

**$\rho(1700)$** 

$$J^{PC} = 1^{+}(1^{-}-)$$

**THE  $\rho(1450)$  AND THE  $\rho(1700)$** 

Updated April 2006 by S. Eidelman (Novosibirsk) and J.J Hernandez-Rey (Valencia).

In our 1988 edition, we replaced the  $\rho(1600)$  entry with two new ones, the  $\rho(1450)$  and the  $\rho(1700)$ , because there was emerging evidence that the 1600-MeV region actually contains two  $\rho$ -like resonances. ERKAL 86 had pointed out this possibility with a theoretical analysis on the consistency of  $2\pi$  and  $4\pi$  electromagnetic form factors and the  $\pi\pi$  scattering length. DONNACHIE 87, with a full analysis of data on the  $2\pi$  and  $4\pi$  final states in  $e^+e^-$  annihilation and photoproduction reactions, had also argued that in order to obtain a consistent picture, two resonances were necessary. The existence of  $\rho(1450)$  was supported by the analysis of  $\eta\rho^0$  mass spectra obtained in photoproduction and  $e^+e^-$  annihilation (DONNACHIE 87B), as well as that of  $e^+e^- \rightarrow \omega\pi$  (DONNACHIE 91).

The analysis of DONNACHIE 87 was further extended by CLEGG 88, 94 to include new data on  $4\pi$  systems produced in  $e^+e^-$  annihilation, and in  $\tau$  decays ( $\tau$  decays to  $4\pi$  and  $e^+e^-$  annihilation to  $4\pi$  can be related by the Conserved Vector Current assumption). These systems were successfully analyzed using interfering contributions from two  $\rho$ -like states, and from the tail of the  $\rho(770)$  decaying into two-body states. While specific conclusions on  $\rho(1450) \rightarrow 4\pi$  were obtained, little could be said about the  $\rho(1700)$ .

Independent evidence for two  $1^-$  states is provided by KILLIAN 80 in  $4\pi$  electroproduction at  $\langle Q^2 \rangle = 1$  (GeV/c)<sup>2</sup>, and by FUKUI 88 in a high-statistics sample of the  $\eta\pi\pi$  system in  $\pi^-p$  charge exchange.

This scenario with two overlapping resonances is supported by other data. BISELLO 89 measured the pion form factor in the interval 1.35–2.4 GeV and observed a deep minimum around 1.6 GeV. The best fit was obtained with the hypothesis of  $\rho$ -like resonances at 1420 and 1770 MeV, with widths of about 250 MeV. ANTONELLI 88 found that the  $e^+e^- \rightarrow \eta \pi^+ \pi^-$  cross section is better fitted with two fully interfering Breit-Wigners, with parameters in fair agreement with those of DONNACHIE 87 and BISELLO 89. These results can be considered as a confirmation of the  $\rho(1450)$ .

Decisive evidence for the  $\pi\pi$  decay mode of both  $\rho(1450)$  and  $\rho(1700)$  came from recent results in  $\bar{p}p$  annihilation at rest (ABELE 97). It was shown that these resonances also possess a  $K\bar{K}$  decay mode (ABELE 98, BERTIN 98B, ABELE 99D). High statistics studies of the decays  $\tau \rightarrow \pi\pi\nu_\tau$  (BARATE 97M, URHEIM 97), and  $\tau \rightarrow 4\pi\nu_\tau$  (EDWARDS 00A), also require the  $\rho(1450)$ , but are not sensitive to the  $\rho(1700)$ , because it is too close to the  $\tau$  mass. Recently in a very high statistics study of the  $\tau \rightarrow \pi\pi\nu_\tau$  decay performed at Belle (ABE 05H) both  $\rho(1450)$  and  $\rho(1700)$  were observed for the first time in  $\tau$  decays.

The structure of these  $\rho$  states is not yet completely clear. BARNES 97 and CLOSE 97C claim that  $\rho(1450)$  has a mass consistent with radial  $2S$ , but its decays show characteristics of hybrids, and suggest that this state may be a  $2S$ -hybrid mixture. DONNACHIE 99 argues that hybrid states could have a  $4\pi$  decay mode dominated by the  $a_1\pi$ . Such behavior has recently been observed by AKHMETSHIN 99E in  $e^+e^- \rightarrow 4\pi$  in the energy range 1.05–1.38 GeV, and by EDWARDS 00A in  $\tau \rightarrow 4\pi$  decays. ALEXANDER 01B observed the  $\rho(1450) \rightarrow \omega\pi$  decay mode in B-meson decays, however, didn't find  $\rho(1700) \rightarrow \omega\pi^0$ . A similar conclusion is made by AKHMETSHIN 03B who studied

the process  $e^+e^- \rightarrow \omega\pi^0$ . Various decay modes of the  $\rho(1450)$  and  $\rho(1700)$  were observed in  $\bar{p}n$  and  $\bar{p}p$  annihilation (ABELE 01B, BARGIOTTI 03B), but no definite conclusions could be drawn. More data should be collected to clarify the nature of the  $\rho$  states, particularly in the energy range above 1.6 GeV.

We also list under the  $\rho(1450)$  the  $\phi\pi$  state with  $J^{PC} = 1^{--}$  or  $C(1480)$  observed by BITYUKOV 87. While ACHASOV 96B shows that it may be a threshold effect, CLEGG 88 and LANDSBERG 92 suggest two independent vector states with this decay mode. Note, however, that  $C(1480)$  in its  $\phi\pi$  decay mode was not confirmed by  $e^+e^-$  (DOLINSKY 91, BISELLO 91C) and  $\bar{p}p$  (ABELE 97H) experiments.

Several observations on the  $\omega\pi$  system in the 1200-MeV region (FRENKIEL 72, COSME 76, BARBER 80C, ASTON 80C, ATKINSON 84C, BRAU 88, AMSLER 93B) may be interpreted in terms of either  $J^P = 1^-$   $\rho(770) \rightarrow \omega\pi$  production (LAYSSAC 71), or  $J^P = 1^+$   $b_1(1235)$  production (BRAU 88, AMSLER 93B). We argue that no special entry for a  $\rho(1250)$  is needed. The LASS amplitude analysis (ASTON 91B) showing evidence for  $\rho(1270)$  is preliminary and needs confirmation. For completeness, the relevant observations are listed under the  $\rho(1450)$ .

Evidence for  $\rho$ -like mesons decaying into  $6\pi$  states was first noted by CLEGG 90 in the analysis of  $6\pi$  mass spectra from  $e^+e^-$  annihilation (BISELLO 81, CASTRO 88) and diffractive photoproduction (ATKINSON 85). CLEGG 90 argued that two states at about 2.1 and 1.8 GeV exist: while the former is a candidate for a new resonance ( $\rho(2150)$ ), the latter could be a manifestation of the  $\rho(1700)$  distorted by threshold effects. Recently, the E687 Collaboration at Fermilab reported an observation of a narrow dip structure at 1.9 GeV in the  $3\pi^+3\pi^-$  diffractive photoproduction (FRABETTI 01). A similar

effect of the dip in the cross section of  $e^+e^- \rightarrow 6\pi$  around 1.9 GeV has been earlier reported by DM2 (CASTRO 88), where  $6\pi$  included both  $3\pi^+3\pi^-$  and  $2\pi^+2\pi^-2\pi^0$ . Later the dip in the R value (the total cross section of  $e^+e^- \rightarrow$  hadrons divided by the cross section of  $e^+e^- \rightarrow \mu^+\mu^-$ ) was observed by ANTONELLI 96, again around 1.9 GeV. This energy is close to the  $N\bar{N}$  threshold which hints to the possible relation between the dip and  $N\bar{N}$ , e.g., the frequently discussed narrow  $N\bar{N}$  resonance or just a threshold effect. Such behaviour is also characteristic of exotic objects like vector  $q\bar{q}$  hybrids. Note that AGNELLO 02 failed to find this state in the reaction  $\bar{n}p \rightarrow 3\pi^+2\pi^-\pi^0$ . A reanalysis of the E687 data by FRABETTI 04 shows that a dip may arise due to interference of a narrow object with a broad  $\rho(1700)$  independently of the nature of the former. Recently BaBar studied the processes  $e^+e^- \rightarrow 3\pi^+3\pi^-$  and  $e^+e^- \rightarrow 2\pi^+2\pi^-2\pi^0$  using the radiative return and observed a structure around 1.9 GeV in both final states (AUBERT 06D). The data are not well described by a single Breit-Wigner state, and a good fit is achieved while taking into account the interference of such a structure with a Jacob-Slansky amplitude for continuum. The mass of this state obtained by BaBar is consistent with ANTONELLI 96 and FRABETTI 01, but the width is substantially larger. We list these observations under a separate particle  $\rho(1900)$ , which needs confirmation.

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### **$\rho(1700)$ MASS**

#### **$\eta\rho^0$ AND $\pi^+\pi^-$ MODES**

VALUE (MeV)

**$1720 \pm 20$  OUR ESTIMATE**

DOCUMENT ID

**$\eta\rho^0$  MODE**

<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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The data in this block is included in the average printed for a previous datablock.

● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●

1740±20	ANTONELLI	88	DM2	$e^+e^- \rightarrow \eta\pi^+\pi^-$
1701±15	<sup>2</sup> FUKUI	88	SPEC	$8.95\pi^-p \rightarrow \eta\pi^+\pi^-n$

 **$\pi\pi$  MODE**

<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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The data in this block is included in the average printed for a previous datablock.

● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●

1780 $\begin{smallmatrix} +37 \\ -29 \end{smallmatrix}$	<sup>3</sup> ABELE	97	CBAR	$\bar{p}n \rightarrow \pi^-\pi^0\pi^0$
1719 ±15	<sup>3</sup> BERTIN	97C	OBLX	$0.0\bar{p}p \rightarrow \pi^+\pi^-\pi^0$
1730 ±30	CLEGG	94	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
1768 ±21	BISELLO	89	DM2	$e^+e^- \rightarrow \pi^+\pi^-$
1745.7±91.9	DUBNICKA	89	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
1546 ±26	GESHKEN...	89	RVUE	
1650	<sup>4</sup> ERKAL	85	RVUE	$20-70\gamma p \rightarrow \gamma\pi$
1550 ±70	ABE	84B	HYBR	$20\gamma p \rightarrow \pi^+\pi^-p$
1590 ±20	<sup>5</sup> ASTON	80	OMEG	$20-70\gamma p \rightarrow p2\pi$
1600 ±10	<sup>6</sup> ATIYA	79B	SPEC	$50\gamma C \rightarrow C2\pi$
1598 $\begin{smallmatrix} +24 \\ -22 \end{smallmatrix}$	BECKER	79	ASPK	$17\pi^-p$ polarized
1659 ±25	<sup>4</sup> LANG	79	RVUE	
1575	<sup>4</sup> MARTIN	78C	RVUE	$17\pi^-p \rightarrow \pi^+\pi^-n$
1610 ±30	<sup>4</sup> FROGGATT	77	RVUE	$17\pi^-p \rightarrow \pi^+\pi^-n$
1590 ±20	<sup>7</sup> HYAMS	73	ASPK	$17\pi^-p \rightarrow \pi^+\pi^-n$

 **$\pi\omega$  MODE**

<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●

1550 to 1620	<sup>8</sup> ACHASOV	00I	SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
1580 to 1710	<sup>9</sup> ACHASOV	00I	SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
1710±90	ACHASOV	97	RVUE	$e^+e^- \rightarrow \omega\pi^0$

 **$K\bar{K}$  MODE**

<u>VALUE (MeV)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
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● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●

1740.8±22.2	27k	<sup>1</sup> ABELE	99D	CBAR	$\pm 0.0\bar{p}p \rightarrow K^+K^-\pi^0$
1582 ±36	1600	CLELAND	82B	SPEC	$\pm 50\pi p \rightarrow K_S^0 K^\pm p$

<sup>1</sup> K-matrix pole. Isospin not determined, could be  $\omega(1650)$  or  $\phi(1680)$ .

**2( $\pi^+\pi^-$ ) MODE**

<u>VALUE (MeV)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
1851 <sup>+27</sup> <sub>-24</sub>		ACHASOV	97 RVUE	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
1570 ± 20		<sup>10</sup> CORDIER	82 DM1	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
1520 ± 30		<sup>5</sup> ASTON	81E OMEG	20–70 $\gamma p \rightarrow p4\pi$
1654 ± 25		<sup>11</sup> DIBIANCA	81 DBC	$\pi^+d \rightarrow pp2(\pi^+\pi^-)$
1666 ± 39		<sup>10</sup> BACCI	80 FRAG	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
1780	34	KILLIAN	80 SPEC	11 $e^-p \rightarrow 2(\pi^+\pi^-)$
1500		<sup>12</sup> ATIYA	79B SPEC	50 $\gamma C \rightarrow C4\pi^\pm$
1570 ± 60	65	<sup>13</sup> ALEXANDER	75 HBC	7.5 $\gamma p \rightarrow p4\pi$
1550 ± 60		<sup>5</sup> CONVERSI	74 OSPK	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
1550 ± 50	160	SCHACHT	74 STRC	5.5–9 $\gamma p \rightarrow p4\pi$
1450 ± 100	340	SCHACHT	74 STRC	9–18 $\gamma p \rightarrow p4\pi$
1430 ± 50	400	BINGHAM	72B HBC	9.3 $\gamma p \rightarrow p4\pi$

 **$\pi^+\pi^-\pi^0\pi^0$  MODE**

<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
1660 ± 30	ATKINSON	85B OMEG	20–70 $\gamma p$

**3( $\pi^+\pi^-$ ) AND 2( $\pi^+\pi^-\pi^0$ ) MODES**

<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
1730 ± 34	<sup>14</sup> FRABETTI	04 E687	$\gamma p \rightarrow 3\pi^+3\pi^-p$
1783 ± 15	CLEGG	90 RVUE	$e^+e^- \rightarrow 3(\pi^+\pi^-)2(\pi^+\pi^-\pi^0)$

<sup>2</sup> Assuming  $\rho^+ f_0(1370)$  decay mode interferes with  $a_1(1260)^+\pi$  background. From a two Breit-Wigner fit.

<sup>3</sup> T-matrix pole.

<sup>4</sup> From phase shift analysis of HYAMS 73 data.

<sup>5</sup> Simple relativistic Breit-Wigner fit with constant width.

<sup>6</sup> An additional 40 MeV uncertainty in both the mass and width is present due to the choice of the background shape.

<sup>7</sup> Included in BECKER 79 analysis.

<sup>8</sup> Taking into account both  $\rho(1450)$  and  $\rho(1700)$  contributions. Using the data of ACHASOV 00I on  $e^+e^- \rightarrow \omega\pi^0$  and of EDWARDS 00A on  $\tau^- \rightarrow \omega\pi^-\nu_\tau$ .  $\rho(1450)$  mass and width fixed at 1400 MeV and 500 MeV respectively.

<sup>9</sup> Taking into account the  $\rho(1700)$  contribution only. Using the data of ACHASOV 00I on  $e^+e^- \rightarrow \omega\pi^0$  and of EDWARDS 00A on  $\tau^- \rightarrow \omega\pi^-\nu_\tau$ .

<sup>10</sup> Simple relativistic Breit-Wigner fit with model dependent width.

<sup>11</sup> One peak fit result.

<sup>12</sup> Parameters roughly estimated, not from a fit.

<sup>13</sup> Skew mass distribution compensated by Ross-Stodolsky factor.

<sup>14</sup> From a fit with two resonances with the JACOB 72 continuum.

## $\rho(1700)$ WIDTH

### $\eta\rho^0$ AND $\pi^+\pi^-$ MODES

VALUE (MeV)	DOCUMENT ID
<b>250±100 OUR ESTIMATE</b>	

### $\eta\rho^0$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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The data in this block is included in the average printed for a previous datablock.

• • • We do not use the following data for averages, fits, limits, etc. • • •

150±30	ANTONELLI	88	DM2	$e^+e^- \rightarrow \eta\pi^+\pi^-$
282±44	<sup>16</sup> FUKUI	88	SPEC	8.95 $\pi^-p \rightarrow \eta\pi^+\pi^-n$

### $\pi\pi$ MODE

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
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The data in this block is included in the average printed for a previous datablock.

• • • We do not use the following data for averages, fits, limits, etc. • • •

275 ± 45	<sup>17</sup> ABELE	97	CBAR	$\bar{p}n \rightarrow \pi^-\pi^0\pi^0$
310 ± 40	<sup>17</sup> BERTIN	97C	OBLX	0.0 $\bar{p}p \rightarrow \pi^+\pi^-\pi^0$
400 ± 100	CLEGG	94	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
224 ± 22	BISELLO	89	DM2	$e^+e^- \rightarrow \pi^+\pi^-$
242.5±163.0	DUBNICKA	89	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
620 ± 60	GESHKEN...	89	RVUE	
<315	<sup>18</sup> ERKAL	85	RVUE	20-70 $\gamma p \rightarrow \gamma\pi$
280 + 30 - 80	ABE	84B	HYBR	20 $\gamma p \rightarrow \pi^+\pi^-p$
230 ± 80	<sup>19</sup> ASTON	80	OMEG	20-70 $\gamma p \rightarrow p2\pi$
283 ± 14	<sup>20</sup> ATIYA	79B	SPEC	50 $\gamma C \rightarrow C2\pi$
175 + 98 - 53	BECKER	79	ASPK	17 $\pi^-p$ polarized
232 ± 34	<sup>18</sup> LANG	79	RVUE	
340	<sup>18</sup> MARTIN	78C	RVUE	17 $\pi^-p \rightarrow \pi^+\pi^-n$
300 ± 100	<sup>18</sup> FROGGATT	77	RVUE	17 $\pi^-p \rightarrow \pi^+\pi^-n$
180 ± 50	<sup>21</sup> HYAMS	73	ASPK	17 $\pi^-p \rightarrow \pi^+\pi^-n$

### $K\bar{K}$ MODE

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

187.2± 26.7	27k	<sup>15</sup> ABELE	99D	CBAR	± 0.0 $\bar{p}p \rightarrow K^+K^-\pi^0$
265 ± 120	1600	CLELAND	82B	SPEC	± 50 $\pi p \rightarrow K_S^0 K^\pm p$

<sup>15</sup> K-matrix pole. Isospin not determined, could be  $\omega(1650)$  or  $\phi(1680)$ .

**2( $\pi^+\pi^-$ ) MODE**

<u>VALUE (MeV)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
510 ± 40		22 CORDIER	82 DM1	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
400 ± 50		19 ASTON	81E OMEG	20–70 $\gamma p \rightarrow p4\pi$
400 ± 146		23 DIBIANCA	81 DBC	$\pi^+d \rightarrow pp2(\pi^+\pi^-)$
700 ± 160		22 BACCI	80 FRAG	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
100	34	KILLIAN	80 SPEC	11 $e^-p \rightarrow 2(\pi^+\pi^-)$
600		24 ATIYA	79B SPEC	50 $\gamma C \rightarrow C4\pi^\pm$
340 ± 160	65	25 ALEXANDER	75 HBC	7.5 $\gamma p \rightarrow p4\pi$
360 ± 100		19 CONVERSI	74 OSPK	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
400 ± 120	160	26 SCHACHT	74 STRC	5.5–9 $\gamma p \rightarrow p4\pi$
850 ± 200	340	26 SCHACHT	74 STRC	9–18 $\gamma p \rightarrow p4\pi$
650 ± 100	400	BINGHAM	72B HBC	9.3 $\gamma p \rightarrow p4\pi$

 **$\pi^+\pi^-\pi^0\pi^0$  MODE**

<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
300 ± 50	ATKINSON	85B OMEG	20–70 $\gamma p$

 **$\omega\pi^0$  MODE**

<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
350 to 580	27 ACHASOV	00i SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
490 to 1040	28 ACHASOV	00i SND	$e^+e^- \rightarrow \pi^0\pi^0\gamma$

**3( $\pi^+\pi^-$ ) AND 2( $\pi^+\pi^-\pi^0$ ) MODES**

<u>VALUE (MeV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●			
315 ± 100	29 FRABETTI	04 E687	$\gamma p \rightarrow 3\pi^+3\pi^-p$
285 ± 20	CLEGG	90 RVUE	$e^+e^- \rightarrow 3(\pi^+\pi^-)2(\pi^+\pi^-\pi^0)$

<sup>16</sup> Assuming  $\rho^+ f_0(1370)$  decay mode interferes with  $a_1(1260)^+\pi$  background. From a two Breit-Wigner fit.

<sup>17</sup> T-matrix pole.

<sup>18</sup> From phase shift analysis of HYAMS 73 data.

<sup>19</sup> Simple relativistic Breit-Wigner fit with constant width.

<sup>20</sup> An additional 40 MeV uncertainty in both the mass and width is present due to the choice of the background shape.

<sup>21</sup> Included in BECKER 79 analysis.

<sup>22</sup> Simple relativistic Breit-Wigner fit with model-dependent width.

<sup>23</sup> One peak fit result.

<sup>24</sup> Parameters roughly estimated, not from a fit.

<sup>25</sup> Skew mass distribution compensated by Ross-Stodolsky factor.

<sup>26</sup> Width errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

<sup>27</sup> Taking into account both  $\rho(1450)$  and  $\rho(1700)$  contributions. Using the data of ACHASOV 00i on  $e^+e^- \rightarrow \omega\pi^0$  and of EDWARDS 00A on  $\tau^- \rightarrow \omega\pi^-\nu_\tau$ .  $\rho(1450)$  mass and width fixed at 1400 MeV and 500 MeV respectively.

<sup>28</sup> Taking into account the  $\rho(1700)$  contribution only. Using the data of ACHASOV 00i on  $e^+e^- \rightarrow \omega\pi^0$  and of EDWARDS 00A on  $\tau^- \rightarrow \omega\pi^-\nu_\tau$ .

<sup>29</sup> From a fit with two resonances with the JACOB 72 continuum.



## $\rho(1700)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $4\pi$	
$\Gamma_2$ $2(\pi^+\pi^-)$	large
$\Gamma_3$ $\rho\pi\pi$	dominant
$\Gamma_4$ $\rho^0\pi^+\pi^-$	large
$\Gamma_5$ $\rho^0\pi^0\pi^0$	
$\Gamma_6$ $\rho^\pm\pi^\mp\pi^0$	large
$\Gamma_7$ $a_1(1260)\pi$	seen
$\Gamma_8$ $h_1(1170)\pi$	seen
$\Gamma_9$ $\pi(1300)\pi$	seen
$\Gamma_{10}$ $\rho\rho$	seen
$\Gamma_{11}$ $\pi^+\pi^-$	seen
$\Gamma_{12}$ $\pi\pi$	seen
$\Gamma_{13}$ $K\bar{K}^*(892) + \text{c.c.}$	seen
$\Gamma_{14}$ $\eta\rho$	seen
$\Gamma_{15}$ $a_2(1320)\pi$	not seen
$\Gamma_{16}$ $K\bar{K}$	seen
$\Gamma_{17}$ $e^+e^-$	seen
$\Gamma_{18}$ $\pi^0\omega$	seen

### $\rho(1700) \Gamma(i)\Gamma(e^+e^-)/\Gamma(\text{total})$

This combination of a partial width with the partial width into  $e^+e^-$  and with the total width is obtained from the cross-section into channel<sub>i</sub> in  $e^+e^-$  annihilation.

#### $\Gamma(2(\pi^+\pi^-)) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$ $\Gamma_2\Gamma_{17}/\Gamma$

<u>VALUE (keV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
••• We do not use the following data for averages, fits, limits, etc. •••			
2.6 ± 0.2	DELCOURT	81B DM1	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
2.83 ± 0.42	BACCI	80 FRAG	$e^+e^- \rightarrow 2(\pi^+\pi^-)$

#### $\Gamma(\pi^+\pi^-) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$ $\Gamma_{11}\Gamma_{17}/\Gamma$

<u>VALUE (keV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
••• We do not use the following data for averages, fits, limits, etc. •••			
0.13	<sup>30</sup> DIEKMAN	88 RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
0.029 <sup>+0.016</sup> <sub>-0.012</sub>	KURDADZE	83 OLYA	0.64–1.4 $e^+e^- \rightarrow \pi^+\pi^-$

<sup>30</sup> Using total width = 220 MeV.

#### $\Gamma(K\bar{K}^*(892) + \text{c.c.}) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$ $\Gamma_{13}\Gamma_{17}/\Gamma$

<u>VALUE (keV)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
••• We do not use the following data for averages, fits, limits, etc. •••			
0.305 ± 0.071	<sup>31</sup> BIZOT	80 DM1	$e^+e^-$

$\Gamma(\eta\rho) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{14}\Gamma_{17}/\Gamma$

VALUE (eV)	DOCUMENT ID	TECN	COMMENT
$7 \pm 3$	ANTONELLI	88	DM2 $e^+e^- \rightarrow \eta\pi^+\pi^-$

$\Gamma(K\bar{K}) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{16}\Gamma_{17}/\Gamma$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
$0.035 \pm 0.029$	<sup>31</sup> BIZOT	80	DM1 $e^+e^-$

$\Gamma(\rho\pi\pi) \times \Gamma(e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_3\Gamma_{17}/\Gamma$

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
$3.510 \pm 0.090$	<sup>31</sup> BIZOT	80	DM1 $e^+e^-$
<sup>31</sup> Model dependent.			

**$\rho(1700)$  BRANCHING RATIOS**

$\Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{11}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.287^{+0.043}_{-0.042}$	BECKER	79	ASPK $17 \pi^- p$ polarized
0.15 to 0.30	<sup>32</sup> MARTIN	78C	RVUE $17 \pi^- p \rightarrow \pi^+\pi^-n$
<0.20	<sup>33</sup> COSTA...	77B	RVUE $e^+e^- \rightarrow 2\pi, 4\pi$
$0.30 \pm 0.05$	<sup>32</sup> FROGGATT	77	RVUE $17 \pi^- p \rightarrow \pi^+\pi^-n$
<0.15	<sup>34</sup> EISENBERG	73	HBC $5 \pi^+ p \rightarrow \Delta^{++}2\pi$
$0.25 \pm 0.05$	<sup>35</sup> HYAMS	73	ASPK $17 \pi^- p \rightarrow \pi^+\pi^-n$

<sup>32</sup> From phase shift analysis of HYAMS 73 data.

<sup>33</sup> Estimate using unitarity, time reversal invariance, Breit-Wigner.

<sup>34</sup> Estimated using one-pion-exchange model.

<sup>35</sup> Included in BECKER 79 analysis.

$\Gamma(\pi^+\pi^-)/\Gamma(2(\pi^+\pi^-))$   $\Gamma_{11}/\Gamma_2$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.13 \pm 0.05$	ASTON	80	OMEG $20-70 \gamma p \rightarrow p2\pi$
<0.14	<sup>36</sup> DAVIER	73	STRC $6-18 \gamma p \rightarrow p4\pi$
<0.2	<sup>37</sup> BINGHAM	72B	HBC $9.3 \gamma p \rightarrow p2\pi$

<sup>36</sup> Upper limit is estimate.

<sup>37</sup>  $2\sigma$  upper limit.

$\Gamma(\pi\pi)/\Gamma(4\pi)$   $\Gamma_{12}/\Gamma_1$

VALUE	DOCUMENT ID	TECN	COMMENT
$0.16 \pm 0.04$	<sup>42,43</sup> ABELE	01B	CBAR $0.0 \bar{p}n \rightarrow 5\pi$

$\Gamma(K\bar{K}^*(892)+c.c.)/\Gamma(2(\pi^+\pi^-))$   $\Gamma_{13}/\Gamma_2$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.15 \pm 0.03$	<sup>38</sup> DELCOURT	81B DM1	$e^+e^- \rightarrow \bar{K}K\pi$
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<sup>38</sup> Assuming  $\rho(1700)$  and  $\omega$  radial excitations to be degenerate in mass.

 $\Gamma(\eta\rho)/\Gamma_{\text{total}}$   $\Gamma_{14}/\Gamma$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

possibly seen		AKHMETSHIN 00D	CMD2	$e^+e^- \rightarrow \eta\pi^+\pi^-$
$< 0.04$		DONNACHIE	87B RVUE	
$< 0.02$	58	ATKINSON	86B OMEG	20–70 $\gamma p$

 $\Gamma(a_2(1320)\pi)/\Gamma_{\text{total}}$   $\Gamma_{15}/\Gamma$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

not seen	AMELIN	00 VES	37 $\pi^- p \rightarrow \eta\pi^+\pi^- n$
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 $\Gamma(\eta\rho)/\Gamma(2(\pi^+\pi^-))$   $\Gamma_{14}/\Gamma_2$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.123 \pm 0.027$	DELCOURT	82 DM1	$e^+e^- \rightarrow \pi^+\pi^- \text{MM}$
$\sim 0.1$	ASTON	80 OMEG	20–70 $\gamma p$

 $\Gamma(\pi^+\pi^- \text{ neutrals})/\Gamma(2(\pi^+\pi^-))$   $(\Gamma_5+\Gamma_6+0.714\Gamma_{14})/\Gamma_2$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$2.6 \pm 0.4$	<sup>39</sup> BALLAM	74 HBC	9.3 $\gamma p$
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<sup>39</sup> Upper limit. Background not subtracted.

 $\Gamma(\pi^0\omega)/\Gamma_{\text{total}}$   $\Gamma_{18}/\Gamma$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

not seen	2382	AKHMETSHIN 03B	CMD2	$e^+e^- \rightarrow \pi^0\pi^0\gamma$
seen		ACHASOV	97 RVUE	$e^+e^- \rightarrow \omega\pi^0$

 $\Gamma(a_1(1260)\pi)/\Gamma(4\pi)$   $\Gamma_7/\Gamma_1$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.16 \pm 0.05$	<sup>42</sup> ABELE	01B CBAR	$0.0 \bar{p}n \rightarrow 5\pi$
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 $\Gamma(h_1(1170)\pi)/\Gamma(4\pi)$   $\Gamma_8/\Gamma_1$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.17 \pm 0.06$	<sup>42</sup> ABELE	01B CBAR	$0.0 \bar{p}n \rightarrow 5\pi$
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$\Gamma(\pi(1300)\pi)/\Gamma(4\pi)$  $\Gamma_9/\Gamma_1$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.30±0.10	<sup>42</sup> ABELE	01B CBAR	0.0 $\bar{p}n \rightarrow 5\pi$
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 $\Gamma(\rho\rho)/\Gamma(4\pi)$  $\Gamma_{10}/\Gamma_1$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.09±0.03	<sup>42</sup> ABELE	01B CBAR	0.0 $\bar{p}n \rightarrow 5\pi$
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 $\Gamma(\rho\pi\pi)/\Gamma(4\pi)$  $\Gamma_3/\Gamma_1$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.28±0.06	<sup>42</sup> ABELE	01B CBAR	0.0 $\bar{p}n \rightarrow 5\pi$
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 $\Gamma(K\bar{K})/\Gamma(2(\pi^+\pi^-))$  $\Gamma_{16}/\Gamma_2$ 

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.015±0.010		<sup>40</sup> DELCOURT	81B DM1		$e^+e^- \rightarrow \bar{K}K$
<0.04	95	BINGHAM	72B HBC	0	9.3 $\gamma p$

<sup>40</sup> Assuming  $\rho(1700)$  and  $\omega$  radial excitations to be degenerate in mass. $\Gamma(K\bar{K})/\Gamma(K\bar{K}^*(892)+c.c.)$  $\Gamma_{16}/\Gamma_{13}$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.052±0.026	BUON	82 DM1	$e^+e^- \rightarrow$ hadrons
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 $\Gamma(\rho^0\pi^+\pi^-)/\Gamma(2(\pi^+\pi^-))$  $\Gamma_4/\Gamma_2$ 

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

~ 1.0		DELCOURT	81B DM1	$e^+e^- \rightarrow 2(\pi^+\pi^-)$
0.7 ±0.1	500	SCHACHT	74 STRC	5.5–18 $\gamma p \rightarrow p4\pi$
0.80		<sup>41</sup> BINGHAM	72B HBC	9.3 $\gamma p \rightarrow p4\pi$

<sup>41</sup> The  $\pi\pi$  system is in  $S$ -wave. $\Gamma(\rho^0\pi^0\pi^0)/\Gamma(\rho^\pm\pi^\mp\pi^0)$  $\Gamma_5/\Gamma_6$ 

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.10	ATKINSON	85B OMEG		20–70 $\gamma p$
<0.15	ATKINSON	82 OMEG 0		20–70 $\gamma p \rightarrow p4\pi$

<sup>42</sup>  $\omega\pi$  not included.<sup>43</sup> Using ABELE 97.

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