-94, \tau \rightarrow$ $3\pi \nu$
$446 \pm 21 \begin{smallmatrix} +140 \\ -0 \end{smallmatrix}$		<sup>22</sup> ALBRECHT	93C	ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
$239 \pm 11$		ANDO	92	SPEC	$8 \pi^- p \rightarrow$ $\pi^+ \pi^- \pi^0 n$
$266 \pm 13 \pm 4$		<sup>26</sup> ANDO	92	SPEC	$8 \pi^- p \rightarrow$ $\pi^+ \pi^- \pi^0 n$
$465 \begin{smallmatrix} +228 \\ -143 \end{smallmatrix}$		<sup>27</sup> IVANOV	91	RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
$298 \begin{smallmatrix} +40 \\ -34 \end{smallmatrix}$		<sup>28</sup> IVANOV	91	RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
$488 \pm 32$		<sup>29</sup> IVANOV	91	RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
$430 \pm 50$		ARMSTRONG	90	OMEG 0	$300.0 pp \rightarrow$ $pp \pi^+ \pi^- \pi^0$
$420 \pm 40$		<sup>30</sup> ISGUR	89	RVUE	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
$396 \pm 43$		<sup>31</sup> BOWLER	88	RVUE	
$405 \pm 75 \pm 25$		BAND	87	MAC	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
$419 \pm 108 \pm 57$		BAND	87	MAC	$\tau^+ \rightarrow \pi^+ \pi^0 \pi^0 \nu$
$521 \pm 27$		ALBRECHT	86B	ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$

$476^{+132}_{-120} \pm 54$	RUCKSTUHL 86	DLCO	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
$462 \pm 56 \pm 30$	SCHMIDKE 86	MRK2	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
$292 \pm 40$	BELLINI 85	SPEC	$40 \pi^- A \rightarrow \pi^- \pi^+ \pi^- A$
$380 \pm 100$	<sup>32</sup> DANKOWY... 81	SPEC 0	$8.45 \pi^- p \rightarrow n 3\pi$
$300 \pm 50$	<sup>32</sup> DAUM 81B	CNTR	$63,94 \pi^- p \rightarrow p 3\pi$
$230 \pm 50$	<sup>33</sup> GAVILLET 77	HBC +	$4.2 K^- p \rightarrow \Sigma 3\pi$

<sup>16</sup> Using the data of BARATE 98R.

<sup>17</sup> From a fit of the  $K^- K^{*0}$  distribution assuming  $m_{a_1} = 1230$  MeV and purely resonant production of the  $K^- K^{*0}$  system.

<sup>18</sup> From a fit to the  $3\pi$  mass spectrum including the  $K \bar{K}^*(892)$  threshold.

<sup>19</sup> Using the  $a_1(1260)$  mass of 1230 MeV.

<sup>20</sup> From AKHMETSHIN 99E and ASNER 00 data using the  $a_1(1260)$  mass of 1230 MeV.

<sup>21</sup> Uses the model of KUHN 90.

<sup>22</sup> Uses the model of ISGUR 89.

<sup>23</sup> Includes the effect of a possible  $a_1'$  state.

<sup>24</sup> Uses the model of FEINDT 90.

<sup>25</sup> Supersedes AKERS 95P.

<sup>26</sup> Average and spread of values using 2 variants of the model of BOWLER 75.

<sup>27</sup> Reanalysis of RUCKSTUHL 86.

<sup>28</sup> Reanalysis of SCHMIDKE 86.

<sup>29</sup> Reanalysis of ALBRECHT 86B.

<sup>30</sup> From a combined reanalysis of ALBRECHT 86B, SCHMIDKE 86, and RUCKSTUHL 86.

<sup>31</sup> From a combined reanalysis of ALBRECHT 86B and DAUM 81B.

<sup>32</sup> Uses the model of BOWLER 75.

<sup>33</sup> Produced in  $K^-$  backward scattering.

### **$a_1(1260)$ DECAY MODES**

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\pi^+ \pi^- \pi^0$	
$\Gamma_2$ $\pi^0 \pi^0 \pi^0$	
$\Gamma_3$ $(\rho\pi)_{S\text{-wave}}$	seen
$\Gamma_4$ $(\rho\pi)_{D\text{-wave}}$	seen
$\Gamma_5$ $(\rho(1450)\pi)_{S\text{-wave}}$	seen
$\Gamma_6$ $(\rho(1450)\pi)_{D\text{-wave}}$	seen
$\Gamma_7$ $\sigma\pi$	seen
$\Gamma_8$ $f_0(980)\pi$	not seen
$\Gamma_9$ $f_0(1370)\pi$	seen
$\Gamma_{10}$ $f_2(1270)\pi$	seen
$\Gamma_{11}$ $K \bar{K}^*(892) + \text{c.c.}$	seen
$\Gamma_{12}$ $\pi\gamma$	seen

## $a_1(1260)$ PARTIAL WIDTHS

$\Gamma(\pi\gamma)$				$\Gamma_{12}$
VALUE (keV)	DOCUMENT ID	TECN	COMMENT	
<b>640±246</b>	ZIELINSKI	84C	SPEC	200 $\pi^+ Z \rightarrow Z 3\pi$

### D-wave/S-wave AMPLITUDE RATIO IN DECAY OF $a_1(1260) \rightarrow \rho\pi$

VALUE	DOCUMENT ID	TECN	COMMENT	
<b>-0.108±0.016 OUR AVERAGE</b>				
-0.14 ±0.04 ±0.07	<sup>36</sup> CHUNG	02	B852	18.3 $\pi^- p \rightarrow \pi^+ \pi^- \pi^- p$
-0.10 ±0.02 ±0.02	<sup>34,35</sup> ACKERSTAFF	97R	OPAL	$E_{cm}^{ee} = 88-94, \tau \rightarrow 3\pi\nu$
-0.11 ±0.02	<sup>34</sup> ALBRECHT	93C	ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$

<sup>34</sup> Uses the model of ISGUR 89.

<sup>35</sup> Supersedes AKERS 95P.

<sup>36</sup> Deck-type background not subtracted.

## $a_1(1260)$ BRANCHING RATIOS

$\Gamma((\rho\pi)_S\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_3/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
60.19	37k	<sup>37</sup> ASNER	00	CLE2 10.6 $e^+ e^- \rightarrow \tau^+ \tau^-$ , $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

$\Gamma((\rho\pi)_D\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_4/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
1.30±0.60±0.22	37k	<sup>37</sup> ASNER	00	CLE2 10.6 $e^+ e^- \rightarrow \tau^+ \tau^-$ , $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

$\Gamma((\rho(1450)\pi)_S\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_5/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
0.56±0.84±0.32	37k	<sup>37,38</sup> ASNER	00	CLE2 10.6 $e^+ e^- \rightarrow \tau^+ \tau^-$ , $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

$\Gamma((\rho(1450)\pi)_D\text{-wave})/\Gamma_{\text{total}}$				$\Gamma_6/\Gamma$
VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
2.04±1.20±0.28	37k	<sup>37,38</sup> ASNER	00	CLE2 10.6 $e^+ e^- \rightarrow \tau^+ \tau^-$ , $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

**$\Gamma(\sigma\pi)/\Gamma_{\text{total}}$**   **$\Gamma_7/\Gamma$**

<u>VALUE (units <math>10^{-2}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
seen		CHUNG	02	B852 $18.3 \pi^- p \rightarrow$ $\pi^+ \pi^- \pi^- p$
$18.76 \pm 4.29 \pm 1.48$	37k 37,39	ASNER	00	CLE2 $10.6 e^+ e^- \rightarrow \tau^+ \tau^-$ , $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

**$\Gamma(f_0(980)\pi)/\Gamma_{\text{total}}$**   **$\Gamma_8/\Gamma$**

<u>VALUE (units <math>10^{-2}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
not seen	37k	ASNER	00	CLE2 $10.6 e^+ e^- \rightarrow \tau^+ \tau^-$ , $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

**$\Gamma(f_0(1370)\pi)/\Gamma_{\text{total}}$**   **$\Gamma_9/\Gamma$**

<u>VALUE (units <math>10^{-2}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$7.40 \pm 2.71 \pm 1.26$	37k 37,40	ASNER	00	CLE2 $10.6 e^+ e^- \rightarrow \tau^+ \tau^-$ , $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

**$\Gamma(f_2(1270)\pi)/\Gamma_{\text{total}}$**   **$\Gamma_{10}/\Gamma$**

<u>VALUE (units <math>10^{-2}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$1.19 \pm 0.49 \pm 0.17$	37k 37,41	ASNER	00	CLE2 $10.6 e^+ e^- \rightarrow \tau^+ \tau^-$ , $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

**$\Gamma(K\bar{K}^*(892)+\text{c.c.})/\Gamma_{\text{total}}$**   **$\Gamma_{11}/\Gamma$**

<u>VALUE (units <math>10^{-2}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$2.2 \pm 0.5$	2255	42 COAN	04	CLEO $\tau^- \rightarrow K^- \pi^- K^+ \nu_\tau$
8 to 15	205	43 DRUTSKOY	02	BELL $B \rightarrow D^{(*)} K^- K^{*0}$
$3.3 \pm 0.5 \pm 0.1$	37k	44 ASNER	00	CLE2 $10.6 e^+ e^- \rightarrow \tau^+ \tau^-$ , $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$
$2.6 \pm 0.3$		45 BARATE	99R	ALEP $\tau \rightarrow K\bar{K}\pi\nu_\tau$

**$\Gamma(\sigma\pi)/\Gamma((\rho\pi)_{S\text{-wave}})$**   **$\Gamma_7/\Gamma_3$**

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$\sim 0.3$	28k	AKHMETSHIN 99E	CMD2	$1.05\text{--}1.38 e^+ e^- \rightarrow$ $\pi^+ \pi^- \pi^+ \pi^-$
$0.003 \pm 0.003$		46 LONGACRE	82	RVUE

$$\Gamma(\pi^0\pi^0\pi^0)/\Gamma(\pi^+\pi^-\pi^0)$$

 $\Gamma_2/\Gamma_1$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
<0.008	90	47 BARBERIS 01	450 $p\bar{p} \rightarrow p_f 3\pi^0 p_s$
37 From a fit to the Dalitz plot.			
38 Assuming for $\rho(1450)$ mass and width of 1370 and 386 MeV respectively.			
39 Assuming for $\sigma$ mass and width of 860 and 880 MeV respectively.			
40 Assuming for $f_0(1370)$ mass and width of 1186 and 350 MeV respectively.			
41 Assuming for $f_2(1270)$ mass and width of 1275 and 185 MeV respectively.			
42 Using structure functions from KUHN 92 and DECKER 93A and $B(\tau^- \rightarrow K^- \pi^- K^+ \nu_\tau) = (0.155 \pm 0.006 \pm 0.009)\%$ from BRIERE 03.			
43 From a comparison to ALAM 94 assuming purely resonant production of the $K^- K^*0$ system.			
44 From a fit to the $3\pi$ mass spectrum including the $K\bar{K}^*(892)$ threshold.			
45 Assuming $a_1(1260)$ dominance and taking $B(\tau \rightarrow a_1(1260)\nu_\tau)$ from BUSKULIC 96.			
46 Uses multichannel Aitchison-Bowler model (BOWLER 75). Uses data from GAVILLET 77, DAUM 80, and DANKOWYCH 81.			
47 Inconsistent with observations of $\sigma\pi$ , $f_0(1370)\pi$ , and $f_2(1270)\pi$ decay modes.			

### $a_1(1260)$ REFERENCES

COAN 04	PRL 92 232001	T.E. Coan <i>et al.</i>	(CLEO Collab.)
GOMEZ-DUMM04	PR D69 073002	D. Gomez Dumm, A. Pich, J. Portoles	
BRIERE 03	PRL 90 181802	R. A. Briere <i>et al.</i>	(CLEO Collab.)
CHUNG 02	PR D65 072001	S.U. Chung <i>et al.</i>	(BNL E852 Collab.)
DRUTSKOY 02	PL B542 171	A. Drutskoy <i>et al.</i>	(BELLE Collab.)
BARBERIS 01	PL B507 14	D. Barberis <i>et al.</i>	
ASNER 00	PR D61 012002	D.M. Asner <i>et al.</i>	(CLEO Collab.)
AKHMETSHIN 99E	PL B466 392	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
BARATE 99R	EPJ C11 599	R. Barate <i>et al.</i>	(ALEPH Collab.)
BONDAR 99	PL B466 403	A.E. Bondar <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
ABREU 98G	PL B426 411	P. Abreu <i>et al.</i>	(DELPHI Collab.)
BARATE 98R	EPJ C4 409	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARBERIS 98B	PL B422 399	D. Barberis <i>et al.</i>	(WA 102 Collab.)
ACKERSTAFF 97R	ZPHY C75 593	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
BUSKULIC 96	ZPHY C70 579	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
AKERS 95P	ZPHY C67 45	R. Akers <i>et al.</i>	(OPAL Collab.)
ALAM 94	PR D50 43	M.S. Alam <i>et al.</i>	(CLEO Collab.)
ALBRECHT 93C	ZPHY C58 61	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
DECKER 93A	ZPHY C58 445	R. Decker <i>et al.</i>	
ANDO 92	PL B291 496	A. Ando <i>et al.</i>	(KEK, KYOT, NIRS, SAGA+)
KUHN 92	ZPHY C56 661	J.H. Kuhn, E. Mirkes	
IVANOV 91	ZPHY C49 563	Y.P. Ivanov, A.A. Osipov, M.K. Volkov	(JINR)
ARMSTRONG 90	ZPHY C48 213	T.A. Armstrong, M. Benayoun, W. Beusch	
FEINDT 90	ZPHY C48 681	M. Feindt	(HAMB)
KUHN 90	ZPHY C48 445	J.H. Kuhn <i>et al.</i>	(MPIM)
ISGUR 89	PR D39 1357	N. Isgur, C. Morningstar, C. Reader	(TNTO)
BOWLER 88	PL B209 99	M.G. Bowler	(OXF)
BAND 87	PL B198 297	H.R. Band <i>et al.</i>	(MAC Collab.)
TORNQVIST 87	ZPHY C36 695	N.A. Tornqvist	(HELS)
ALBRECHT 86B	ZPHY C33 7	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
RUCKSTUHL 86	PRL 56 2132	W. Ruckstuhl <i>et al.</i>	(DELCO Collab.)
SCHMIDKE 86	PRL 57 527	W.B. Schmidke <i>et al.</i>	(Mark II Collab.)
BELLINI 85	SJNP 41 781	D. Bellini <i>et al.</i>	
	Translated from YAF 41 1223.		
ZIELINSKI 84C	PRL 52 1195	M. Zielinski <i>et al.</i>	(ROCH, MINN, FNAL)
LONGACRE 82	PR D26 83	R.S. Longacre	(BNL)
DANKOWY... 81	PRL 46 580	J.A. Dankowych <i>et al.</i>	(TNTO, BNL, CARL+)
DAUM 81B	NP B182 269	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+)
DAUM 80	PL 89B 281	C. Daum <i>et al.</i>	(AMST, CERN, CRAC, MPIM+) JP
GAVILLET 77	PL 69B 119	P. Gavillet <i>et al.</i>	(AMST, CERN, NIJM+) JP
BOWLER 75	NP B97 227	M.G. Bowler <i>et al.</i>	(OXFTP, DARE)

————— OTHER RELATED PAPERS —————

DZIERBA	06	PR D73 072001	A.R. Dzierba <i>et al.</i>	(BNL E852 Collab.)
BAKER	03	PL B563 140	C.A. Baker <i>et al.</i>	
CHUNG	02	PR D65 072001	S.U. Chung <i>et al.</i>	(BNL E852 Collab.)
FEUILLAT	01	PL B501 37	M. Feuillat, J.L. Lucio, M.J. Pestieau	
MOLCHANOV	01	PL B521 171	V.V. Molchanov <i>et al.</i>	(FNAL SELEX Collab.)
BAKER	99	PL B449 114	C.A. Baker <i>et al.</i>	
ZAIMIDOROGA	99	PAN 30 1	O.A. Zaimidoroga	
		Translated from SJPN 30 5.		
BARNES	97	PR D55 4157	T. Barnes <i>et al.</i>	(ORNL, RAL, MCHS)
AMELIN	95B	PL B356 595	D.V. Amelin <i>et al.</i>	(SERP, TBIL)
BOLONKIN	95	PAN 58 1535	B.V. Bolonkin <i>et al.</i>	(ITEP)
		Translated from YAF 58 1628.		
WINGATE	95	PRL 74 4596	M. Wingate, T. de Grand	(COLO, FSU)
CONDO	93	PR D48 3045	G.T. Condo <i>et al.</i>	(SLAC Hybrid Collab.)
GOUZ	92	Dallas HEP 92, p. 572	Yu.P. Gouz <i>et al.</i>	(VES Collab.)
		Proceedings XXVI Int. Conf. on High Energy Physics		
IIZUKA	89	PR D39 3357	J. Iizuka, H. Koibuchi, F. Masuda	(NAGO, IBAR+)
BOWLER	86	PL B182 400	M.G. Bowler	(OXF)
BASDEVANT	78	PRL 40 994	J.L. Basdevant, E.L. Berger	(FNAL, ANL) JP
BASDEVANT	77	PR D16 657	J.L. Basdevant, E.L. Berger	(FNAL, ANL) JP
ADERHOLZ	64	PL 10 226	M. Aderholz <i>et al.</i>	(AACH3, BERL, BIRM+)
GOLDHABER	64	PRL 12 336	G. Goldhaber <i>et al.</i>	(LRL, UCB)
LANDER	64	PRL 13 346A	R.L. Lander <i>et al.</i>	(UCSD) JP
BELLINI	63	NC 29 896	G. Bellini <i>et al.</i>	(MILA)