

**$a_1(1260)$** 

$$I^G(J^{PC}) = 1^-(1^{++})$$

**THE  $a_1(1260)$  AND  $a_1(1640)$** 

Updated December 2003 by S. Eidelman (Novosibirsk).

The main experimental data on the  $a_1(1260)$  may be grouped into two classes:

**(1) Hadronic Production:** This comprises diffractive production with incident  $\pi^-$  (DAUM 80, 81B) and charge-exchange production with low-energy  $\pi^-$  (DANKOWYCH 81, ANDO 92). The 1980's experiments explain the  $I^G L J^P = 1^+ S 0^+$  data using a phenomenological amplitude consisting of a rescattered Deck amplitude plus a direct resonance-production term. They agree on an  $a_1(1260)$  mass of about 1270 MeV and a width of 300–380 MeV. ANDO 92 finds rather lower values for the mass (1121 MeV) and width (239 MeV) in a partial-wave analysis based on the isobar model of the  $\pi^+\pi^-\pi^0$  system. However, in this analysis, only Breit-Wigner terms were considered. Recently BARBERIS 98B studied central production of the  $\pi^+\pi^-\pi^0$  system, and observed the  $a_1(1260)$  meson with a mass of 1240 MeV and a width of about 400 MeV.

CONDO 93 found no evidence for charge-exchange photoproduction of the  $a_1(1260)$  (but found a clear signal of  $a_2(1320)$  photoproduction). Similarly, MOLCHANOV 01 found no evidence for Coulomb production of  $a_1(1260)$  studying coherent production of three pions (but observed a prominent signal of  $a_2(1320)$ ). They show that it is consistent with either an extremely large  $a_1(1260)$  hadronic width, or with a small radiative width to  $\pi\gamma$ , which could be accommodated if the  $a_1$  mass is somewhat below 1260 MeV.

**(2)  $\tau$  Decay:** Various experiments reported good data on  $\tau \rightarrow a_1(1260)\nu_\tau \rightarrow \rho\pi\nu_\tau$  (RUCKSTUHL 86, SCHMIDKE 86,

ALBRECHT 86B, BAND 87, ACKERSTAFF 97R, ABREU 98G, and ASNER 00). They are somewhat inconsistent concerning the  $a_1(1260)$  mass, which can, however, be attributed to model-dependent systematic uncertainties (BOWLER 86, ALBRECHT 93C, ACKERSTAFF 97R). They all find a width greater than 400 MeV.

The discrepancies between the hadronic- and  $\tau$ -decay results have stimulated several reanalyses. BASDEVANT 77, 78 used the early diffractive dissociation and  $\tau$  decay data, and showed that they could be well reproduced with an  $a_1$  resonance mass of  $1180 \pm 50$  MeV and width of  $400 \pm 50$  MeV. Later, BOWLER 86, TORNQVIST 87, ISGUR 89, and IVANOV 91 have studied the process  $\tau \rightarrow 3\pi\nu_\tau$ . Despite quite different approaches, they all found a good overall description of the  $\tau$ -decay data with an  $a_1(1260)$  mass near 1230 MeV, consistent with the hadronic data. However, their widths remain significantly larger (400–600 MeV) than those extracted from diffractive-hadronic data. This is also the case with the later OPAL experiment (ACKERSTAFF 97R). In the high statistics analysis of ACKERSTAFF 97R, the models of ISGUR 89 and KUHN 90 are used to fit distributions of the  $3\pi$  invariant mass, as well as the  $2\pi$  invariant mass projections of the Dalitz plot, and neither model is found to provide a completely satisfactory description of the data. Another recent high statistics analysis of ABREU 98G obtains a good description of the  $\tau \rightarrow 3\pi$  data using the model of FEINDT 90, which includes the  $a_1(1640)$  meson, most probably a radial excitation of the  $a_1(1260)$  meson (BARNES 97), with a mass of 1700 MeV and a width of 300 MeV. A similar signal has been observed in various hadroproduction processes: by AMELIN 95B and CHUNG 02 in the  $D$ -wave of the  $\rho\pi$  state, by GOUZ 92 in the  $f_1(1285)\pi$  state, and by BAKER 99 in the  $f_0(600)\pi$  state, as well as by BAKER 03 in the  $\omega\pi^+\pi^-$

state. The existence of such a resonance is also suggested by the very big data sample of ASNER 00, which shows an excess of events at high  $3\pi$  mass. Their data are better described by the  $a_1'$  contribution, though at a level below that reported by ABREU 98G. Since the statistical significance of the  $a_1'$  contribution is 2–3 standard deviations only, they conclude that more data is needed to establish the  $a_1'$  existence.

ASNER 00 has also performed the analysis of the substructures in the Dalitz plot and found significant contributions of the  $a_1$  decay to  $f_0(600)\pi$ ,  $f_0(1370)\pi$ , and  $f_2(1270)\pi$ . The contribution of the  $a_1 \rightarrow f_0(600)\pi$  at a similar level has independently been observed in  $e^+e^- \rightarrow 4\pi$  annihilation (AKHMETSHIN 99E), where the  $2\pi^+2\pi^-$  final state was shown to be dominated by the  $a_1(1260)\pi$  mechanism. Note that existence of the isoscalar contributions to the two-pion state, in addition to the isovector one ( $\rho\pi$ ), will influence the ratio  $B(a_1^- \rightarrow \pi^-\pi^+\pi^-)/B(a_1^- \rightarrow \pi^-\pi^0\pi^0)$ , which should be equal to 1 for the pure  $\rho\pi$  state. The fit of ASNER 00 improves when the  $K\bar{K}^*$  (892) threshold is included. Recently DRUTSKOY 02 found direct evidence for the decay mode  $a_1(1260) \rightarrow K\bar{K}^*(892) + c.c.$  in  $B$ -meson decays.

BOWLER 88 showed that good fits to both the hadronic and the  $\tau$ -decay data could be obtained with a width of about 400 MeV. However, applying the same type of analysis to the ANDO 92 data, the low mass and narrow width they obtained with the Breit-Wigner PWA do not change appreciably.

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### **$a_1(1260)$ MASS**

<u>VALUE (MeV)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>CHG</u>	<u>COMMENT</u>
<b><math>1230 \pm 40</math></b>					<b>OUR ESTIMATE</b>

• • • We do not use the following data for averages, fits, limits, etc. • • •

1331±10± 3	37k	<sup>1</sup> ASNER	00	CLE2	10.6 $e^+ e^- \rightarrow \tau^+ \tau^-, \tau^- \tau^+$ $\pi^- \pi^0 \pi^0 \nu_\tau$
1255± 7± 6	5904	<sup>2</sup> ABREU	98G	DLPH	$e^+ e^-$
1207± 5± 8	5904	<sup>3</sup> ABREU	98G	DLPH	$e^+ e^-$
1196± 4± 5	5904	<sup>4,5</sup> ABREU	98G	DLPH	$e^+ e^-$
1240±10		BARBERIS	98B		450 $p p \rightarrow p_f \pi^+ \pi^- \pi^0 p_s$
1262± 9± 7		<sup>2,6</sup> ACKERSTAFF	97R	OPAL	$E_{cm}^{ee} = 88-94, \tau \rightarrow 3\pi \nu$
1210± 7± 2		<sup>3,6</sup> ACKERSTAFF	97R	OPAL	$E_{cm}^{ee} = 88-94, \tau \rightarrow 3\pi \nu$
1211± 7 <sup>+50</sup> <sub>-0</sub>		<sup>3</sup> ALBRECHT	93C	ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1121± 8		<sup>7</sup> ANDO	92	SPEC	8 $\pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
1242±37		<sup>8</sup> IVANOV	91	RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
1260±14		<sup>9</sup> IVANOV	91	RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
1250± 9		<sup>10</sup> IVANOV	91	RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
1208±15		ARMSTRONG	90	OMEG 0	300.0 $p p \rightarrow p p \pi^+ \pi^- \pi^0$
1220±15		<sup>11</sup> ISGUR	89	RVUE	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1260±25		<sup>12</sup> BOWLER	88	RVUE	
1166±18±11		BAND	87	MAC	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1164±41±23		BAND	87	MAC	$\tau^+ \rightarrow \pi^+ \pi^0 \pi^0 \nu$
1250±40		<sup>11</sup> TORNQVIST	87	RVUE	
1046±11		ALBRECHT	86B	ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1056±20±15		RUCKSTUHL	86	DLCO	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1194±14±10		SCHMIDKE	86	MRK2	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
1255±23		BELLINI	85	SPEC	40 $\pi^- A \rightarrow \pi^- \pi^+ \pi^- A$
1240±80		<sup>13</sup> DANKOWY...	81	SPEC 0	8.45 $\pi^- p \rightarrow n 3\pi$
1280±30		<sup>13</sup> DAUM	81B	CNTR	63,94 $\pi^- p \rightarrow p 3\pi$
1041±13		<sup>14</sup> GAVILLET	77	HBC +	4.2 $K^- p \rightarrow \Sigma 3\pi$

<sup>1</sup> From a fit to the  $3\pi$  mass spectrum including the  $K\bar{K}^*(892)$  threshold.

<sup>2</sup> Uses the model of KUHN 90.

<sup>3</sup> Uses the model of ISGUR 89.

<sup>4</sup> Includes the effect of a possible  $a_1'$  state.

<sup>5</sup> Uses the model of FEINDT 90.

<sup>6</sup> Supersedes AKERS 95P.

<sup>7</sup> Average and spread of values using 2 variants of the model of BOWLER 75.

<sup>8</sup> Reanalysis of RUCKSTUHL 86.

<sup>9</sup> Reanalysis of SCHMIDKE 86.

<sup>10</sup> Reanalysis of ALBRECHT 86B.

<sup>11</sup> From a combined reanalysis of ALBRECHT 86B, SCHMIDKE 86, and RUCKSTUHL 86.

<sup>12</sup> From a combined reanalysis of ALBRECHT 86B and DAUM 81B.

<sup>13</sup> Uses the model of BOWLER 75.

<sup>14</sup> Produced in  $K^-$  backward scattering.

## $a_1(1260)$ WIDTH

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
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### 250 to 600 OUR ESTIMATE

• • • We do not use the following data for averages, fits, limits, etc. • • •

460 ± 85	205	15 DRUTSKOY	02	BELL	$B \rightarrow D^{(*)} K^- K^{*0}$
814 ± 36 ± 13	37k	16 ASNER	00	CLE2	$10.6 e^+ e^- \rightarrow \tau^+ \tau^-, \tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$
450 ± 50	22k	17 AKHMETSHIN	99E	CMD2	$1.05-1.38 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0 \pi^0$
570 ± 10		18 BONDAR	99	RVUE	$e^+ e^- \rightarrow 4\pi, \tau \rightarrow 3\pi \nu_\tau$
587 ± 27 ± 21	5904	19 ABREU	98G	DLPH	$e^+ e^-$
478 ± 3 ± 15	5904	20 ABREU	98G	DLPH	$e^+ e^-$
425 ± 14 ± 8	5904	21,22 ABREU	98G	DLPH	$e^+ e^-$
400 ± 35		BARBERIS	98B		$450 p p \rightarrow p_f \pi^+ \pi^- \pi^0 p_s$
621 ± 32 ± 58		19,23 ACKERSTAFF	97R	OPAL	$E_{cm}^{ee} = 88-94, \tau \rightarrow 3\pi \nu$
457 ± 15 ± 17		20,23 ACKERSTAFF	97R	OPAL	$E_{cm}^{ee} = 88-94, \tau \rightarrow 3\pi \nu$
446 ± 21 <sup>+140</sup> / <sub>0</sub>		20 ALBRECHT	93C	ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
239 ± 11		ANDO	92	SPEC	$8 \pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
266 ± 13 ± 4		24 ANDO	92	SPEC	$8 \pi^- p \rightarrow \pi^+ \pi^- \pi^0 n$
465 <sup>+228</sup> / <sub>-143</sub>		25 IVANOV	91	RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
298 <sup>+40</sup> / <sub>-34</sub>		26 IVANOV	91	RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
488 ± 32		27 IVANOV	91	RVUE	$\tau \rightarrow \pi^+ \pi^+ \pi^- \nu$
430 ± 50		ARMSTRONG	90	OMEG 0	$300.0 p p \rightarrow p p \pi^+ \pi^- \pi^0$
420 ± 40		28 ISGUR	89	RVUE	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
396 ± 43		29 BOWLER	88	RVUE	
405 ± 75 ± 25		BAND	87	MAC	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
419 ± 108 ± 57		BAND	87	MAC	$\tau^+ \rightarrow \pi^+ \pi^0 \pi^0 \nu$
521 ± 27		ALBRECHT	86B	ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
476 <sup>+132</sup> / <sub>-120</sub> ± 54		RUCKSTUHL	86	DLCO	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
462 ± 56 ± 30		SCHMIDKE	86	MRK2	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$
292 ± 40		BELLINI	85	SPEC	$40 \pi^- A \rightarrow \pi^- \pi^+ \pi^- A$
380 ± 100		30 DANKOWY...	81	SPEC 0	$8.45 \pi^- p \rightarrow n 3\pi$
300 ± 50		30 DAUM	81B	CNTR	$63,94 \pi^- p \rightarrow p 3\pi$
230 ± 50		31 GAVILLET	77	HBC +	$4.2 K^- p \rightarrow \Sigma 3\pi$

<sup>15</sup> From a fit of the  $K^- K^{*0}$  distribution assuming  $m_{a_1} = 1230$  MeV and purely resonant production of the  $K^- K^{*0}$  system.

<sup>16</sup> From a fit to the  $3\pi$  mass spectrum including the  $K\bar{K}^*(892)$  threshold.

<sup>17</sup> Using the  $a_1(1260)$  mass of 1230 MeV.

<sup>18</sup> From AKHMETSHIN 99E and ASNER 00 data using the  $a_1(1260)$  mass of 1230 MeV.

- 19 Uses the model of KUHN 90.  
 20 Uses the model of ISGUR 89.  
 21 Includes the effect of a possible  $a_1'$  state.  
 22 Uses the model of FEINDT 90.  
 23 Supersedes AKERS 95P.  
 24 Average and spread of values using 2 variants of the model of BOWLER 75.  
 25 Reanalysis of RUCKSTUHL 86.  
 26 Reanalysis of SCHMIDKE 86.  
 27 Reanalysis of ALBRECHT 86B.  
 28 From a combined reanalysis of ALBRECHT 86B, SCHMIDKE 86, and RUCKSTUHL 86.  
 29 From a combined reanalysis of ALBRECHT 86B and DAUM 81B.  
 30 Uses the model of BOWLER 75.  
 31 Produced in  $K^-$  backward scattering.

### $a_1(1260)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )
$\Gamma_1$ $\pi^+ \pi^- \pi^0$	
$\Gamma_2$ $\pi^0 \pi^0 \pi^0$	
$\Gamma_3$ $(\rho\pi)_{S\text{-wave}}$	seen
$\Gamma_4$ $(\rho\pi)_{D\text{-wave}}$	seen
$\Gamma_5$ $(\rho(1450)\pi)_{S\text{-wave}}$	seen
$\Gamma_6$ $(\rho(1450)\pi)_{D\text{-wave}}$	seen
$\Gamma_7$ $\sigma\pi$	seen
$\Gamma_8$ $f_0(980)\pi$	not seen
$\Gamma_9$ $f_0(1370)\pi$	seen
$\Gamma_{10}$ $f_2(1270)\pi$	seen
$\Gamma_{11}$ $K\bar{K}^*(892) + \text{c.c.}$	seen
$\Gamma_{12}$ $\pi\gamma$	seen

### $a_1(1260)$ PARTIAL WIDTHS

$\Gamma(\pi\gamma)$				$\Gamma_{12}$
VALUE (keV)	DOCUMENT ID	TECN	COMMENT	
<b><math>640 \pm 246</math></b>	ZIELINSKI	84C SPEC	$200 \pi^+ Z \rightarrow Z 3\pi$	

### D-wave/S-wave AMPLITUDE RATIO IN DECAY OF $a_1(1260) \rightarrow \rho\pi$

VALUE	DOCUMENT ID	TECN	COMMENT
<b><math>-0.108 \pm 0.016</math> OUR AVERAGE</b>			
$-0.14 \pm 0.04 \pm 0.07$	34 CHUNG	02 MPS	$18.3 \pi^- p \rightarrow \pi^+ \pi^- \pi^- p$
$-0.10 \pm 0.02 \pm 0.02$	32,33 ACKERSTAFF	97R OPAL	$E_{\text{cm}}^{\text{e}^{\text{e}}} = 88\text{--}94, \tau \rightarrow 3\pi\nu$
$-0.11 \pm 0.02$	32 ALBRECHT	93C ARG	$\tau^+ \rightarrow \pi^+ \pi^+ \pi^- \nu$

32 Uses the model of ISGUR 89.

33 Supersedes AKERS 95P.

34 Deck-type background not subtracted.

## $a_1(1260)$ BRANCHING RATIOS

### $\Gamma((\rho\pi)_{S\text{-wave}})/\Gamma_{\text{total}}$ $\Gamma_3/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

60.19	37k	<sup>35</sup> ASNER	00 CLE2	10.6 $e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$
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### $\Gamma((\rho\pi)_{D\text{-wave}})/\Gamma_{\text{total}}$ $\Gamma_4/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.30 \pm 0.60 \pm 0.22$	37k	<sup>35</sup> ASNER	00 CLE2	10.6 $e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$
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### $\Gamma((\rho(1450)\pi)_{S\text{-wave}})/\Gamma_{\text{total}}$ $\Gamma_5/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.56 \pm 0.84 \pm 0.32$	37k	<sup>35,36</sup> ASNER	00 CLE2	10.6 $e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$
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### $\Gamma((\rho(1450)\pi)_{D\text{-wave}})/\Gamma_{\text{total}}$ $\Gamma_6/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$2.04 \pm 1.20 \pm 0.28$	37k	<sup>35,36</sup> ASNER	00 CLE2	10.6 $e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$
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### $\Gamma(\sigma\pi)/\Gamma_{\text{total}}$ $\Gamma_7/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

seen		CHUNG	02 MPS	$18.3 \pi^- p \rightarrow \pi^+\pi^-\pi^-p$
$18.76 \pm 4.29 \pm 1.48$	37k	<sup>35,37</sup> ASNER	00 CLE2	10.6 $e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$

### $\Gamma(f_0(980)\pi)/\Gamma_{\text{total}}$ $\Gamma_8/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

not seen	37k	ASNER	00 CLE2	10.6 $e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$
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### $\Gamma(f_0(1370)\pi)/\Gamma_{\text{total}}$ $\Gamma_9/\Gamma$

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$7.40 \pm 2.71 \pm 1.26$	37k	<sup>35,38</sup> ASNER	00 CLE2	10.6 $e^+e^- \rightarrow \tau^+\tau^-$ , $\tau^- \rightarrow \pi^-\pi^0\pi^0\nu_\tau$
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**$\Gamma(f_2(1270)\pi)/\Gamma_{\text{total}}$   $\Gamma_{10}/\Gamma$**

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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••• We do not use the following data for averages, fits, limits, etc. •••

$1.19 \pm 0.49 \pm 0.17$	37k	<sup>35,39</sup> ASNER	00 CLE2	$10.6 e^+ e^- \rightarrow \tau^+ \tau^-$ , $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$
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**$\Gamma(K\bar{K}^*(892)+\text{c.c.})/\Gamma_{\text{total}}$   $\Gamma_{11}/\Gamma$**

VALUE (units $10^{-2}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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••• We do not use the following data for averages, fits, limits, etc. •••

8 to 15	205	<sup>40</sup> DRUTSKOY	02 BELL	$B \rightarrow D^{(*)} K^- K^{*0}$
$3.3 \pm 0.5 \pm 0.1$	37k	<sup>41</sup> ASNER	00 CLE2	$10.6 e^+ e^- \rightarrow \tau^+ \tau^-$ , $\tau^- \rightarrow \pi^- \pi^0 \pi^0 \nu_\tau$

**$\Gamma(\sigma\pi)/\Gamma((\rho\pi)_{S\text{-wave}})$   $\Gamma_7/\Gamma_3$**

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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••• We do not use the following data for averages, fits, limits, etc. •••

$\sim 0.3$	28k	AKHMETSHIN 99E	CMD2	$1.05\text{--}1.38 e^+ e^- \rightarrow \pi^+ \pi^- \pi^+ \pi^-$
$0.003 \pm 0.003$		<sup>42</sup> LONGACRE	82 RVUE	

**$\Gamma(\pi^0 \pi^0 \pi^0)/\Gamma(\pi^+ \pi^- \pi^0)$   $\Gamma_2/\Gamma_1$**

VALUE	CL%	DOCUMENT ID	COMMENT
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••• We do not use the following data for averages, fits, limits, etc. •••

$< 0.008$	90	<sup>43</sup> BARBERIS	01 450 $p p \rightarrow p_f 3\pi^0 p_s$
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<sup>35</sup> From a fit to the Dalitz plot.

<sup>36</sup> Assuming for  $\rho(1450)$  mass and width of 1370 and 386 MeV respectively.

<sup>37</sup> Assuming for  $\sigma$  mass and width of 860 and 880 MeV respectively.

<sup>38</sup> Assuming for  $f_0(1370)$  mass and width of 1186 and 350 MeV respectively.

<sup>39</sup> Assuming for  $f_2(1270)$  mass and width of 1275 and 185 MeV respectively.

<sup>40</sup> From a comparison to ALAM 94 assuming purely resonant production of the  $K^- K^{*0}$  system.

<sup>41</sup> From a fit to the  $3\pi$  mass spectrum including the  $K\bar{K}^*(892)$  threshold.

<sup>42</sup> Uses multichannel Aitchison-Bowler model (BOWLER 75). Uses data from GAVILLET 77, DAUM 80, and DANKOWYCH 81.

<sup>43</sup> Inconsistent with observations of  $\sigma\pi$ ,  $f_0(1370)\pi$ , and  $f_2(1270)\pi$  decay modes.

**$a_1(1260)$  REFERENCES**

CHUNG	02	PR D65 072001	S.U. Chung <i>et al.</i>	
DRUTSKOY	02	PL B542 171	A. Drutskoy <i>et al.</i>	(BELLE Collab.)
BARBERIS	01	PL B507 14	D. Barberis <i>et al.</i>	
ASNER	00	PR D61 012002	D.M. Asner <i>et al.</i>	(CLEO Collab.)
AKHMETSHIN	99E	PL B466 392	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
BONDAR	99	PL B466 403	A.E. Bondar <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
ABREU	98G	PL B426 411	P. Abreu <i>et al.</i>	(DELPHI Collab.)
BARBERIS	98B	PL B422 399	D. Barberis <i>et al.</i>	(WA 102 Collab.)
ACKERSTAFF	97R	ZPHY C75 593	K. Ackerstaff <i>et al.</i>	(OPAL Collab.)
AKERS	95P	ZPHY C67 45	R. Akers <i>et al.</i>	(OPAL Collab.)
ALAM	94	PR D50 43	M.S. Alam <i>et al.</i>	(CLEO Collab.)
ALBRECHT	93C	ZPHY C58 61	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ANDO	92	PL B291 496	A. Ando <i>et al.</i>	(KEK, KYOT, NIRS, SAGA+)
IVANOV	91	ZPHY C49 563	Y.P. Ivanov, A.A. Osipov, M.K. Volkov	(JINR)
ARMSTRONG	90	ZPHY C48 213	T.A. Armstrong, M. Benayoun, W. Beusch	



FEINDT	90	ZPHY C48 681	M. Feindt	(HAMB)
KUHN	90	ZPHY C48 445	J.H. Kuhn <i>et al.</i>	(MPIM)
ISGUR	89	PR D39 1357	N. Isgur, C. Morningstar, C. Reader	(TNTO)
BOWLER	88	PL B209 99	M.G. Bowler	(OXF)
BAND	87	PL B198 297	H.R. Band <i>et al.</i>	(MAC Collab.)
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ALBRECHT	86B	ZPHY C33 7	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
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