



$$J = \frac{1}{2}$$

$\tau$  discovery paper was PERL 75.  $e^+e^- \rightarrow \tau^+\tau^-$  cross-section threshold behavior and magnitude are consistent with pointlike spin-1/2 Dirac particle. BRANDELIK 78 ruled out pointlike spin-0 or spin-1 particle. FELDMAN 78 ruled out  $J = 3/2$ . KIRKBY 79 also ruled out  $J=\text{integer}$ ,  $J = 3/2$ .

### $\tau$ MASS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1776.99<sup>+0.29</sup><sub>-0.26</sub> OUR AVERAGE</b>				
1775.1 ±1.6 ±1.0	13.3k	<sup>1</sup> ABBIENDI	00A OPAL	1990–1995 LEP runs
1778.2 ±0.8 ±1.2		ANASTASSOV 97	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
1776.96 <sup>+0.18</sup> <sub>-0.21</sub> <sup>+0.25</sup> <sub>-0.17</sub>	65	<sup>2</sup> BAI	96 BES	$E_{\text{cm}}^{ee} = 3.54\text{--}3.57$ GeV
1776.3 ±2.4 ±1.4	11k	<sup>3</sup> ALBRECHT	92M ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
1783 <sup>+3</sup> <sub>-4</sub>	692	<sup>4</sup> BACINO	78B DLCO	$E_{\text{cm}}^{ee} = 3.1\text{--}7.4$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1777.8 ±0.7 ±1.7	35k	<sup>5</sup> BALEST	93 CLEO	Repl. by ANASTASSOV 97
1776.9 <sup>+0.4</sup> <sub>-0.5</sub> ±0.2	14	<sup>6</sup> BAI	92 BES	Repl. by BAI 96

<sup>1</sup> ABBIENDI 00A fit  $\tau$  pseudomass spectrum in  $\tau \rightarrow \pi^\pm \leq 2\pi^0 \nu_\tau$  and  $\tau \rightarrow \pi^\pm \pi^+ \pi^- \leq 1\pi^0 \nu_\tau$  decays. Result assumes  $m_{\nu_\tau} = 0$ .

<sup>2</sup> BAI 96 fit  $\sigma(e^+e^- \rightarrow \tau^+\tau^-)$  at different energies near threshold.

<sup>3</sup> ALBRECHT 92M fit  $\tau$  pseudomass spectrum in  $\tau^- \rightarrow 2\pi^- \pi^+ \nu_\tau$  decays. Result assumes  $m_{\nu_\tau} = 0$ .

<sup>4</sup> BACINO 78B value comes from  $e^\pm X^\mp$  threshold. Published mass 1782 MeV increased by 1 MeV using the high precision  $\psi(2S)$  mass measurement of ZHOLENTZ 80 to eliminate the absolute SPEAR energy calibration uncertainty.

<sup>5</sup> BALEST 93 fit spectra of minimum kinematically allowed  $\tau$  mass in events of the type  $e^+e^- \rightarrow \tau^+\tau^- \rightarrow (\pi^+ n\pi^0 \nu_\tau)(\pi^- m\pi^0 \nu_\tau)$   $n \leq 2, m \leq 2, 1 \leq n+m \leq 3$ . If  $m_{\nu_\tau} \neq 0$ , result increases by  $(m_{\nu_\tau}^2 / 1100 \text{ MeV})$ .

<sup>6</sup> BAI 92 fit  $\sigma(e^+e^- \rightarrow \tau^+\tau^-)$  near threshold using  $e\mu$  events.

$$(m_{\tau^+} - m_{\tau^-})/m_{\text{average}}$$

A test of CPT invariance.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.0 × 10<sup>-3</sup></b>	90	ABBIENDI	00A OPAL	1990–1995 LEP runs

## $\tau$ MEAN LIFE

VALUE ( $10^{-15}$ s)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>290.6<math>\pm</math> 1.1</b>	<b>OUR AVERAGE</b>			
293.2 $\pm$ 2.0 $\pm$ 1.5		ACCIARRI	00B L3	1991–1995 LEP runs
290.1 $\pm$ 1.5 $\pm$ 1.1		BARATE	97R ALEP	1989–1994 LEP runs
291.4 $\pm$ 3.0		ABREU	96B DLPH	1991–1993 LEP runs
289.2 $\pm$ 1.7 $\pm$ 1.2		ALEXANDER	96E OPAL	1990–1994 LEP runs
289.0 $\pm$ 2.8 $\pm$ 4.0	57.4k	BALEST	96 CLEO	$E_{cm}^{ee} = 10.6$ GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
291.2 $\pm$ 2.0 $\pm$ 1.2		BARATE	97I ALEP	Repl. by BARATE 97R
290.1 $\pm$ 4.0	34k	ACCIARRI	96K L3	Repl. by ACCIARRI 00B
297 $\pm$ 9 $\pm$ 5	1671	ABE	95Y SLD	1992–1993 SLC runs
304 $\pm$ 14 $\pm$ 7	4100	BATTLE	92 CLEO	$E_{cm}^{ee} = 10.6$ GeV
301 $\pm$ 29	3780	KLEINWORT	89 JADE	$E_{cm}^{ee} = 35$ –46 GeV
288 $\pm$ 16 $\pm$ 17	807	AMIDEI	88 MRK2	$E_{cm}^{ee} = 29$ GeV
306 $\pm$ 20 $\pm$ 14	695	BRAUNSCH...	88C TASS	$E_{cm}^{ee} = 36$ GeV
299 $\pm$ 15 $\pm$ 10	1311	ABACHI	87C HRS	$E_{cm}^{ee} = 29$ GeV
295 $\pm$ 14 $\pm$ 11	5696	ALBRECHT	87P ARG	$E_{cm}^{ee} = 9.3$ –10.6 GeV
309 $\pm$ 17 $\pm$ 7	3788	BAND	87B MAC	$E_{cm}^{ee} = 29$ GeV
325 $\pm$ 14 $\pm$ 18	8470	BEBEK	87C CLEO	$E_{cm}^{ee} = 10.5$ GeV
460 $\pm$ 190	102	FELDMAN	82 MRK2	$E_{cm}^{ee} = 29$ GeV

## $\tau$ MAGNETIC MOMENT ANOMALY

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

$$\mu_\tau / (e\hbar/2m_\tau) - 1 = (g_\tau - 2)/2$$

For a theoretical calculation [ $(g_\tau - 2)/2 = 11773(3) \times 10^{-7}$ ], see SAMUEL 91B.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&gt; -0.052 and &lt; 0.058 (CL = 95%) OUR LIMIT</b>				
> -0.052 and < 0.058	95	ACCIARRI	98E L3	1991–1995 LEP runs
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
> -0.007 and < 0.005	95	<sup>7</sup> GONZALEZ-S..00	RVUE	$e^+e^- \rightarrow \tau^+\tau^-$ and $W \rightarrow \tau\nu_\tau$
> -0.068 and < 0.065	95	<sup>8</sup> ACKERSTAFF	98N OPAL	1990–1995 LEP runs
> -0.004 and < 0.006	95	<sup>9</sup> ESCRIBANO	97 RVUE	$Z \rightarrow \tau^+\tau^-$ at LEP
< 0.01	95	<sup>10</sup> ESCRIBANO	93 RVUE	$Z \rightarrow \tau^+\tau^-$ at LEP
< 0.12	90	GRIFOLS	91 RVUE	$Z \rightarrow \tau\tau\gamma$ at LEP
< 0.023	95	<sup>11</sup> SILVERMAN	83 RVUE	$e^+e^- \rightarrow \tau^+\tau^-$ at PETRA

<sup>7</sup> GONZALEZ-SPRINBERG 00 use data on tau lepton production at LEP1, SLC, and LEP2, and data from colliders and LEP2 to determine limits. Assume imaginary component is zero.

<sup>8</sup> ACKERSTAFF 98N use  $Z \rightarrow \tau^+\tau^-\gamma$  events. The limit applies to an average of the form factor for off-shell  $\tau$ 's having  $p^2$  ranging from  $m_\tau^2$  to  $(M_Z - m_\tau)^2$ .

<sup>9</sup> ESCRIBANO 97 use preliminary experimental results.

- <sup>10</sup> ESCRIBANO 93 limit derived from  $\Gamma(Z \rightarrow \tau^+ \tau^-)$ , and is on the absolute value of the magnetic moment anomaly.  
<sup>11</sup> SILVERMAN 83 limit is derived from  $e^+ e^- \rightarrow \tau^+ \tau^-$  total cross-section measurements for  $q^2$  up to  $(37 \text{ GeV})^2$ .

## $\tau$ ELECTRIC DIPOLE MOMENT ( $d_\tau$ )

A nonzero value is forbidden by both  $T$  invariance and  $P$  invariance.

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

### Re( $d_\tau$ )

VALUE ( $10^{-16}$ ecm)	CL%	DOCUMENT ID	TECN	COMMENT
<b>-0.22 to 0.45</b>	95	<sup>12</sup> INAMI	03 BELL	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
< 4.6	95	<sup>13</sup> ALBRECHT	00 ARG	$E_{\text{cm}}^{ee} = 10.4 \text{ GeV}$
> -3.1 and < 3.1	95	ACCIARRI	98E L3	1991-1995 LEP
> -3.8 and < 3.6	95	<sup>14</sup> ACKERSTAFF	98N OPAL	1990-1995 LEP runs
< 0.11	95	<sup>15,16</sup> ESCRIBANO	97 RVUE	$Z \rightarrow \tau^+ \tau^-$ at LEP
< 0.5	95	<sup>17</sup> ESCRIBANO	93 RVUE	$Z \rightarrow \tau^+ \tau^-$ at LEP
< 7	90	GRIFOLS	91 RVUE	$Z \rightarrow \tau \tau \gamma$ at LEP
< 1.6	90	DELAGUILA	90 RVUE	$e^+ e^- \rightarrow \tau^+ \tau^-$ $E_{\text{cm}}^{ee} = 35 \text{ GeV}$

<sup>12</sup> INAMI 03 use  $e^+ e^- \rightarrow \tau^+ \tau^-$  events.

<sup>13</sup> ALBRECHT 00 use  $e^+ e^- \rightarrow \tau^+ \tau^-$  events. Limit is on the absolute value of Re( $d_\tau$ ).

<sup>14</sup> ACKERSTAFF 98N use  $Z \rightarrow \tau^+ \tau^- \gamma$  events. The limit applies to an average of the form factor for off-shell  $\tau$ 's having  $p^2$  ranging from  $m_\tau^2$  to  $(M_Z - m_\tau)^2$ .

<sup>15</sup> ESCRIBANO 97 derive the relationship  $|d_\tau| = \cot \theta_W |d_\tau^W|$  using effective Lagrangian methods, and use a conference result  $|d_\tau^W| < 5.8 \times 10^{-18} \text{ ecm}$  at 95% CL (L. Silvestris, ICHEP96) to obtain this result.

<sup>16</sup> ESCRIBANO 97 use preliminary experimental results.

<sup>17</sup> ESCRIBANO 93 limit derived from  $\Gamma(Z \rightarrow \tau^+ \tau^-)$ , and is on the absolute value of the electric dipole moment.

### Im( $d_\tau$ )

VALUE ( $10^{-16}$ ecm)	CL%	DOCUMENT ID	TECN	COMMENT
<b>= 0.25 to 0.008</b>	95	<sup>18</sup> INAMI	03 BELL	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
< 1.8	95	<sup>19</sup> ALBRECHT	00 ARG	$E_{\text{cm}}^{ee} = 10.4 \text{ GeV}$
<sup>18</sup> INAMI 03 use $e^+ e^- \rightarrow \tau^+ \tau^-$ events.				
<sup>19</sup> ALBRECHT 00 use $e^+ e^- \rightarrow \tau^+ \tau^-$ events. Limit is on the absolute value of Im( $d_\tau$ ).				

## $\tau$ WEAK DIPOLE MOMENT ( $d_{\tau}^W$ )

A nonzero value is forbidden by  $CP$  invariance.

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

### Re( $d_{\tau}^W$ )

VALUE ( $10^{-17}$ ecm)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.50</b>	95	20 HEISTER	03F ALEP	1990–1995 LEP runs
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
<3.0	90	20 ACCIARRI	98C L3	1991–1995 LEP runs
<0.56	95	ACKERSTAFF	97L OPAL	1991–1995 LEP runs
<0.78	95	21 AKERS	95F OPAL	Repl. by ACKER-STAFF 97L
<1.5	95	21 BUSKULIC	95C ALEP	Repl. by HEISTER 03F
<7.0	95	21 ACTON	92F OPAL	$Z \rightarrow \tau^+ \tau^-$ at LEP
<3.7	95	21 BUSKULIC	92J ALEP	Repl. by BUSKULIC 95C

<sup>20</sup> Limit is on the absolute value of the real part of the weak dipole moment.

<sup>21</sup> Limit is on the absolute value of the real part of the weak dipole moment, and applies for  $q^2 = m_Z^2$ .

### Im( $d_{\tau}^W$ )

VALUE ( $10^{-17}$ ecm)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.1</b>	95	22 HEISTER	03F ALEP	1990–1995 LEP runs
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
<1.5	95	ACKERSTAFF	97L OPAL	1991–1995 LEP runs
<4.5	95	23 AKERS	95F OPAL	Repl. by ACKER-STAFF 97L

<sup>22</sup> HEISTER 03F limit is on the absolute value of the imaginary part of the weak dipole moment.

<sup>23</sup> Limit is on the absolute value of the imaginary part of the weak dipole moment, and applies for  $q^2 = m_Z^2$ .

## $\tau$ WEAK ANOMALOUS MAGNETIC DIPOLE MOMENT ( $\alpha_{\tau}^W$ )

Electroweak radiative corrections are expected to contribute at the  $10^{-6}$  level. See BERNABEU 95.

The  $q^2$  dependence is expected to be small providing no thresholds are nearby.

### Re( $\alpha_{\tau}^W$ )

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;1.1 <math>\times 10^{-3}</math></b>	95	24 HEISTER	03F ALEP	1990–1995 LEP runs
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
> -0.0024 and < 0.0025	95	25 GONZALEZ-S..00	RVUE	$e^+ e^- \rightarrow \tau^+ \tau^-$ and $W \rightarrow \tau \nu_{\tau}$
<4.5 $\times 10^{-3}$	90	24 ACCIARRI	98C L3	1991–1995 LEP runs

<sup>24</sup> Limit is on the absolute value of the real part of the weak anomalous magnetic dipole moment.

<sup>25</sup> GONZALEZ-SPRINBERG 00 use data on tau lepton production at LEP1, SLC, and LEP2, and data from colliders and LEP2 to determine limits. Assume imaginary component is zero.

### $\text{Im}(\alpha_\tau^W)$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.7 \times 10^{-3}$	95	<sup>26</sup> HEISTER	03F ALEP	1990–1995 LEP runs
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$<9.9 \times 10^{-3}$	90	<sup>26</sup> ACCIARRI	98C L3	1991–1995 LEP runs

<sup>26</sup> Limit is on the absolute value of the imaginary part of the weak anomalous magnetic dipole moment.

### $\tau^-$ DECAY MODES

$\tau^+$  modes are charge conjugates of the modes below. “ $h^\pm$ ” stands for  $\pi^\pm$  or  $K^\pm$ . “ $\ell$ ” stands for e or  $\mu$ . “Neutrals” stands for  $\gamma$ 's and/or  $\pi^0$ 's.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
<b>Modes with one charged particle</b>		
$\Gamma_1$ particle <sup>-</sup> $\geq 0$ neutrals $\geq 0 K^0 \nu_\tau$ (“1-prong”)	(85.35±0.07) %	S=1.1
$\Gamma_2$ particle <sup>-</sup> $\geq 0$ neutrals $\geq 0 K_L^0 \nu_\tau$	(84.72±0.07) %	S=1.1
$\Gamma_3$ $\mu^- \bar{\nu}_\mu \nu_\tau$	[a] (17.36±0.06) %	
$\Gamma_4$ $\mu^- \bar{\nu}_\mu \nu_\tau \gamma$	[b] ( 3.6 ±0.4 ) × 10 <sup>-3</sup>	
$\Gamma_5$ $e^- \bar{\nu}_e \nu_\tau$	[a] (17.84±0.06) %	
$\Gamma_6$ $e^- \bar{\nu}_e \nu_\tau \gamma$	[b] ( 1.75±0.18 ) %	
$\Gamma_7$ $h^- \geq 0 K_L^0 \nu_\tau$	(12.30±0.11) %	S=1.4
$\Gamma_8$ $h^- \nu_\tau$	(11.75±0.11) %	S=1.4
$\Gamma_9$ $\pi^- \nu_\tau$	[a] (11.06±0.11) %	S=1.4
$\Gamma_{10}$ $K^- \nu_\tau$	[a] ( 6.86±0.23 ) × 10 <sup>-3</sup>	
$\Gamma_{11}$ $h^- \geq 1$ neutrals $\nu_\tau$	(36.92±0.14) %	S=1.1
$\Gamma_{12}$ $h^- \pi^0 \nu_\tau$	(25.87±0.13) %	S=1.1
$\Gamma_{13}$ $\pi^- \pi^0 \nu_\tau$	[a] (25.42±0.14) %	S=1.1
$\Gamma_{14}$ $\pi^- \pi^0 \text{non-}\rho(770) \nu_\tau$	( 3.0 ±3.2 ) × 10 <sup>-3</sup>	
$\Gamma_{15}$ $K^- \pi^0 \nu_\tau$	[a] ( 4.50±0.30 ) × 10 <sup>-3</sup>	
$\Gamma_{16}$ $h^- \geq 2 \pi^0 \nu_\tau$	(10.77±0.15) %	S=1.1
$\Gamma_{17}$ $h^- 2 \pi^0 \nu_\tau$	( 9.39±0.14 ) %	S=1.1
$\Gamma_{18}$ $h^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ )	( 9.23±0.14 ) %	S=1.1
$\Gamma_{19}$ $\pi^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ )	[a] ( 9.17±0.14 ) %	S=1.1
$\Gamma_{20}$ $\pi^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ ),	< 9 × 10 <sup>-3</sup>	CL=95%
$\Gamma_{21}$ scalar $\pi^- 2 \pi^0 \nu_\tau$ (ex. $K^0$ ), vector	< 7 × 10 <sup>-3</sup>	CL=95%

$\Gamma_{22}$	$K^- 2\pi^0 \nu_\tau$ (ex. $K^0$ )	[a]	$(5.8 \pm 2.3) \times 10^{-4}$	
$\Gamma_{23}$	$h^- \geq 3\pi^0 \nu_\tau$		$(1.37 \pm 0.11) \%$	S=1.1
$\Gamma_{24}$	$h^- 3\pi^0 \nu_\tau$		$(1.21 \pm 0.10) \%$	
$\Gamma_{25}$	$\pi^- 3\pi^0 \nu_\tau$ (ex. $K^0$ )	[a]	$(1.08 \pm 0.10) \%$	
$\Gamma_{26}$	$K^- 3\pi^0 \nu_\tau$ (ex. $K^0, \eta$ )	[a]	$(3.8^{+2.2}_{-2.0}) \times 10^{-4}$	
$\Gamma_{27}$	$h^- 4\pi^0 \nu_\tau$ (ex. $K^0$ )		$(1.6 \pm 0.6) \times 10^{-3}$	
$\Gamma_{28}$	$h^- 4\pi^0 \nu_\tau$ (ex. $K^0, \eta$ )	[a]	$(1.0^{+0.6}_{-0.5}) \times 10^{-3}$	
$\Gamma_{29}$	$K^- \geq 0\pi^0 \geq 0K^0 \geq 0\gamma \nu_\tau$		$(1.56 \pm 0.04) \%$	
$\Gamma_{30}$	$K^- \geq 1 (\pi^0 \text{ or } K^0 \text{ or } \gamma) \nu_\tau$		$(8.74 \pm 0.35) \times 10^{-3}$	

### Modes with $K^0$ 's

$\Gamma_{31}$	$K_S^0$ (particles) $^- \nu_\tau$		$(9.2 \pm 0.4) \times 10^{-3}$	S=1.1
$\Gamma_{32}$	$h^- \bar{K}^0 \nu_\tau$		$(1.05 \pm 0.04) \%$	S=1.1
$\Gamma_{33}$	$\pi^- \bar{K}^0 \nu_\tau$	[a]	$(8.9 \pm 0.4) \times 10^{-3}$	S=1.1
$\Gamma_{34}$	$\pi^- \bar{K}^0$ (non- $K^*(892)^-$ ) $\nu_\tau$	<	$1.7 \times 10^{-3}$	CL=95%
$\Gamma_{35}$	$K^- K^0 \nu_\tau$	[a]	$(1.54 \pm 0.16) \times 10^{-3}$	
$\Gamma_{36}$	$K^- K^0 \geq 0\pi^0 \nu_\tau$		$(3.09 \pm 0.24) \times 10^{-3}$	
$\Gamma_{37}$	$h^- \bar{K}^0 \pi^0 \nu_\tau$		$(5.2 \pm 0.4) \times 10^{-3}$	
$\Gamma_{38}$	$\pi^- \bar{K}^0 \pi^0 \nu_\tau$	[a]	$(3.7 \pm 0.4) \times 10^{-3}$	
$\Gamma_{39}$	$\bar{K}^0 \rho^- \nu_\tau$		$(2.2 \pm 0.5) \times 10^{-3}$	
$\Gamma_{40}$	$K^- K^0 \pi^0 \nu_\tau$	[a]	$(1.55 \pm 0.20) \times 10^{-3}$	
$\Gamma_{41}$	$\pi^- \bar{K}^0 \geq 1\pi^0 \nu_\tau$		$(3.2 \pm 1.0) \times 10^{-3}$	
$\Gamma_{42}$	$\pi^- \bar{K}^0 \pi^0 \pi^0 \nu_\tau$		$(2.6 \pm 2.4) \times 10^{-4}$	
$\Gamma_{43}$	$K^- K^0 \pi^0 \pi^0 \nu_\tau$	<	$1.6 \times 10^{-4}$	CL=95%
$\Gamma_{44}$	$\pi^- K^0 \bar{K}^0 \nu_\tau$		$(1.59 \pm 0.29) \times 10^{-3}$	S=1.1
$\Gamma_{45}$	$\pi^- K_S^0 K_S^0 \nu_\tau$	[a]	$(2.4 \pm 0.5) \times 10^{-4}$	
$\Gamma_{46}$	$\pi^- K_S^0 K_L^0 \nu_\tau$	[a]	$(1.10 \pm 0.28) \times 10^{-3}$	S=1.1
$\Gamma_{47}$	$\pi^- K^0 \bar{K}^0 \pi^0 \nu_\tau$		$(3.1 \pm 2.3) \times 10^{-4}$	
$\Gamma_{48}$	$\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau$	<	$2.0 \times 10^{-4}$	CL=95%
$\Gamma_{49}$	$\pi^- K_S^0 K_L^0 \pi^0 \nu_\tau$		$(3.1 \pm 1.2) \times 10^{-4}$	
$\Gamma_{50}$	$K^0 h^+ h^- h^- \geq 0$ neutrals $\nu_\tau$	<	$1.7 \times 10^{-3}$	CL=95%
$\Gamma_{51}$	$K^0 h^+ h^- h^- \nu_\tau$		$(2.3 \pm 2.0) \times 10^{-4}$	

### Modes with three charged particles

$\Gamma_{52}$	$h^- h^- h^+ \geq 0$ neutrals $\geq 0K_L^0 \nu_\tau$		$(15.19 \pm 0.07) \%$	S=1.1
$\Gamma_{53}$	$h^- h^- h^+ \geq 0$ neutrals $\nu_\tau$ (ex. $K_S^0 \rightarrow \pi^+ \pi^-$ ) ("3-prong")		$(14.57 \pm 0.07) \%$	S=1.1
$\Gamma_{54}$	$h^- h^- h^+ \nu_\tau$		$(10.01 \pm 0.09) \%$	S=1.2
$\Gamma_{55}$	$h^- h^- h^+ \nu_\tau$ (ex. $K^0$ )		$(9.65 \pm 0.09) \%$	S=1.2
$\Gamma_{56}$	$h^- h^- h^+ \nu_\tau$ (ex. $K^0, \omega$ )		$(9.60 \pm 0.09) \%$	S=1.2
$\Gamma_{57}$	$\pi^- \pi^+ \pi^- \nu_\tau$		$(9.47 \pm 0.10) \%$	S=1.2
$\Gamma_{58}$	$\pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ )		$(9.16 \pm 0.10) \%$	S=1.2

Г59	$\pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ ), non-axial vector	< 2.4	%	CL=95%
Г60	$\pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0, \omega$ )	[a]	( 9.12 ± 0.10 ) %	S=1.2
Г61	$h^- h^- h^+ \geq 1$ neutrals $\nu_\tau$		( 5.19 ± 0.10 ) %	S=1.3
Г62	$h^- h^- h^+ \geq 1$ neutrals $\nu_\tau$ (ex. $K_S^0 \rightarrow \pi^+ \pi^-$ )		( 4.92 ± 0.09 ) %	S=1.3
Г63	$h^- h^- h^+ \pi^0 \nu_\tau$		( 4.53 ± 0.09 ) %	S=1.3
Г64	$h^- h^- h^+ \pi^0 \nu_\tau$ (ex. $K^0$ )		( 4.35 ± 0.09 ) %	S=1.3
Г65	$h^- h^- h^+ \pi^0 \nu_\tau$ (ex. $K^0, \omega$ )		( 2.62 ± 0.09 ) %	S=1.2
Г66	$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$		( 4.37 ± 0.09 ) %	S=1.3
Г67	$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0$ )		( 4.25 ± 0.09 ) %	S=1.3
Г68	$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \omega$ )	[a]	( 2.51 ± 0.09 ) %	S=1.2
Г69	$h^- \rho \pi^0 \nu_\tau$			
Г70	$h^- \rho^+ h^- \nu_\tau$			
Г71	$h^- \rho^- h^+ \nu_\tau$			
Г72	$h^- h^- h^+ 2\pi^0 \nu_\tau$		( 5.5 ± 0.4 ) × 10 <sup>-3</sup>	
Г73	$h^- h^- h^+ 2\pi^0 \nu_\tau$ (ex. $K^0$ )		( 5.4 ± 0.4 ) × 10 <sup>-3</sup>	
Г74	$h^- h^- h^+ 2\pi^0 \nu_\tau$ (ex. $K^0, \omega, \eta$ )	[a]	( 1.1 ± 0.4 ) × 10 <sup>-3</sup>	
Г75	$h^- h^- h^+ 3\pi^0 \nu_\tau$	[a]	( 2.3 ± 0.8 ) × 10 <sup>-4</sup>	S=1.5
Г76	$K^- h^+ h^- \geq 0$ neutrals $\nu_\tau$		( 6.9 ± 0.4 ) × 10 <sup>-3</sup>	S=1.3
Г77	$K^- h^+ \pi^- \nu_\tau$ (ex. $K^0$ )		( 4.8 ± 0.4 ) × 10 <sup>-3</sup>	S=1.5
Г78	$K^- h^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0$ )		( 1.07 ± 0.22 ) × 10 <sup>-3</sup>	
Г79	$K^- \pi^+ \pi^- \geq 0$ neutrals $\nu_\tau$		( 5.0 ± 0.4 ) × 10 <sup>-3</sup>	S=1.3
Г80	$K^- \pi^+ \pi^- \geq 0 \pi^0 \nu_\tau$ (ex. $K^0$ )		( 3.9 ± 0.4 ) × 10 <sup>-3</sup>	S=1.3
Г81	$K^- \pi^+ \pi^- \nu_\tau$		( 3.8 ± 0.4 ) × 10 <sup>-3</sup>	S=1.6
Г82	$K^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ )	[a]	( 3.3 ± 0.4 ) × 10 <sup>-3</sup>	S=1.6
Г83	$K^- \rho^0 \nu_\tau \rightarrow$ $K^- \pi^+ \pi^- \nu_\tau$		( 1.6 ± 0.6 ) × 10 <sup>-3</sup>	
Г84	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$		( 1.18 ± 0.25 ) × 10 <sup>-3</sup>	
Г85	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0$ )		( 6.5 ± 2.4 ) × 10 <sup>-4</sup>	
Г86	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \eta$ )	[a]	( 5.9 ± 2.4 ) × 10 <sup>-4</sup>	
Г87	$K^- \pi^+ K^- \geq 0$ neut. $\nu_\tau$		< 9 × 10 <sup>-4</sup>	CL=95%
Г88	$K^- K^+ \pi^- \geq 0$ neut. $\nu_\tau$		( 1.97 ± 0.18 ) × 10 <sup>-3</sup>	S=1.1
Г89	$K^- K^+ \pi^- \nu_\tau$	[a]	( 1.55 ± 0.07 ) × 10 <sup>-3</sup>	
Г90	$K^- K^+ \pi^- \pi^0 \nu_\tau$	[a]	( 4.2 ± 1.6 ) × 10 <sup>-4</sup>	S=1.1
Г91	$K^- K^+ K^- \geq 0$ neut. $\nu_\tau$		< 2.1 × 10 <sup>-3</sup>	CL=95%
Г92	$K^- K^+ K^- \nu_\tau$		< 3.7 × 10 <sup>-5</sup>	CL=90%
Г93	$\pi^- K^+ \pi^- \geq 0$ neut. $\nu_\tau$		< 2.5 × 10 <sup>-3</sup>	CL=95%
Г94	$e^- e^- e^+ \bar{\nu}_e \nu_\tau$		( 2.8 ± 1.5 ) × 10 <sup>-5</sup>	
Г95	$\mu^- e^- e^+ \bar{\nu}_\mu \nu_\tau$		< 3.6 × 10 <sup>-5</sup>	CL=90%

### Modes with five charged particles

$\Gamma_{96}$	$3h^- 2h^+ \geq 0$ neutrals $\nu_\tau$ (ex. $K_S^0 \rightarrow \pi^- \pi^+$ ) ("5-prong")		$(1.00 \pm 0.06) \times 10^{-3}$	
$\Gamma_{97}$	$3h^- 2h^+ \nu_\tau$ (ex. $K^0$ )	[a]	$(8.2 \pm 0.6) \times 10^{-4}$	
$\Gamma_{98}$	$3h^- 2h^+ \pi^0 \nu_\tau$ (ex. $K^0$ )	[a]	$(1.81 \pm 0.27) \times 10^{-4}$	
$\Gamma_{99}$	$3h^- 2h^+ 2\pi^0 \nu_\tau$		$< 1.1 \times 10^{-4}$	CL=90%

### Miscellaneous other allowed modes

$\Gamma_{100}$	$(5\pi)^- \nu_\tau$		$(8.0 \pm 0.7) \times 10^{-3}$	
$\Gamma_{101}$	$4h^- 3h^+ \geq 0$ neutrals $\nu_\tau$ ("7-prong")		$< 2.4 \times 10^{-6}$	CL=90%
$\Gamma_{102}$	$X^-(S=-1) \nu_\tau$		$(2.91 \pm 0.08) \%$	S=1.1
$\Gamma_{103}$	$K^*(892)^- \geq 0$ neutrals $\geq$ $0K_L^0 \nu_\tau$		$(1.42 \pm 0.18) \%$	S=1.4
$\Gamma_{104}$	$K^*(892)^- \nu_\tau$		$(1.29 \pm 0.05) \%$	
$\Gamma_{105}$	$K^*(892)^0 K^- \geq 0$ neutrals $\nu_\tau$		$(3.2 \pm 1.4) \times 10^{-3}$	
$\Gamma_{106}$	$K^*(892)^0 K^- \nu_\tau$		$(2.1 \pm 0.4) \times 10^{-3}$	
$\Gamma_{107}$	$\bar{K}^*(892)^0 \pi^- \geq 0$ neutrals $\nu_\tau$		$(3.8 \pm 1.7) \times 10^{-3}$	
$\Gamma_{108}$	$\bar{K}^*(892)^0 \pi^- \nu_\tau$		$(2.2 \pm 0.5) \times 10^{-3}$	
$\Gamma_{109}$	$(\bar{K}^*(892)\pi)^- \nu_\tau \rightarrow$ $\pi^- \bar{K}^0 \pi^0 \nu_\tau$		$(1.0 \pm 0.4) \times 10^{-3}$	
$\Gamma_{110}$	$K_1(1270)^- \nu_\tau$		$(4.7 \pm 1.1) \times 10^{-3}$	
$\Gamma_{111}$	$K_1(1400)^- \nu_\tau$		$(1.7 \pm 2.6) \times 10^{-3}$	S=1.7
$\Gamma_{112}$	$K^*(1410)^- \nu_\tau$		$(1.5^{+1.4}_{-1.0}) \times 10^{-3}$	
$\Gamma_{113}$	$K_0^*(1430)^- \nu_\tau$		$< 5 \times 10^{-4}$	CL=95%
$\Gamma_{114}$	$K_2^*(1430)^- \nu_\tau$		$< 3 \times 10^{-3}$	CL=95%
$\Gamma_{115}$	$a_0(980)^- \geq 0$ neutrals $\nu_\tau$			
$\Gamma_{116}$	$\eta \pi^- \nu_\tau$		$< 1.4 \times 10^{-4}$	CL=95%
$\Gamma_{117}$	$\eta \pi^- \pi^0 \nu_\tau$	[a]	$(1.74 \pm 0.24) \times 10^{-3}$	
$\Gamma_{118}$	$\eta \pi^- \pi^0 \pi^0 \nu_\tau$		$(1.5 \pm 0.5) \times 10^{-4}$	
$\Gamma_{119}$	$\eta K^- \nu_\tau$	[a]	$(2.7 \pm 0.6) \times 10^{-4}$	
$\Gamma_{120}$	$\eta K^*(892)^- \nu_\tau$		$(2.9 \pm 0.9) \times 10^{-4}$	
$\Gamma_{121}$	$\eta K^- \pi^0 \nu_\tau$		$(1.8 \pm 0.9) \times 10^{-4}$	
$\Gamma_{122}$	$\eta \bar{K}^0 \pi^- \nu_\tau$		$(2.2 \pm 0.7) \times 10^{-4}$	
$\Gamma_{123}$	$\eta \pi^+ \pi^- \pi^- \geq 0$ neutrals $\nu_\tau$		$< 3 \times 10^{-3}$	CL=90%
$\Gamma_{124}$	$\eta \pi^- \pi^+ \pi^- \nu_\tau$		$(2.3 \pm 0.5) \times 10^{-4}$	
$\Gamma_{125}$	$\eta a_1(1260)^- \nu_\tau \rightarrow \eta \pi^- \rho^0 \nu_\tau$		$< 3.9 \times 10^{-4}$	CL=90%
$\Gamma_{126}$	$\eta \eta \pi^- \nu_\tau$		$< 1.1 \times 10^{-4}$	CL=95%
$\Gamma_{127}$	$\eta \eta \pi^- \pi^0 \nu_\tau$		$< 2.0 \times 10^{-4}$	CL=95%
$\Gamma_{128}$	$\eta'(958) \pi^- \nu_\tau$		$< 7.4 \times 10^{-5}$	CL=90%
$\Gamma_{129}$	$\eta'(958) \pi^- \pi^0 \nu_\tau$		$< 8.0 \times 10^{-5}$	CL=90%
$\Gamma_{130}$	$\phi \pi^- \nu_\tau$		$< 2.0 \times 10^{-4}$	CL=90%



$\Gamma_{131}$	$\phi K^- \nu_\tau$		$< 6.7$	$\times 10^{-5}$	CL=90%
$\Gamma_{132}$	$f_1(1285) \pi^- \nu_\tau$		$(5.8 \pm 2.3)$	$\times 10^{-4}$	
$\Gamma_{133}$	$f_1(1285) \pi^- \nu_\tau \rightarrow$ $\eta \pi^- \pi^+ \pi^- \nu_\tau$		$(1.3 \pm 0.4)$	$\times 10^{-4}$	
$\Gamma_{134}$	$\pi(1300)^- \nu_\tau \rightarrow (\rho \pi)^- \nu_\tau \rightarrow$ $(3\pi)^- \nu_\tau$		$< 1.0$	$\times 10^{-4}$	CL=90%
$\Gamma_{135}$	$\pi(1300)^- \nu_\tau \rightarrow$ $((\pi \pi)_{S\text{-wave}} \pi)^- \nu_\tau \rightarrow$ $(3\pi)^- \nu_\tau$		$< 1.9$	$\times 10^{-4}$	CL=90%
$\Gamma_{136}$	$h^- \omega \geq 0$ neutrals $\nu_\tau$		$(2.38 \pm 0.08)$	%	
$\Gamma_{137}$	$h^- \omega \nu_\tau$	[a]	$(1.94 \pm 0.07)$	%	
$\Gamma_{138}$	$h^- \omega \pi^0 \nu_\tau$	[a]	$(4.4 \pm 0.5)$	$\times 10^{-3}$	
$\Gamma_{139}$	$h^- \omega 2\pi^0 \nu_\tau$		$(1.4 \pm 0.5)$	$\times 10^{-4}$	
$\Gamma_{140}$	$2h^- h^+ \omega \nu_\tau$		$(1.20 \pm 0.22)$	$\times 10^{-4}$	

### Lepton Family number (*LF*), Lepton number (*L*), or Baryon number (*B*) violating modes

*L* means lepton number violation (e.g.  $\tau^- \rightarrow e^+ \pi^- \pi^-$ ). Following common usage, *LF* means lepton family violation *and not* lepton number violation (e.g.  $\tau^- \rightarrow e^- \pi^+ \pi^-$ ). *B* means baryon number violation.

$\Gamma_{141}$	$e^- \gamma$	<i>LF</i>	$< 2.7$	$\times 10^{-6}$	CL=90%
$\Gamma_{142}$	$\mu^- \gamma$	<i>LF</i>	$< 1.1$	$\times 10^{-6}$	CL=90%
$\Gamma_{143}$	$e^- \pi^0$	<i>LF</i>	$< 3.7$	$\times 10^{-6}$	CL=90%
$\Gamma_{144}$	$\mu^- \pi^0$	<i>LF</i>	$< 4.0$	$\times 10^{-6}$	CL=90%
$\Gamma_{145}$	$e^- K_S^0$	<i>LF</i>	$< 9.1$	$\times 10^{-7}$	CL=90%
$\Gamma_{146}$	$\mu^- K_S^0$	<i>LF</i>	$< 9.5$	$\times 10^{-7}$	CL=90%
$\Gamma_{147}$	$e^- \eta$	<i>LF</i>	$< 8.2$	$\times 10^{-6}$	CL=90%
$\Gamma_{148}$	$\mu^- \eta$	<i>LF</i>	$< 9.6$	$\times 10^{-6}$	CL=90%
$\Gamma_{149}$	$e^- \rho^0$	<i>LF</i>	$< 2.0$	$\times 10^{-6}$	CL=90%
$\Gamma_{150}$	$\mu^- \rho^0$	<i>LF</i>	$< 6.3$	$\times 10^{-6}$	CL=90%
$\Gamma_{151}$	$e^- K^*(892)^0$	<i>LF</i>	$< 5.1$	$\times 10^{-6}$	CL=90%
$\Gamma_{152}$	$\mu^- K^*(892)^0$	<i>LF</i>	$< 7.5$	$\times 10^{-6}$	CL=90%
$\Gamma_{153}$	$e^- \bar{K}^*(892)^0$	<i>LF</i>	$< 7.4$	$\times 10^{-6}$	CL=90%
$\Gamma_{154}$	$\mu^- \bar{K}^*(892)^0$	<i>LF</i>	$< 7.5$	$\times 10^{-6}$	CL=90%
$\Gamma_{155}$	$e^- \phi$	<i>LF</i>	$< 6.9$	$\times 10^{-6}$	CL=90%
$\Gamma_{156}$	$\mu^- \phi$	<i>LF</i>	$< 7.0$	$\times 10^{-6}$	CL=90%
$\Gamma_{157}$	$e^- e^+ e^-$	<i>LF</i>	$< 2.9$	$\times 10^{-6}$	CL=90%
$\Gamma_{158}$	$e^- \mu^+ \mu^-$	<i>LF</i>	$< 1.8$	$\times 10^{-6}$	CL=90%
$\Gamma_{159}$	$e^+ \mu^- \mu^-$	<i>LF</i>	$< 1.5$	$\times 10^{-6}$	CL=90%
$\Gamma_{160}$	$\mu^- e^+ e^-$	<i>LF</i>	$< 1.7$	$\times 10^{-6}$	CL=90%
$\Gamma_{161}$	$\mu^+ e^- e^-$	<i>LF</i>	$< 1.5$	$\times 10^{-6}$	CL=90%
$\Gamma_{162}$	$\mu^- \mu^+ \mu^-$	<i>LF</i>	$< 1.9$	$\times 10^{-6}$	CL=90%
$\Gamma_{163}$	$e^- \pi^+ \pi^-$	<i>LF</i>	$< 2.2$	$\times 10^{-6}$	CL=90%
$\Gamma_{164}$	$e^+ \pi^- \pi^-$	<i>L</i>	$< 1.9$	$\times 10^{-6}$	CL=90%

$\Gamma_{165}$	$\mu^- \pi^+ \pi^-$	<i>LF</i>	< 8.2	$\times 10^{-6}$	CL=90%
$\Gamma_{166}$	$\mu^+ \pi^- \pi^-$	<i>L</i>	< 3.4	$\times 10^{-6}$	CL=90%
$\Gamma_{167}$	$e^- \pi^+ K^-$	<i>LF</i>	< 6.4	$\times 10^{-6}$	CL=90%
$\Gamma_{168}$	$e^- \pi^- K^+$	<i>LF</i>	< 3.8	$\times 10^{-6}$	CL=90%
$\Gamma_{169}$	$e^+ \pi^- K^-$	<i>L</i>	< 2.1	$\times 10^{-6}$	CL=90%
$\Gamma_{170}$	$e^- K_S^0 K_S^0$	<i>LF</i>	< 2.2	$\times 10^{-6}$	CL=90%
$\Gamma_{171}$	$e^- K^+ K^-$	<i>LF</i>	< 6.0	$\times 10^{-6}$	CL=90%
$\Gamma_{172}$	$e^+ K^- K^-$	<i>L</i>	< 3.8	$\times 10^{-6}$	CL=90%
$\Gamma_{173}$	$\mu^- \pi^+ K^-$	<i>LF</i>	< 7.5	$\times 10^{-6}$	CL=90%
$\Gamma_{174}$	$\mu^- \pi^- K^+$	<i>LF</i>	< 7.4	$\times 10^{-6}$	CL=90%
$\Gamma_{175}$	$\mu^+ \pi^- K^-$	<i>L</i>	< 7.0	$\times 10^{-6}$	CL=90%
$\Gamma_{176}$	$\mu^- K_S^0 K_S^0$	<i>LF</i>	< 3.4	$\times 10^{-6}$	CL=90%
$\Gamma_{177}$	$\mu^- K^+ K^-$	<i>LF</i>	< 1.5	$\times 10^{-5}$	CL=90%
$\Gamma_{178}$	$\mu^+ K^- K^-$	<i>L</i>	< 6.0	$\times 10^{-6}$	CL=90%
$\Gamma_{179}$	$e^- \pi^0 \pi^0$	<i>LF</i>	< 6.5	$\times 10^{-6}$	CL=90%
$\Gamma_{180}$	$\mu^- \pi^0 \pi^0$	<i>LF</i>	< 1.4	$\times 10^{-5}$	CL=90%
$\Gamma_{181}$	$e^- \eta \eta$	<i>LF</i>	< 3.5	$\times 10^{-5}$	CL=90%
$\Gamma_{182}$	$\mu^- \eta \eta$	<i>LF</i>	< 6.0	$\times 10^{-5}$	CL=90%
$\Gamma_{183}$	$e^- \pi^0 \eta$	<i>LF</i>	< 2.4	$\times 10^{-5}$	CL=90%
$\Gamma_{184}$	$\mu^- \pi^0 \eta$	<i>LF</i>	< 2.2	$\times 10^{-5}$	CL=90%
$\Gamma_{185}$	$\bar{p} \gamma$	<i>L,B</i>	< 3.5	$\times 10^{-6}$	CL=90%
$\Gamma_{186}$	$\bar{p} \pi^0$	<i>L,B</i>	< 1.5	$\times 10^{-5}$	CL=90%
$\Gamma_{187}$	$\bar{p} 2\pi^0$	<i>L,B</i>	< 3.3	$\times 10^{-5}$	CL=90%
$\Gamma_{188}$	$\bar{p} \eta$	<i>L,B</i>	< 8.9	$\times 10^{-6}$	CL=90%
$\Gamma_{189}$	$\bar{p} \pi^0 \eta$	<i>L,B</i>	< 2.7	$\times 10^{-5}$	CL=90%
$\Gamma_{190}$	$e^-$ light boson	<i>LF</i>	< 2.7	$\times 10^{-3}$	CL=95%
$\Gamma_{191}$	$\mu^-$ light boson	<i>LF</i>	< 5	$\times 10^{-3}$	CL=95%

[a] Basis mode for the  $\tau$ .

[b] See the Particle Listings below for the energy limits used in this measurement.

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### CONSTRAINED FIT INFORMATION

An overall fit to 65 branching ratios uses 128 measurements and one constraint to determine 31 parameters. The overall fit has a  $\chi^2 = 62.5$  for 98 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \cdot \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_5$	3										
$x_9$	-5	-5									
$x_{10}$	0	0	-20								
$x_{13}$	-13	-13	-25	1							
$x_{15}$	0	0	1	-3	-22						
$x_{19}$	-13	-14	-25	0	-36	4					
$x_{22}$	0	0	1	-2	5	-21	-16				
$x_{25}$	-8	-8	-15	0	-18	5	-24	3			
$x_{26}$	0	0	1	-2	5	-19	3	-13	-22		
$x_{28}$	-4	-4	-7	0	-10	0	-12	0	-7	0	
$x_{33}$	-1	-2	-13	0	-3	1	-8	1	-1	1	
$x_{35}$	0	0	-4	-1	2	-12	-1	-8	2	-8	
$x_{38}$	-2	-2	-2	0	-5	3	-5	2	-9	2	
$x_{40}$	-1	-1	0	-1	2	-16	1	-11	-1	-10	
$x_{45}$	0	0	-2	0	0	0	0	0	0	0	
$x_{46}$	-2	-2	-2	0	-4	1	-4	1	-2	1	
$x_{60}$	-1	-1	-2	0	-4	0	-3	0	-2	0	
$x_{68}$	-1	-1	-3	0	-3	0	-4	0	-2	0	
$x_{74}$	0	0	0	0	0	0	0	0	0	0	
$x_{75}$	0	0	0	0	0	0	0	0	0	0	
$x_{82}$	0	0	0	0	0	0	0	0	0	0	
$x_{86}$	0	0	0	0	0	0	0	0	0	0	
$x_{89}$	0	0	0	0	0	0	0	0	0	0	
$x_{90}$	0	0	0	0	0	0	0	0	0	0	
$x_{97}$	0	0	-1	0	-1	0	-1	0	0	0	
$x_{98}$	0	0	0	0	0	0	0	0	0	0	
$x_{117}$	-1	-1	-1	0	-2	0	-2	0	-1	0	
$x_{119}$	0	0	0	0	0	-3	0	-2	-2	-2	
$x_{137}$	-1	-1	-2	0	-3	0	-3	0	-1	0	
$x_{138}$	-1	-1	-1	0	-2	0	-2	0	-1	0	
	$x_3$	$x_5$	$x_9$	$x_{10}$	$x_{13}$	$x_{15}$	$x_{19}$	$x_{22}$	$x_{25}$	$x_{26}$	

x33	-1									
x35	0	-5								
x38	-1	-7	0							
x40	0	-2	-15	-19						
x45	0	-2	-1	-2	0					
x46	-1	-12	-4	-10	-3	-3				
x60	-1	-7	-3	3	2	0	0			
x68	-1	4	2	-5	-2	0	0	-47		
x74	2	1	0	1	0	0	0	-8	-7	
x75	0	0	0	0	0	0	0	-3	-3	0
x82	0	-1	0	0	0	0	0	-31	-1	0
x86	0	0	0	0	0	0	0	2	-14	0
x89	0	0	0	0	0	0	0	-3	-1	0
x90	0	0	0	0	0	0	0	-6	-1	-1
x97	0	0	0	0	0	0	0	-1	-1	0
x98	0	0	0	0	0	0	0	0	0	0
x117	-14	0	0	0	0	0	0	-1	-1	-14
x119	0	0	-1	0	-2	0	0	0	0	0
x137	-1	1	1	-3	-1	0	0	-23	-29	-3
x138	-1	1	0	1	0	0	0	-9	-10	-44
	x28	x33	x35	x38	x40	x45	x46	x60	x68	x74

x82	0									
x86	0	-14								
x89	0	-11	1							
x90	0	6	-47	-1						
x97	0	0	0	0	0					
x98	0	0	0	0	0	-19				
x117	0	0	0	0	0	0	0			
x119	0	0	-6	0	0	0	0	0		
x137	-1	-2	2	0	2	0	0	0	0	
x138	-1	-1	0	0	-1	0	0	0	0	-4
	x75	x82	x86	x89	x90	x97	x98	x117	x119	x137

## $\tau$ BRANCHING FRACTIONS

Revised April 2004 by K.G. Hayes (Hillsdale College).

*The constrained fit to  $\tau$  branching fractions:* The Lep-ton Summary Table and the List of  $\tau$ -Decay Modes contain

branching fractions for 109 conventional  $\tau$ -decay modes and upper limits on the branching fractions for 27 other conventional  $\tau$ -decay modes. Of the 109 modes with branching fractions, 79 are derived from a constrained fit to  $\tau$  branching fraction data. The goal of the constrained fit is to make optimal use of the experimental data to determine  $\tau$  branching fractions. For example, the branching fractions for the decay modes  $\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau$  and  $\tau^- \rightarrow \pi^- \pi^+ \pi^- \pi^0 \nu_\tau$  are determined mostly from experimental measurements of the branching fractions for  $\tau^- \rightarrow h^- h^- h^+ \nu_\tau$  and  $\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau$  and recent measurements of exclusive branching fractions for 3-prong modes containing charged kaons and 0 or 1  $\pi^0$ 's.

Branching fractions from the constrained fit are derived from a set of basis modes. The basis modes form an exclusive set whose branching fractions are constrained to sum exactly to one. The set of selected basis modes expands as branching fraction measurements for new  $\tau$ -decay modes are published. The number of basis modes has expanded from 12 in the year 1994 fit to 31 in the 2002 and 2004 fits. The 31 basis modes selected for the 2004 fit are listed in Table 1. See the 1996 edition of this *Review* [1] for a complete description of our notation for naming  $\tau$ -decay modes and the selection of the basis modes. For each edition since the 1996 edition, the changes in the selected basis modes from the previous edition are described in the  $\tau$  Branching Fractions Review.

In selecting the basis modes, assumptions and choices must be made. For example, we assume the decays  $\tau^- \rightarrow \pi^- K^+ \pi^- \geq 0 \pi^0 \nu_\tau$  and  $\tau^- \rightarrow \pi^+ K^- K^- \geq 0 \pi^0 \nu_\tau$  have negligible branching fractions. This is consistent with standard model predictions for  $\tau$  decay, although the experimental limits for these branching fractions are not very stringent. The 95% confidence

**Table 1:** Basis modes for the 2004 fit to  $\tau$  branching fraction data.

$e^- \bar{\nu}_e \nu_\tau$	$K^- K^0 \pi^0 \nu_\tau$
$\mu^- \bar{\nu}_\mu \nu_\tau$	$\pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0, \omega$ )
$\pi^- \nu_\tau$	$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \omega$ )
$\pi^- \pi^0 \nu_\tau$	$K^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ )
$\pi^- 2\pi^0 \nu_\tau$ (ex. $K^0$ )	$K^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \eta$ )
$\pi^- 3\pi^0 \nu_\tau$ (ex. $K^0$ )	$K^- K^+ \pi^- \nu_\tau$
$h^- 4\pi^0 \nu_\tau$ (ex. $K^0, \eta$ )	$K^- K^+ \pi^- \pi^0 \nu_\tau$
$K^- \nu_\tau$	$h^- h^- h^+ 2\pi^0 \nu_\tau$ (ex. $K^0, \omega, \eta$ )
$K^- \pi^0 \nu_\tau$	$h^- h^- h^+ 3\pi^0 \nu_\tau$
$K^- 2\pi^0 \nu_\tau$ (ex. $K^0$ )	$3h^- 2h^+ \nu_\tau$ (ex. $K^0$ )
$K^- 3\pi^0 \nu_\tau$ (ex. $K^0, \eta$ )	$3h^- 2h^+ \pi^0 \nu_\tau$ (ex. $K^0$ )
$\pi^- \bar{K}^0 \nu_\tau$	$h^- \omega \nu_\tau$
$\pi^- \bar{K}^0 \pi^0 \nu_\tau$	$h^- \omega \pi^0 \nu_\tau$
$\pi^- K_S^0 K_S^0 \nu_\tau$	$\eta \pi^- \pi^0 \nu_\tau$
$\pi^- K_S^0 K_L^0 \nu_\tau$	$\eta K^- \nu_\tau$
$K^- K^0 \nu_\tau$	

level upper limits for these branching fractions in the current Listings are  $B(\tau^- \rightarrow \pi^- K^+ \pi^- \geq 0\pi^0 \nu_\tau) < 0.25\%$  and  $B(\tau^- \rightarrow \pi^+ K^- K^- \geq 0\pi^0 \nu_\tau) < 0.09\%$ , values not so different from measured branching fractions for allowed 3-prong modes containing charged kaons. Although our usual goal is to impose as few theoretical constraints as possible so that the world averages and fit results can be used to test the theoretical constraints (*i.e.*, we do not make use of the theoretical constraint from lepton universality on the ratio of the  $\tau$ -leptonic branching fractions  $B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau) / B(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau) = 0.9726$ ), the experimental challenge to identify charged prongs in 3-prong  $\tau$  decays is sufficiently difficult that experimenters have

been forced to make these assumptions when measuring the branching fractions of the allowed decays.

There are several recently measured modes with small but well-measured ( $> 2.5$  sigma from zero) branching fractions [2] which cannot be expressed in terms of the selected basis modes and are therefore left out of the fit:

$$\begin{aligned} B(\tau^- \rightarrow \pi^- K_S^0 K_L^0 \pi^0 \nu_\tau) &= (3.1 \pm 1.2) \times 10^{-4} \\ B(\tau^- \rightarrow h^- \omega \pi^0 \pi^0 \nu_\tau) &= (1.4 \pm 0.5) \times 10^{-4} \\ B(\tau^- \rightarrow 2h^- h^+ \omega \nu_\tau) &= (1.20 \pm 0.22) \times 10^{-4} \end{aligned}$$

plus the  $\eta \rightarrow \gamma\gamma$  and  $\eta \rightarrow \pi^+ \pi^- \gamma$  components of the branching fractions

$$\begin{aligned} B(\tau^- \rightarrow \eta \pi^- \pi^+ \pi^- \nu_\tau) &= (2.3 \pm 0.5) \times 10^{-4} , \\ B(\tau^- \rightarrow \eta \pi^- \pi^0 \pi^0 \nu_\tau) &= (1.5 \pm 0.5) \times 10^{-4} , \\ B(\tau^- \rightarrow \eta \bar{K}^0 \pi^- \nu_\tau) &= (2.2 \pm 0.7) \times 10^{-4} . \end{aligned}$$

The sum of these excluded branching fractions is  $(0.08 \pm 0.01)\%$ . This is near our goal of 0.1% for the internal consistency of the  $\tau$  Listings for this edition, and thus for simplicity we do not include these small branching fraction decay modes in the basis set.

Beginning with the 2002 edition, the fit algorithm has been improved to allow for correlations between branching fraction measurements used in the fit. In this edition, correlations between measurements contained in Refs. [3,4,5,6] have been included. In the  $\tau$  Listings, the correlation coefficients are listed in the footnote for each measurement. Sometimes experimental papers contain correlation coefficients between measurements using only statistical errors without including systematic errors. We usually cannot make use of these correlation coefficients.

The constrained fit has a  $\chi^2$  of 62.5 for 99 degrees of freedom. Only one of the year 2004 basis mode branching

fractions shifted by more than 1 sigma from its 2002 value:  $B(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  changed from  $(0.28 \pm 0.05)\%$  to  $(0.33 \pm 0.04)\%$ .

***Overconsistency of Leptonic Branching Fraction Measurements:*** To minimize the effects of older experiments which often have larger systematic errors and sometimes make assumptions that have later been shown to be invalid, we exclude old measurements in decay modes which contain at least several newer data of much higher precision. As a rule, we exclude those experiments with large errors which together would contribute no more than 5% of the weight in the average. This procedure leaves six measurements for  $B_e \equiv B(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  and five measurements for  $B_\mu \equiv B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)$ . For both  $B_e$  and  $B_\mu$ , the six measurements are considerably more consistent with each other than should be expected from the quoted errors on the individual measurements. The  $\chi^2$  from the calculation of the average of the selected measurements is 0.49 for  $B_e$  and 0.09 for  $B_\mu$ .

## References

1. R.M. Barnett *et al.* (Particle Data Group), *Review of Particle Physics*, Phys. Rev. **D54**, 1 (1996).
2. See the  $\tau$  Listings for references.
3. P. Abreu *et al.* (**DELPHI** Collaboration), Eur. Phys. J. **C20**, 617 (2001).
4. P. Achard *et al.* (**L3** Collaboration), Phys. Lett. **B519**, 189 (2001).
5. A. Anastassov *et al.* (**CLEO** Collaboration), Phys. Rev. **D55**, 2559 (1997) and Phys. Rev. **D58**, 119903 (1998) (erratum).
6. M. Acciarri *et al.* (**L3** Collaboration), Phys. Lett. **B507**, 47 (2001).



## $\tau^-$ BRANCHING RATIOS

$$\Gamma(\text{particle}^- \geq 0 \text{ neutrals} \geq 0 K^0 \nu_\tau (\text{"1-prong"})) / \Gamma_{\text{total}} \quad \Gamma_1 / \Gamma$$

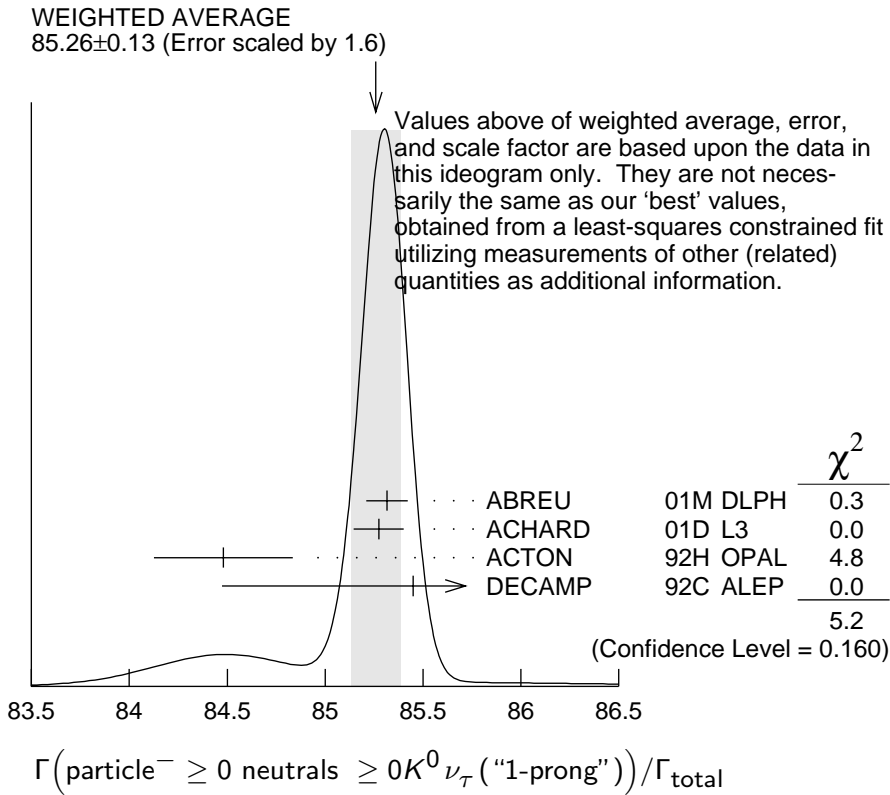
$$\Gamma_1 / \Gamma = (\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10} + \Gamma_{13} + \Gamma_{15} + \Gamma_{19} + \Gamma_{22} + \Gamma_{25} + \Gamma_{26} + \Gamma_{28} + \Gamma_{33} + \Gamma_{35} + \Gamma_{38} + \Gamma_{40} + 2\Gamma_{45} + \Gamma_{46} + 0.708\Gamma_{117} + 0.715\Gamma_{119} + 0.09\Gamma_{137} + 0.09\Gamma_{138}) / \Gamma$$

The charged particle here can be  $e$ ,  $\mu$ , or hadron. In many analyses, the sum of the topological branching fractions (1, 3, and 5 prongs) is constrained to be unity. Since the 5-prong fraction is very small, the measured 1-prong and 3-prong fractions are highly correlated and cannot be treated as independent quantities in our overall fit. We arbitrarily choose to use the 3-prong fraction in our fit, and leave the 1-prong fraction out. We do, however, use these 1-prong measurements in our average below. The measurements used only for the average are marked "avg," whereas "f&a" marks a result used for the fit and the average.

<u>VALUE (%)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>85.35 ± 0.07</b>	<b>OUR FIT</b>	Error includes scale factor of 1.1.		
<b>85.26 ± 0.13</b>	<b>OUR AVERAGE</b>	Error includes scale factor of 1.6. See the ideogram below.		
85.316 ± 0.093 ± 0.049	avg	78k	<sup>27</sup> ABREU	01M DLPH 1992–1995 LEP runs
85.274 ± 0.105 ± 0.073	avg		<sup>28</sup> ACHARD	01D L3 1992–1995 LEP runs
84.48 ± 0.27 ± 0.23	avg		ACTON	92H OPAL 1990–1991 LEP runs
85.45 $\begin{smallmatrix} +0.69 \\ -0.73 \end{smallmatrix}$ ± 0.65	f&a		DECAMP	92C ALEP 1989–1990 LEP runs

<sup>27</sup> The correlation coefficients between this measurement and the ABREU 01M measurements of  $B(\tau \rightarrow 3\text{-prong})$  and  $B(\tau \rightarrow 5\text{-prong})$  are  $-0.98$  and  $-0.08$  respectively.

<sup>28</sup> The correlation coefficients between this measurement and the ACHARD 01D measurements of  $B(\tau \rightarrow 3\text{-prong})$  and  $B(\tau \rightarrow 5\text{-prong})$  are  $-0.978$  and  $-0.082$  respectively.



$\Gamma(\text{particle}^- \geq 0 \text{ neutrals} \geq 0K_L^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_2 / \Gamma$

$$\Gamma_2 / \Gamma = (\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10} + \Gamma_{13} + \Gamma_{15} + \Gamma_{19} + \Gamma_{22} + \Gamma_{25} + \Gamma_{26} + \Gamma_{28} + 0.6569\Gamma_{33} + 0.6569\Gamma_{35} + 0.6569\Gamma_{38} + 0.6569\Gamma_{40} + 1.0985\Gamma_{45} + 0.3139\Gamma_{46} + 0.708\Gamma_{117} + 0.715\Gamma_{119} + 0.09\Gamma_{137} + 0.09\Gamma_{138}) / \Gamma$$

VALUE (%)		EVTS	DOCUMENT ID	TECN	COMMENT
<b>84.72±0.07 OUR FIT</b>					Error includes scale factor of 1.1.
<b>85.1 ±0.4 OUR AVERAGE</b>					
85.6 ±0.6 ±0.3	avg	3300	<sup>29</sup> ADEVA	91F L3	$E_{\text{cm}}^{ee} = 88.3\text{--}94.3 \text{ GeV}$
84.9 ±0.4 ±0.3	avg		BEHREND	89B CELL	$E_{\text{cm}}^{ee} = 14\text{--}47 \text{ GeV}$
84.7 ±0.8 ±0.6	avg		<sup>30</sup> AIHARA	87B TPC	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
86.4 ±0.3 ±0.3			ABACHI	89B HRS	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
87.1 ±1.0 ±0.7			<sup>31</sup> BURCHAT	87 MRK2	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
87.2 ±0.5 ±0.8			SCHMIDKE	86 MRK2	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
84.7 ±1.1 $\begin{smallmatrix} +1.6 \\ -1.3 \end{smallmatrix}$		169	<sup>32</sup> ALTHOFF	85 TASS	$E_{\text{cm}}^{ee} = 34.5 \text{ GeV}$
86.1 ±0.5 ±0.9			BARTEL	85F JADE	$E_{\text{cm}}^{ee} = 34.6 \text{ GeV}$
87.8 ±1.3 ±3.9			<sup>33</sup> BERGER	85 PLUT	$E_{\text{cm}}^{ee} = 34.6 \text{ GeV}$
86.7 ±0.3 ±0.6			FERNANDEZ	85 MAC	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$

<sup>29</sup> Not independent of ADEVA 91F  $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0K_L^0 \nu_\tau) / \Gamma_{\text{total}}$  value.  
<sup>30</sup> Not independent of AIHARA 87B  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma_{\text{total}}$ ,  $\Gamma(e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}$ , and  $\Gamma_{\text{total}}$  values.  
<sup>31</sup> Not independent of SCHMIDKE 86 value (also not independent of BURCHAT 87 value for  $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0K_L^0 \nu_\tau) / \Gamma_{\text{total}}$ ).

<sup>32</sup> Not independent of ALTHOFF 85  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma_{\text{total}}$ ,  $\Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$ ,  $\Gamma_{\text{total}}$ , and  $\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals } \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$  values.

<sup>33</sup> Not independent of (1-prong +  $0\pi^0$ ) and (1-prong +  $\geq 1\pi^0$ ) values.

$\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma_{\text{total}}$

$\Gamma_3/\Gamma$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

To minimize the effect of experiments with large systematic errors, we exclude experiments which together would contribute 5% of the weight in the average.

VALUE (%)		EVTS	DOCUMENT ID	TECN	COMMENT
<b>17.36 ±0.06</b>	<b>OUR FIT</b>				
<b>17.33 ±0.06</b>	<b>OUR AVERAGE</b>				
17.34 ±0.09 ±0.06	f&a	31.4k	ABBIENDI	03 OPAL	1990–1995 LEP runs
17.342±0.110±0.067	f&a	21.5k	<sup>34</sup> ACCIARRI	01F L3	1991–1995 LEP runs
17.325±0.095±0.077	f&a	27.7k	ABREU	99X DLPH	1991–1995 LEP runs
17.37 ±0.08 ±0.18	avg		<sup>35</sup> ANASTASSOV	97 CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
17.31 ±0.11 ±0.05	f&a	20.7k	BUSKULIC	96C ALEP	1991–1993 LEP runs
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
17.02 ±0.19 ±0.24		6586	ABREU	95T DLPH	Repl.. by ABREU 99X
17.36 ±0.27		7941	AKERS	95I OPAL	Repl. by ABBIENDI 03
17.6 ±0.4 ±0.4		2148	ADRIANI	93M L3	Repl. by ACCIARRI 01F
17.4 ±0.3 ±0.5			<sup>36</sup> ALBRECHT	93G ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
17.35 ±0.41 ±0.37	f&a		DECAMP	92C ALEP	1989–1990 LEP runs
17.7 ±0.8 ±0.4		568	BEHREND	90 CELL	$E_{\text{cm}}^{ee} = 35$ GeV
17.4 ±1.0		2197	ADEVA	88 MRKJ	$E_{\text{cm}}^{ee} = 14\text{--}16$ GeV
17.7 ±1.2 ±0.7			AIHARA	87B TPC	$E_{\text{cm}}^{ee} = 29$ GeV
18.3 ±0.9 ±0.8			BURCHAT	87 MRK2	$E_{\text{cm}}^{ee} = 29$ GeV
18.6 ±0.8 ±0.7		558	<sup>37</sup> BARTEL	86D JADE	$E_{\text{cm}}^{ee} = 34.6$ GeV
12.9 ±1.7 $\begin{smallmatrix} +0.7 \\ -0.5 \end{smallmatrix}$			ALTHOFF	85 TASS	$E_{\text{cm}}^{ee} = 34.5$ GeV
18.0 ±0.9 ±0.5		473	<sup>37</sup> ASH	85B MAC	$E_{\text{cm}}^{ee} = 29$ GeV
18.0 ±1.0 ±0.6			<sup>38</sup> BALTRUSAIT..	85 MRK3	$E_{\text{cm}}^{ee} = 3.77$ GeV
19.4 ±1.6 ±1.7		153	BERGER	85 PLUT	$E_{\text{cm}}^{ee} = 34.6$ GeV
17.6 ±2.6 ±2.1		47	BEHREND	83C CELL	$E_{\text{cm}}^{ee} = 34$ GeV
17.8 ±2.0 ±1.8			BERGER	81B PLUT	$E_{\text{cm}}^{ee} = 9\text{--}32$ GeV

<sup>34</sup> The correlation coefficient between this measurement and the ACCIARRI 01F measurement of  $B(\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau)$  is 0.08.

<sup>35</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(e^- \bar{\nu}_e \nu_\tau)$ ,  $B(\mu^- \bar{\nu}_\mu \nu_\tau)/B(e^- \bar{\nu}_e \nu_\tau)$ ,  $B(h^- \nu_\tau)$ , and  $B(h^- \nu_\tau)/B(e^- \bar{\nu}_e \nu_\tau)$  are 0.50, 0.58, 0.50, and 0.08 respectively.

<sup>36</sup> Not independent of ALBRECHT 92D  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$  and ALBRECHT 93G  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \times \Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}^2$  values.

<sup>37</sup> Modified using  $B(e^- \bar{\nu}_e \nu_\tau)/B(\text{"1 prong"})$  and  $B(\text{"1 prong"}) = 0.855$ .

<sup>38</sup> Error correlated with BALTRUSAITIS 85  $e\nu\bar{\nu}$  value.

$\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma)/\Gamma_{\text{total}}$   $\Gamma_4/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.361 ± 0.016 ± 0.035</b>		<sup>39</sup> BERGFELD 00	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.30 ± 0.04 ± 0.05	116	<sup>40</sup> ALEXANDER 96S	OPAL	1991–1994 LEP runs
0.23 ± 0.10	10	<sup>41</sup> WU 90	MRK2	$E_{\text{cm}}^{ee} = 29$ GeV

<sup>39</sup> BERGFELD 00 impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10$  MeV. For  $E_\gamma^* > 20$  MeV, they quote  $(3.04 \pm 0.14 \pm 0.30) \times 10^{-3}$ .

<sup>40</sup> ALEXANDER 96S impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma > 20$  MeV.

<sup>41</sup> WU 90 reports  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma)/\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) = 0.013 \pm 0.006$ , which is converted to  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma)/\Gamma_{\text{total}}$  using  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau \gamma)/\Gamma_{\text{total}} = 17.35\%$ . Requirements on detected  $\gamma$ 's correspond to a  $\tau$  rest frame energy cutoff  $E_\gamma > 37$  MeV.

$\Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_5/\Gamma$

To minimize the effect of experiments with large systematic errors, we exclude experiments which together would contribute 5% of the weight in the average.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>17.84 ± 0.06 OUR FIT</b>				
<b>17.81 ± 0.06 OUR AVERAGE</b>				

17.806 ± 0.104 ± 0.076	24.7k	<sup>42</sup> ACCIARRI 01F	L3	1991–1995 LEP runs
17.81 ± 0.09 ± 0.06	33.1k	ABBIENDI 99H	OPAL	1991–1995 LEP runs
17.877 ± 0.109 ± 0.110	23.3k	ABREU 99X	DLPH	1991–1995 LEP runs
17.76 ± 0.06 ± 0.17		<sup>43</sup> ANASTASSOV 97	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
17.79 ± 0.12 ± 0.06	20.6k	BUSKULIC 96C	ALEP	1991–1993 LEP runs
18.09 ± 0.45 ± 0.45		DECAMP 92C	ALEP	1989–1990 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

17.78 ± 0.10 ± 0.09	25.3k	ALEXANDER 96D	OPAL	Repl. by ABBI- ENDI 99H
17.51 ± 0.23 ± 0.31	5059	ABREU 95T	DLPH	Repl. by ABREU 99X
17.9 ± 0.4 ± 0.4	2892	ADRIANI 93M	L3	Repl. by ACCIA- RRI 01F
17.5 ± 0.3 ± 0.5		<sup>44</sup> ALBRECHT 93G	ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
17.97 ± 0.14 ± 0.23	3970	AKERIB 92	CLEO	Repl. by ANAS- TASSOV 97
19.1 ± 0.4 ± 0.6	2960	<sup>45</sup> AMMAR 92	CLEO	$E_{\text{cm}}^{ee} = 10.5\text{--}10.9$ GeV
17.0 ± 0.5 ± 0.6	1.7k	ABACHI 90	HRS	$E_{\text{cm}}^{ee} = 29$ GeV
18.4 ± 0.8 ± 0.4	644	BEHREND 90	CELL	$E_{\text{cm}}^{ee} = 35$ GeV
16.3 ± 0.3 ± 3.2		JANSSEN 89	CBAL	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
18.4 ± 1.2 ± 1.0		AIHARA 87B	TPC	$E_{\text{cm}}^{ee} = 29$ GeV
19.1 ± 0.8 ± 1.1		BURCHAT 87	MRK2	$E_{\text{cm}}^{ee} = 29$ GeV
16.8 ± 0.7 ± 0.9	515	<sup>45</sup> BARTEL 86D	JADE	$E_{\text{cm}}^{ee} = 34.6$ GeV
20.4 ± 3.0 $\begin{smallmatrix} +1.4 \\ -0.9 \end{smallmatrix}$		ALTHOFF 85	TASS	$E_{\text{cm}}^{ee} = 34.5$ GeV

17.8	$\pm 0.9$	$\pm 0.6$	390	45	ASH	85B	MAC	$E_{cm}^{ee} = 29$ GeV
18.2	$\pm 0.7$	$\pm 0.5$		46	BALTRUSAITIS	85	MRK3	$E_{cm}^{ee} = 3.77$ GeV
13.0	$\pm 1.9$	$\pm 2.9$			BERGER	85	PLUT	$E_{cm}^{ee} = 34.6$ GeV
18.3	$\pm 2.4$	$\pm 1.9$	60		BEHREND	83C	CELL	$E_{cm}^{ee} = 34$ GeV
16.0	$\pm 1.3$		459	47	BACINO	78B	DLCO	$E_{cm}^{ee} = 3.1-7.4$ GeV

<sup>42</sup> The correlation coefficient between this measurement and the ACCIARRI 01F measurement of  $B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)$  is 0.08.

<sup>43</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu \bar{\nu}_\mu \nu_\tau)$ ,  $B(\mu \bar{\nu}_\mu \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$ ,  $B(h^- \nu_\tau)$ , and  $B(h^- \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$  are 0.50, -0.42, 0.48, and -0.39 respectively.

<sup>44</sup> Not independent of ALBRECHT 92D  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$  and ALBRECHT 93G  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \times \Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{total}^2$  values.

<sup>45</sup> Modified using  $B(e^- \bar{\nu}_e \nu_\tau)/B(\text{"1 prong"})$  and  $B(\text{"1 prong"})$ , = 0.855.

<sup>46</sup> Error correlated with BALTRUSAITIS 85  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma_{total}$ .

<sup>47</sup> BACINO 78B value comes from fit to events with  $e^\pm$  and one other nonelectron charged prong.

**$\Gamma(e^- \bar{\nu}_e \nu_\tau \gamma)/\Gamma_{total}$   $\Gamma_6/\Gamma$**

<u>VALUE (%)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.75<math>\pm</math>0.06<math>\pm</math>0.17</b>	48 BERGFELD	00	CLEO $E_{cm}^{ee} = 10.6$ GeV

<sup>48</sup> BERGFELD 00 impose requirements on detected  $\gamma$ 's corresponding to a  $\tau$ -rest-frame energy cutoff  $E_\gamma^* > 10$  MeV.

**$\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma(e^- \bar{\nu}_e \nu_\tau)$   $\Gamma_3/\Gamma_5$**

Standard Model prediction including mass effects is 0.9726.

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**0.974  $\pm$  0.004 OUR FIT**  
**0.978  $\pm$  0.011 OUR AVERAGE**

0.9777 $\pm$ 0.0063 $\pm$ 0.0087	f&a	49	ANASTASSOV 97 CLEO $E_{cm}^{ee} = 10.6$ GeV
0.997 $\pm$ 0.035 $\pm$ 0.040	f&a	ALBRECHT 92D	ARG $E_{cm}^{ee} = 9.4-10.6$ GeV

<sup>49</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu \bar{\nu}_\mu \nu_\tau)$ ,  $B(e \bar{\nu}_e \nu_\tau)$ ,  $B(h^- \nu_\tau)$ , and  $B(h^- \nu_\tau)/B(e \bar{\nu}_e \nu_\tau)$  are 0.58, -0.42, 0.07, and 0.45 respectively.

**$\Gamma(h^- \geq 0K_L^0 \nu_\tau)/\Gamma_{total}$   $\Gamma_7/\Gamma$**

$$\Gamma_7/\Gamma = (\Gamma_9 + \Gamma_{10} + \frac{1}{2}\Gamma_{33} + \frac{1}{2}\Gamma_{35} + \Gamma_{45})/\Gamma$$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

<u>VALUE (%)</u>	<u>EVTs</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**12.30 $\pm$ 0.11 OUR FIT** Error includes scale factor of 1.4.

**12.44 $\pm$ 0.14 OUR AVERAGE**

12.44 $\pm$ 0.11 $\pm$ 0.11	f&a	15k	50	BUSKULIC 96 ALEP 1991-1993 LEP run
12.47 $\pm$ 0.26 $\pm$ 0.43	f&a	2967	51	ACCIARRI 95 L3 1992 LEP run
12.4 $\pm$ 0.7 $\pm$ 0.7	f&a	283	52	ABREU 92N DLPH 1990 LEP run

12.98 ± 0.44 ± 0.33	f&a		<sup>53</sup> DECAMP	92C ALEP	1989–1990 LEP runs
12.1 ± 0.7 ± 0.5	f&a	309	ALEXANDER	91D OPAL	1990 LEP run
11.3 ± 0.5 ± 0.8	avg	798	<sup>54</sup> FORD	87 MAC	$E_{cm}^{ee} = 29$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

11.7 ± 0.6 ± 0.8			<sup>55</sup> ALBRECHT	92D ARG	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV
12.3 ± 0.9 ± 0.5		1338	BEHREND	90 CELL	$E_{cm}^{ee} = 35$ GeV
11.1 ± 1.1 ± 1.4			<sup>56</sup> BURCHAT	87 MRK2	$E_{cm}^{ee} = 29$ GeV
12.3 ± 0.6 ± 1.1		328	<sup>57</sup> BARTEL	86D JADE	$E_{cm}^{ee} = 34.6$ GeV
13.0 ± 2.0 ± 4.0			BERGER	85 PLUT	$E_{cm}^{ee} = 34.6$ GeV
11.2 ± 1.7 ± 1.2		34	<sup>58</sup> BEHREND	83C CELL	$E_{cm}^{ee} = 34$ GeV

<sup>50</sup> BUSKULIC 96 quote  $11.78 \pm 0.11 \pm 0.13$  We add 0.66 to undo their correction for unseen  $K_L^0$  and modify the systematic error accordingly.

<sup>51</sup> ACCIARRI 95 with 0.65% added to remove their correction for  $\pi^- K_L^0$  backgrounds.

<sup>52</sup> ABREU 92N with 0.5% added to remove their correction for  $K^*(892)^-$  backgrounds.

<sup>53</sup> DECAMP 92C quote  $B(h^- \geq 0 K_L^0 \geq 0 (K_S^0 \rightarrow \pi^+ \pi^-) \nu_\tau) = 13.32 \pm 0.44 \pm 0.33$ .

We subtract 0.35 to correct for their inclusion of the  $K_S^0$  decays.

<sup>54</sup> FORD 87 result for  $B(\pi^- \nu_\tau)$  with 0.67% added to remove their  $K^-$  correction and adjusted for 1992 B(“1 prong”).

<sup>55</sup> Not independent of ALBRECHT 92D  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) / \Gamma(e^- \bar{\nu}_e \nu_\tau)$ ,  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \times \Gamma(e^- \bar{\nu}_e \nu_\tau)$ , and  $\Gamma(h^- \geq 0 K_L^0 \nu_\tau) / \Gamma(e^- \bar{\nu}_e \nu_\tau)$  values.

<sup>56</sup> BURCHAT 87 with 1.1% added to remove their correction for  $K^-$  and  $K^*(892)^-$  backgrounds.

<sup>57</sup> BARTEL 86D result for  $B(\pi^- \nu_\tau)$  with 0.59% added to remove their  $K^-$  correction and adjusted for 1992 B(“1 prong”).

<sup>58</sup> BEHREND 83C quote  $B(\pi^- \nu_\tau) = 9.9 \pm 1.7 \pm 1.3$  after subtracting  $1.3 \pm 0.5$  to correct for  $B(K^- \nu_\tau)$ .

### $\Gamma(h^- \nu_\tau) / \Gamma_{\text{total}}$

$$\Gamma_8 / \Gamma = (\Gamma_9 + \Gamma_{10}) / \Gamma$$

Data marked “avg” are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. “f&a” marks results used for the fit and the average.

<u>VALUE (%)</u>		<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>11.75 ± 0.11 OUR FIT</b>				Error includes scale factor of 1.4.
<b>11.65 ± 0.21 OUR AVERAGE</b>				Error includes scale factor of 1.9.
11.98 ± 0.13 ± 0.16	f&a	ACKERSTAFF	98M OPAL	1991–1995 LEP runs
11.52 ± 0.05 ± 0.12	f&a	<sup>59</sup> ANASTASSOV	97 CLEO	$E_{cm}^{ee} = 10.6$ GeV

<sup>59</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu^- \bar{\nu}_\mu \nu_\tau)$ ,  $B(e^- \bar{\nu}_e \nu_\tau)$ ,  $B(\mu^- \bar{\nu}_\mu \nu_\tau) / B(e^- \bar{\nu}_e \nu_\tau)$ , and  $B(h^- \nu_\tau) / B(e^- \bar{\nu}_e \nu_\tau)$  are 0.50, 0.48, 0.07, and 0.63 respectively.

### $\Gamma(h^- \nu_\tau) / \Gamma(e^- \bar{\nu}_e \nu_\tau)$

$$\Gamma_8 / \Gamma_5 = (\Gamma_9 + \Gamma_{10}) / \Gamma_5$$

Data marked “avg” are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. “f&a” marks results used for the fit and the average.

<u>VALUE</u>		<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.659 ± 0.007 OUR FIT</b>				Error includes scale factor of 1.4.
<b>0.6484 ± 0.0041 ± 0.0060 avg</b>		<sup>60</sup> ANASTASSOV	97 CLEO	$E_{cm}^{ee} = 10.6$ GeV

<sup>60</sup> The correlation coefficients between this measurement and the ANASTASSOV 97 measurements of  $B(\mu\bar{\nu}_\mu\nu_\tau)$ ,  $B(e\bar{\nu}_e\nu_\tau)$ ,  $B(\mu\bar{\nu}_\mu\nu_\tau)/B(e\bar{\nu}_e\nu_\tau)$ , and  $B(h^-\nu_\tau)$  are 0.08, -0.39, 0.45, and 0.63 respectively.

### $\Gamma(\pi^-\nu_\tau)/\Gamma_{\text{total}}$ $\Gamma_9/\Gamma$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**11.06±0.11 OUR FIT** Error includes scale factor of 1.4.

**11.07±0.18 OUR AVERAGE**

11.06±0.11±0.14	avg	<sup>61</sup> BUSKULIC	96 ALEP	LEP 1991–1993 data
11.7 ±0.4 ±1.8	f&a	1138 BLOCKER	82D MRK2	$E_{\text{cm}}^{ee} = 3.5\text{--}6.7$ GeV

<sup>61</sup> Not independent of BUSKULIC 96  $B(h^-\nu_\tau)$  and  $B(K^-\nu_\tau)$  values.

### $\Gamma(K^-\nu_\tau)/\Gamma_{\text{total}}$ $\Gamma_{10}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.686±0.023 OUR FIT**

**0.685±0.023 OUR AVERAGE**

0.658±0.027±0.029		<sup>62</sup> ABBIENDI	01J OPAL	1990–1995 LEP runs
0.696±0.025±0.014	2032	BARATE	99K ALEP	1991–1995 LEP runs
0.85 ±0.18	27	ABREU	94K DLPH	LEP 1992 Z data
0.66 ±0.07 ±0.09	99	BATTLE	94 CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.72 ±0.04 ±0.04	728	BUSKULIC	96 ALEP	Repl. by BARATE 99K
0.59 ±0.18	16	MILLS	84 DLCO	$E_{\text{cm}}^{ee} = 29$ GeV
1.3 ±0.5	15	BLOCKER	82B MRK2	$E_{\text{cm}}^{ee} = 3.9\text{--}6.7$ GeV

<sup>62</sup> The correlation coefficient between this measurement and the ABBIENDI 01J  $B(\tau^- \rightarrow K^- \geq 0\pi^0 \geq 0K^0 \geq 0\gamma\nu_\tau)$  is 0.60.

### $\Gamma(h^- \geq 1 \text{ neutrals}\nu_\tau)/\Gamma_{\text{total}}$ $\Gamma_{11}/\Gamma$

$\Gamma_{11}/\Gamma = (\Gamma_{13} + \Gamma_{15} + \Gamma_{19} + \Gamma_{22} + \Gamma_{25} + \Gamma_{26} + \Gamma_{28} + 0.157\Gamma_{33} + 0.157\Gamma_{35} + 0.157\Gamma_{38} + 0.157\Gamma_{40} + 0.0985\Gamma_{45} + 0.708\Gamma_{117} + 0.715\Gamma_{119} + 0.09\Gamma_{137} + 0.09\Gamma_{138})/\Gamma$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
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**36.92±0.14 OUR FIT** Error includes scale factor of 1.1.

• • • We do not use the following data for averages, fits, limits, etc. • • •

36.14±0.33±0.58	<sup>63</sup> AKERS	94E OPAL	1991–1992 LEP runs
38.4 ±1.2 ±1.0	<sup>64</sup> BURCHAT	87 MRK2	$E_{\text{cm}}^{ee} = 29$ GeV
42.7 ±2.0 ±2.9	BERGER	85 PLUT	$E_{\text{cm}}^{ee} = 34.6$ GeV

<sup>63</sup> Not independent of ACKERSTAFF 98M  $B(h^-\pi^0\nu_\tau)$  and  $B(h^-\geq 2\pi^0\nu_\tau)$  values.

<sup>64</sup> BURCHAT 87 quote for  $B(\pi^\pm \geq 1 \text{ neutral}\nu_\tau) = 0.378 \pm 0.012 \pm 0.010$ . We add 0.006 to account for contribution from  $(K^{*-}\nu_\tau)$  which they fixed at BR = 0.013.

$\Gamma(h^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{12}/\Gamma = (\Gamma_{13} + \Gamma_{15})/\Gamma$

VALUE (%)      EVTS      DOCUMENT ID      TECN      COMMENT

**25.87±0.13 OUR FIT** Error includes scale factor of 1.1.

**25.76±0.15 OUR AVERAGE**

25.89±0.17±0.29		ACKERSTAFF	98M	OPAL	1991–1995 LEP runs
25.76±0.15±0.13	31k	BUSKULIC	96	ALEP	LEP 1991–1993 data
25.05±0.35±0.50	6613	ACCIARRI	95	L3	1992 LEP run
25.87±0.12±0.42	51k	<sup>65</sup> ARTUSO	94	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

25.98±0.36±0.52		<sup>66</sup> AKERS	94E	OPAL	Repl. by ACKER-STAFF 98M
22.9 ±0.8 ±1.3	283	<sup>67</sup> ABREU	92N	DLPH	$E_{\text{cm}}^{ee} = 88.2\text{--}94.2$ GeV
23.1 ±0.4 ±0.9	1249	<sup>68</sup> ALBRECHT	92Q	ARG	$E_{\text{cm}}^{ee} = 10$ GeV
25.02±0.64±0.88	1849	DECAMP	92C	ALEP	1989–1990 LEP runs
22.0 ±0.8 ±1.9	779	ANTREASYAN	91	CBAL	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
22.6 ±1.5 ±0.7	1101	BEHREND	90	CELL	$E_{\text{cm}}^{ee} = 35$ GeV
23.1 ±1.9 ±1.6		BEHREND	84	CELL	$E_{\text{cm}}^{ee} = 14,22$ GeV

<sup>65</sup>ARTUSO 94 reports the combined result from three independent methods, one of which (23% of the  $\tau^- \rightarrow h^- \pi^0 \nu_\tau$ ) is normalized to the inclusive one-prong branching fraction, taken as  $0.854 \pm 0.004$ . Renormalization to the present value causes negligible change.

<sup>66</sup>AKERS 94E quote  $(26.25 \pm 0.36 \pm 0.52) \times 10^{-2}$ ; we subtract 0.27% from their number to correct for  $\tau^- \rightarrow h^- K_L^0 \nu_\tau$ .

<sup>67</sup>ABREU 92N with 0.5% added to remove their correction for  $K^*(892)^-$  backgrounds.

<sup>68</sup>ALBRECHT 92Q with 0.5% added to remove their correction for  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  background.

$\Gamma(\pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{13}/\Gamma$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%)      EVTS      DOCUMENT ID      TECN      COMMENT

**25.42±0.14 OUR FIT** Error includes scale factor of 1.1.

**25.31±0.18 OUR AVERAGE**

25.30±0.15±0.13	avg	<sup>69</sup> BUSKULIC	96	ALEP	LEP 1991–1993 data
25.36±0.44	avg	<sup>70</sup> ARTUSO	94	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

21.5 ±0.4 ±1.9	4400	<sup>71,72</sup> ALBRECHT	88L	ARG	$E_{\text{cm}}^{ee} = 10$ GeV
23.0 ±1.3 ±1.7	582	ADLER	87B	MRK3	$E_{\text{cm}}^{ee} = 3.77$ GeV
25.8 ±1.7 ±2.5		<sup>73</sup> BURCHAT	87	MRK2	$E_{\text{cm}}^{ee} = 29$ GeV
22.3 ±0.6 ±1.4	629	<sup>72</sup> YELTON	86	MRK2	$E_{\text{cm}}^{ee} = 29$ GeV

<sup>69</sup>Not independent of BUSKULIC 96 B( $h^- \pi^0 \nu_\tau$ ) and B( $K^- \pi^0 \nu_\tau$ ) values.

<sup>70</sup>Not independent of ARTUSO 94 B( $h^- \pi^0 \nu_\tau$ ) and BATTLE 94 B( $K^- \pi^0 \nu_\tau$ ) values.

<sup>71</sup>The authors divide by  $(\Gamma_3 + \Gamma_5 + \Gamma_9 + \Gamma_{10})/\Gamma = 0.467$  to obtain this result.

<sup>72</sup>Experiment had no hadron identification. Kaon corrections were made, but insufficient information is given to permit their removal.

<sup>73</sup>BURCHAT 87 value is not independent of YELTON 86 value. Nonresonant decays included.



$\Gamma(\pi^- \pi^0 \text{non-}\rho(770) \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{14} / \Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.3 ± 0.1 ± 0.3</b>	74	BEHREND	84	CELL $E_{\text{cm}}^{ee} = 14,22 \text{ GeV}$

<sup>74</sup> BEHREND 84 assume a flat nonresonant mass distribution down to the  $\rho(770)$  mass, using events with mass above 1300 to set the level.

$\Gamma(K^- \pi^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{15} / \Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.450 ± 0.030 OUR FIT**

**0.449 ± 0.034 OUR AVERAGE**

0.444 ± 0.026 ± 0.024	923	BARATE	99K	ALEP 1991–1995 LEP runs
0.51 ± 0.10 ± 0.07	37	BATTLE	94	CLEO $E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.52 ± 0.04 ± 0.05	395	BUSKULIC	96	ALEP Repl. by BARATE 99K
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$\Gamma(h^- \geq 2\pi^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{16} / \Gamma$

$$\Gamma_{16} / \Gamma = (\Gamma_{19} + \Gamma_{22} + \Gamma_{25} + \Gamma_{26} + \Gamma_{28} + 0.157\Gamma_{33} + 0.157\Gamma_{35} + 0.157\Gamma_{38} + 0.157\Gamma_{40} + 0.0985\Gamma_{45} + 0.319\Gamma_{117} + 0.322\Gamma_{119}) / \Gamma$$

Data marked “avg” are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. “f&a” marks results used for the fit and the average.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**10.77 ± 0.15 OUR FIT** Error includes scale factor of 1.1.

**10.0 ± 0.4 OUR AVERAGE**

9.91 ± 0.31 ± 0.27 f&a		ACKERSTAFF	98M	OPAL 1991–1995 LEP runs
12.0 ± 1.4 ± 2.5 f&a	<sup>75</sup>	BURCHAT	87	MRK2 $E_{\text{cm}}^{ee} = 29 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

9.89 ± 0.34 ± 0.55	<sup>76</sup>	AKERS	94E	OPAL Repl. by ACKER-STAFF 98M
14.0 ± 1.2 ± 0.6	938	<sup>77</sup> BEHREND	90	CELL $E_{\text{cm}}^{ee} = 35 \text{ GeV}$
13.9 ± 2.0 $\begin{smallmatrix} +1.9 \\ -2.2 \end{smallmatrix}$	<sup>78</sup>	AIHARA	86E	TPC $E_{\text{cm}}^{ee} = 29 \text{ GeV}$

<sup>75</sup> Error correlated with BURCHAT 87  $\Gamma(\rho^- \nu_e) / \Gamma(\text{total})$  value.

<sup>76</sup> AKERS 94E not independent of AKERS 94E  $B(h^- \geq 1\pi^0 \nu_\tau)$  and  $B(h^- \pi^0 \nu_\tau)$  measurements.

<sup>77</sup> No independent of BEHREND 90  $\Gamma(h^- 2\pi^0 \nu_\tau (\text{exp. } K^0))$  and  $\Gamma(h^- \geq 3\pi^0 \nu_\tau)$ .

<sup>78</sup> AIHARA 86E (TPC) quote  $B(2\pi^0 \pi^- \nu_\tau) + 1.6B(3\pi^0 \pi^- \nu_\tau) + 1.1B(\pi^0 \eta \pi^- \nu_\tau)$ .

$\Gamma(h^- 2\pi^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{17} / \Gamma$

$$\Gamma_{17} / \Gamma = (\Gamma_{19} + \Gamma_{22} + 0.157\Gamma_{33} + 0.157\Gamma_{35}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**9.39 ± 0.14 OUR FIT** Error includes scale factor of 1.1.

**9.48 ± 0.13 ± 0.10** <sup>79</sup> BUSKULIC 96 ALEP LEP 1991–1993 data

<sup>79</sup> BUSKULIC 96 quote  $9.29 \pm 0.13 \pm 0.10$ . We add 0.19 to undo their correction for  $\tau^- \rightarrow h^- K^0 \nu_\tau$ .

$\Gamma(h^- 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   $\Gamma_{18}/\Gamma$   
 $\Gamma_{18}/\Gamma = (\Gamma_{19} + \Gamma_{22})/\Gamma$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. f&a marks results used for the fit and the average.

VALUE (%)		EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.23±0.14 OUR FIT</b>	Error includes scale factor of 1.1.				
<b>9.08±0.34 OUR AVERAGE</b>					
8.88±0.37±0.42	f&a	1060	ACCIARRI	95 L3	1992 LEP run
8.96±0.16±0.44	avg		<sup>80</sup> PROCARIO	93 CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV
10.38±0.66±0.82	f&a	809	<sup>81</sup> DECAMP	92C ALEP	1989–1990 LEP runs
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
5.7 ±0.5 <sup>+1.7</sup> / <sub>-1.0</sub>		133	<sup>82</sup> ANTREASYAN	91 CBAL	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
10.0 ±1.5 ±1.1		333	<sup>83</sup> BEHREND	90 CELL	$E_{\text{cm}}^{ee} = 35$ GeV
8.7 ±0.4 ±1.1		815	<sup>84</sup> BAND	87 MAC	$E_{\text{cm}}^{ee} = 29$ GeV
6.2 ±0.6 ±1.2			<sup>85</sup> GAN	87 MRK2	$E_{\text{cm}}^{ee} = 29$ GeV
6.0 ±3.0 ±1.8			BEHREND	84 CELL	$E_{\text{cm}}^{ee} = 14,22$ GeV

<sup>80</sup> PROCARIO 93 entry is obtained from  $B(h^- 2\pi^0 \nu_\tau)/B(h^- \pi^0 \nu_\tau)$  using ARTUSO 94 result for  $B(h^- \pi^0 \nu_\tau)$ .

<sup>81</sup> We subtract 0.0015 to account for  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  contribution.

<sup>82</sup> ANTREASYAN 91 subtract 0.001 to account for the  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  contribution.

<sup>83</sup> BEHREND 90 subtract 0.002 to account for the  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  contribution.

<sup>84</sup> BAND 87 assume  $B(\pi^- 3\pi^0 \nu_\tau) = 0.01$  and  $B(\pi^- \pi^0 \eta \nu_\tau) = 0.005$ .

<sup>85</sup> GAN 87 analysis use photon multiplicity distribution.

$\Gamma(h^- 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma(h^- \pi^0 \nu_\tau)$   $\Gamma_{18}/\Gamma_{12}$   
 $\Gamma_{18}/\Gamma_{12} = (\Gamma_{19} + \Gamma_{22})/(\Gamma_{13} + \Gamma_{15})$

VALUE		DOCUMENT ID	TECN	COMMENT
<b>0.357±0.006 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>0.342±0.006±0.016</b>		<sup>86</sup> PROCARIO	93 CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV

<sup>86</sup> PROCARIO 93 quote  $0.345 \pm 0.006 \pm 0.016$  after correction for 2 kaon backgrounds assuming  $B(K^{*-} \nu_\tau) = 1.42 \pm 0.18\%$  and  $B(h^- K^0 \pi^0 \nu_\tau) = 0.48 \pm 0.48\%$ . We multiply by  $0.990 \pm 0.010$  to remove these corrections to  $B(h^- \pi^0 \nu_\tau)$ .

$\Gamma(\pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   $\Gamma_{19}/\Gamma$   
 Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%)		DOCUMENT ID	TECN	COMMENT
<b>9.17±0.14 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>9.21±0.13±0.11</b>	avg	<sup>87</sup> BUSKULIC	96 ALEP	LEP 1991–1993 data

<sup>87</sup> Not independent of BUSKULIC 96  $B(h^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$  and  $B(K^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$  values.

$\Gamma(\pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0), \text{ scalar})/\Gamma(\pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$   $\Gamma_{20}/\Gamma_{19}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.094</b>	95	<sup>88</sup> BROWDER	00 CLEO	$4.7 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

<sup>88</sup> Model-independent limit from structure function analysis on contribution to  $B(\tau^- \rightarrow \pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$  from scalars.

$\Gamma(\pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0), \text{ vector})/\Gamma(\pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$   $\Gamma_{21}/\Gamma_{19}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.073</b>	95	<sup>89</sup> BROWDER	00 CLEO	$4.7 \text{ fb}^{-1} E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

<sup>89</sup> Model-independent limit from structure function analysis on contribution to  $B(\tau^- \rightarrow \pi^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$  from vectors.

$\Gamma(K^- 2\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   $\Gamma_{22}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.058±0.023 OUR FIT**

**0.058±0.024 OUR AVERAGE**

0.056±0.020±0.015    131    BARATE    99K ALEP    1991–1995 LEP runs

0.09 ±0.10 ±0.03    3    <sup>90</sup> BATTLE    94 CLEO     $E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.08 ±0.02 ±0.02    59    BUSKULIC    96 ALEP    Repl. by BARATE 99K

<sup>90</sup> BATTLE 94 quote  $0.14 \pm 0.10 \pm 0.03$  or  $< 0.3\%$  at 90% CL. We subtract  $(0.05 \pm 0.02)\%$  to account for  $\tau^- \rightarrow K^- (K^0 \rightarrow \pi^0 \pi^0) \nu_\tau$  background.

$\Gamma(h^- \geq 3\pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{23}/\Gamma$

$$\Gamma_{23}/\Gamma = (\Gamma_{25} + \Gamma_{26} + \Gamma_{28} + 0.157\Gamma_{38} + 0.157\Gamma_{40} + 0.0985\Gamma_{45} + 0.319\Gamma_{117} + 0.322\Gamma_{119})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.37±0.11 OUR FIT** Error includes scale factor of 1.1.

**1.53±0.40±0.46**    186    DECAMP    92C ALEP    1989–1990 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

3.2 ±1.0 ±1.0    BEHREND    90 CELL     $E_{\text{cm}}^{ee} = 35 \text{ GeV}$

$\Gamma(h^- 3\pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{24}/\Gamma$

$$\Gamma_{24}/\Gamma = (\Gamma_{25} + \Gamma_{26} + 0.157\Gamma_{38} + 0.157\Gamma_{40} + 0.322\Gamma_{119})/\Gamma$$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.21±0.10 OUR FIT**

**1.22±0.10 OUR AVERAGE**

1.24±0.09±0.11    f&a    2.3k    <sup>91</sup> BUSKULIC    96 ALEP    LEP 1991–1993 data

1.70±0.24±0.38    f&a    293    ACCIARRI    95 L3    1992 LEP run

1.15±0.08±0.13    avg    <sup>92</sup> PROCARIO    93 CLEO     $E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.0 <sup>+1.4</sup> <sub>-0.1</sub> <sup>+1.1</sup> <sub>-0.1</sub>    <sup>93</sup> GAN    87 MRK2     $E_{\text{cm}}^{ee} = 29 \text{ GeV}$

<sup>91</sup> BUSKULIC 96 quote  $B(h^- 3\pi^0 \nu_\tau (\text{ex. } K^0)) = 1.17 \pm 0.09 \pm 0.11$ . We add 0.07 to remove their correction for  $K^0$  backgrounds.

<sup>92</sup> PROCARIO 93 entry is obtained from  $B(h^- 3\pi^0 \nu_\tau)/B(h^- \pi^0 \nu_\tau)$  using ARTUSO 94 result for  $B(h^- \pi^0 \nu_\tau)$ .

<sup>93</sup> Highly correlated with GAN 87  $\Gamma(\eta \pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$  value. Authors quote  $B(\pi^\pm 3\pi^0 \nu_\tau) + 0.67B(\pi^\pm \eta \pi^0 \nu_\tau) = 0.047 \pm 0.010 \pm 0.011$ .

$$\frac{\Gamma(h^- 3\pi^0 \nu_\tau)/\Gamma(h^- \pi^0 \nu_\tau)}{\Gamma_{24}/\Gamma_{12}} = \frac{(\Gamma_{25} + \Gamma_{26} + 0.157\Gamma_{38} + 0.157\Gamma_{40} + 0.322\Gamma_{119})/(\Gamma_{13} + \Gamma_{15})}{\Gamma_{24}/\Gamma_{12}}$$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.047±0.004 OUR FIT</b>			Error includes scale factor of 1.1.
<b>0.044±0.003±0.005</b>	<sup>94</sup> PROCARIO 93	CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV

<sup>94</sup> PROCARIO 93 quote  $0.041 \pm 0.003 \pm 0.005$  after correction for 2 kaon backgrounds assuming  $B(K^{*-} \nu_\tau) = 1.42 \pm 0.18\%$  and  $B(h^- K^0 \pi^0 \nu_\tau) = 0.48 \pm 0.48\%$ . We add  $0.003 \pm 0.003$  and multiply the sum by  $0.990 \pm 0.010$  to remove these corrections.

$$\frac{\Gamma(\pi^- 3\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}}{\Gamma_{25}/\Gamma}$$

VALUE (%)	DOCUMENT ID
<b>1.08±0.10 OUR FIT</b>	

$$\frac{\Gamma(K^- 3\pi^0 \nu_\tau (\text{ex. } K^0, \eta))/\Gamma_{\text{total}}}{\Gamma_{26}/\Gamma}$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.038<sup>+0.022</sup><sub>-0.020</sub> OUR FIT</b>				
<b>0.037±0.021±0.011</b>	22	BARATE	99K ALEP	1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •  
 $0.05 \pm 0.13$  <sup>95</sup> BUSKULIC 94E ALEP Repl. by BARATE 99K  
<sup>95</sup> BUSKULIC 94E quote  $B(K^- \geq 0\pi^0 \geq 0K^0 \nu_\tau) - [B(K^- \nu_\tau) + B(K^- \pi^0 \nu_\tau) + B(K^- K^0 \nu_\tau) + B(K^- \pi^0 \pi^0 \nu_\tau) + B(K^- \pi^0 K^0 \nu_\tau)] = 0.05 \pm 0.13\%$  accounting for common systematic errors in BUSKULIC 94E and BUSKULIC 94F measurements of these modes. We assume  $B(K^- \geq 2K^0 \nu_\tau)$  and  $B(K^- \geq 4\pi^0 \nu_\tau)$  are negligible.

$$\frac{\Gamma(h^- 4\pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}}{\Gamma_{27}/\Gamma} = \frac{\Gamma_{27}}{\Gamma_{28} + 0.319\Gamma_{117}}$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.16±0.06 OUR FIT</b>				
<b>0.16±0.06 OUR AVERAGE</b>				
$0.16 \pm 0.04 \pm 0.09$	232	<sup>96</sup> BUSKULIC 96	ALEP	LEP 1991–1993 data
$0.16 \pm 0.05 \pm 0.05$		<sup>97</sup> PROCARIO 93	CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV

<sup>96</sup> BUSKULIC 96 quote result for  $\tau^- \rightarrow h^- \geq 4\pi^0 \nu_\tau$ . We assume  $B(h^- \geq 5\pi^0 \nu_\tau)$  is negligible.

<sup>97</sup> PROCARIO 93 quotes  $B(h^- 4\pi^0 \nu_\tau)/B(h^- \pi^0 \nu_\tau) = 0.006 \pm 0.002 \pm 0.002$ . We multiply by the ARTUSO 94 result for  $B(h^- \pi^0 \nu_\tau)$  to obtain  $B(h^- 4\pi^0 \nu_\tau)$ . PROCARIO 93 assume  $B(h^- \geq 5\pi^0 \nu_\tau)$  is small and do not correct for it.

$\Gamma(h^- 4\pi^0 \nu_\tau (\text{ex. } K^0, \eta))/\Gamma_{\text{total}}$   $\Gamma_{28}/\Gamma$

VALUE (%) \_\_\_\_\_ DOCUMENT ID \_\_\_\_\_

**0.10<sup>+0.06</sup><sub>-0.05</sub> OUR FIT**

$\Gamma(K^- \geq 0\pi^0 \geq 0K^0 \geq 0\gamma \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{29}/\Gamma$

$$\Gamma_{29}/\Gamma = (\Gamma_{10} + \Gamma_{15} + \Gamma_{22} + \Gamma_{26} + \Gamma_{35} + \Gamma_{40} + 0.715\Gamma_{119})/\Gamma$$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%) \_\_\_\_\_ EVTS \_\_\_\_\_ DOCUMENT ID \_\_\_\_\_ TECN \_\_\_\_\_ COMMENT \_\_\_\_\_

**1.56 ± 0.04 OUR FIT**

**1.53 ± 0.04 OUR AVERAGE**

1.528 ± 0.039 ± 0.040	f&a		<sup>98</sup> ABBIENDI	01J OPAL	1990–1995 LEP runs
1.520 ± 0.040 ± 0.041	avg	4006	<sup>99</sup> BARATE	99K ALEP	1991–1995 LEP runs
1.54 ± 0.24	f&a		ABREU	94K DLPH	LEP 1992 Z data
1.70 ± 0.12 ± 0.19	f&a	202	<sup>100</sup> BATTLE	94 CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV
1.6 ± 0.4 ± 0.2	f&a	35	AIHARA	87B TPC	$E_{\text{cm}}^{ee} = 29$ GeV
1.71 ± 0.29	f&a	53	MILLS	84 DLCO	$E_{\text{cm}}^{ee} = 29$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.70 ± 0.05 ± 0.06		1610	<sup>101</sup> BUSKULIC	96 ALEP	Repl. by BARATE 99K
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<sup>98</sup> The correlation coefficient between this measurement and the ABBIENDI 01J  $B(\tau^- \rightarrow K^- \nu_\tau)$  is 0.60.

<sup>99</sup> Not independent of BARATE 99K  $B(K^- \nu_\tau)$ ,  $B(K^- \pi^0 \nu_\tau)$ ,  $B(K^- 2\pi^0 \nu_\tau (\text{ex. } K^0))$ ,  $B(K^- 3\pi^0 \nu_\tau (\text{ex. } K^0))$ ,  $B(K^- K^0 \nu_\tau)$ , and  $B(K^- K^0 \pi^0 \nu_\tau)$  values.

<sup>100</sup> BATTLE 94 quote  $1.60 \pm 0.12 \pm 0.19$ . We add  $0.10 \pm 0.02$  to correct for their rejection of  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

<sup>101</sup> Not independent of BUSKULIC 96  $B(K^- \nu_\tau)$ ,  $B(K^- \pi^0 \nu_\tau)$ ,  $B(K^- 2\pi^0 \nu_\tau)$ ,  $B(K^- K^0 \nu_\tau)$ , and  $B(K^- K^0 \pi^0 \nu_\tau)$  values.

$\Gamma(K^- \geq 1(\pi^0 \text{ or } K^0 \text{ or } \gamma) \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{30}/\Gamma$

$$\Gamma_{30}/\Gamma = (\Gamma_{15} + \Gamma_{22} + \Gamma_{26} + \Gamma_{35} + \Gamma_{40} + 0.715\Gamma_{119})/\Gamma$$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%) \_\_\_\_\_ EVTS \_\_\_\_\_ DOCUMENT ID \_\_\_\_\_ TECN \_\_\_\_\_ COMMENT \_\_\_\_\_

**0.874 ± 0.035 OUR FIT**

**0.86 ± 0.05 OUR AVERAGE**

0.869 ± 0.031 ± 0.034	avg		<sup>102</sup> ABBIENDI	01J OPAL	1990–1995 LEP runs
0.69 ± 0.25	avg		<sup>103</sup> ABREU	94K DLPH	LEP 1992 Z data
1.2 ± 0.5 <sup>+0.2</sup> <sub>-0.4</sub>	f&a	9	AIHARA	87B TPC	$E_{\text{cm}}^{ee} = 29$ GeV

<sup>102</sup> Not independent of ABBIENDI 01J  $B(\tau^- \rightarrow K^- \nu_\tau)$  and  $B(\tau^- \rightarrow K^- \geq 0\pi^0 \geq 0K^0 \geq 0\gamma \nu_\tau)$  values.

<sup>103</sup> Not independent of ABREU 94K  $B(K^- \nu_\tau)$  and  $B(K^- \geq 0 \text{ neutrals } \nu_\tau)$  measurements.

$\Gamma(K_S^0(\text{particles})^- \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{31}/\Gamma$

$$\Gamma_{31}/\Gamma = (\frac{1}{2}\Gamma_{33} + \frac{1}{2}\Gamma_{35} + \frac{1}{2}\Gamma_{38} + \frac{1}{2}\Gamma_{40} + \Gamma_{45} + \Gamma_{46})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.92 ± 0.04 OUR FIT</b>				Error includes scale factor of 1.1.
<b>0.97 ± 0.07 OUR AVERAGE</b>				
0.970 ± 0.058 ± 0.062	929	BARATE	98E ALEP	1991–1995 LEP runs
0.97 ± 0.09 ± 0.06	141	AKERS	94G OPAL	$E_{\text{cm}}^{ee} = 88\text{--}94$ GeV

$\Gamma(h^- \bar{K}^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{32}/\Gamma = (\Gamma_{33} + \Gamma_{35})/\Gamma$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.05 ± 0.04 OUR FIT</b>				Error includes scale factor of 1.1.
<b>0.90 ± 0.07 OUR AVERAGE</b>				
1.01 ± 0.11 ± 0.07 avg	555	<sup>104</sup> BARATE	98E ALEP	1991–1995 LEP runs
0.855 ± 0.036 ± 0.073 f&a	1242	COAN	96 CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV

<sup>104</sup> Not independent of BARATE 98E  $B(\tau^- \rightarrow \pi^- \bar{K}^0 \nu_\tau)$  and  $B(\tau^- \rightarrow K^- K^0 \nu_\tau)$  values.

$\Gamma(\pi^- \bar{K}^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{33}/\Gamma$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.89 ± 0.04 OUR FIT</b>				Error includes scale factor of 1.1.
<b>0.88 ± 0.05 OUR AVERAGE</b>				Error includes scale factor of 1.2.
0.933 ± 0.068 ± 0.049 f&a	377	ABBIENDI	00C OPAL	1991–1995 LEP runs
0.928 ± 0.045 ± 0.034 f&a	937	<sup>105</sup> BARATE	99K ALEP	1991–1995 LEP runs
0.855 ± 0.117 ± 0.066 avg	509	<sup>106</sup> BARATE	98E ALEP	1991–1995 LEP runs
0.704 ± 0.041 ± 0.072 avg		<sup>107</sup> COAN	96 CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV
0.95 ± 0.15 ± 0.06 f&a		<sup>108</sup> ACCIARRI	95F L3	1991–1993 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.79 ± 0.10 ± 0.09 98 <sup>109</sup> BUSKULIC 96 ALEP Repl. by BARATE 99K

<sup>105</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

<sup>106</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays. Not independent of BARATE 98E  $B(K^0 \text{ particles}^- \nu_\tau)$  value.

<sup>107</sup> Not independent of COAN 96  $B(h^- K^0 \nu_\tau)$  and  $B(K^- K^0 \nu_\tau)$  measurements.

<sup>108</sup> ACCIARRI 95F do not identify  $\pi^-/K^-$  and assume  $B(K^- K^0 \nu_\tau) = (0.29 \pm 0.12)\%$ .

<sup>109</sup> BUSKULIC 96 measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

$\Gamma(\pi^- \bar{K}^0(\text{non-}K^*(892)^-) \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{34}/\Gamma$

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.17</b>	95	ACCIARRI	95F L3	1991–1993 LEP runs

$\Gamma(K^- K^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{35}/\Gamma$

VALUE (%)      EVTS      DOCUMENT ID      TECN      COMMENT

**0.154±0.016 OUR FIT**

**0.158±0.017 OUR AVERAGE**

0.162±0.021±0.011      150      <sup>110</sup> BARATE      99K ALEP      1991–1995 LEP runs

0.158±0.042±0.017      46      <sup>111</sup> BARATE      98E ALEP      1991–1995 LEP runs

0.151±0.021±0.022      111      COAN      96 CLEO       $E_{\text{cm}}^{ee} \approx 10.6$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.26 ±0.09 ±0.02      13      <sup>112</sup> BUSKULIC      96 ALEP      Repl. by BARATE 99K

<sup>110</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

<sup>111</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

<sup>112</sup> BUSKULIC 96 measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

$\Gamma(K^- K^0 \geq 0\pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{36}/\Gamma = (\Gamma_{35} + \Gamma_{40})/\Gamma$

VALUE (%)      EVTS      DOCUMENT ID      TECN      COMMENT

**0.309±0.024 OUR FIT**

**0.330±0.055±0.039**      124      ABBIENDI      00C OPAL      1991–1995 LEP runs

$\Gamma(h^- \bar{K}^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{37}/\Gamma = (\Gamma_{38} + \Gamma_{40})/\Gamma$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%)      EVTS      DOCUMENT ID      TECN      COMMENT

**0.52 ±0.04 OUR FIT**

**0.50 ±0.06 OUR AVERAGE**      Error includes scale factor of 1.2.

0.446±0.052±0.046      avg      157      <sup>113</sup> BARATE      98E ALEP      1991–1995 LEP runs

0.562±0.050±0.048      f&a      264      COAN      96 CLEO       $E_{\text{cm}}^{ee} \approx 10.6$  GeV

<sup>113</sup> Not independent of BARATE 98E  $B(\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau)$  and  $B(\tau^- \rightarrow K^- K^0 \pi^0 \nu_\tau)$  values.

$\Gamma(\pi^- \bar{K}^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{38}/\Gamma$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%)      EVTS      DOCUMENT ID      TECN      COMMENT

**0.37 ±0.04 OUR FIT**

**0.36 ±0.04 OUR AVERAGE**

0.347±0.053±0.037      f&a      299      <sup>114</sup> BARATE      99K ALEP      1991–1995 LEP runs

0.294±0.073±0.037      f&a      142      <sup>115</sup> BARATE      98E ALEP      1991–1995 LEP runs

0.417±0.058±0.044      avg      <sup>116</sup> COAN      96 CLEO       $E_{\text{cm}}^{ee} \approx 10.6$  GeV

0.41 ±0.12 ±0.03      f&a      <sup>117</sup> ACCIARRI      95F L3      1991–1993 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.32 ±0.11 ±0.05      23      <sup>118</sup> BUSKULIC      96 ALEP      Repl. by BARATE 99K

<sup>114</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

<sup>115</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

<sup>116</sup> Not independent of COAN 96  $B(h^- K^0 \pi^0 \nu_\tau)$  and  $B(K^- K^0 \pi^0 \nu_\tau)$  measurements.

<sup>117</sup> ACCIARRI 95F do not identify  $\pi^-/K^-$  and assume  $B(K^- K^0 \pi^0 \nu_\tau) = (0.05 \pm 0.05)\%$ .

<sup>118</sup> BUSKULIC 96 measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

$\Gamma(\bar{K}^0 \rho^- \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{39}/\Gamma$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
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**0.22 ± 0.05 OUR AVERAGE**

0.250 ± 0.057 ± 0.044	119 BARATE	99K ALEP	1991–1995 LEP runs
0.188 ± 0.054 ± 0.038	120 BARATE	98E ALEP	1991–1995 LEP runs

119 BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in hadron calorimeter. They determine the  $\bar{K}^0 \rho^-$  fraction in  $\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$  decays to be  $(0.72 \pm 0.12 \pm 0.10)$  and multiply their  $B(\pi^- \bar{K}^0 \pi^0 \nu_\tau)$  measurement by this fraction to obtain the quoted result.

120 BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays. They determine the  $\bar{K}^0 \rho^-$  fraction in  $\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$  decays to be  $(0.64 \pm 0.09 \pm 0.10)$  and multiply their  $B(\pi^- \bar{K}^0 \pi^0 \nu_\tau)$  measurement by this fraction to obtain the quoted result.

$\Gamma(K^- K^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{40}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.155 ± 0.020 OUR FIT**

**0.144 ± 0.023 OUR AVERAGE**

0.143 ± 0.025 ± 0.015	78	121 BARATE	99K ALEP	1991–1995 LEP runs
0.152 ± 0.076 ± 0.021	15	122 BARATE	98E ALEP	1991–1995 LEP runs
0.145 ± 0.036 ± 0.020	32	COAN	96 CLEO	$E_{\text{cm}}^{e^+e^-} \approx 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.10 ± 0.05 ± 0.03	5	123 BUSKULIC	96 ALEP	Repl. by BARATE 99K
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121 BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

122 BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

123 BUSKULIC 96 measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

$\Gamma(\pi^- \bar{K}^0 \geq 1\pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{41}/\Gamma = (\Gamma_{38} + \Gamma_{42})/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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<b>0.324 ± 0.074 ± 0.066</b>	148	ABBIENDI	00C OPAL	1991–1995 LEP runs
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$\Gamma(\pi^- \bar{K}^0 \pi^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{42}/\Gamma$

VALUE (units $10^{-3}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
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<b>0.26 ± 0.24</b>			124 BARATE	99R ALEP	1991–1995 LEP runs
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 0.66$	95	17	125 BARATE	99K ALEP	1991–1995 LEP runs
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0.58 ± 0.33 ± 0.14		5	126 BARATE	98E ALEP	1991–1995 LEP runs
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124 BARATE 99R combine the BARATE 98E and BARATE 99K measurements to obtain this value.

125 BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter.

126 BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.



$\Gamma(K^- K^0 \pi^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{43}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<0.16 \times 10^{-3}$	95	127 BARATE	99R ALEP	1991–1995 LEP runs
$<0.18 \times 10^{-3}$	95	128 BARATE	99K ALEP	1991–1995 LEP runs
$<0.39 \times 10^{-3}$	95	129 BARATE	98E ALEP	1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •  
 127 BARATE 99R combine the BARATE 98E and BARATE 99K bounds to obtain this value.  
 128 BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in hadron calorimeter.  
 129 BARATE 98E reconstruct  $K^0$ 's by using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays.

$\Gamma(\pi^- K^0 \bar{K}^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{44}/\Gamma = (2\Gamma_{45} + \Gamma_{46})/\Gamma$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.159 ± 0.029 OUR FIT</b>				Error includes scale factor of 1.1.
<b>0.153 ± 0.030 ± 0.016 avg</b>	74	130 BARATE	98E ALEP	1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •  
 0.31 ± 0.12 ± 0.04 131 ACCIARRI 95F L3 1991–1993 LEP runs

130 BARATE 98E obtain this value by adding twice their  $B(\pi^- K_S^0 K_S^0 \nu_\tau)$  value to their  $B(\pi^- K_S^0 K_L^0 \nu_\tau)$  value.

131 ACCIARRI 95F assume  $B(\pi^- K_S^0 K_S^0 \nu) = B(\pi^- K_S^0 K_L^0 \nu) = 1/2 B(\pi^- K_S^0 K_L^0 \nu)$ .

$\Gamma(\pi^- K_S^0 K_S^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{45}/\Gamma$

Bose-Einstein correlations might make the mixing fraction different than 1/4.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.024 ± 0.005 OUR FIT</b>				
<b>0.024 ± 0.005 OUR AVERAGE</b>				
0.026 ± 0.010 ± 0.005	6	BARATE	98E ALEP	1991–1995 LEP runs
0.023 ± 0.005 ± 0.003	42	COAN	96 CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV

$\Gamma(\pi^- K_S^0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{46}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.110 ± 0.028 OUR FIT</b>				Error includes scale factor of 1.1.
<b>0.101 ± 0.023 ± 0.013</b>	68	BARATE	98E ALEP	1991–1995 LEP runs

$\Gamma(\pi^- K^0 \bar{K}^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{47}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>(0.31 ± 0.23) × 10<sup>-3</sup></b>	132 BARATE	99R ALEP	1991–1995 LEP runs

132 BARATE 99R combine BARATE 98E  $\Gamma(\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$  and  $\Gamma(\pi^- K_S^0 K_L^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$  measurements to obtain this value.

$\Gamma(\pi^- K_S^0 K_S^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{48}/\Gamma$

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.020</b>	95	BARATE	98E ALEP	1991–1995 LEP runs

$\Gamma(\pi^- K_S^0 K_L^0 \pi^0 \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{49}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.031±0.011±0.005</b>	11	BARATE	98E ALEP	1991–1995 LEP runs	

$\Gamma(K^0 h^+ h^- h^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{50}/\Gamma$
VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT	
<b>&lt;0.17</b>	95	TSCHIRHART	88 HRS	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$	
• • • We do not use the following data for averages, fits, limits, etc. • • •					
<0.27	90	BELTRAMI	85 HRS	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$	

$\Gamma(K^0 h^+ h^- h^- \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{51}/\Gamma$
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.023±0.019±0.007</b>	6	<sup>133</sup> BARATE	98E ALEP	1991–1995 LEP runs	
<sup>133</sup> BARATE 98E reconstruct $K^0$ 's using $K_S^0 \rightarrow \pi^+ \pi^-$ decays.					

$\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals } \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{52}/\Gamma$
$\Gamma_{52}/\Gamma = (0.3431\Gamma_{33} + 0.3431\Gamma_{35} + 0.3431\Gamma_{38} + 0.3431\Gamma_{40} + 0.4307\Gamma_{45} + 0.6861\Gamma_{46} + \Gamma_{60} + \Gamma_{68} + \Gamma_{74} + \Gamma_{75} + \Gamma_{82} + \Gamma_{86} + \Gamma_{89} + \Gamma_{90} + 0.285\Gamma_{117} + 0.285\Gamma_{119} + 0.9101\Gamma_{137} + 0.9101\Gamma_{138})/\Gamma$					

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>15.19± 0.07 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>14.8 ± 0.4 OUR AVERAGE</b>				
14.4 ± 0.6 ± 0.3		ADEVA	91F L3	$E_{\text{cm}}^{ee} = 88.3\text{--}94.3 \text{ GeV}$
15.0 ± 0.4 ± 0.3		BEHREND	89B CELL	$E_{\text{cm}}^{ee} = 14\text{--}47 \text{ GeV}$
15.1 ± 0.8 ± 0.6		AIHARA	87B TPC	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
13.5 ± 0.3 ± 0.3		ABACHI	89B HRS	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
12.8 ± 1.0 ± 0.7	<sup>134</sup>	BURCHAT	87 MRK2	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
12.1 ± 0.5 ± 1.2		RUCKSTUHL	86 DLCO	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
12.8 ± 0.5 ± 0.8	1420	SCHMIDKE	86 MRK2	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
15.3 ± 1.1 $\begin{smallmatrix} +1.3 \\ -1.6 \end{smallmatrix}$	367	ALTHOFF	85 TASS	$E_{\text{cm}}^{ee} = 34.5 \text{ GeV}$
13.6 ± 0.5 ± 0.8		BARTEL	85F JADE	$E_{\text{cm}}^{ee} = 34.6 \text{ GeV}$
12.2 ± 1.3 ± 3.9	<sup>135</sup>	BERGER	85 PLUT	$E_{\text{cm}}^{ee} = 34.6 \text{ GeV}$
13.3 ± 0.3 ± 0.6		FERNANDEZ	85 MAC	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
24 ± 6	35	BRANDELIK	80 TASS	$E_{\text{cm}}^{ee} = 30 \text{ GeV}$
32 ± 5	692	<sup>136</sup> BACINO	78B DLCO	$E_{\text{cm}}^{ee} = 3.1\text{--}7.4 \text{ GeV}$
35 ± 11		<sup>136</sup> BRANDELIK	78 DASP	Assumes $V\text{--}A$ decay
18 ± 6.5	33	<sup>136</sup> JAROS	78 MRK1	$E_{\text{cm}}^{ee} > 6 \text{ GeV}$

<sup>134</sup>BURCHAT 87 value is not independent of SCHMIDKE 86 value.

<sup>135</sup>Not independent of BERGER 85  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau)/\Gamma_{\text{total}}$ ,  $\Gamma(e^- \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$ ,  $\Gamma(h^- \geq 1 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$ , and  $\Gamma(h^- \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}}$ , and therefore not used in the fit.

<sup>136</sup>Low energy experiments are not in average or fit because the systematic errors in background subtraction are judged to be large.

$$\Gamma(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau (\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-) (\text{"3-prong"})) / \Gamma_{\text{total}} \quad \Gamma_{53} / \Gamma$$

$$\Gamma_{53} / \Gamma = (\Gamma_{60} + \Gamma_{68} + \Gamma_{74} + \Gamma_{75} + \Gamma_{82} + \Gamma_{86} + \Gamma_{89} + \Gamma_{90} + 0.285\Gamma_{117} + 0.285\Gamma_{119} + 0.9101\Gamma_{137} + 0.9101\Gamma_{138}) / \Gamma$$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%)		EVTS	DOCUMENT ID	TECN	COMMENT
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**14.57 ± 0.07 OUR FIT** Error includes scale factor of 1.1.

**14.59 ± 0.08 OUR AVERAGE** Error includes scale factor of 1.1.

14.569 ± 0.093 ± 0.048	f&a	23k	<sup>137</sup> ABREU	01M DLPH	1992–1995 LEP runs
14.556 ± 0.105 ± 0.076	f&a		<sup>138</sup> ACHARD	01D L3	1992–1995 LEP runs
14.96 ± 0.09 ± 0.22	f&a	10.4k	AKERS	95Y OPAL	1991–1994 LEP runs
14.22 ± 0.10 ± 0.37	avg		<sup>139</sup> BALEST	95C CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •					
15.26 ± 0.26 ± 0.22			ACTON	92H OPAL	Repl. by AK-ERS 95Y
13.3 ± 0.3 ± 0.8			<sup>140</sup> ALBRECHT	92D ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
14.35 $\begin{smallmatrix} +0.40 \\ -0.45 \end{smallmatrix}$ ± 0.24			DECAMP	92C ALEP	1989–1990 LEP runs

<sup>137</sup> The correlation coefficients between this measurement and the ABREU 01M measurements of  $B(\tau \rightarrow 1\text{-prong})$  and  $B(\tau \rightarrow 5\text{-prong})$  are  $-0.98$  and  $-0.08$  respectively.

<sup>138</sup> The correlation coefficients between this measurement and the ACHARD 01D measurements of  $B(\tau \rightarrow \text{"1-prong"})$  and  $B(\tau \rightarrow \text{"5-prong"})$  are  $-0.978$  and  $-0.19$  respectively.

<sup>139</sup> Not independent of BALEST 95C  $B(h^- h^- h^+ \nu_\tau)$  and  $B(h^- h^- h^+ \pi^0 \nu_\tau)$  values, and BORTOLETTO 93  $B(h^- h^- h^+ 2\pi^0 \nu_\tau) / B(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau)$  value.

<sup>140</sup> This ALBRECHT 92D value is not independent of their  $\Gamma(\mu^- \bar{\nu}_\mu \nu_\tau) \Gamma(e^- \bar{\nu}_e \nu_\tau) / \Gamma_{\text{total}}^2$  value.

$$\Gamma(h^- h^- h^+ \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{54} / \Gamma$$

$$\Gamma_{54} / \Gamma = (0.3431\Gamma_{33} + 0.3431\Gamma_{35} + \Gamma_{60} + \Gamma_{82} + \Gamma_{89} + 0.0221\Gamma_{137}) / \Gamma$$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

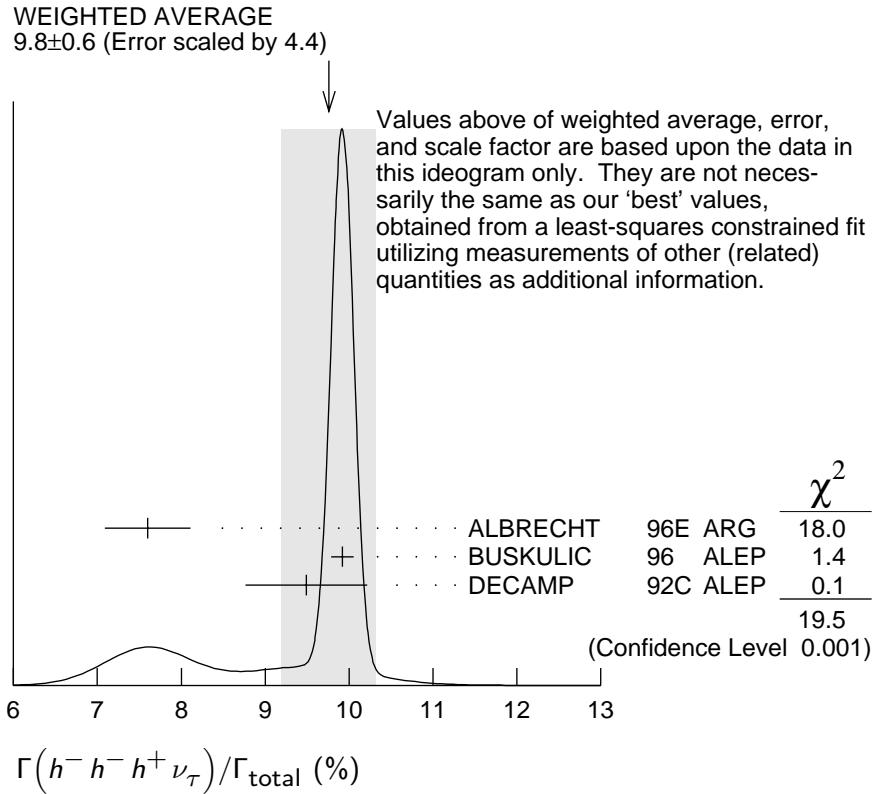
VALUE (%)		EVTS	DOCUMENT ID	TECN	COMMENT
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**10.01 ± 0.09 OUR FIT** Error includes scale factor of 1.2.

**9.8 ± 0.6 OUR AVERAGE** Error includes scale factor of 4.4. See the ideogram below.

7.6 ± 0.1 ± 0.5	avg	7.5k	<sup>141</sup> ALBRECHT	96E ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
9.92 ± 0.10 ± 0.09	f&a	11.2k	<sup>142</sup> BUSKULIC	96 ALEP	LEP 1991–1993 data
9.49 ± 0.36 ± 0.63	f&a		DECAMP	92C ALEP	1989–1990 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •					
8.7 ± 0.7 ± 0.3		694	<sup>143</sup> BEHREND	90 CELL	$E_{\text{cm}}^{ee} = 35$ GeV
7.0 ± 0.3 ± 0.7		1566	<sup>144</sup> BAND	87 MAC	$E_{\text{cm}}^{ee} = 29$ GeV
6.7 ± 0.8 ± 0.9			<sup>145</sup> BURCHAT	87 MRK2	$E_{\text{cm}}^{ee} = 29$ GeV
6.4 ± 0.4 ± 0.9			<sup>146</sup> RUCKSTUHL	86 DLCO	$E_{\text{cm}}^{ee} = 29$ GeV
7.8 ± 0.5 ± 0.8		890	SCHMIDKE	86 MRK2	$E_{\text{cm}}^{ee} = 29$ GeV
8.4 ± 0.4 ± 0.7		1255	<sup>146</sup> FERNANDEZ	85 MAC	$E_{\text{cm}}^{ee} = 29$ GeV
9.7 ± 2.0 ± 1.3			BEHREND	84 CELL	$E_{\text{cm}}^{ee} = 14,22$ GeV

- 141 ALBRECHT 96E not independent of ALBRECHT 93C  $\Gamma(h^- h^- h^+ \nu_\tau(\text{ex. } K^0)) \times \Gamma(\text{particle}^- \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) / \Gamma_{\text{total}}^2$  value.
- 142 BUSKULIC 96 quote  $B(h^- h^- h^+ \nu_\tau(\text{ex. } K^0)) = 9.50 \pm 0.10 \pm 0.11$ . We add 0.42 to remove their  $K^0$  correction and reduce the systematic error accordingly.
- 143 BEHREND 90 subtract 0.3% to account for the  $\tau^- \rightarrow K^*(892)^- \nu_\tau$  contribution to measured events.
- 144 BAND 87 subtract for charged kaon modes; not independent of FERNANDEZ 85 value.
- 145 BURCHAT 87 value is not independent of SCHMIDKE 86 value.
- 146 Value obtained by multiplying paper's  $R = B(h^- h^- h^+ \nu_\tau) / B(3\text{-prong})$  by  $B(3\text{-prong}) = 0.143$  and subtracting 0.3% for  $K^*(892)$  background.



$\Gamma(h^- h^- h^+ \nu_\tau(\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{55} / \Gamma$   
 $\Gamma_{55} / \Gamma = (\Gamma_{60} + \Gamma_{82} + \Gamma_{89} + 0.0221 \Gamma_{137}) / \Gamma$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%)		EVTS	DOCUMENT ID	TECN	COMMENT
<b>9.65±0.09 OUR FIT</b>	Error includes scale factor of 1.2.				
<b>9.57±0.11 OUR AVERAGE</b>					
9.50±0.10±0.11	avg	11.2k	147 BUSKULIC	96 ALEP	LEP 1991–1993 data
9.87±0.10±0.24	avg		148 AKERS	95Y OPAL	1991–1994 LEP runs
9.51±0.07±0.20	f&a	37.7k	BALEST	95C CLEO	$E_{\text{cm}}^e \approx 10.6 \text{ GeV}$

- 147 Not independent of BUSKULIC 96  $B(h^- h^- h^+ \nu_\tau)$  value.
- 148 Not independent of AKERS 95Y  $B(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau(\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  and  $B(h^- h^- h^+ \nu_\tau(\text{ex. } K^0)) / B(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau(\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  values.

$$\Gamma(h^- h^- h^+ \nu_\tau (\text{ex. } K^0)) / \Gamma(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau (\text{ex. } K_S^0 \rightarrow \pi^+ \pi^-))$$

**Γ<sub>55</sub>/Γ<sub>53</sub>**

$$\Gamma_{55}/\Gamma_{53} = (\Gamma_{60} + \Gamma_{82} + \Gamma_{89} + 0.0221\Gamma_{137}) / (\Gamma_{60} + \Gamma_{68} + \Gamma_{74} + \Gamma_{75} + \Gamma_{82} + \Gamma_{86} + \Gamma_{89} + \Gamma_{90} + 0.285\Gamma_{117} + 0.285\Gamma_{119} + 0.9101\Gamma_{137} + 0.9101\Gamma_{138})$$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.662 ± 0.006 OUR FIT</b>			Error includes scale factor of 1.3.
<b>0.660 ± 0.004 ± 0.014</b>	AKERS	95Y OPAL	1991–1994 LEP runs

$$\Gamma(h^- h^- h^+ \nu_\tau (\text{ex. } K^0, \omega)) / \Gamma_{\text{total}} \quad \Gamma_{56}/\Gamma = (\Gamma_{60} + \Gamma_{82} + \Gamma_{89}) / \Gamma$$

VALUE (%)      DOCUMENT ID

**9.60 ± 0.09 OUR FIT**      Error includes scale factor of 1.2.

$$\Gamma(\pi^- \pi^+ \pi^- \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{57}/\Gamma = (0.3431\Gamma_{33} + \Gamma_{60} + 0.0221\Gamma_{137}) / \Gamma$$

VALUE (%)      DOCUMENT ID

**9.47 ± 0.10 OUR FIT**      Error includes scale factor of 1.2.

$$\Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{58}/\Gamma = (\Gamma_{60} + 0.0221\Gamma_{137}) / \Gamma$$

VALUE (%)      EVTS      DOCUMENT ID      TECN      COMMENT

**9.16 ± 0.10 OUR FIT**      Error includes scale factor of 1.2.

**9.13 ± 0.05 ± 0.46**      43k      149 BRIERE      03      CLE3       $E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

<sup>149</sup> 47% correlated with BRIERE 03  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$  and 71% correlated with  $\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau$  because of a common 5% normalization error.

$$\Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0), \text{ non-axial vector}) / \Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$$

**Γ<sub>59</sub>/Γ<sub>58</sub> = Γ<sub>59</sub> / (Γ<sub>60</sub> + 0.0221Γ<sub>137</sub>)**

VALUE      CL%      DOCUMENT ID      TECN      COMMENT

**< 0.261**      95      150 ACKERSTAFF 97R OPAL      1992–1994 LEP runs

<sup>150</sup> Model-independent limit from structure function analysis on contribution to  $B(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  from non-axial vectors.

$$\Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0, \omega)) / \Gamma_{\text{total}} \quad \Gamma_{60}/\Gamma$$

VALUE (%)      DOCUMENT ID

**9.12 ± 0.10 OUR FIT**      Error includes scale factor of 1.2.

$$\Gamma(h^- h^- h^+ \geq 1 \text{ neutrals } \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{61}/\Gamma$$

$$\Gamma_{61}/\Gamma = (0.3431\Gamma_{38} + 0.3431\Gamma_{40} + 0.4307\Gamma_{45} + 0.6861\Gamma_{46} + \Gamma_{68} + \Gamma_{74} + \Gamma_{75} + \Gamma_{86} + \Gamma_{90} + 0.285\Gamma_{117} + 0.285\Gamma_{119} + 0.888\Gamma_{137} + 0.9101\Gamma_{138}) / \Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>5.19 ± 0.10 OUR FIT</b>				Error includes scale factor of 1.3.

• • • We do not use the following data for averages, fits, limits, etc. • • •

5.6 ± 0.7 ± 0.3	352	<sup>151</sup> BEHREND	90	CELL	$E_{\text{cm}}^{ee} = 35 \text{ GeV}$
4.2 ± 0.5 ± 0.9	203	<sup>152</sup> ALBRECHT	87L	ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
6.1 ± 0.8 ± 0.9		<sup>153</sup> BURCHAT	87	MRK2	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
7.6 ± 0.4 ± 0.9		<sup>154,155</sup> RUCKSTUHL	86	DLCO	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
4.7 ± 0.5 ± 0.8	530	<sup>156</sup> SCHMIDKE	86	MRK2	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
5.6 ± 0.4 ± 0.7		<sup>155</sup> FERNANDEZ	85	MAC	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
6.2 ± 2.3 ± 1.7		BEHREND	84	CELL	$E_{\text{cm}}^{ee} = 14, 22 \text{ GeV}$

- 151 BEHREND 90 value is not independent of BEHREND 90  $B(3h\nu_\tau \geq 1 \text{ neutrals}) + B(5\text{-prong})$ .
- 152 ALBRECHT 87L measure the product of branching ratios  $B(3\pi^\pm \pi^0 \nu_\tau) B((e\bar{\nu} \text{ or } \mu\bar{\nu} \text{ or } \pi \text{ or } K \text{ or } \rho)\nu_\tau) = 0.029$  and use the PDG 86 values for the second branching ratio which sum to  $0.69 \pm 0.03$  to get the quoted value.
- 153 BURCHAT 87 value is not independent of SCHMIDKE 86 value.
- 154 Contributions from kaons and from  $>1\pi^0$  are subtracted. Not independent of  $(3\text{-prong} + 0\pi^0)$  and  $(3\text{-prong} + \geq 0\pi^0)$  values.
- 155 Value obtained using paper's  $R = B(h^- h^- h^+ \nu_\tau)/B(3\text{-prong})$  and current  $B(3\text{-prong}) = 0.143$ .
- 156 Not independent of SCHMIDKE 86  $h^- h^- h^+ \nu_\tau$  and  $h^- h^- h^+ (\geq 0\pi^0)\nu_\tau$  values.

$$\Gamma(h^- h^- h^+ \geq 1 \text{ neutrals } \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^+ \pi^-))/\Gamma_{\text{total}} \quad \Gamma_{62}/\Gamma$$

$$\Gamma_{62}/\Gamma = (\Gamma_{68} + \Gamma_{74} + \Gamma_{75} + \Gamma_{86} + \Gamma_{90} + 0.285\Gamma_{117} + 0.285\Gamma_{119} + 0.888\Gamma_{137} + 0.9101\Gamma_{138})/\Gamma$$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.92±0.09 OUR FIT</b>				Error includes scale factor of 1.3.
<b>5.07±0.24 OUR AVERAGE</b>				
5.09±0.10±0.23	avg	157 AKERS	95Y OPAL	1991–1994 LEP runs
4.95±0.29±0.65	f&a	570 DECAMP	92C ALEP	1989–1990 LEP runs

- 157 Not independent of AKERS 95Y  $B(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  and  $B(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau \text{ (ex. } K^0))/B(h^- h^- h^+ \geq 0 \text{ neutrals } \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^+ \pi^-))$  values.

$$\Gamma(h^- h^- h^+ \pi^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{63}/\Gamma$$

$$\Gamma_{63}/\Gamma = (0.3431\Gamma_{38} + 0.3431\Gamma_{40} + \Gamma_{68} + \Gamma_{86} + \Gamma_{90} + 0.231\Gamma_{119} + 0.888\Gamma_{137} + 0.0221\Gamma_{138})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.53±0.09 OUR FIT</b>				Error includes scale factor of 1.3.
<b>4.45±0.09±0.07</b>	6.1k	158 BUSKULIC	96 ALEP	LEP 1991–1993 data

- 158 BUSKULIC 96 quote  $B(h^- h^- h^+ \pi^0 \nu_\tau \text{ (ex. } K^0)) = 4.30 \pm 0.09 \pm 0.09$ . We add 0.15 to remove their  $K^0$  correction and reduce the systematic error accordingly.

$$\Gamma(h^- h^- h^+ \pi^0 \nu_\tau \text{ (ex. } K^0))/\Gamma_{\text{total}} \quad \Gamma_{64}/\Gamma$$

$$\Gamma_{64}/\Gamma = (\Gamma_{68} + \Gamma_{86} + \Gamma_{90} + 0.231\Gamma_{119} + 0.888\Gamma_{137} + 0.0221\Gamma_{138})/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.35±0.09 OUR FIT</b>				Error includes scale factor of 1.3.
<b>4.23±0.06±0.22</b>	7.2k	BALEST	95C CLEO	$E_{\text{cm}}^{ee} \approx 10.6 \text{ GeV}$

$$\Gamma(h^- h^- h^+ \pi^0 \nu_\tau \text{ (ex. } K^0, \omega))/\Gamma_{\text{total}} \quad \Gamma_{65}/\Gamma = (\Gamma_{68} + \Gamma_{86} + \Gamma_{90} + 0.231\Gamma_{119})/\Gamma$$

VALUE (%)	DOCUMENT ID
<b>2.62±0.09 OUR FIT</b>	Error includes scale factor of 1.2.

$$\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{66} / \Gamma = (0.3431\Gamma_{38} + \Gamma_{68} + 0.888\Gamma_{137} + 0.0221\Gamma_{138}) / \Gamma$$

VALUE (%) DOCUMENT ID  
**4.37 ± 0.09 OUR FIT** Error includes scale factor of 1.3.

$$\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{67} / \Gamma = (\Gamma_{68} + 0.888\Gamma_{137} + 0.0221\Gamma_{138}) / \Gamma$$

VALUE (%) DOCUMENT ID TECN COMMENT  
**4.25 ± 0.09 OUR FIT** Error includes scale factor of 1.3.  
**4.19 ± 0.10 ± 0.21** <sup>159</sup> EDWARDS 00A CLEO 4.7 fb<sup>-1</sup> E<sub>cm</sub><sup>ee</sup> = 10.6 GeV  
<sup>159</sup> EDWARDS 00A quote (4.19 ± 0.10) × 10<sup>-2</sup> with a 5% systematic error.

$$\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0, \omega)) / \Gamma_{\text{total}} \quad \Gamma_{68} / \Gamma$$

VALUE (%) DOCUMENT ID  
**2.51 ± 0.09 OUR FIT** Error includes scale factor of 1.2.

$$\Gamma(h^- \rho \pi^0 \nu_\tau) / \Gamma(h^- h^- h^+ \pi^0 \nu_\tau) \quad \Gamma_{69} / \Gamma_{63}$$

VALUE EVTS DOCUMENT ID TECN COMMENT  
 • • • We do not use the following data for averages, fits, limits, etc. • • •  
 0.30 ± 0.04 ± 0.02 393 ALBRECHT 91D ARG E<sub>cm</sub><sup>ee</sup> = 9.4–10.6 GeV

$$\Gamma(h^- \rho^+ h^- \nu_\tau) / \Gamma(h^- h^- h^+ \pi^0 \nu_\tau) \quad \Gamma_{70} / \Gamma_{63}$$

VALUE EVTS DOCUMENT ID TECN COMMENT  
 • • • We do not use the following data for averages, fits, limits, etc. • • •  
 0.10 ± 0.03 ± 0.04 142 ALBRECHT 91D ARG E<sub>cm</sub><sup>ee</sup> = 9.4–10.6 GeV

$$\Gamma(h^- \rho^- h^+ \nu_\tau) / \Gamma(h^- h^- h^+ \pi^0 \nu_\tau) \quad \Gamma_{71} / \Gamma_{63}$$

VALUE EVTS DOCUMENT ID TECN COMMENT  
 • • • We do not use the following data for averages, fits, limits, etc. • • •  
 0.26 ± 0.05 ± 0.01 370 ALBRECHT 91D ARG E<sub>cm</sub><sup>ee</sup> = 9.4–10.6 GeV

$$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{72} / \Gamma$$

$$\Gamma_{72} / \Gamma = (0.4307\Gamma_{45} + \Gamma_{74} + 0.236\Gamma_{117} + 0.888\Gamma_{138}) / \Gamma$$

VALUE (%) DOCUMENT ID  
**0.55 ± 0.04 OUR FIT**

$$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{73} / \Gamma$$

$$\Gamma_{73} / \Gamma = (\Gamma_{74} + 0.236\Gamma_{117} + 0.888\Gamma_{138}) / \Gamma$$

VALUE (%) EVTS DOCUMENT ID TECN COMMENT  
**0.54 ± 0.04 OUR FIT**  
**0.50 ± 0.07 ± 0.07** 1.8k BUSKULIC 96 ALEP LEP 1991–1993 data

$$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) \quad \Gamma_{73} / \Gamma_{52}$$

$$\Gamma_{73} / \Gamma_{52} = (\Gamma_{74} + 0.236\Gamma_{117} + 0.888\Gamma_{138}) / (0.3431\Gamma_{33} + 0.3431\Gamma_{35} + 0.3431\Gamma_{38} + 0.3431\Gamma_{40} + 0.4307\Gamma_{45} + 0.6861\Gamma_{46} + \Gamma_{60} + \Gamma_{68} + \Gamma_{74} + \Gamma_{75} + \Gamma_{82} + \Gamma_{86} + \Gamma_{89} + \Gamma_{90} + 0.285\Gamma_{117} + 0.285\Gamma_{119} + 0.9101\Gamma_{137} + 0.9101\Gamma_{138})$$

VALUE EVTS DOCUMENT ID TECN COMMENT  
**0.0355 ± 0.0028 OUR FIT**  
**0.034 ± 0.002 ± 0.003** 668 BORTOLETTO93 CLEO E<sub>cm</sub><sup>ee</sup> ≈ 10.6 GeV

$\Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau (\text{ex. } K^0, \omega, \eta))/\Gamma_{\text{total}}$   $\Gamma_{74}/\Gamma$   
VALUE (%) DOCUMENT ID  
**0.11±0.04 OUR FIT**

$\Gamma(h^- h^- h^+ 3\pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{75}/\Gamma$   
VALUE (%) EVTS DOCUMENT ID TECN COMMENT  
**0.023 ±0.008 OUR FIT** Error includes scale factor of 1.5.  
**0.023 ±0.005 OUR AVERAGE**

0.022 ±0.003 ±0.004	139	ANASTASSOV 01	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
0.11 ±0.04 ±0.05	440	<sup>160</sup> BUSKULIC 96	ALEP	LEP 1991–1993 data

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.0285±0.0056±0.0051	57	ANDERSON 97	CLEO	Repl. by ANASTASSOV 01
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<sup>160</sup>BUSKULIC 96 state their measurement is for  $B(h^- h^- h^+ \geq 3\pi^0 \nu_\tau)$ . We assume that  $B(h^- h^- h^+ \geq 4\pi^0 \nu_\tau)$  is very small.

$\Gamma(K^- h^+ h^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$   
 $\Gamma_{76}/\Gamma = (0.3431\Gamma_{35} + 0.3431\Gamma_{40} + \Gamma_{82} + \Gamma_{86} + \Gamma_{89} + \Gamma_{90} + 0.285\Gamma_{119})/\Gamma$   
VALUE (%) CL% DOCUMENT ID TECN COMMENT  
**0.69±0.04 OUR FIT** Error includes scale factor of 1.3.  
**<0.6** 90 AIHARA 84C TPC  $E_{\text{cm}}^{ee} = 29$  GeV

$\Gamma(K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   $\Gamma_{77}/\Gamma = (\Gamma_{82} + \Gamma_{89})/\Gamma$   
VALUE (%) DOCUMENT ID  
**0.48±0.04 OUR FIT** Error includes scale factor of 1.5.

$\Gamma(K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$   
 $\Gamma_{77}/\Gamma_{58} = (\Gamma_{82} + \Gamma_{89})/(\Gamma_{60} + 0.0221\Gamma_{137})$   
VALUE (%) EVTS DOCUMENT ID TECN COMMENT  
**5.2 ±0.4 OUR FIT** Error includes scale factor of 1.6.  
**5.44±0.21±0.53** 7.9k RICHICHI 99 CLEO  $E_{\text{cm}}^{ee} = 10.6$  GeV

$\Gamma(K^- h^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   $\Gamma_{78}/\Gamma = (\Gamma_{86} + \Gamma_{90} + 0.231\Gamma_{119})/\Gamma$   
VALUE (%) DOCUMENT ID  
**0.107±0.022 OUR FIT**

$\Gamma(K^- h^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0))$   
 $\Gamma_{78}/\Gamma_{67} = (\Gamma_{86} + \Gamma_{90} + 0.231\Gamma_{119})/(\Gamma_{68} + 0.888\Gamma_{137} + 0.0221\Gamma_{138})$   
VALUE (%) EVTS DOCUMENT ID TECN COMMENT  
**2.5 ±0.5 OUR FIT**  
**2.61±0.45±0.42** 719 RICHICHI 99 CLEO  $E_{\text{cm}}^{ee} = 10.6$  GeV

$\Gamma(K^- \pi^+ \pi^- \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$   
 $\Gamma_{79}/\Gamma = (0.3431\Gamma_{35} + 0.3431\Gamma_{40} + \Gamma_{82} + \Gamma_{86} + 0.285\Gamma_{119})/\Gamma$   
VALUE (%) EVTS DOCUMENT ID TECN COMMENT  
**0.50±0.04 OUR FIT** Error includes scale factor of 1.3.

<b>0.58<sup>+0.15</sup><sub>-0.13</sub>±0.12</b>	20	<sup>161</sup> BAUER	94	TPC	$E_{\text{cm}}^{ee} = 29$ GeV
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.22 <sup>+0.16</sup> <sub>-0.13</sub> ±0.05	9	<sup>162</sup> MILLS	85	DLCO	$E_{\text{cm}}^{ee} = 29$ GeV
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<sup>161</sup>We multiply 0.58% by 0.20, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

<sup>162</sup>Error correlated with MILLS 85 ( $K K \pi \nu$ ) value. We multiply 0.22% by 0.23, the relative systematic error quoted by MILLS 85, to obtain the systematic error.

$$\Gamma(K^- \pi^+ \pi^- \geq 0 \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{80} / \Gamma = (\Gamma_{82} + \Gamma_{86} + 0.231 \Gamma_{119}) / \Gamma$$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%)		DOCUMENT ID	TECN	COMMENT
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**0.39 ± 0.04 OUR FIT** Error includes scale factor of 1.3.

**0.30 ± 0.05 OUR AVERAGE**

0.343 ± 0.073 ± 0.031	f&a	ABBIENDI	00D OPAL	1990–1995 LEP runs
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0.275 ± 0.064	avg	<sup>163</sup> BARATE	98 ALEP	1991–1995 LEP runs
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<sup>163</sup>Not independent of BARATE 98  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow K^- \pi^+ \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}}$  values.

$$\Gamma(K^- \pi^+ \pi^- \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{81} / \Gamma = (0.3431 \Gamma_{35} + \Gamma_{82}) / \Gamma$$

VALUE (%)	DOCUMENT ID
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**0.38 ± 0.04 OUR FIT** Error includes scale factor of 1.6.

$$\Gamma(K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}} \quad \Gamma_{82} / \Gamma$$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%)		EVTS	DOCUMENT ID	TECN	COMMENT
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**0.33 ± 0.04 OUR FIT** Error includes scale factor of 1.6.

**0.32 ± 0.04 OUR AVERAGE** Error includes scale factor of 1.6. See the ideogram below.

0.384 ± 0.014 ± 0.038	f&a	3.5k	<sup>164</sup> BRIERE	03 CLE3	$E_{\text{cm}}^{e e} = 10.6 \text{ GeV}$
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0.360 ± 0.082 ± 0.048	avg		ABBIENDI	00D OPAL	1990–1995 LEP runs
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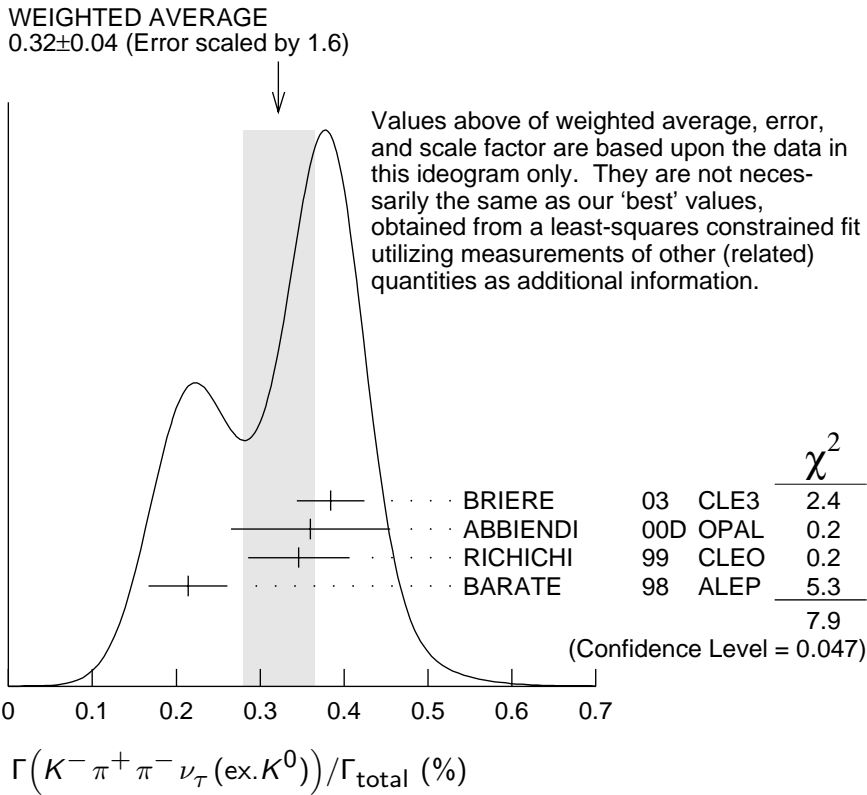
0.346 ± 0.023 ± 0.056	avg	158	<sup>165</sup> RICHICHI	99 CLEO	$E_{\text{cm}}^{e e} = 10.6 \text{ GeV}$
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0.214 ± 0.037 ± 0.029	f&a		BARATE	98 ALEP	1991–1995 LEP runs
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<sup>164</sup>47% correlated with BRIERE 03  $\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau$  and 34% correlated with  $\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau$  because of a common 5% normalization error.

<sup>165</sup>Not independent of RICHICHI 99

$\Gamma(\tau^- \rightarrow K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0)) / \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$ ,  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  and BALEST 95C  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  values.



**$\Gamma(K^- \rho^0 \nu_\tau \rightarrow K^- \pi^+ \pi^- \nu_\tau) / \Gamma(K^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$   $\Gamma_{83} / \Gamma_{82}$**

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>0.48 \pm 0.14 \pm 0.10</math></b>	166 ASNER	00B CLEO	$E_{\text{cm}}^e = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.39 \pm 0.14$	167 BARATE	99R ALEP	1991–1995 LEP runs
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166 ASNER 00B assume  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$  (ex.  $K^0$ ) decays proceed only through  $K\rho$  and  $K^* \pi$  intermediate states. They assume the resonance structure of  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$  (ex.  $K^0$ ) decays is dominated by  $K_1(1270)^-$  and  $K_1(1400)^-$  resonances, and assume  $B(K_1(1270) \rightarrow K^*(892)\pi) = (16 \pm 5)\%$ ,  $B(K_1(1270) \rightarrow K\rho) = (42 \pm 6)\%$ , and  $B(K_1(1400) \rightarrow K\rho) = 0$ .

167 BARATE 99R assume  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$  (ex.  $K^0$ ) decays proceed only through  $K\rho$  and  $K^* \pi$  intermediate states. The quoted error is statistical only.

**$\Gamma(K^- \pi^+ \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{84} / \Gamma = (0.3431\Gamma_{40} + \Gamma_{86} + 0.231\Gamma_{119}) / \Gamma$**

VALUE (units $10^{-4}$ )	DOCUMENT ID
<b><math>11.8 \pm 2.5</math> OUR FIT</b>	

**$\Gamma(K^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$   $\Gamma_{85} / \Gamma = (\Gamma_{86} + 0.231\Gamma_{119}) / \Gamma$**

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**6.5±2.4 OUR FIT**

**7.0±2.5 OUR AVERAGE**

7.5±2.6±1.8	avg	168 RICHICHI	99 CLEO	$E_{cm}^{ee} = 10.6$ GeV
6.1±3.9±1.8	f&a	BARATE	98 ALEP	1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

<17	95	ABBIENDI	00D OPAL	1990–1995 LEP runs
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<sup>168</sup> Not independent of RICHICHI 99

$\Gamma(\tau^- \rightarrow K^- h^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$ ,  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau)/\Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  and BALEST 95C  $\Gamma(\tau^- \rightarrow h^- h^+ \pi^- \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$  values.

**$\Gamma(K^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0, \eta))/\Gamma_{\text{total}}$**

**$\Gamma_{86}/\Gamma$**

Test of lepton family number conservation.

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>DOCUMENT ID</u>
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**5.9±2.4 OUR FIT**

**$\Gamma(K^- \pi^+ K^- \geq 0 \text{ neut. } \nu_\tau)/\Gamma_{\text{total}}$**

**$\Gamma_{87}/\Gamma$**

<u>VALUE (%)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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<0.09	95	BAUER	94 TPC	$E_{cm}^{ee} = 29$ GeV
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**$\Gamma(K^- K^+ \pi^- \geq 0 \text{ neut. } \nu_\tau)/\Gamma_{\text{total}}$**

**$\Gamma_{88}/\Gamma = (\Gamma_{89} + \Gamma_{90})/\Gamma$**

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

<u>VALUE (%)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**0.197±0.018 OUR FIT** Error includes scale factor of 1.1.

**0.203±0.031 OUR AVERAGE**

0.159±0.053±0.020	f&a	ABBIENDI	00D OPAL	1990–1995 LEP runs
0.238±0.042	avg	<sup>169</sup> BARATE	98 ALEP	1991–1995 LEP runs
0.15 $\begin{smallmatrix} +0.09 \\ -0.07 \end{smallmatrix}$ ±0.03	f&a	4 <sup>170</sup> BAUER	94 TPC	$E_{cm}^{ee} = 29$ GeV

<sup>169</sup> Not independent of BARATE 98  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$  and  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \pi^0 \nu_\tau)/\Gamma_{\text{total}}$  values.

<sup>170</sup> We multiply 0.15% by 0.20, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

**$\Gamma(K^- K^+ \pi^- \nu_\tau)/\Gamma_{\text{total}}$**

**$\Gamma_{89}/\Gamma$**

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

<u>VALUE (%)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**0.155±0.007 OUR FIT**

**0.154±0.009 OUR AVERAGE**

0.155±0.006±0.009	f&a	932 <sup>171</sup> BRIERE	03 CLE3	$E_{cm}^{ee} = 10.6$ GeV
0.087±0.056±0.040	avg	ABBIENDI	00D OPAL	1990–1995 LEP runs
0.145±0.013±0.028	avg	2.3k <sup>172</sup> RICHICHI	99 CLEO	$E_{cm}^{ee} = 10.6$ GeV

0.163±0.021±0.017 f&a BARATE 98 ALEP 1991–1995 LEP runs  
 0.22  $^{+0.17}_{-0.11}$  ±0.05 f&a 9 <sup>173</sup> MILLS 85 DLCO  $E_{cm}^{ee} = 29$  GeV

<sup>171</sup> 71% correlated with BRIERE 03  $\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau$  and 34% correlated with  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$  because of a common 5% normalization error.

<sup>172</sup> Not independent of RICHICHI 99  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  and BALEST 95C  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  values.

<sup>173</sup> Error correlated with MILLS 85 ( $K \pi \pi \pi^0 \nu$ ) value. We multiply 0.22% by 0.23, the relative systematic error quoted by MILLS 85, to obtain obtain the systematic error.

$\Gamma(K^- K^+ \pi^- \nu_\tau) / \Gamma(\pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{89} / \Gamma_{58} = \Gamma_{89} / (\Gamma_{60} + 0.0221 \Gamma_{137})$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.69±0.08 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>1.60±0.15±0.30</b>	2.3k	RICHICHI	99	CLEO $E_{cm}^{ee} = 10.6$ GeV

$\Gamma(K^- K^+ \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{90} / \Gamma$

Data marked “avg” are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. “f&a” marks results used for the fit and the average.

VALUE (units 10 <sup>-4</sup> )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>4.2±1.6 OUR FIT</b>	Error includes scale factor of 1.1.				
<b>4.4±1.8 OUR AVERAGE</b>	Error includes scale factor of 1.1.				
3.3±1.8±0.7	avg	158	<sup>174</sup> RICHICHI	99	CLEO $E_{cm}^{ee} = 10.6$ GeV
7.5±2.9±1.5	f&a		BARATE	98	ALEP 1991–1995 LEP runs

• • • We do not use the following data for averages, fits, limits, etc. • • •

<27 95 ABBIENDI 00D OPAL 1990–1995 LEP runs

<sup>174</sup> Not independent of RICHICHI 99  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma(\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau (\text{ex. } K^0))$  and BALEST 95C  $\Gamma(\tau^- \rightarrow h^- h^- h^+ \nu_\tau (\text{ex. } K^0)) / \Gamma_{\text{total}}$  values.

$\Gamma(K^- K^+ \pi^- \pi^0 \nu_\tau) / \Gamma(\pi^- \pi^+ \pi^- \pi^0 \nu_\tau (\text{ex. } K^0)) \quad \Gamma_{90} / \Gamma_{67} = \Gamma_{90} / (\Gamma_{68} + 0.888 \Gamma_{137} + 0.0221 \Gamma_{138})$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.0 ±0.4 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>0.79±0.44±0.16</b>	158	<sup>175</sup> RICHICHI	99	CLEO $E_{cm}^{ee} = 10.6$ GeV

<sup>175</sup> RICHICHI 99 also quote a 95%CL upper limit of 0.0157 for this measurement.

$\Gamma(K^- K^+ K^- \geq 0 \text{ neut. } \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{91} / \Gamma$

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.21</b>	95	BAUER	94	TPC $E_{cm}^{ee} = 29$ GeV

$\Gamma(K^- K^+ K^- \nu_\tau) / \Gamma_{\text{total}} \quad \Gamma_{92} / \Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;3.7 × 10<sup>-5</sup></b>	90	BRIERE	03	CLE3 $E_{cm}^{ee} = 10.6$ GeV
<1.9 × 10 <sup>-4</sup>	90	BARATE	98	ALEP 1991–1995 LEP runs

$\Gamma(\pi^- K^+ \pi^- \geq 0 \text{ neut. } \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{93}/\Gamma$

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
<0.25	95	BAUER	94 TPC	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$

$\Gamma(e^- e^- e^+ \bar{\nu}_e \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{94}/\Gamma$

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
$2.8 \pm 1.4 \pm 0.4$	5	ALAM	96 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(\mu^- e^- e^+ \bar{\nu}_\mu \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{95}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<3.6	90	ALAM	96 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(3h^- 2h^+ \geq 0 \text{ neutrals } \nu_\tau \text{ (ex. } K_S^0 \rightarrow \pi^- \pi^+ \text{) ("5-prong")})/\Gamma_{\text{total}}$   $\Gamma_{96}/\Gamma$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.  $\Gamma_{96}/\Gamma = (\Gamma_{97} + \Gamma_{98})/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.100 ± 0.006 OUR FIT**

**0.111 ± 0.008 OUR AVERAGE** Error includes scale factor of 1.1.

$0.115 \pm 0.013 \pm 0.006$	f&a	112	<sup>176</sup> ABREU	01M DLPH	1992–1995 LEP runs
$0.170 \pm 0.022 \pm 0.026$	f&a		<sup>177</sup> ACHARD	01D L3	1992–1995 LEP runs
$0.119 \pm 0.013 \pm 0.008$	avg	119	<sup>178</sup> ACKERSTAFF	99E OPAL	1991–1995 LEP runs
$0.097 \pm 0.005 \pm 0.011$	f&a	419	GIBAUT	94B CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$0.102 \pm 0.029$	f&a	13	BYLSMA	87 HRS	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.26 \pm 0.06 \pm 0.05$			ACTON	92H OPAL	$E_{\text{cm}}^{ee} = 88.2\text{--}94.2 \text{ GeV}$
$0.10 \begin{smallmatrix} +0.05 \\ -0.04 \end{smallmatrix} \pm 0.03$			DECAMP	92C ALEP	1989–1990 LEP runs
$0.16 \pm 0.13 \pm 0.04$			BEHREND	89B CELL	$E_{\text{cm}}^{ee} = 14\text{--}47 \text{ GeV}$
$0.3 \pm 0.1 \pm 0.2$			BARTEL	85F JADE	$E_{\text{cm}}^{ee} = 34.6 \text{ GeV}$
$0.13 \pm 0.04$		10	BELTRAMI	85 HRS	Repl. by BYLSMA 87
$0.16 \pm 0.08 \pm 0.04$		4	BURCHAT	85 MRK2	$E_{\text{cm}}^{ee} = 29 \text{ GeV}$
$1.0 \pm 0.4$		10	BEHREND	82 CELL	Repl. by BEHREND 89B

<sup>176</sup> The correlation coefficients between this measurement and the ABREU 01M measurements of  $B(\tau \rightarrow 1\text{-prong})$  and  $B(\tau \rightarrow 3\text{-prong})$  are  $-0.08$  and  $-0.08$  respectively.

<sup>177</sup> The correlation coefficients between this measurement and the ACHARD 01D measurements of  $B(\tau \rightarrow "1\text{-prong}")$  and  $B(\tau \rightarrow "3\text{-prong}")$  are  $-0.082$  and  $-0.19$  respectively.

<sup>178</sup> Not independent of ACKERSTAFF 99E  $B(\tau^- \rightarrow 3h^- 2h^+ \nu_\tau \text{ (ex. } K^0))$  and  $B(\tau^- \rightarrow 3h^- 2h^+ \pi^0 \nu_\tau \text{ (ex. } K^0))$  measurements.

$\Gamma(3h^- 2h^+ \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   $\Gamma_{97}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.082±0.006 OUR FIT</b>				
<b>0.076±0.007 OUR AVERAGE</b>				
0.091±0.014±0.006	97	ACKERSTAFF 99E	OPAL	1991–1995 LEP runs
0.080±0.011±0.013	58	BUSKULIC 96	ALEP	LEP 1991–1993 data
0.077±0.005±0.009	295	GIBAUT 94B	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
0.064±0.023±0.01	12	ALBRECHT 88B	ARG	$E_{\text{cm}}^{ee} = 10$ GeV
0.051±0.020	7	BYLSMA 87	HRS	$E_{\text{cm}}^{ee} = 29$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.067±0.030	5	<sup>179</sup> BELTRAMI 85	HRS	Repl. by BYLSMA 87
<sup>179</sup> The error quoted is statistical only.				

$\Gamma(3h^- 2h^+ \pi^0 \nu_\tau (\text{ex. } K^0))/\Gamma_{\text{total}}$   $\Gamma_{98}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0181±0.0027 OUR FIT</b>				
<b>0.0172±0.0027 OUR AVERAGE</b>				
0.017 ±0.002 ±0.002	231	ANASTASSOV 01	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
0.027 ±0.018 ±0.009	23	ACKERSTAFF 99E	OPAL	1991–1995 LEP runs
0.018 ±0.007 ±0.012	18	BUSKULIC 96	ALEP	LEP 1991–1993 data
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.019 ±0.004 ±0.004	31	GIBAUT 94B	CLEO	Repl. by ANAS-TASSOV 01
0.051 ±0.022	6	BYLSMA 87	HRS	$E_{\text{cm}}^{ee} = 29$ GeV
0.067 ±0.030	5	<sup>180</sup> BELTRAMI 85	HRS	Repl. by BYLSMA 87
<sup>180</sup> The error quoted is statistical only.				

$\Gamma(3h^- 2h^+ 2\pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{99}/\Gamma$

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
<0.011	90	GIBAUT 94B	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

$\Gamma((5\pi)^- \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{100}/\Gamma$

$$\Gamma_{100}/\Gamma = (\Gamma_{28} + \Gamma_{45} + \Gamma_{74} + \Gamma_{97} + 0.553\Gamma_{117} + 0.888\Gamma_{138})/\Gamma$$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.80±0.07 OUR FIT</b>				
<b>0.61±0.06±0.08</b>	<b>avg</b>	<sup>181</sup> GIBAUT 94B	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

<sup>181</sup> Not independent of GIBAUT 94B  $B(3h^- 2h^+ \nu_\tau)$ , PROCARIO 93  $B(h^- 4\pi^0 \nu_\tau)$ , and BORTOLETTO 93  $B(2h^- h^+ 2\pi^0 \nu_\tau)/B(\text{"3prong"})$  measurements. Result is corrected for  $\eta$  contributions.

$\Gamma(4h^- 3h^+ \geq 0 \text{ neutrals } \nu_\tau (\text{"7-prong"}))/\Gamma_{\text{total}}$   $\Gamma_{101}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<2.4 × 10 <sup>-6</sup>	90	EDWARDS 97B	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<1.8 × 10 <sup>-5</sup>	95	ACKERSTAFF 97J	OPAL	1990–1995 LEP runs
<2.9 × 10 <sup>-4</sup>	90	BYLSMA 87	HRS	$E_{\text{cm}}^{ee} = 29$ GeV

$$\Gamma(X^-(S=-1)\nu_\tau)/\Gamma_{\text{total}}$$

$$\Gamma_{102}/\Gamma = (\Gamma_{10} + \Gamma_{15} + \Gamma_{22} + \Gamma_{26} + \Gamma_{33} + \Gamma_{38} + \Gamma_{82} + \Gamma_{86} + \Gamma_{119})/\Gamma$$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%)		DOCUMENT ID	TECN	COMMENT
<b>2.91±0.08 OUR FIT</b>				Error includes scale factor of 1.1.
<b>2.87±0.12</b>	avg	<sup>182</sup> BARATE	99R ALEP	1991–1995 LEP runs

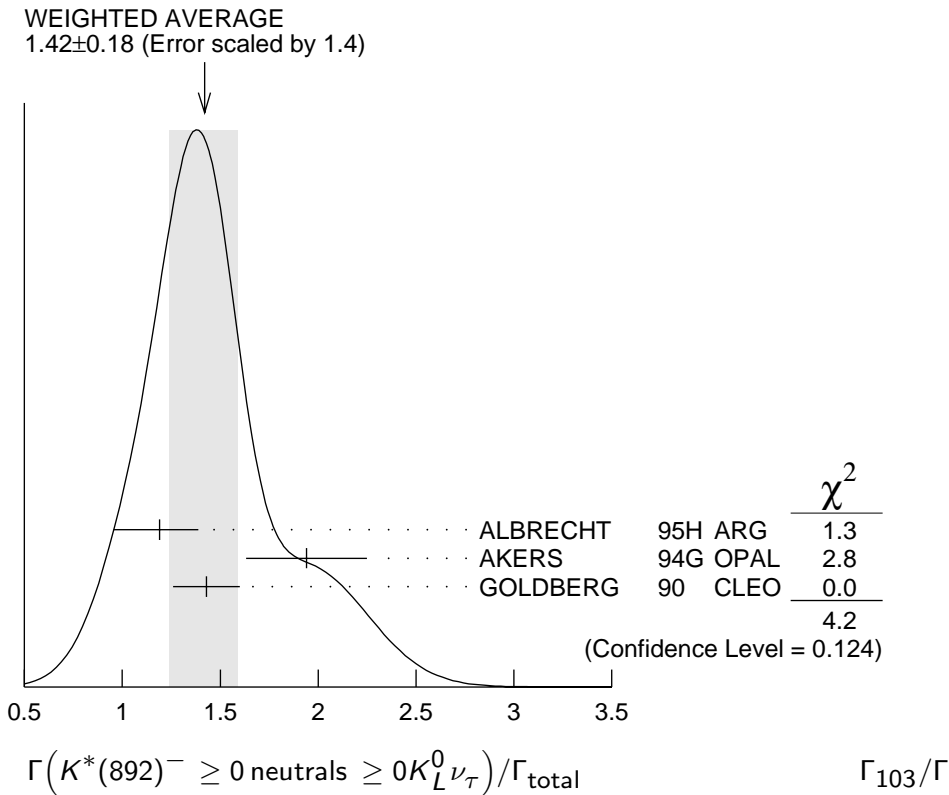
<sup>182</sup> BARATE 99R perform a combined analysis of all ALEPH LEP 1 data on  $\tau$  branching fraction measurements for decay modes having total strangeness equal to  $-1$ .

$$\Gamma(K^*(892)^- \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau)/\Gamma_{\text{total}} \quad \Gamma_{103}/\Gamma$$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.42±0.18 OUR AVERAGE</b>				Error includes scale factor of 1.4. See the ideogram below.
1.19±0.15 <sup>+0.13</sup> <sub>-0.18</sub>	104	ALBRECHT	95H ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV
1.94±0.27±0.15	74	<sup>183</sup> AKERS	94G OPAL	$E_{\text{cm}}^{ee} = 88\text{--}94$ GeV
1.43±0.11±0.13	475	<sup>184</sup> GOLDBERG	90 CLEO	$E_{\text{cm}}^{ee} = 9.4\text{--}10.9$ GeV

<sup>183</sup> AKERS 94G reject events in which a  $K_S^0$  accompanies the  $K^*(892)^-$ . We do not correct for them.

<sup>184</sup> GOLDBERG 90 estimates that 10% of observed  $K^*(892)$  are accompanied by a  $\pi^0$ .



$\Gamma(K^*(892)^-\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{104}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.29 ± 0.05 OUR AVERAGE</b>				
1.326 ± 0.063		BARATE	99R ALEP	1991–1995 LEP runs
1.11 ± 0.12		185 COAN	96 CLEO	$E_{\text{cm}}^{ee} \approx 10.6$ GeV
1.42 ± 0.22 ± 0.09		186 ACCIARRI	95F L3	1991–1993 LEP runs
1.23 ± 0.21 $\begin{smallmatrix} +0.11 \\ -0.21 \end{smallmatrix}$	54	187 ALBRECHT	88L ARG	$E_{\text{cm}}^{ee} = 10$ GeV
1.9 ± 0.3 ± 0.4	44	188 TSCHIRHART	88 HRS	$E_{\text{cm}}^{ee} = 29$ GeV
1.5 ± 0.4 ± 0.4	15	189 AIHARA	87C TPC	$E_{\text{cm}}^{ee} = 29$ GeV
1.3 ± 0.3 ± 0.3	31	YELTON	86 MRK2	$E_{\text{cm}}^{ee} = 29$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.39 ± 0.09 ± 0.10		190 BUSKULIC	96 ALEP	Repl. by BARATE 99R
1.45 ± 0.13 ± 0.11	273	191 BUSKULIC	94F ALEP	Repl. by BUSKULIC 96
1.7 ± 0.7	11	DORFAN	81 MRK2	$E_{\text{cm}}^{ee} = 4.2\text{--}6.7$ GeV

185 Not independent of COAN 96  $B(\pi^-\bar{K}^0\nu_\tau)$  and BATTLE 94  $B(K^-\pi^0\nu_\tau)$  measurements.  $K\pi$  final states are consistent with and assumed to originate from  $K^*(892)^-$  production.

186 This result is obtained from their  $B(\pi^-\bar{K}^0\nu_\tau)$  assuming all those decays originate in  $K^*(892)^-$  decays.

187 The authors divide by  $\Gamma_2/\Gamma = 0.865$  to obtain this result.

188 Not independent of TSCHIRHART 88  $\Gamma(\tau^- \rightarrow h^-\bar{K}^0 \geq 0 \text{ neutrals} \geq 0K_L^0\nu_\tau)/\Gamma(\text{total})$ .

189 Decay  $\pi^-$  identified in this experiment, is assumed in the others.

190 Not independent of BUSKULIC 96  $B(\pi^-\bar{K}^0\nu_\tau)$  and  $B(K^-\pi^0\nu_\tau)$  measurements.

191 BUSKULIC 94F obtain this result from BUSKULIC 94F  $B(\bar{K}^0\pi^-\nu_\tau)$  and BUSKULIC 94E  $B(K^-\pi^0\nu_\tau)$  assuming all of those decays originate in  $K^*(892)^-$  decays.

$\Gamma(K^*(892)^-\nu_\tau)/\Gamma(\pi^-\pi^0\nu_\tau)$   $\Gamma_{104}/\Gamma_{13}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.075 ± 0.027</b>	192 ABREU	94K DLPH	LEP 1992 Z data

192 ABREU 94K quote  $B(\tau^- \rightarrow K^*(892)^-\nu_\tau)B(K^*(892)^- \rightarrow K^-\pi^0)/B(\tau^- \rightarrow \rho^-\nu_\tau) = 0.025 \pm 0.009$ . We divide by  $B(K^*(892)^- \rightarrow K^-\pi^0) = 0.333$  to obtain this result.

$\Gamma(K^*(892)^0 K^- \geq 0 \text{ neutrals} \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{105}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.32 ± 0.08 ± 0.12</b>	119	GOLDBERG	90 CLEO	$E_{\text{cm}}^{ee} = 9.4\text{--}10.9$ GeV

$\Gamma(K^*(892)^0 K^-\nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{106}/\Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.21 ± 0.04 OUR AVERAGE</b>				

0.213 ± 0.048		193 BARATE	98 ALEP	1991–1995 LEP runs
0.20 ± 0.05 ± 0.04	47	ALBRECHT	95H ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6$ GeV

193 BARATE 98 measure the  $K^-(\rho^0 \rightarrow \pi^+\pi^-)$  fraction in  $\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau$  decays to be  $(35 \pm 11)\%$  and derive this result from their measurement of  $\Gamma(\tau^- \rightarrow K^-\pi^+\pi^-\nu_\tau)/\Gamma_{\text{total}}$  assuming the intermediate states are all  $K^-\rho$  and  $K^-K^*(892)^0$ .



$\Gamma(\bar{K}^*(892)^0 \pi^- \geq 0 \text{ neutrals } \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{107} / \Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.38 ± 0.11 ± 0.13</b>	105	GOLDBERG	90	CLEO $E_{\text{cm}}^{ee} = 9.4\text{--}10.9 \text{ GeV}$

$\Gamma(\bar{K}^*(892)^0 \pi^- \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{108} / \Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.22 ± 0.05 OUR AVERAGE</b>				
0.209 ± 0.058		<sup>194</sup> BARATE	98	ALEP 1991–1995 LEP runs
0.25 ± 0.10 ± 0.05	27	ALBRECHT	95H	ARG $E_{\text{cm}}^{ee} = 9.4\text{--}10.6 \text{ GeV}$

<sup>194</sup> BARATE 98 measure the  $K^- K^*(892)^0$  fraction in  $\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau$  decays to be  $(87 \pm 13)\%$  and derive this result from their measurement of  $\Gamma(\tau^- \rightarrow K^- K^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$ .

$\Gamma((\bar{K}^*(892)\pi)^- \nu_\tau \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{109} / \Gamma$

VALUE (%)	DOCUMENT ID	TECN	COMMENT
<b>0.10 ± 0.04 OUR AVERAGE</b>			
0.097 ± 0.044 ± 0.036	<sup>195</sup> BARATE	99K	ALEP 1991–1995 LEP runs
0.106 ± 0.037 ± 0.032	<sup>196</sup> BARATE	98E	ALEP 1991–1995 LEP runs

<sup>195</sup> BARATE 99K measure  $K^0$ 's by detecting  $K_L^0$ 's in their hadron calorimeter. They determine the  $\bar{K}^0 \rho^-$  fraction in  $\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$  decays to be  $(0.72 \pm 0.12 \pm 0.10)$  and multiply their  $B(\pi^- \bar{K}^0 \pi^0 \nu_\tau)$  measurement by one minus this fraction to obtain the quoted result.

<sup>196</sup> BARATE 98E reconstruct  $K^0$ 's using  $K_S^0 \rightarrow \pi^+ \pi^-$  decays. They determine the  $\bar{K}^0 \rho^-$  fraction in  $\tau^- \rightarrow \pi^- \bar{K}^0 \pi^0 \nu_\tau$  decays to be  $(0.64 \pm 0.09 \pm 0.10)$  and multiply their  $B(\pi^- \bar{K}^0 \pi^0 \nu_\tau)$  measurement by one minus this fraction to obtain the quoted result.

$\Gamma(K_1(1270)^- \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{110} / \Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.47 ± 0.11 OUR AVERAGE</b>				
0.48 ± 0.11		BARATE	99R	ALEP 1991–1995 LEP runs
0.41 <sup>+0.41</sup> <sub>-0.35</sub> ± 0.10	5	<sup>197</sup> BAUER	94	TPC $E_{\text{cm}}^{ee} = 29 \text{ GeV}$

<sup>197</sup> We multiply 0.41% by 0.25, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

$\Gamma(K_1(1400)^- \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{111} / \Gamma$

VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.17 ± 0.26 OUR AVERAGE</b>				Error includes scale factor of 1.7.
0.05 ± 0.17		BARATE	99R	ALEP 1991–1995 LEP runs
0.76 <sup>+0.40</sup> <sub>-0.33</sub> ± 0.20	11	<sup>198</sup> BAUER	94	TPC $E_{\text{cm}}^{ee} = 29 \text{ GeV}$

<sup>198</sup> We multiply 0.76% by 0.25, the relative systematic error quoted by BAUER 94, to obtain the systematic error.

$[\Gamma(K_1(1270)^- \nu_\tau) + \Gamma(K_1(1400)^- \nu_\tau)]/\Gamma_{\text{total}}$		$(\Gamma_{110} + \Gamma_{111})/\Gamma$			
VALUE (%)	EVTS	DOCUMENT ID	TECN	COMMENT	
$1.17^{+0.41}_{-0.37} \pm 0.29$	16	<sup>199</sup> BAUER	94	TPC	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

<sup>199</sup> We multiply 1.17% by 0.25, the relative systematic error quoted by BAUER 94, to obtain the systematic error. Not independent of BAUER 94  $B(K_1(1270)^- \nu_\tau)$  and BAUER 94  $B(K_1(1400)^- \nu_\tau)$  measurements.

$\Gamma(K_1(1270)^- \nu_\tau) / [\Gamma(K_1(1270)^- \nu_\tau) + \Gamma(K_1(1400)^- \nu_\tau)]$		$\Gamma_{110} / (\Gamma_{110} + \Gamma_{111})$			
VALUE		DOCUMENT ID	TECN	COMMENT	
<b>0.69 ± 0.15 OUR AVERAGE</b>					
$0.71 \pm 0.16 \pm 0.11$		<sup>200</sup> ABBIENDI	00D	OPAL	1990–1995 LEP runs
$0.66 \pm 0.19 \pm 0.13$		<sup>201</sup> ASNER	00B	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6 \text{ GeV}$

<sup>200</sup> ABBIENDI 00D assume the resonance structure of  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$  decays is dominated by the  $K_1(1270)^-$  and  $K_1(1400)^-$  resonances.

<sup>201</sup> ASNER 00B assume the resonance structure of  $\tau^- \rightarrow K^- \pi^+ \pi^- \nu_\tau$  (ex.  $K^0$ ) decays is dominated by  $K_1(1270)^-$  and  $K_1(1400)^-$  resonances.

$\Gamma(K^*(1410)^- \nu_\tau) / \Gamma_{\text{total}}$		$\Gamma_{112} / \Gamma$			
VALUE (units $10^{-3}$ )		DOCUMENT ID	TECN	COMMENT	
$1.5^{+1.4}_{-1.0}$		BARATE	99R	ALEP	1991–1995 LEP runs

$\Gamma(K_0^*(1430)^- \nu_\tau) / \Gamma_{\text{total}}$		$\Gamma_{113} / \Gamma$			
VALUE (units $10^{-3}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
< 0.5	95	BARATE	99R	ALEP	1991–1995 LEP runs

$\Gamma(K_2^*(1430)^- \nu_\tau) / \Gamma_{\text{total}}$		$\Gamma_{114} / \Gamma$			
VALUE (%)	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
< 0.3	95		TSCHIRHART 88	HRS	$E_{\text{cm}}^{\text{ee}} = 29 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 0.33	95	<sup>202</sup> ACCIARRI	95F	L3	1991–1993 LEP runs
< 0.9	95	0	DORFAN	81	MRK2 $E_{\text{cm}}^{\text{ee}} = 4.2\text{--}6.7 \text{ GeV}$

<sup>202</sup> ACCIARRI 95F quote  $B(\tau^- \rightarrow K^*(1430)^- \rightarrow \pi^- \bar{K}^0 \nu_\tau) < 0.11\%$ . We divide by  $B(K^*(1430)^- \rightarrow \pi^- \bar{K}^0) = 0.33$  to obtain the limit shown.

$\Gamma(a_0(980)^- \geq 0 \text{ neutrals } \nu_\tau) / \Gamma_{\text{total}} \times B(a_0(980) \rightarrow K^0 K^-)$		$\Gamma_{115} / \Gamma \times B$			
VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT	
< 2.8	90	GOLDBERG	90	CLEO	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.9 \text{ GeV}$

$\Gamma(\eta \pi^- \nu_\tau) / \Gamma_{\text{total}}$		$\Gamma_{116} / \Gamma$			
VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
< 1.4	95	0	BARTELT	96	CLEO $E_{\text{cm}}^{\text{ee}} \approx 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 6.2	95		BUSKULIC	97C ALEP	1991–1994 LEP
< 3.4	95		ARTUSO	92 CLEO	$E_{cm}^{ee} \approx 10.6$ GeV runs
< 90	95		ALBRECHT	88M ARG	$E_{cm}^{ee} \approx 10$ GeV
<140	90		BEHREND	88 CELL	$E_{cm}^{ee} = 14–46.8$ GeV
<180	95		BARINGER	87 CLEO	$E_{cm}^{ee} = 10.5$ GeV
<250	90	0	COFFMAN	87 MRK3	$E_{cm}^{ee} = 3.77$ GeV
510 $\pm 100 \pm 120$		65	DERRICK	87 HRS	$E_{cm}^{ee} = 29$ GeV
<100	95		GAN	87B MRK2	$E_{cm}^{ee} = 29$ GeV

$\Gamma(\eta\pi^-\pi^0\nu_\tau)/\Gamma_{total}$   $\Gamma_{117}/\Gamma$

<u>VALUE (%)</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.174±0.024 OUR FIT</b>					
<b>0.173±0.024 OUR AVERAGE</b>					
0.18 ±0.04 ±0.02			BUSKULIC	97C ALEP	1991–1994 LEP
0.17 ±0.02 ±0.02		125	ARTUSO	92 CLEO	$E_{cm}^{ee} \approx 10.6$ runs GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<1.10	95		ALBRECHT	88M ARG	$E_{cm}^{ee} \approx 10$ GeV
<2.10	95		BARINGER	87 CLEO	$E_{cm}^{ee} = 10.5$ GeV
4.20 $\begin{smallmatrix} +0.70 \\ -1.20 \end{smallmatrix} \pm 1.60$		203	GAN	87 MRK2	$E_{cm}^{ee} = 29$ GeV

<sup>203</sup> Highly correlated with GAN 87  $\Gamma(\pi^- 3\pi^0\nu_\tau)/\Gamma(total)$  value.

$\Gamma(\eta\pi^-\pi^0\pi^0\nu_\tau)/\Gamma_{total}$   $\Gamma_{118}/\Gamma$

<u>VALUE (units 10<sup>-4</sup>)</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.5±0.5</b>					
		30	<sup>204</sup> ANASTASSOV 01	CLEO	$E_{cm}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.4±0.6±0.3		15	<sup>205</sup> BERGFELD	97 CLEO	Repl. by ANAS- TASSOV 01
< 4.3	95		ARTUSO	92 CLEO	$E_{cm}^{ee} \approx 10.6$ GeV
<120	95		ALBRECHT	88M ARG	$E_{cm}^{ee} \approx 10$ GeV

<sup>204</sup> Weighted average of BERGFELD 97 and ANASTASSOV 01 value of  $(1.5 \pm 0.6 \pm 0.3) \times 10^{-4}$  obtained using  $\eta$ 's reconstructed from  $\eta \rightarrow \pi^+\pi^-\pi^0$  decays.

<sup>205</sup> BERGFELD 97 reconstruct  $\eta$ 's using  $\eta \rightarrow \gamma\gamma$  decays.

$\Gamma(\eta K^- \nu_\tau)/\Gamma_{total}$   $\Gamma_{119}/\Gamma$

<u>VALUE (units 10<sup>-4</sup>)</u>	<u>CL%</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>2.7±0.6 OUR FIT</b>					
<b>2.7±0.6 OUR AVERAGE</b>					
2.9 $\begin{smallmatrix} +1.3 \\ -1.2 \end{smallmatrix} \pm 0.7$			BUSKULIC	97C ALEP	1991–1994 LEP runs
2.6±0.5±0.5		85	BARTELT	96 CLEO	$E_{cm}^{ee} \approx 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<4.7	95		ARTUSO	92 CLEO	$E_{cm}^{ee} \approx 10.6$ GeV
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$\Gamma(\eta K^*(892)^- \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{120} / \Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.90 ± 0.80 ± 0.42</b>	25	BISHAI	99	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV

$\Gamma(\eta K^- \pi^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{121} / \Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.77 ± 0.56 ± 0.71</b>	36	BISHAI	99	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV

$\Gamma(\eta \bar{K}^0 \pi^- \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{122} / \Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.20 ± 0.70 ± 0.22</b>	15	<sup>206</sup> BISHAI	99	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV

<sup>206</sup>We multiply the BISHAI 99 measurement  $B(\tau^- \rightarrow \eta K_S^0 \pi^- \nu_\tau) = (1.10 \pm 0.35 \pm 0.11) \times 10^{-4}$  by 2 to obtain the listed value.

$\Gamma(\eta \pi^+ \pi^- \pi^- \geq 0 \text{ neutrals } \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{123} / \Gamma$

VALUE (%)	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.3</b>	90	ABACHI	87B	HRS $E_{\text{cm}}^{ee} = 29$ GeV

$\Gamma(\eta \pi^- \pi^+ \pi^- \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{124} / \Gamma$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>2.3 ± 0.5</b>	170	<sup>207</sup> ANASTASSOV 01	CLEO	$E_{\text{cm}}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$3.4^{+0.6}_{-0.5} \pm 0.6$	89	<sup>208</sup> BERGFELD	97	CLEO Repl. by ANASTASSOV 01
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<sup>207</sup>Weighted average of BERGFELD 97 and ANASTASSOV 01 measurements using  $\eta$ 's reconstructed from  $\eta \rightarrow \pi^+ \pi^- \pi^0$  and  $\eta \rightarrow 3\pi^0$  decays.

<sup>208</sup>BERGFELD 97 reconstruct  $\eta$ 's using  $\eta \rightarrow \gamma\gamma$  and  $\eta \rightarrow 3\pi^0$  decays.

$\Gamma(\eta a_1(1260)^- \nu_\tau \rightarrow \eta \pi^- \rho^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{125} / \Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 3.9 × 10<sup>-4</sup></b>	90	BERGFELD	97	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV

$\Gamma(\eta \eta \pi^- \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{126} / \Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 1.1</b>	95	ARTUSO	92	CLEO $E_{\text{cm}}^{ee} \approx 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 83	95	ALBRECHT	88M	ARG $E_{\text{cm}}^{ee} \approx 10$ GeV
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$\Gamma(\eta \eta \pi^- \pi^0 \nu_\tau) / \Gamma_{\text{total}}$   $\Gamma_{127} / \Gamma$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 2.0</b>	95	ARTUSO	92	CLEO $E_{\text{cm}}^{ee} \approx 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

< 90	95	ALBRECHT	88M	ARG $E_{\text{cm}}^{ee} \approx 10$ GeV
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$\Gamma(\eta'(958)\pi^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{128}/\Gamma$
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
$<7.4 \times 10^{-5}$	90	BERGFELD	97	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$\Gamma(\eta'(958)\pi^-\pi^0\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{129}/\Gamma$
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
$<8.0 \times 10^{-5}$	90	BERGFELD	97	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$\Gamma(\phi\pi^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{130}/\Gamma$
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
$<2.0 \times 10^{-4}$	90	<sup>209</sup> AVERY	97	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$<3.5 \times 10^{-4}$	90	ALBRECHT	95H	ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6 \text{ GeV}$
<sup>209</sup> AVERY 97 limit varies from $(1.2\text{--}2.0) \times 10^{-4}$ depending on decay model assumptions.					
$\Gamma(\phi K^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{131}/\Gamma$
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
$<6.7 \times 10^{-5}$	90	<sup>210</sup> AVERY	97	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
<sup>210</sup> AVERY 97 limit varies from $(5.4\text{--}6.7) \times 10^{-5}$ depending on decay model assumptions.					
$\Gamma(f_1(1285)\pi^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{132}/\Gamma$
<u>VALUE (units <math>10^{-4}</math>)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
$5.8^{+1.4}_{-1.3} \pm 1.8$	54	BERGFELD	97	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$\Gamma(f_1(1285)\pi^-\nu_\tau \rightarrow \eta\pi^-\pi^+\pi^-\nu_\tau)/\Gamma(\eta\pi^-\pi^+\pi^-\nu_\tau)$					$\Gamma_{133}/\Gamma_{124}$
<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>		
$0.55 \pm 0.14$	BERGFELD	97	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$	
$\Gamma(\pi(1300)^-\nu_\tau \rightarrow (\rho\pi)^-\nu_\tau \rightarrow (3\pi)^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{134}/\Gamma$
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
$<1.0 \times 10^{-4}$	90	ASNER	00	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$\Gamma(\pi(1300)^-\nu_\tau \rightarrow ((\pi\pi)_{\text{S-wave}}\pi)^-\nu_\tau \rightarrow (3\pi)^-\nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{135}/\Gamma$
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
$<1.9 \times 10^{-4}$	90	ASNER	00	CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
$\Gamma(h^-\omega \geq 0 \text{ neutrals } \nu_\tau)/\Gamma_{\text{total}}$					$\Gamma_{136}/\Gamma$
$\Gamma_{136}/\Gamma = (\Gamma_{137} + \Gamma_{138})/\Gamma$					
Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.					
<u>VALUE (%)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>2.38 ± 0.08 OUR FIT</b>					
<b>1.65 ± 0.3 ± 0.2 avg</b>	1513	ALBRECHT	88M	ARG	$E_{\text{cm}}^{ee} \approx 10 \text{ GeV}$

$\Gamma(h^- \omega \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{137}/\Gamma$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE (%)		EVTS	DOCUMENT ID	TECN	COMMENT
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**1.94±0.07 OUR FIT**

**1.92±0.07 OUR AVERAGE**

1.91±0.07±0.06	f&a	5803	BUSKULIC	97C ALEP	1991–1994 LEP
1.95±0.07±0.11	avg	2223	<sup>211</sup> BALEST	95C CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6$ GeV
1.60±0.27±0.41	f&a	139	BARINGER	87 CLEO	$E_{\text{cm}}^{\text{ee}} = 10.5$ GeV

<sup>211</sup> Not independent of BALEST 95C  $B(\tau^- \rightarrow h^- \omega \nu_\tau)/B(\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau)$  value.

$\Gamma(h^- \omega \nu_\tau)/\Gamma(h^- h^- h^+ \pi^0 \nu_\tau (\text{ex. } K^0))$   $\Gamma_{137}/\Gamma_{64}$

$$\Gamma_{137}/\Gamma_{64} = \Gamma_{137}/(\Gamma_{68} + \Gamma_{86} + \Gamma_{90} + 0.231\Gamma_{119} + 0.888\Gamma_{137} + 0.0221\Gamma_{138})$$

VALUE		EVTS	DOCUMENT ID	TECN	COMMENT
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**0.446±0.015 OUR FIT**

**0.453±0.019 OUR AVERAGE**

0.431±0.033		2350	<sup>212</sup> BUSKULIC	96 ALEP	LEP 1991–1993 data
0.464±0.016±0.017		2223	<sup>213</sup> BALEST	95C CLEO	$E_{\text{cm}}^{\text{ee}} \approx 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.37 ±0.05 ±0.02		458	<sup>214</sup> ALBRECHT	91D ARG	$E_{\text{cm}}^{\text{ee}} = 9.4\text{--}10.6$ GeV
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<sup>212</sup> BUSKULIC 96 quote the fraction of  $\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau$  (ex.  $K^0$ ) decays which originate in a  $h^- \omega$  final state =  $0.383 \pm 0.029$ . We divide this by the  $\omega(782) \rightarrow \pi^+ \pi^- \pi^0$  branching fraction (0.888).

<sup>213</sup> BALEST 95C quote the fraction of  $\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau$  (ex.  $K^0$ ) decays which originate in a  $h^- \omega$  final state equals  $0.412 \pm 0.014 \pm 0.015$ . We divide this by the  $\omega(782) \rightarrow \pi^+ \pi^- \pi^0$  branching fraction (0.888).

<sup>214</sup> ALBRECHT 91D quote the fraction of  $\tau^- \rightarrow h^- h^- h^+ \pi^0 \nu_\tau$  decays which originate in a  $\pi^- \omega$  final state equals  $0.33 \pm 0.04 \pm 0.02$ . We divide this by the  $\omega(782) \rightarrow \pi^+ \pi^- \pi^0$  branching fraction (0.888).

$\Gamma(h^- \omega \pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{138}/\Gamma$

VALUE (%)		EVTS	DOCUMENT ID	TECN	COMMENT
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**0.44±0.05 OUR FIT**

<b>0.43±0.06±0.05</b>		7283	BUSKULIC	97C ALEP	1991–1994 LEP runs
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$\Gamma(h^- \omega 2\pi^0 \nu_\tau)/\Gamma_{\text{total}}$   $\Gamma_{139}/\Gamma$

VALUE (units $10^{-4}$ )		EVTS	DOCUMENT ID	TECN	COMMENT
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<b>1.4 ±0.4 ±0.3</b>		53	ANASTASSOV 01	CLEO	$E_{\text{cm}}^{\text{ee}} = 10.6$ GeV
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• • • We do not use the following data for averages, fits, limits, etc. • • •

1.89 <sup>+0.74</sup> <sub>-0.67</sub> ±0.40		19	ANDERSON	97 CLEO	Repl. by ANASTASSOV 01
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$$\Gamma(h^- \omega \pi^0 \nu_\tau) / \Gamma(h^- h^- h^+ \geq 0 \text{ neutrals} \geq 0 K_L^0 \nu_\tau) \quad \Gamma_{138} / \Gamma_{52}$$

$$\Gamma_{138} / \Gamma_{52} = \Gamma_{138} / (0.3431\Gamma_{33} + 0.3431\Gamma_{35} + 0.3431\Gamma_{38} + 0.3431\Gamma_{40} + 0.4307\Gamma_{45} + 0.6861\Gamma_{46} + \Gamma_{60} + \Gamma_{68} + \Gamma_{74} + \Gamma_{75} + \Gamma_{82} + \Gamma_{86} + \Gamma_{89} + \Gamma_{90} + 0.285\Gamma_{117} + 0.285\Gamma_{119} + 0.9101\Gamma_{137} + 0.9101\Gamma_{138})$$

Data marked "avg" are highly correlated with data appearing elsewhere in the Listings, and are therefore used for the average given below but not in the overall fits. "f&a" marks results used for the fit and the average.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0286 ± 0.0031 OUR FIT</b>				
<b>0.028 ± 0.003 ± 0.003</b>	avg	430	215 BORTOLETTO93	CLEO $E_{cm}^{ee} \approx 10.6$ GeV

215 Not independent of BORTOLETTO 93  $\Gamma(\tau^- \rightarrow h^- h^- h^+ 2\pi^0 \nu_\tau \text{ (ex. } K^0))$  value.

$$\Gamma(h^- \omega \pi^0 \nu_\tau) / \Gamma(h^- h^- h^+ 2\pi^0 \nu_\tau \text{ (ex. } K^0)) \quad \Gamma_{138} / \Gamma_{73}$$

$$\Gamma_{138} / \Gamma_{73} = \Gamma_{138} / (\Gamma_{74} + 0.236\Gamma_{117} + 0.888\Gamma_{138})$$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.81 ± 0.08 OUR FIT</b>			
<b>0.81 ± 0.06 ± 0.06</b>	BORTOLETTO93	CLEO	$E_{cm}^{ee} \approx 10.6$ GeV

$$\Gamma(2h^- h^+ \omega \nu_\tau) / \Gamma_{total} \quad \Gamma_{140} / \Gamma$$

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.2 ± 0.2 ± 0.1</b>	110	ANASTASSOV 01	CLEO	$E_{cm}^{ee} = 10.6$ GeV

$$\Gamma(e^- \gamma) / \Gamma_{total} \quad \Gamma_{141} / \Gamma$$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 2.7 × 10<sup>-6</sup></b>	90	EDWARDS	97 CLEO	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
< 1.1 × 10 <sup>-4</sup>	90	ABREU	95U DLPH	1990–1993 LEP runs
< 1.2 × 10 <sup>-4</sup>	90	ALBRECHT	92K ARG	$E_{cm}^{ee} = 10$ GeV
< 2.0 × 10 <sup>-4</sup>	90	KEH	88 CBAL	$E_{cm}^{ee} = 10$ GeV
< 6.4 × 10 <sup>-4</sup>	90	HAYES	82 MRK2	$E_{cm}^{ee} = 3.8–6.8$ GeV

$$\Gamma(\mu^- \gamma) / \Gamma_{total} \quad \Gamma_{142} / \Gamma$$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 1.1 × 10<sup>-6</sup></b>	90	AHMED	00 CLEO	$E_{cm}^{ee} = 10.6$ GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
< 3.0 × 10 <sup>-6</sup>	90	EDWARDS	97 CLEO	
< 6.2 × 10 <sup>-5</sup>	90	ABREU	95U DLPH	1990–1993 LEP runs
< 0.42 × 10 <sup>-5</sup>	90	BEAN	93 CLEO	$E_{cm}^{ee} = 10.6$ GeV
< 3.4 × 10 <sup>-5</sup>	90	ALBRECHT	92K ARG	$E_{cm}^{ee} = 10$ GeV
< 55 × 10 <sup>-5</sup>	90	HAYES	82 MRK2	$E_{cm}^{ee} = 3.8–6.8$ GeV

$\Gamma(e^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{143}/\Gamma$

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$< 3.7 \times 10^{-6}$	90	BONVICINI	97 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 17 \times 10^{-5}$	90	ALBRECHT	92k ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 14 \times 10^{-5}$	90	KEH	88 CBAL	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 210 \times 10^{-5}$	90	HAYES	82 MRK2	$E_{\text{cm}}^{ee} = 3.8\text{--}6.8 \text{ GeV}$

$\Gamma(\mu^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{144}/\Gamma$

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$< 4.0 \times 10^{-6}$	90	BONVICINI	97 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 4.4 \times 10^{-5}$	90	ALBRECHT	92k ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 82 \times 10^{-5}$	90	HAYES	82 MRK2	$E_{\text{cm}}^{ee} = 3.8\text{--}6.8 \text{ GeV}$

$\Gamma(e^- K_S^0)/\Gamma_{\text{total}}$   $\Gamma_{145}/\Gamma$

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$< 9.1 \times 10^{-7}$	90	CHEN	02c CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.3 \times 10^{-3}$	90	HAYES	82 MRK2	$E_{\text{cm}}^{ee} = 3.8\text{--}6.8 \text{ GeV}$

$\Gamma(\mu^- K_S^0)/\Gamma_{\text{total}}$   $\Gamma_{146}/\Gamma$

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$< 9.5 \times 10^{-7}$	90	CHEN	02c CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.0 \times 10^{-3}$	90	HAYES	82 MRK2	$E_{\text{cm}}^{ee} = 3.8\text{--}6.8 \text{ GeV}$

$\Gamma(e^- \eta)/\Gamma_{\text{total}}$   $\Gamma_{147}/\Gamma$

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$< 8.2 \times 10^{-6}$	90	BONVICINI	97 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 6.3 \times 10^{-5}$	90	ALBRECHT	92k ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 24 \times 10^{-5}$	90	KEH	88 CBAL	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$

$\Gamma(\mu^- \eta)/\Gamma_{\text{total}}$   $\Gamma_{148}/\Gamma$

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$< 9.6 \times 10^{-6}$	90	BONVICINI	97 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 7.3 \times 10^{-5}$	90	ALBRECHT	92k ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$



$\Gamma(e^- \rho^0)/\Gamma_{\text{total}}$   $\Gamma_{149}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 2.0 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 0.42 \times 10^{-5}$	90	<sup>216</sup> BARTELT	94	CLEO Repl. by BLISS 98
$< 1.9 \times 10^{-5}$	90	ALBRECHT	92k	ARG $E_{\text{cm}}^{ee} = 10$ GeV
$< 37 \times 10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{ee} = 3.8\text{--}6.8$ GeV

<sup>216</sup> BARTELT 94 assume phase space decays.

$\Gamma(\mu^- \rho^0)/\Gamma_{\text{total}}$   $\Gamma_{150}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 6.3 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 0.57 \times 10^{-5}$	90	<sup>217</sup> BARTELT	94	CLEO Repl. by BLISS 98
$< 2.9 \times 10^{-5}$	90	ALBRECHT	92k	ARG $E_{\text{cm}}^{ee} = 10$ GeV
$< 44 \times 10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{ee} = 3.8\text{--}6.8$ GeV

<sup>217</sup> BARTELT 94 assume phase space decays.

$\Gamma(e^- K^*(892)^0)/\Gamma_{\text{total}}$   $\Gamma_{151}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 5.1 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 0.63 \times 10^{-5}$	90	<sup>218</sup> BARTELT	94	CLEO Repl. by BLISS 98
$< 3.8 \times 10^{-5}$	90	ALBRECHT	92k	ARG $E_{\text{cm}}^{ee} = 10$ GeV

<sup>218</sup> BARTELT 94 assume phase space decays.

$\Gamma(\mu^- K^*(892)^0)/\Gamma_{\text{total}}$   $\Gamma_{152}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 7.5 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 0.94 \times 10^{-5}$	90	<sup>219</sup> BARTELT	94	CLEO Repl. by BLISS 98
$< 4.5 \times 10^{-5}$	90	ALBRECHT	92k	ARG $E_{\text{cm}}^{ee} = 10$ GeV

<sup>219</sup> BARTELT 94 assume phase space decays.

$\Gamma(e^- \bar{K}^*(892)^0)/\Gamma_{\text{total}}$   $\Gamma_{153}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 7.4 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 1.1 \times 10^{-5}$	90	<sup>220</sup> BARTELT	94	CLEO Repl. by BLISS 98

<sup>220</sup> BARTELT 94 assume phase space decays.

$\Gamma(\mu^- \bar{K}^*(892)^0)/\Gamma_{\text{total}}$   $\Gamma_{154}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<7.5 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<0.87 \times 10^{-5}$	90	<sup>221</sup> BARTELT	94	CLEO Repl. by BLISS 98
<sup>221</sup> BARTELT 94 assume phase space decays.				

$\Gamma(e^- \phi)/\Gamma_{\text{total}}$   $\Gamma_{155}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<6.9 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV

$\Gamma(\mu^- \phi)/\Gamma_{\text{total}}$   $\Gamma_{156}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<7.0 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV

$\Gamma(e^- e^+ e^-)/\Gamma_{\text{total}}$   $\Gamma_{157}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 2.9 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 0.33 \times 10^{-5}$	90	<sup>222</sup> BARTELT	94	CLEO Repl. by BLISS 98
$< 1.3 \times 10^{-5}$	90	ALBRECHT	92k	ARG $E_{\text{cm}}^{ee} = 10$ GeV
$< 2.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{ee} = 10.4-10.9$
$<40 \times 10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{ee} = 3.8-6.8$ GeV
<sup>222</sup> BARTELT 94 assume phase space decays.				

$\Gamma(e^- \mu^+ \mu^-)/\Gamma_{\text{total}}$   $\Gamma_{158}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 1.8 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$< 0.36 \times 10^{-5}$	90	<sup>223</sup> BARTELT	94	CLEO Repl. by BLISS 98
$< 1.9 \times 10^{-5}$	90	ALBRECHT	92k	ARG $E_{\text{cm}}^{ee} = 10$ GeV
$< 2.7 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{ee} = 10.4-10.9$
$<33 \times 10^{-5}$	90	HAYES	82	MRK2 $E_{\text{cm}}^{ee} = 3.8-6.8$ GeV
<sup>223</sup> BARTELT 94 assume phase space decays.				

$\Gamma(e^+ \mu^- \mu^-)/\Gamma_{\text{total}}$   $\Gamma_{159}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<1.5 \times 10^{-6}$	90	BLISS	98	CLEO $E_{\text{cm}}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<0.35 \times 10^{-5}$	90	<sup>224</sup> BARTELT	94	CLEO Repl. by BLISS 98
$<1.8 \times 10^{-5}$	90	ALBRECHT	92k	ARG $E_{\text{cm}}^{ee} = 10$ GeV
$<1.6 \times 10^{-5}$	90	BOWCOCK	90	CLEO $E_{\text{cm}}^{ee} = 10.4-10.9$
<sup>224</sup> BARTELT 94 assume phase space decays.				

**$\Gamma(\mu^- e^+ e^-)/\Gamma_{\text{total}}$**   **$\Gamma_{160}/\Gamma$**

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt; 1.7 \times 10^{-6}</math></b>	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$< 0.34 \times 10^{-5}$	90	<sup>225</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$< 1.4 \times 10^{-5}$	90	ALBRECHT	92k ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 2.7 \times 10^{-5}$	90	BOWCOCK	90 CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$
$< 44 \times 10^{-5}$	90	HAYES	82 MRK2	$E_{\text{cm}}^{ee} = 3.8\text{--}6.8 \text{ GeV}$

<sup>225</sup> BARTELT 94 assume phase space decays.

**$\Gamma(\mu^+ e^- e^-)/\Gamma_{\text{total}}$**   **$\Gamma_{161}/\Gamma$**

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt; 1.5 \times 10^{-6}</math></b>	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$< 0.34 \times 10^{-5}$	90	<sup>226</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$< 1.4 \times 10^{-5}$	90	ALBRECHT	92k ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 1.6 \times 10^{-5}$	90	BOWCOCK	90 CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>226</sup> BARTELT 94 assume phase space decays.

**$\Gamma(\mu^- \mu^+ \mu^-)/\Gamma_{\text{total}}$**   **$\Gamma_{162}/\Gamma$**

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt; 1.9 \times 10^{-6}</math></b>	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$< 0.43 \times 10^{-5}$	90	<sup>227</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$< 1.9 \times 10^{-5}$	90	ALBRECHT	92k ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 1.7 \times 10^{-5}$	90	BOWCOCK	90 CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$
$< 49 \times 10^{-5}$	90	HAYES	82 MRK2	$E_{\text{cm}}^{ee} = 3.8\text{--}6.8 \text{ GeV}$

<sup>227</sup> BARTELT 94 assume phase space decays.

**$\Gamma(e^- \pi^+ \pi^-)/\Gamma_{\text{total}}$**   **$\Gamma_{163}/\Gamma$**

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt; 2.2 \times 10^{-6}</math></b>	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$< 0.44 \times 10^{-5}$	90	<sup>228</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$< 2.7 \times 10^{-5}$	90	ALBRECHT	92k ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$< 6.0 \times 10^{-5}$	90	BOWCOCK	90 CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>228</sup> BARTELT 94 assume phase space decays.

**$\Gamma(e^+ \pi^- \pi^-)/\Gamma_{\text{total}}$**   **$\Gamma_{164}/\Gamma$**

Test of lepton number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt;1.9 \times 10^{-6}</math></b>	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$<0.44 \times 10^{-5}$	90	<sup>229</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$<1.8 \times 10^{-5}$	90	ALBRECHT	92k ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$<1.7 \times 10^{-5}$	90	BOWCOCK	90 CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>229</sup> BARTELT 94 assume phase space decays.

**$\Gamma(\mu^- \pi^+ \pi^-)/\Gamma_{\text{total}}$**   **$\Gamma_{165}/\Gamma$**

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt;8.2 \times 10^{-6}</math></b>	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$<0.74 \times 10^{-5}$	90	<sup>230</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$<3.6 \times 10^{-5}$	90	ALBRECHT	92k ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$<3.9 \times 10^{-5}$	90	BOWCOCK	90 CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>230</sup> BARTELT 94 assume phase space decays.

**$\Gamma(\mu^+ \pi^- \pi^-)/\Gamma_{\text{total}}$**   **$\Gamma_{166}/\Gamma$**

Test of lepton number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt;3.4 \times 10^{-6}</math></b>	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$<0.69 \times 10^{-5}$	90	<sup>231</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$<6.3 \times 10^{-5}$	90	ALBRECHT	92k ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$<3.9 \times 10^{-5}$	90	BOWCOCK	90 CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>231</sup> BARTELT 94 assume phase space decays.

**$\Gamma(e^- \pi^+ K^-)/\Gamma_{\text{total}}$**   **$\Gamma_{167}/\Gamma$**

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt;6.4 \times 10^{-6}</math></b>	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$<0.77 \times 10^{-5}$	90	<sup>232</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$<2.9 \times 10^{-5}$	90	ALBRECHT	92k ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$<5.8 \times 10^{-5}$	90	BOWCOCK	90 CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>232</sup> BARTELT 94 assume phase space decays.

**$\Gamma(e^- \pi^- K^+)/\Gamma_{\text{total}}$**   **$\Gamma_{168}/\Gamma$**

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b><math>&lt;3.8 \times 10^{-6}</math></b>	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$<0.46 \times 10^{-5}$	90	<sup>233</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$<5.8 \times 10^{-5}$	90	BOWCOCK	90 CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>233</sup> BARTELT 94 assume phase space decays.

$\Gamma(e^+ \pi^- K^-)/\Gamma_{\text{total}}$   $\Gamma_{169}/\Gamma$

Test of lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.1 \times 10^{-6}$	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<0.45 \times 10^{-5}$	90	<sup>234</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$<2.0 \times 10^{-5}$	90	ALBRECHT	92k ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$<4.9 \times 10^{-5}$	90	BOWCOCK	90 CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>234</sup> BARTELT 94 assume phase space decays.

$\Gamma(e^- K_S^0 K_S^0)/\Gamma_{\text{total}}$   $\Gamma_{170}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<2.2 \times 10^{-6}$	90	CHEN	02c CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(e^- K^+ K^-)/\Gamma_{\text{total}}$   $\Gamma_{171}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<6.0 \times 10^{-6}$	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(e^+ K^- K^-)/\Gamma_{\text{total}}$   $\Gamma_{172}/\Gamma$

Test of lepton number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<3.8 \times 10^{-6}$	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(\mu^- \pi^+ K^-)/\Gamma_{\text{total}}$   $\Gamma_{173}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<7.5 \times 10^{-6}$	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<0.87 \times 10^{-5}$	90	<sup>235</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$<11 \times 10^{-5}$	90	ALBRECHT	92k ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$<7.7 \times 10^{-5}$	90	BOWCOCK	90 CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>235</sup> BARTELT 94 assume phase space decays.

$\Gamma(\mu^- \pi^- K^+)/\Gamma_{\text{total}}$   $\Gamma_{174}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<7.4 \times 10^{-6}$	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<1.5 \times 10^{-5}$	90	<sup>236</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$<7.7 \times 10^{-5}$	90	BOWCOCK	90 CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>236</sup> BARTELT 94 assume phase space decays.

$\Gamma(\mu^+ \pi^- K^-)/\Gamma_{\text{total}}$   $\Gamma_{175}/\Gamma$

Test of lepton number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<7.0 \times 10^{-6}$	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$<2.0 \times 10^{-5}$	90	<sup>237</sup> BARTELT	94 CLEO	Repl. by BLISS 98
$<5.8 \times 10^{-5}$	90	ALBRECHT	92k ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
$<4.0 \times 10^{-5}$	90	BOWCOCK	90 CLEO	$E_{\text{cm}}^{ee} = 10.4\text{--}10.9$

<sup>237</sup> BARTELT 94 assume phase space decays.

$\Gamma(\mu^- K_S^0 K_S^0)/\Gamma_{\text{total}}$   $\Gamma_{176}/\Gamma$

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<3.4 \times 10^{-6}$	90	CHEN	02c CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(\mu^- K^+ K^-)/\Gamma_{\text{total}}$   $\Gamma_{177}/\Gamma$

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<15 \times 10^{-6}$	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(\mu^+ K^- K^-)/\Gamma_{\text{total}}$   $\Gamma_{178}/\Gamma$

Test of lepton number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<6.0 \times 10^{-6}$	90	BLISS	98 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(e^- \pi^0 \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{179}/\Gamma$

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<6.5 \times 10^{-6}$	90	BONVICINI	97 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(\mu^- \pi^0 \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{180}/\Gamma$

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<14 \times 10^{-6}$	90	BONVICINI	97 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(e^- \eta \eta)/\Gamma_{\text{total}}$   $\Gamma_{181}/\Gamma$

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<35 \times 10^{-6}$	90	BONVICINI	97 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(\mu^- \eta \eta)/\Gamma_{\text{total}}$   $\Gamma_{182}/\Gamma$

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<60 \times 10^{-6}$	90	BONVICINI	97 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(e^- \pi^0 \eta)/\Gamma_{\text{total}}$   $\Gamma_{183}/\Gamma$

Test of lepton family number conservation.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$<24 \times 10^{-6}$	90	BONVICINI	97 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(\mu^- \pi^0 \eta)/\Gamma_{\text{total}}$   $\Gamma_{184}/\Gamma$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 22 \times 10^{-6}$	90	BONVICINI	97 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(\bar{p}\gamma)/\Gamma_{\text{total}}$   $\Gamma_{185}/\Gamma$

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 3.5 \times 10^{-6}$	90	GODANG	99 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 29 \times 10^{-5}$	90	ALBRECHT	92K ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
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$\Gamma(\bar{p}\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{186}/\Gamma$

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 15 \times 10^{-6}$	90	GODANG	99 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 66 \times 10^{-5}$	90	ALBRECHT	92K ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
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$\Gamma(\bar{p}2\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{187}/\Gamma$

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 33 \times 10^{-6}$	90	GODANG	99 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(\bar{p}\eta)/\Gamma_{\text{total}}$   $\Gamma_{188}/\Gamma$

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 8.9 \times 10^{-6}$	90	GODANG	99 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 130 \times 10^{-5}$	90	ALBRECHT	92K ARG	$E_{\text{cm}}^{ee} = 10 \text{ GeV}$
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$\Gamma(\bar{p}\pi^0\eta)/\Gamma_{\text{total}}$   $\Gamma_{189}/\Gamma$

Test of lepton number and baryon number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 27 \times 10^{-6}$	90	GODANG	99 CLEO	$E_{\text{cm}}^{ee} = 10.6 \text{ GeV}$

$\Gamma(e^- \text{light boson})/\Gamma(e^- \bar{\nu}_e \nu_\tau)$   $\Gamma_{190}/\Gamma_5$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$< 0.015$	95	238 ALBRECHT	95G ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6 \text{ GeV}$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 0.018$	95	239 ALBRECHT	90E ARG	$E_{\text{cm}}^{ee} = 9.4\text{--}10.6 \text{ GeV}$
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$< 0.040$	95	240 BALTRUSAITIS	85 MRK3	$E_{\text{cm}}^{ee} = 3.77 \text{ GeV}$
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<sup>238</sup> ALBRECHT 95G limit holds for bosons with mass  $< 0.4 \text{ GeV}$ . The limit rises to 0.036 for a mass of 1.0 GeV, then falls to 0.006 at the upper mass limit of 1.6 GeV.

<sup>239</sup> ALBRECHT 90E limit applies for spinless boson with mass  $< 100 \text{ MeV}$ , and rises to 0.050 for mass = 500 MeV.

<sup>240</sup> BALTRUSAITIS 85 limit applies for spinless boson with mass  $< 100 \text{ MeV}$ .

$\Gamma(\mu^- \text{ light boson})/\Gamma(e^- \bar{\nu}_e \nu_\tau)$   $\Gamma_{191}/\Gamma_5$

Test of lepton family number conservation.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.026</b>	95	<sup>241</sup> ALBRECHT	95G ARG	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.033	95	<sup>242</sup> ALBRECHT	90E ARG	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV
<0.125	95	<sup>243</sup> BALTRUSAITIS	85 MRK3	$E_{cm}^{ee} = 3.77$ GeV

<sup>241</sup> ALBRECHT 95G limit holds for bosons with mass < 1.3 GeV. The limit rises to 0.034 for a mass of 1.4 GeV, then falls to 0.003 at the upper mass limit of 1.6 GeV.

<sup>242</sup> ALBRECHT 90E limit applies for spinless boson with mass < 100 MeV, and rises to 0.071 for mass = 500 MeV.

<sup>243</sup> BALTRUSAITIS 85 limit applies for spinless boson with mass < 100 MeV.

**$\tau$ -DECAY PARAMETERS**

**$\tau$ -LEPTON DECAY PARAMETERS**

Written April 2002 by A. Stahl (DESY).

The purpose of the measurements of the decay parameters (*i.e.*, Michel parameters) of the  $\tau$  is to determine the structure (spin and chirality) of the current mediating its decays.

**Leptonic Decays:** The Michel parameters are extracted from the energy spectrum of the charged daughter lepton  $\ell = e, \mu$  in the decays  $\tau \rightarrow \ell \nu_\ell \nu_\tau$ . Ignoring radiative corrections, neglecting terms of order  $(m_\ell/m_\tau)^2$  and  $(m_\tau/\sqrt{s})^2$ , and setting the neutrino masses to zero, the spectrum in the laboratory frame reads

$$\frac{d\Gamma}{dx} = \frac{G_{\tau\ell}^2 m_\tau^5}{192 \pi^3} \times \left\{ f_0(x) + \rho f_1(x) + \eta \frac{m_\ell}{m_\tau} f_2(x) - P_\tau [\xi g_1(x) + \xi \delta g_2(x)] \right\}, \quad (1)$$

with

$$\begin{aligned} f_0(x) &= 2 - 6x^2 + 4x^3 \\ f_1(x) &= -\frac{4}{9} + 4x^2 - \frac{32}{9}x^3 & g_1(x) &= -\frac{2}{3} + 4x - 6x^2 + \frac{8}{3}x^3 \\ f_2(x) &= 12(1-x)^2 & g_2(x) &= \frac{4}{9} - \frac{16}{3}x + 12x^2 - \frac{64}{9}x^3. \end{aligned}$$



The integrated decay width is given by

$$\Gamma = \frac{G_{\tau\ell}^2 m_\tau^5}{192 \pi^3} \left( 1 + 4\eta \frac{m_\ell}{m_\tau} \right) . \quad (2)$$

The situation is similar to muon decays  $\mu \rightarrow e\nu_e\nu_\mu$ . The generalized matrix element with the couplings  $g_{\varepsilon\mu}^\gamma$  and their relations to the Michel parameters  $\rho$ ,  $\eta$ ,  $\xi$ , and  $\delta$  have been described in the “Note on Muon Decay Parameters”. The Standard Model expectations are 3/4, 0, 1, and 3/4, respectively. For more details, see Ref. 1.

**Hadronic Decays:** In the case of hadronic decays  $\tau \rightarrow h\nu_\tau$ , with  $h = \pi, \rho$ , or  $a_1$ , the ansatz is restricted to purely vectorial currents. The matrix element is

$$\frac{G_{\tau h}}{\sqrt{2}} \sum_{\lambda=R,L} g_\lambda \langle \bar{\Psi}_\omega(\nu_\tau) | \gamma^\mu | \Psi_\lambda(\tau) \rangle J_\mu^h \quad (3)$$

with the hadronic current  $J_\mu^h$ . The neutrino chirality  $\omega$  is uniquely determined from  $\lambda$ . The spectrum depends only on a single parameter  $\xi_h$

$$\frac{d\Gamma}{d\vec{x}} = f(\vec{x}) + \xi_h P_\tau g(\vec{x}) , \quad (4)$$

with  $f$  and  $g$  being channel-dependent functions of the observables  $\vec{x}$  (see Ref. 2). The parameter  $\xi_h$  is related to the couplings through

$$\xi_h = |g_L|^2 - |g_R|^2 . \quad (5)$$

$\xi_h$  is the negative of the chirality of the  $\tau$  neutrino in these decays. In the Standard Model,  $\xi_h = 1$ . Also included are measurements of the neutrino helicity which coincide with  $\xi_h$ , if the neutrino is massless (ASNER 00, ACKERSTAFF 97R, AKERS 95P, ALBRECHT 93C, and ALBRECHT 90I).

**Combination of Measurements:** The individual measurements are combined, taking into account the correlations between the parameters. There is one fit, assuming universality between the two leptonic decays, and between all hadronic decays and a second fit without these assumptions. These are the values labeled 'OUR FIT' in the tables. The measurements show good agreement with the Standard Model. The  $\chi^2$  values with respect to the Standard model predictions are 24.1 for 41 degrees of freedom and 26.8 for 56 degrees of freedom, respectively. The correlations are reduced through this combination to less than 20%, with the exception of  $\rho$  and  $\eta$  which are correlated by +23%, for the fit with universality and by +70% for  $\tau \rightarrow \mu\nu_\mu\nu_\tau$ .

**Model-independent Analysis:** From the Michel parameters, limits can be derived on the couplings  $g_{\epsilon\lambda}^k$  without further module assumptions. In the Standard model  $g_{LL}^V = 1$  (leptonic decays), and  $g_L = 1$  (hadronic decays) and all other couplings vanish. First, the partial decay widths have to be compared to the Standard Model predictions to derive limits on the normalization of the couplings  $A_x = G_{\tau x}^2/G_F^2$  with Fermi's constant  $G_F$ :

$$\begin{aligned} A_e &= 1.0012 \pm 0.0053 , \\ A_\mu &= 0.981 \pm 0.018 , \\ A_\pi &= 1.018 \pm 0.012 . \end{aligned} \tag{6}$$

Then limits on the couplings (95% CL) can be extracted (see Ref. 3 and Ref. 4). Without the assumption of universality, the limits given in Table 1 are derived.

**Table 1:** Coupling constants  $g_{\varepsilon\mu}^\gamma$ . 95% confidence level experimental limits. The limits include the quoted values of  $A_e$ ,  $A_\mu$ , and  $A_\pi$  and assume  $A_\rho = A_{a_1} = 1$ .

$\tau \rightarrow e\nu_e\nu_\tau$		
$ g_{RR}^S  < 0.70$	$ g_{RR}^V  < 0.17$	$ g_{RR}^T  \equiv 0$
$ g_{LR}^S  < 0.99$	$ g_{LR}^V  < 0.13$	$ g_{LR}^T  < 0.082$
$ g_{RL}^S  < 2.01$	$ g_{RL}^V  < 0.52$	$ g_{RL}^T  < 0.51$
$ g_{LL}^S  < 2.01$	$ g_{LL}^V  < 1.005$	$ g_{LL}^T  \equiv 0$
$\tau \rightarrow \mu\nu_\mu\nu_\tau$		
$ g_{RR}^S  < 0.72$	$ g_{RR}^V  < 0.18$	$ g_{RR}^T  \equiv 0$
$ g_{LR}^S  < 0.95$	$ g_{LR}^V  < 0.12$	$ g_{LR}^T  < 0.079$
$ g_{RL}^S  < 2.01$	$ g_{RL}^V  < 0.52$	$ g_{RL}^T  < 0.51$
$ g_{LL}^S  < 2.01$	$ g_{LL}^V  < 1.005$	$ g_{LL}^T  \equiv 0$
$\tau \rightarrow \pi\nu_\tau$		
$ g_R^V  < 0.15$	$ g_L^V  > 0.992$	
$\tau \rightarrow \rho\nu_\tau$		
$ g_R^V  < 0.10$	$ g_L^V  > 0.995$	
$\tau \rightarrow a_1\nu_\tau$		
$ g_R^V  < 0.16$	$ g_L^V  > 0.987$	

**Model-dependent Interpretation:** More stringent limits can be derived assuming specific models. For example, in the framework of a two Higgs doublet model, the measurements correspond to a limit of  $m_{H^\pm} > 1.9 \text{ GeV} \times \tan\beta$  on the mass of the

charged Higgs boson, or a limit of 253 GeV on the mass of the second  $W$  boson in left-right symmetric models for arbitrary mixing (both 95% CL). See Ref. 4 and Ref. 5.

## Footnotes and References

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### $\rho^\tau(\mathbf{e}$ or $\mu)$ PARAMETER

( $V-A$ ) theory predicts  $\rho = 0.75$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.745±0.008 OUR FIT</b>				
<b>0.749±0.008 OUR AVERAGE</b>				
0.742±0.014±0.006	81k	HEISTER	01E ALEP	1991–1995 LEP runs
0.775±0.023±0.020	36k	ABREU	00L DLPH	1992–1995 runs
0.781±0.028±0.018	46k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.762±0.035	54k	ACCIARRI	98R L3	1991–1995 LEP runs
0.731±0.031		<sup>244</sup> ALBRECHT	98 ARG	$E_{\text{cm}}^{e^+e^-} = 9.5\text{--}10.6$ GeV
0.72 ±0.09 ±0.03		<sup>245</sup> ABE	97O SLD	1993–1995 SLC runs
0.747±0.010±0.006	55k	ALEXANDER	97F CLEO	$E_{\text{cm}}^{e^+e^-} = 10.6$ GeV
0.79 ±0.10 ±0.10	3732	FORD	87B MAC	$E_{\text{cm}}^{e^+e^-} = 29$ GeV
0.71 ±0.09 ±0.03	1426	BEHRENDIS	85 CLEO	$e^+e^-$ near $\Upsilon(4S)$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
0.735±0.013±0.008	31k	AMMAR	97B CLEO	Repl. by ALEXAN- DER 97F
0.794±0.039±0.031	18k	ACCIARRI	96H L3	Repl. by ACCIARRI 98R
0.732±0.034±0.020	8.2k	<sup>246</sup> ALBRECHT	95 ARG	$E_{\text{cm}}^{e^+e^-} = 9.5\text{--}10.6$ GeV
0.738±0.038		<sup>247</sup> ALBRECHT	95C ARG	Repl. by ALBRECHT 98
0.751±0.039±0.022		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
0.742±0.035±0.020	8000	ALBRECHT	90E ARG	$E_{\text{cm}}^{e^+e^-} = 9.4\text{--}10.6$ GeV

- 244 Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 98, ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.
- 245 ABE 97O assume  $\eta^\tau = 0$  in their fit. Letting  $\eta^\tau$  vary in the fit gives a  $\rho^\tau$  value of  $0.69 \pm 0.13 \pm 0.05$ .
- 246 Value is from a simultaneous fit for the  $\rho^\tau$  and  $\eta^\tau$  decay parameters to the lepton energy spectrum. Not independent of ALBRECHT 90E  $\rho^\tau$  (e or  $\mu$ ) value which assumes  $\eta^\tau = 0$ . Result is strongly correlated with ALBRECHT 95C.
- 247 Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E.

### $\rho^\tau(e)$ PARAMETER

(V-A) theory predicts  $\rho = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.747±0.010 OUR FIT</b>				
<b>0.744±0.010 OUR AVERAGE</b>				
0.747±0.019±0.014	44k	HEISTER	01E ALEP	1991–1995 LEP runs
0.744±0.036±0.037	17k	ABREU	00L DLPH	1992–1995 runs
0.779±0.047±0.029	25k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.68 ±0.04 ±0.07		248 ALBRECHT	98 ARG	$E_{cm}^{ee} = 9.5\text{--}10.6$ GeV
0.71 ±0.14 ±0.05		ABE	97O SLD	1993–1995 SLC runs
0.747±0.012±0.004	34k	ALEXANDER	97F CLEO	$E_{cm}^{ee} = 10.6$ GeV
0.735±0.036±0.020	4.7k	249 ALBRECHT	95 ARG	$E_{cm}^{ee} = 9.5\text{--}10.6$ GeV
0.79 ±0.08 ±0.06	3230	250 ALBRECHT	93G ARG	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV
0.64 ±0.06 ±0.07	2753	JANSSEN	89 CBAL	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV
0.62 ±0.17 ±0.14	1823	FORD	87B MAC	$E_{cm}^{ee} = 29$ GeV
0.60 ±0.13	699	BEHRENDIS	85 CLEO	$e^+ e^-$ near $\Upsilon(4S)$
0.72 ±0.10 ±0.11	594	BACINO	79B DLCO	$E_{cm}^{ee} = 3.5\text{--}7.4$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.732±0.014±0.009	19k	AMMAR	97B CLEO	Repl. by ALEXANDER 97F
0.793±0.050±0.025		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
0.747±0.045±0.028	5106	ALBRECHT	90E ARG	Repl. by ALBRECHT 95
248 ALBRECHT 98 use tau pair events of the type $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.				
249 ALBRECHT 95 use tau pair events of the type $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(h^+ h^- h^+ (\pi^0) \bar{\nu}_\tau)$ and their charged conjugates.				
250 ALBRECHT 93G use tau pair events of the type $\tau^- \tau^+ \rightarrow (\mu^- \bar{\nu}_\mu \nu_\tau)(e^+ \nu_e \bar{\nu}_\tau)$ and their charged conjugates.				

### $\rho^\tau(\mu)$ PARAMETER

(V-A) theory predicts  $\rho = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.763±0.020 OUR FIT</b>				
<b>0.770±0.022 OUR AVERAGE</b>				
0.776±0.045±0.019	46k	HEISTER	01E ALEP	1991–1995 LEP runs
0.999±0.098±0.045	22k	ABREU	00L DLPH	1992–1995 runs
0.777±0.044±0.016	27k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.69 ±0.06 ±0.06		251 ALBRECHT	98 ARG	$E_{cm}^{ee} = 9.5\text{--}10.6$ GeV
0.54 ±0.28 ±0.14		ABE	97O SLD	1993–1995 SLC runs

$0.750 \pm 0.017 \pm 0.045$	22k	ALEXANDER	97F CLEO	$E_{cm}^{ee} = 10.6$ GeV
$0.76 \pm 0.07 \pm 0.08$	3230	ALBRECHT	93G ARG	$E_{cm}^{ee} = 9.4-10.6$ GeV
$0.734 \pm 0.055 \pm 0.027$	3041	ALBRECHT	90E ARG	$E_{cm}^{ee} = 9.4-10.6$ GeV
$0.89 \pm 0.14 \pm 0.08$	1909	FORD	87B MAC	$E_{cm}^{ee} = 29$ GeV
$0.81 \pm 0.13$	727	BEHREND	85 CLEO	$e^+ e^-$ near $\Upsilon(4S)$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$0.747 \pm 0.048 \pm 0.044$	13k	AMMAR	97B CLEO	Repl. by ALEXANDER 97F
$0.693 \pm 0.057 \pm 0.028$		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
<sup>251</sup> ALBRECHT 98 use tau pair events of the type $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.				

### $\xi^\tau(e \text{ or } \mu)$ PARAMETER

(V-A) theory predicts  $\xi = 1$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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#### **0.985 ± 0.030 OUR FIT**

#### **0.981 ± 0.031 OUR AVERAGE**

$0.986 \pm 0.068 \pm 0.031$	81k	HEISTER	01E ALEP	1991-1995 LEP runs
$0.929 \pm 0.070 \pm 0.030$	36k	ABREU	00L DLPH	1992-1995 runs
$0.98 \pm 0.22 \pm 0.10$	46k	ACKERSTAFF	99D OPAL	1990-1995 LEP runs
$0.70 \pm 0.16$	54k	ACCIARRI	98R L3	1991-1995 LEP runs
$1.03 \pm 0.11$		<sup>252</sup> ALBRECHT	98 ARG	$E_{cm}^{ee} = 9.5-10.6$ GeV
$1.05 \pm 0.35 \pm 0.04$		<sup>253</sup> ABE	97O SLD	1993-1995 SLC runs
$1.007 \pm 0.040 \pm 0.015$	55k	ALEXANDER	97F CLEO	$E_{cm}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.94 \pm 0.21 \pm 0.07$	18k	ACCIARRI	96H L3	Repl. by ACCIARRI 98R
$0.97 \pm 0.14$		<sup>254</sup> ALBRECHT	95C ARG	Repl. by ALBRECHT 98
$1.18 \pm 0.15 \pm 0.16$		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
$0.90 \pm 0.15 \pm 0.10$	3230	<sup>255</sup> ALBRECHT	93G ARG	$E_{cm}^{ee} = 9.4-10.6$ GeV

<sup>252</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 98, ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

<sup>253</sup> ABE 97O assume  $\eta^\tau = 0$  in their fit. Letting  $\eta^\tau$  vary in the fit gives a  $\xi^\tau$  value of  $1.02 \pm 0.36 \pm 0.05$ .

<sup>254</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 95C uses events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(h^+ h^- h^+ \bar{\nu}_\tau)$  and their charged conjugates.

<sup>255</sup> ALBRECHT 93G measurement determines  $|\xi^\tau|$  for the case  $\xi^\tau(e) = \xi^\tau(\mu)$ , but the authors point out that other LEP experiments determine the sign to be positive.

### $\xi^\tau(e)$ PARAMETER

(V-A) theory predicts  $\xi = 1$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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#### **0.994 ± 0.040 OUR FIT**

#### **1.00 ± 0.04 OUR AVERAGE**

$1.011 \pm 0.094 \pm 0.038$	44k	HEISTER	01E ALEP	1991-1995 LEP runs
$1.01 \pm 0.12 \pm 0.05$	17k	ABREU	00L DLPH	1992-1995 runs
$1.13 \pm 0.39 \pm 0.14$	25k	ACKERSTAFF	99D OPAL	1990-1995 LEP runs

1.11 ±0.20 ±0.08		<sup>256</sup> ALBRECHT	98 ARG	$E_{cm}^{ee} = 9.5\text{--}10.6$ GeV
1.16 ±0.52 ±0.06		ABE	97O SLD	1993–1995 SLC runs
0.979±0.048±0.016	34k	ALEXANDER	97F CLEO	$E_{cm}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.03 ±0.23 ±0.09		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
<sup>256</sup> ALBRECHT 98 use tau pair events of the type $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.				

### $\xi^\tau(\mu)$ PARAMETER

(V–A) theory predicts  $\xi = 1$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.030±0.059 OUR FIT</b>				
<b>1.06 ±0.06 OUR AVERAGE</b>				
1.030±0.120±0.050	46k	HEISTER	01E ALEP	1991–1995 LEP runs
1.16 ±0.19 ±0.06	22k	ABREU	00L DLPH	1992–1995 runs
0.79 ±0.41 ±0.09	27k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
1.26 ±0.27 ±0.14		<sup>257</sup> ALBRECHT	98 ARG	$E_{cm}^{ee} = 9.5\text{--}10.6$ GeV
0.75 ±0.50 ±0.14		ABE	97O SLD	1993–1995 SLC runs
1.054±0.069±0.047	22k	ALEXANDER	97F CLEO	$E_{cm}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
1.23 ±0.22 ±0.10		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
<sup>257</sup> ALBRECHT 98 use tau pair events of the type $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.				

### $\eta^\tau(e \text{ or } \mu)$ PARAMETER

(V–A) theory predicts  $\eta = 0$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.013±0.020 OUR FIT</b>				
<b>0.015±0.021 OUR AVERAGE</b>				
0.012±0.026±0.004	81k	HEISTER	01E ALEP	1991–1995 LEP runs
–0.005±0.036±0.037		ABREU	00L DLPH	1992–1995 runs
0.027±0.055±0.005	46k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.27 ±0.14	54k	ACCIARRI	98R L3	1991–1995 LEP runs
–0.13 ±0.47 ±0.15		ABE	97O SLD	1993–1995 SLC runs
–0.015±0.061±0.062	31k	AMMAR	97B CLEO	$E_{cm}^{ee} = 10.6$ GeV
0.03 ±0.18 ±0.12	8.2k	ALBRECHT	95 ARG	$E_{cm}^{ee} = 9.5\text{--}10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.25 ±0.17 ±0.11	18k	ACCIARRI	96H L3	Repl. by ACCIARRI 98R
–0.04 ±0.15 ±0.11		BUSKULIC	95D ALEP	Repl. by HEISTER 01E

### $\eta^\tau(\mu)$ PARAMETER

(V–A) theory predicts  $\eta = 0$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.094±0.073 OUR FIT</b>				
<b>0.17 ±0.15 OUR AVERAGE</b> Error includes scale factor of 1.2.				
0.160±0.150±0.060	46k	HEISTER	01E ALEP	1991–1995 LEP runs
0.72 ±0.32 ±0.15		ABREU	00L DLPH	1992–1995 runs
–0.59 ±0.82 ±0.45		<sup>258</sup> ABE	97O SLD	1993–1995 SLC runs
0.010±0.149±0.171	13k	<sup>259</sup> AMMAR	97B CLEO	$E_{cm}^{ee} = 10.6$ GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				

- 0.010±0.065±0.001    27k    <sup>260</sup>ACKERSTAFF 99D OPAL 1990–1995 LEP runs  
 –0.24 ±0.23 ±0.18                      BUSKULIC    95D ALEP    Repl. by HEISTER 01E
- <sup>258</sup> Highly correlated (corr. = 0.92) with ABE 97O  $\rho^T(\mu)$  measurement.
- <sup>259</sup> Highly correlated (corr. = 0.949) with AMMAR 97B  $\rho^T(\mu)$  value.
- <sup>260</sup> ACKERSTAFF 99D result is dominated by a constraint on  $\eta^T$  from the OPAL measurements of the  $\tau$  lifetime and  $B(\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau)$  assuming lepton universality for the total coupling strength.

**( $\delta\xi$ )<sup>T</sup>(e or  $\mu$ ) PARAMETER**

(V–A) theory predicts ( $\delta\xi$ ) = 0.75.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.746±0.021 OUR FIT</b>				
<b>0.744±0.022 OUR AVERAGE</b>				
0.776±0.045±0.024	81k	HEISTER	01E ALEP	1991–1995 LEP runs
0.779±0.070±0.028	36k	ABREU	00L DLPH	1992–1995 runs
0.65 ±0.14 ±0.07	46k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.70 ±0.11	54k	ACCIARRI	98R L3	1991–1995 LEP runs
0.63 ±0.09		<sup>261</sup> ALBRECHT	98 ARG	$E_{cm}^{ee} = 9.5\text{--}10.6$ GeV
0.88 ±0.27 ±0.04		<sup>262</sup> ABE	97O SLD	1993–1995 SLC runs
0.745±0.026±0.009	55k	ALEXANDER	97F CLEO	$E_{cm}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.81 ±0.14 ±0.06	18k	ACCIARRI	96H L3	Repl. by ACCIARRI 98R
0.65 ±0.12		<sup>263</sup> ALBRECHT	95C ARG	Repl. by ALBRECHT 98
0.88 ±0.11 ±0.07		BUSKULIC	95D ALEP	Repl. by HEISTER 01E

- <sup>261</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 98, ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.
- <sup>262</sup> ABE 97O assume  $\eta^T = 0$  in their fit. Letting  $\eta^T$  vary in the fit gives a ( $\rho\xi$ )<sup>T</sup> value of 0.87 ± 0.27 ± 0.04.
- <sup>263</sup> Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E. ALBRECHT 95C uses events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(h^+ h^- h^+ \bar{\nu}_\tau)$  and their charged conjugates.

**( $\delta\xi$ )<sup>T</sup>(e) PARAMETER**

(V–A) theory predicts ( $\delta\xi$ ) = 0.75.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.734±0.028 OUR FIT</b>				
<b>0.731±0.029 OUR AVERAGE</b>				
0.778±0.066±0.024	44k	HEISTER	01E ALEP	1991–1995 LEP runs
0.85 ±0.12 ±0.04	17k	ABREU	00L DLPH	1992–1995 runs
0.72 ±0.31 ±0.14	25k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.56 ±0.14 ±0.06		<sup>264</sup> ALBRECHT	98 ARG	$E_{cm}^{ee} = 9.5\text{--}10.6$ GeV
0.85 ±0.43 ±0.08		ABE	97O SLD	1993–1995 SLC runs
0.720±0.032±0.010	34k	ALEXANDER	97F CLEO	$E_{cm}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.11 ±0.17 ±0.07		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
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<sup>264</sup> ALBRECHT 98 use tau pair events of the type  $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.



### $(\delta\xi)^T(\mu)$ PARAMETER

( $V-A$ ) theory predicts  $(\delta\xi) = 0.75$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.778±0.037 OUR FIT**

**0.79 ±0.04 OUR AVERAGE**

0.786±0.066±0.028	46k	HEISTER	01E ALEP	1991–1995 LEP runs
0.86 ±0.13 ±0.04	22k	ABREU	00L DLPH	1992–1995 runs
0.63 ±0.23 ±0.05	27k	ACKERSTAFF	99D OPAL	1990–1995 LEP runs
0.73 ±0.18 ±0.10		<sup>265</sup> ALBRECHT	98 ARG	$E_{cm}^{ee} = 9.5\text{--}10.6$ GeV
0.82 ±0.32 ±0.07		ABE	97O SLD	1993–1995 SLC runs
0.786±0.041±0.032	22k	ALEXANDER	97F CLEO	$E_{cm}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.71 ±0.14 ±0.06		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
		<sup>265</sup> ALBRECHT 98		use tau pair events of the type $\tau^- \tau^+ \rightarrow (\ell^- \bar{\nu}_\ell \nu_\tau)(\pi^+ \pi^0 \bar{\nu}_\tau)$ , and their charged conjugates.

### $\xi^T(\pi)$ PARAMETER

( $V-A$ ) theory predicts  $\xi^T(\pi) = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.993±0.022 OUR FIT**

**0.994±0.023 OUR AVERAGE**

0.994±0.020±0.014	27k	HEISTER	01E ALEP	1991–1995 LEP runs
0.81 ±0.17 ±0.02		ABE	97O SLD	1993–1995 SLC runs
1.03 ±0.06 ±0.04	2.0k	COAN	97 CLEO	$E_{cm}^{ee} = 10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.987±0.057±0.027		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
0.95 ±0.11 ±0.05		<sup>266</sup> BUSKULIC	94D ALEP	1990+1991 LEP run
		<sup>266</sup>		Superseded by BUSKULIC 95D.

### $\xi^T(\rho)$ PARAMETER

( $V-A$ ) theory predicts  $\xi^T(\rho) = 1$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.994±0.008 OUR FIT**

**0.994±0.009 OUR AVERAGE**

0.987±0.012±0.011	59k	HEISTER	01E ALEP	1991–1995 LEP runs
0.99 ±0.12 ±0.04		ABE	97O SLD	1993–1995 SLC runs
0.995±0.010±0.003	66k	ALEXANDER	97F CLEO	$E_{cm}^{ee} = 10.6$ GeV
1.022±0.028±0.030	1.7k	<sup>267</sup> ALBRECHT	94E ARG	$E_{cm}^{ee} = 9.4\text{--}10.6$ GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.045±0.058±0.032		BUSKULIC	95D ALEP	Repl. by HEISTER 01E
1.03 ±0.11 ±0.05		<sup>268</sup> BUSKULIC	94D ALEP	1990+1991 LEP run

<sup>267</sup>ALBRECHT 94E measure the square of this quantity and use the sign determined by ALBRECHT 90I to obtain the quoted result.

<sup>268</sup>Superseded by BUSKULIC 95D.

## $\xi^T(a_1)$ PARAMETER

(V-A) theory predicts  $\xi^T(a_1) = 1$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**1.001±0.027 OUR FIT**

**1.002±0.028 OUR AVERAGE**

1.000±0.016±0.024	35k	269 HEISTER	01E ALEP	1991–1995 LEP runs
1.02 ±0.13 ±0.03	17.2k	ASNER	00 CLEO	$E_{cm}^{ee} = 10.6$ GeV
1.29 ±0.26 ±0.11	7.4k	270 ACKERSTAFF	97R OPAL	1992–1994 LEP runs
0.85 $\begin{smallmatrix} +0.15 \\ -0.17 \end{smallmatrix}$ ±0.05		ALBRECHT	95C ARG	$E_{cm}^{ee} = 9.5$ –10.6 GeV
1.25 ±0.23 $\begin{smallmatrix} +0.15 \\ -0.08 \end{smallmatrix}$	7.5k	ALBRECHT	93C ARG	$E_{cm}^{ee} = 9.4$ –10.6 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

1.08 $\begin{smallmatrix} +0.46 \\ -0.41 \end{smallmatrix}$ $\begin{smallmatrix} +0.14 \\ -0.25 \end{smallmatrix}$	2.6k	271 AKERS	95P OPAL	Repl. by ACKER-STAFF 97R
0.937±0.116±0.064		BUSKULIC	95D ALEP	Repl. by HEISTER 01E

269 HEISTER 01E quote  $1.000 \pm 0.016 \pm 0.013 \pm 0.020$  where the errors are statistical, systematic, and an uncertainty due to the final state model. We combine the systematic error and model uncertainty.

270 ACKERSTAFF 97R obtain this result with a model independent fit to the hadronic structure functions. Fitting with the model of Kuhn and Santamaria (ZPHY **C48**, 445 (1990)) gives  $0.87 \pm 0.16 \pm 0.04$ , and with the model of of Isgur *et al.* (PR **D39**,1357 (1989)) they obtain  $1.20 \pm 0.21 \pm 0.14$ .

271 AKERS 95P obtain this result with a model independent fit to the hadronic structure functions. Fitting with the model of Kuhn and Santamaria (ZPHY **C48**, 445 (1990)) gives  $0.87 \pm 0.27 \begin{smallmatrix} +0.05 \\ -0.06 \end{smallmatrix}$ , and with the model of of Isgur *et al.* (PR **D39**,1357 (1989)) they obtain  $1.10 \pm 0.31 \begin{smallmatrix} +0.13 \\ -0.14 \end{smallmatrix}$ .

## $\xi^T(\text{all hadronic modes})$ PARAMETER

(V-A) theory predicts  $\xi^T = 1$ .

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**0.995±0.007 OUR FIT**

**0.997±0.007 OUR AVERAGE**

0.992±0.007±0.008	102k	272 HEISTER	01E ALEP	1991–1995 LEP runs
0.997±0.027±0.011	39k	273 ABREU	00L DLPH	1992–1995 runs
1.02 ±0.13 ±0.03	17.2k	274 ASNER	00 CLEO	$E_{cm}^{ee} = 10.6$ GeV
1.032±0.031	37k	275 ACCIARRI	98R L3	1991–1995 LEP runs
0.93 ±0.10 ±0.04		ABE	97O SLD	1993–1995 SLC runs
1.29 ±0.26 ±0.11	7.4k	276 ACKERSTAFF	97R OPAL	1992–1994 LEP runs
0.995±0.010±0.003	66k	277 ALEXANDER	97F CLEO	$E_{cm}^{ee} = 10.6$ GeV
1.03 ±0.06 ±0.04	2.0k	278 COAN	97 CLEO	$E_{cm}^{ee} = 10.6$ GeV
1.017±0.039		279 ALBRECHT	95C ARG	$E_{cm}^{ee} = 9.5$ –10.6 GeV
1.25 ±0.23 $\begin{smallmatrix} +0.15 \\ -0.08 \end{smallmatrix}$	7.5k	280 ALBRECHT	93C ARG	$E_{cm}^{ee} = 9.4$ –10.6 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.970±0.053±0.011	14k	281 ACCIARRI	96H L3	Repl. by ACCIARRI 98R
1.08 $\begin{smallmatrix} +0.46 \\ -0.41 \end{smallmatrix}$ $\begin{smallmatrix} +0.14 \\ -0.25 \end{smallmatrix}$	2.6k	282 AKERS	95P OPAL	Repl. by ACKER-STAFF 97R
1.006±0.032±0.019		283 BUSKULIC	95D ALEP	Repl. by HEISTER 01E
1.022±0.028±0.030	1.7k	284 ALBRECHT	94E ARG	$E_{cm}^{ee} = 9.4$ –10.6 GeV
0.99 ±0.07 ±0.04		285 BUSKULIC	94D ALEP	1990+1991 LEP run

- 272 HEISTER 01E quote  $0.992 \pm 0.007 \pm 0.006 \pm 0.005$  where the errors are statistical, systematic, and an uncertainty due to the final state model. We combine the systematic error and model uncertainty. They use  $\tau \rightarrow \pi\nu_\tau$ ,  $\tau \rightarrow K\nu_\tau$ ,  $\tau \rightarrow \rho\nu_\tau$ , and  $\tau \rightarrow a_1\nu_\tau$  decays.
- 273 ABREU 00L use  $\tau^- \rightarrow h^- \geq 0\pi^0\nu_\tau$  decays.
- 274 ASNER 00 use  $\tau^- \rightarrow \pi^- 2\pi^0\nu_\tau$  decays.
- 275 ACCIARRI 98R use  $\tau \rightarrow \pi\nu_\tau$ ,  $\tau \rightarrow K\nu_\tau$ , and  $\tau \rightarrow \rho\nu_\tau$  decays.
- 276 ACKERSTAFF 97R use  $\tau \rightarrow a_1\nu_\tau$  decays.
- 277 ALEXANDER 97F use  $\tau \rightarrow \rho\nu_\tau$  decays.
- 278 COAN 97 use  $h^+h^-$  energy correlations.
- 279 Combined fit to ARGUS tau decay parameter measurements in ALBRECHT 95C, ALBRECHT 93G, and ALBRECHT 94E.
- 280 Uses  $\tau \rightarrow a_1\nu_\tau$  decays. Replaced by ALBRECHT 95C.
- 281 ACCIARRI 96H use  $\tau \rightarrow \pi\nu_\tau$ ,  $\tau \rightarrow K\nu_\tau$ , and  $\tau \rightarrow \rho\nu_\tau$  decays.
- 282 AKERS 95P use  $\tau \rightarrow a_1\nu_\tau$  decays.
- 283 BUSKULIC 95D use  $\tau \rightarrow \pi\nu_\tau$ ,  $\tau \rightarrow \rho\nu_\tau$ , and  $\tau \rightarrow a_1\nu_\tau$  decays.
- 284 ALBRECHT 94E measure the square of this quantity and use the sign determined by ALBRECHT 90I to obtain the quoted result. Uses  $\tau \rightarrow a_1\nu_\tau$  decays. Replaced by ALBRECHT 95C.
- 285 BUSKULIC 94D use  $\tau \rightarrow \pi\nu_\tau$  and  $\tau \rightarrow \rho\nu_\tau$  decays. Superseded by BUSKULIC 95D.

## $\tau$ REFERENCES

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BRIERE	03	PRL 90 181802	R. A. Briere <i>et al.</i>	(CLEO Collab.)
HEISTER	03F	EPJ C30 291	A. Heister <i>et al.</i>	(ALEPH Collab.)
INAMI	03	PL B551 16	K. Inami <i>et al.</i>	(BELLE Collab.)
CHEN	02C	PR D66 071101R	S. Chen <i>et al.</i>	(CLEO Collab.)
ABBIENDI	01J	EPJ C19 653	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABREU	01M	EPJ C20 617	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACCIARRI	01F	PL B507 47	M. Acciarri <i>et al.</i>	(L3 Collab.)
ACHARD	01D	PL B519 189	P. Achard <i>et al.</i>	(L3 Collab.)
ANASTASSOV	01	PRL 86 4467	A. Anastassov <i>et al.</i>	(CLEO Collab.)
HEISTER	01E	EPJ C22 217	A. Heister <i>et al.</i>	(ALEPH Collab.)
ABBIENDI	00A	PL B492 23	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABBIENDI	00C	EPJ C13 213	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABBIENDI	00D	EPJ C13 197	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABREU	00L	EPJ C16 229	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACCIARRI	00B	PL B479 67	M. Acciarri <i>et al.</i>	(L3 Collab.)
AHMED	00	PR D61 071101R	S. Ahmed <i>et al.</i>	(CLEO Collab.)
ALBRECHT	00	PL B485 37	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ASNER	00	PR D61 012002	D.M. Asner <i>et al.</i>	(CLEO Collab.)
ASNER	00B	PR D62 072006	D.M. Asner <i>et al.</i>	(CLEO Collab.)
BERGFELD	00	PRL 84 830	T. Bergfeld <i>et al.</i>	(CLEO Collab.)
BROWDER	00	PR D61 052004	T.E. Browder <i>et al.</i>	(CLEO Collab.)
EDWARDS	00A	PR D61 072003	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
GONZALEZ-S...	00	NP B582 3	G.A. Gonzalez-Sprinberg <i>et al.</i>	
ABBIENDI	99H	PL B447 134	G. Abbiendi <i>et al.</i>	(OPAL Collab.)
ABREU	99X	EPJ C10 201	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACKERSTAFF	99D	EPJ C8 3	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
ACKERSTAFF	99E	EPJ C8 183	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
BARATE	99K	EPJ C10 1	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARATE	99R	EPJ C11 599	R. Barate <i>et al.</i>	(ALEPH Collab.)
BISHAI	99	PRL 82 281	M. Bishai <i>et al.</i>	(CLEO Collab.)
GODANG	99	PR D59 091303	R. Godang <i>et al.</i>	(CLEO Collab.)
RICHICHI	99	PR D60 112002	S.J. Richichi <i>et al.</i>	(CLEO Collab.)
ACCIARRI	98C	PL B426 207	M. Acciarri <i>et al.</i>	(L3 Collab.)
ACCIARRI	98E	PL B434 169	M. Aciarri <i>et al.</i>	(L3 Collab.)
ACCIARRI	98R	PL B438 405	M. Acciarri <i>et al.</i>	(L3 Collab.)
ACKERSTAFF	98M	EPJ C4 193	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
ACKERSTAFF	98N	PL B431 188	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
ALBRECHT	98	PL B431 179	H. Albrecht <i>et al.</i>	(ARGUS Collab.)

BARATE	98	EPJ C1 65	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARATE	98E	EPJ C4 29	R. Barate <i>et al.</i>	(ALEPH Collab.)
BLISS	98	PR D57 5903	D.W. Bliss <i>et al.</i>	(CLEO Collab.)
ABE	97O	PRL 78 4691	K. Abe <i>et al.</i>	(SLD Collab.)
ACKERSTAFF	97J	PL B404 213	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
ACKERSTAFF	97L	ZPHY C74 403	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
ACKERSTAFF	97R	ZPHY C75 593	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
ALEXANDER	97F	PR D56 5320	J.P. Alexander <i>et al.</i>	(CLEO Collab.)
AMMAR	97B	PRL 78 4686	R. Ammar <i>et al.</i>	(CLEO Collab.)
ANASTASSOV	97	PR D55 2559	A. Anastassov <i>et al.</i>	(CLEO Collab.)
Also	98B	PR D58 119903 (erratum)	A. Anastassov <i>et al.</i>	(CLEO Collab.)
ANDERSON	97	PRL 79 3814	S. Anderson <i>et al.</i>	(CLEO Collab.)
AVERY	97	PR D55 R1119	P. Avery <i>et al.</i>	(CLEO Collab.)
BARATE	97I	ZPHY C74 387	R. Barate <i>et al.</i>	(ALEPH Collab.)
BARATE	97R	PL B414 362	R. Barate <i>et al.</i>	(ALEPH Collab.)
BERGFELD	97	PRL 79 2406	T. Bergfeld <i>et al.</i>	(CLEO Collab.)
BONVICINI	97	PRL 79 1221	G. Bonvicini <i>et al.</i>	(CLEO Collab.)
BUSKULIC	97C	ZPHY C74 263	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
COAN	97	PR D55 7291	T.E. Coan <i>et al.</i>	(CLEO Collab.)
EDWARDS	97	PR D55 R3919	K.W. Edwards <i>et al.</i>	(CLEO Collab.)
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ESCRIBANO	97	PL B395 369	R. Escribano, E. Masso	(BARC, PARIT)
ABREU	96B	PL B365 448	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACCIARRI	96H	PL B377 313	M. Acciarri <i>et al.</i>	(L3 Collab.)
ACCIARRI	96K	PL B389 187	M. Acciarri <i>et al.</i>	(L3 Collab.)
ALAM	96	PRL 76 2637	M.S. Alam <i>et al.</i>	(CLEO Collab.)
ALBRECHT	96E	PRPL 276 223	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALEXANDER	96D	PL B369 163	G. Alexander <i>et al.</i>	(OPAL Collab.)
ALEXANDER	96E	PL B374 341	G. Alexander <i>et al.</i>	(OPAL Collab.)
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BAI	96	PR D53 20	J.Z. Bai <i>et al.</i>	(BES Collab.)
BALEST	96	PL B388 402	R. Balest <i>et al.</i>	(CLEO Collab.)
BARTELT	96	PRL 76 4119	J.E. Bartelt <i>et al.</i>	(CLEO Collab.)
BUSKULIC	96	ZPHY C70 579	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
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COAN	96	PR D53 6037	T.E. Coan <i>et al.</i>	(CLEO Collab.)
ABE	95Y	PR D52 4828	K. Abe <i>et al.</i>	(SLD Collab.)
ABREU	95T	PL B357 715	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ABREU	95U	PL B359 411	P. Abreu <i>et al.</i>	(DELPHI Collab.)
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AKERS	95F	ZPHY C66 31	R. Akers <i>et al.</i>	(OPAL Collab.)
AKERS	95I	ZPHY C66 543	R. Akers <i>et al.</i>	(OPAL Collab.)
AKERS	95P	ZPHY C67 45	R. Akers <i>et al.</i>	(OPAL Collab.)
AKERS	95Y	ZPHY C68 555	R. Akers <i>et al.</i>	(OPAL Collab.)
ALBRECHT	95	PL B341 441	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	95C	PL B349 576	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	95G	ZPHY C68 25	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
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BALEST	95C	PRL 75 3809	R. Balest <i>et al.</i>	(CLEO Collab.)
BERNABEU	95	NP B436 474	J. Bernabeu <i>et al.</i>	
BUSKULIC	95C	PL B346 371	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	95D	PL B346 379	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
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ABREU	94K	PL B334 435	P. Abreu <i>et al.</i>	(DELPHI Collab.)
AKERS	94E	PL B328 207	R. Akers <i>et al.</i>	(OPAL Collab.)
AKERS	94G	PL B339 278	R. Akers <i>et al.</i>	(OPAL Collab.)
ALBRECHT	94E	PL B337 383	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ARTUSO	94	PRL 72 3762	M. Artuso <i>et al.</i>	(CLEO Collab.)
BARTELT	94	PRL 73 1890	J.E. Bartelt <i>et al.</i>	(CLEO Collab.)
BATTLE	94	PRL 73 1079	M. Battle <i>et al.</i>	(CLEO Collab.)
BAUER	94	PR D50 R13	D.A. Bauer <i>et al.</i>	(TPC/2gamma Collab.)
BUSKULIC	94D	PL B321 168	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
BUSKULIC	94E	PL B332 209	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
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GIBAUT	94B	PRL 73 934	D. Gibaut <i>et al.</i>	(CLEO Collab.)
ADRIANI	93M	PRPL 236 1	O. Adriani <i>et al.</i>	(L3 Collab.)
ALBRECHT	93C	ZPHY C58 61	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	93G	PL B316 608	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BALEST	93	PR D47 R3671	R. Balest <i>et al.</i>	(CLEO Collab.)
BEAN	93	PRL 70 138	A. Bean <i>et al.</i>	(CLEO Collab.)

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ESCRIBANO	93	PL B301 419	R. Escribano, E. Masso	(BARC)
PROCARIO	93	PRL 70 1207	M. Procaro <i>et al.</i>	(CLEO Collab.)
ABREU	92N	ZPHY C55 555	P. Abreu <i>et al.</i>	(DELPHI Collab.)
ACTON	92F	PL B281 405	D.P. Acton <i>et al.</i>	(OPAL Collab.)
ACTON	92H	PL B288 373	P.D. Acton <i>et al.</i>	(OPAL Collab.)
AKERIB	92	PRL 69 3610	D.S. Akerib <i>et al.</i>	(CLEO Collab.)
Also	93B	PRL 71 3395 (erratum)	D.S. Akerib <i>et al.</i>	(CLEO Collab.)
ALBRECHT	92D	ZPHY C53 367	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	92K	ZPHY C55 179	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	92M	PL B292 221	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	92Q	ZPHY C56 339	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
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ARTUSO	92	PRL 69 3278	M. Artuso <i>et al.</i>	(CLEO Collab.)
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BATTLE	92	PL B291 488	M. Battle <i>et al.</i>	(CLEO Collab.)
BUSKULIC	92J	PL B297 459	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
DECAMP	92C	ZPHY C54 211	D. Decamp <i>et al.</i>	(ALEPH Collab.)
ADEVA	91F	PL B265 451	B. Adeva <i>et al.</i>	(L3 Collab.)
ALBRECHT	91D	PL B260 259	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALEXANDER	91D	PL B266 201	G. Alexander <i>et al.</i>	(OPAL Collab.)
ANTREASYAN	91	PL B259 216	D. Antreasyan <i>et al.</i>	(Crystal Ball Collab.)
GRIFOLS	91	PL B255 611	J.A. Grifols, A. Mendez	(BARC)
SAMUEL	91B	PRL 67 668	M.A. Samuel, G.W. Li, R. Mendel	(OKSU, WONT)
Also	92B	PRL 69 995	M.A. Samuel, G.W. Li, R. Mendel	(OKSU, WONT)
Erratum.				
ABACHI	90	PR D41 1414	S. Abachi <i>et al.</i>	(HRS Collab.)
ALBRECHT	90E	PL B246 278	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	90I	PL B250 164	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BEHREND	90	ZPHY C46 537	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
BOWCOCK	90	PR D41 805	T.J.V. Bowcock <i>et al.</i>	(CLEO Collab.)
DELAGUILA	90	PL B252 116	F. del Aguila, M. Sher	(BARC, WILL)
GOLDBERG	90	PL B251 223	M. Goldberg <i>et al.</i>	(CLEO Collab.)
WU	90	PR D41 2339	D.Y. Wu <i>et al.</i>	(Mark II Collab.)
ABACHI	89B	PR D40 902	S. Abachi <i>et al.</i>	(HRS Collab.)
BEHREND	89B	PL B222 163	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
JANSEN	89	PL B228 273	H. Janssen <i>et al.</i>	(Crystal Ball Collab.)
KLEINWORT	89	ZPHY C42 7	C. Kleinwort <i>et al.</i>	(JADE Collab.)
ADEVA	88	PR D38 2665	B. Adeva <i>et al.</i>	(Mark-J Collab.)
ALBRECHT	88B	PL B202 149	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	88L	ZPHY C41 1	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	88M	ZPHY C41 405	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
AMIDEI	88	PR D37 1750	D. Amidei <i>et al.</i>	(Mark II Collab.)
BEHREND	88	PL B200 226	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
BRAUNSCH...	88C	ZPHY C39 331	W. Braunschweig <i>et al.</i>	(TASSO Collab.)
KEH	88	PL B212 123	S. Keh <i>et al.</i>	(Crystal Ball Collab.)
TSCHIRHART	88	PL B205 407	R. Tschirhart <i>et al.</i>	(HRS Collab.)
ABACHI	87B	PL B197 291	S. Abachi <i>et al.</i>	(HRS Collab.)
ABACHI	87C	PRL 59 2519	S. Abachi <i>et al.</i>	(HRS Collab.)
ADLER	87B	PRL 59 1527	J. Adler <i>et al.</i>	(Mark III Collab.)
AIHARA	87B	PR D35 1553	H. Aihara <i>et al.</i>	(TPC Collab.)
AIHARA	87C	PRL 59 751	H. Aihara <i>et al.</i>	(TPC Collab.)
ALBRECHT	87L	PL B185 223	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	87P	PL B199 580	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BAND	87	PL B198 297	H.R. Band <i>et al.</i>	(MAC Collab.)
BAND	87B	PRL 59 415	H.R. Band <i>et al.</i>	(MAC Collab.)
BARINGER	87	PRL 59 1993	P. Baringer <i>et al.</i>	(CLEO Collab.)
BEBEK	87C	PR D36 690	C. Bebek <i>et al.</i>	(CLEO Collab.)
BURCHAT	87	PR D35 27	P.R. Burchat <i>et al.</i>	(Mark II Collab.)
BYLSMA	87	PR D35 2269	B.G. Bylsma <i>et al.</i>	(HRS Collab.)
COFFMAN	87	PR D36 2185	D.M. Coffman <i>et al.</i>	(Mark III Collab.)
DERRICK	87	PL B189 260	M. Derrick <i>et al.</i>	(HRS Collab.)
FORD	87	PR D35 408	W.T. Ford <i>et al.</i>	(MAC Collab.)
FORD	87B	PR D36 1971	W.T. Ford <i>et al.</i>	(MAC Collab.)
GAN	87	PRL 59 411	K.K. Gan <i>et al.</i>	(Mark II Collab.)
GAN	87B	PL B197 561	K.K. Gan <i>et al.</i>	(Mark II Collab.)
AIHARA	86E	PRL 57 1836	H. Aihara <i>et al.</i>	(TPC Collab.)
BARTEL	86D	PL B182 216	W. Bartel <i>et al.</i>	(JADE Collab.)
PDG	86	PL 170B	M. Aguilar-Benitez <i>et al.</i>	(CERN, CIT+)
RUCKSTUHL	86	PRL 56 2132	W. Ruckstuhl <i>et al.</i>	(DELCO Collab.)
SCHMIDKE	86	PRL 57 527	W.B. Schmidke <i>et al.</i>	(Mark II Collab.)

YELTON	86	PRL 56 812	J.M. Yelton <i>et al.</i>	(Mark II Collab.)
ALTHOFF	85	ZPHY C26 521	M. Althoff <i>et al.</i>	(TASSO Collab.)
ASH	85B	PRL 55 2118	W.W. Ash <i>et al.</i>	(MAC Collab.)
BALTRUSAIT...	85	PRL 55 1842	R.M. Baltrusaitis <i>et al.</i>	(Mark III Collab.)
BARTEL	85F	PL 161B 188	W. Bartel <i>et al.</i>	(JADE Collab.)
BEHREND	85	PR D32 2468	S. Behrends <i>et al.</i>	(CLEO Collab.)
BELTRAMI	85	PRL 54 1775	I. Beltrami <i>et al.</i>	(HRS Collab.)
BERGER	85	ZPHY C28 1	C. Berger <i>et al.</i>	(PLUTO Collab.)
BURCHAT	85	PRL 54 2489	P.R. Burchat <i>et al.</i>	(Mark II Collab.)
FERNANDEZ	85	PRL 54 1624	E. Fernandez <i>et al.</i>	(MAC Collab.)
MILLS	85	PRL 54 624	G.B. Mills <i>et al.</i>	(DELCO Collab.)
AIHARA	84C	PR D30 2436	H. Aihara <i>et al.</i>	(TPC Collab.)
BEHREND	84	ZPHY C23 103	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
MILLS	84	PRL 52 1944	G.B. Mills <i>et al.</i>	(DELCO Collab.)
BEHREND	83C	PL 127B 270	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
SILVERMAN	83	PR D27 1196	D.J. Silverman, G.L. Shaw	(UCI)
BEHREND	82	PL 114B 282	H.J. Behrend <i>et al.</i>	(CELLO Collab.)
BLOCKER	82B	PRL 48 1586	C.A. Blocker <i>et al.</i>	(Mark II Collab.)
BLOCKER	82D	PL 109B 119	C.A. Blocker <i>et al.</i>	(Mark II Collab.) J
FELDMAN	82	PRL 48 66	G.J. Feldman <i>et al.</i>	(Mark II Collab.)
HAYES	82	PR D25 2869	K.G. Hayes <i>et al.</i>	(Mark II Collab.)
BERGER	81B	PL 99B 489	C. Berger <i>et al.</i>	(PLUTO Collab.)
DORFAN	81	PRL 46 215	J.M. Dorfan <i>et al.</i>	(Mark II Collab.)
BRANDELIK	80	PL 92B 199	R. Brandelik <i>et al.</i>	(TASSO Collab.)
ZHOLENTZ	80	PL 96B 214	A.A. Zholents <i>et al.</i>	(NOVO)
Also	81	SJNP 34 814	A.A. Zholents <i>et al.</i>	(NOVO)
		Translated from YAF 34 1471.		
BACINO	79B	PRL 42 749	W.J. Bacino <i>et al.</i>	(DELCO Collab.)
KIRKBY	79	SLAC-PUB-2419	J. Kirkby	(SLAC) J
		Batavia Lepton Photon Conference.		
BACINO	78B	PRL 41 13	W.J. Bacino <i>et al.</i>	(DELCO Collab.) J
Also	78	Tokyo Conf. 249	J. Kirz	(STON)
Also	80	PL 96B 214	A.A. Zholents <i>et al.</i>	(NOVO)
BRANDELIK	78	PL 73B 109	R. Brandelik <i>et al.</i>	(DASP Collab.) J
FELDMAN	78	Tokyo Conf. 777	G.J. Feldman	(SLAC) J
JAROS	78	PRL 40 1120	J. Jaros <i>et al.</i>	(SLAC, LBL, NWES, HAWA)
PERL	75	PRL 35 1489	M.L. Perl <i>et al.</i>	(LBL, SLAC)

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PICH	90	MPL A5 1995	A. Pich	(VALE)
BARISH	88	PRPL 157 1	B.C. Barish, R. Stroynowski	(CIT)
GAN	88	IJMP A3 531	K.K. Gan, M.L. Perl	(SLAC)
HAYES	88	PR D38 3351	K.G. Hayes, M.L. Perl	(SLAC)
PERL	80	ARNPS 30 299	M.L. Perl	(SLAC)