



$$I(J^P) = \frac{1}{2}(0^-)$$

## $D^0$ MASS

The fit includes  $D^\pm$ ,  $D^0$ ,  $D_s^\pm$ ,  $D^{*\pm}$ ,  $D^{*0}$ , and  $D_s^{*\pm}$  mass and mass difference measurements.

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1864.5 ± 0.5 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>1864.1 ± 1.0 OUR AVERAGE</b>				
1864.6 ± 0.3 ± 1.0	641	BARLAG	90C ACCM	$\pi^-$ Cu 230 GeV
1852 ± 7	16	ADAMOVICH	87 EMUL	Photoproduction
1861 ± 4		DERRICK	84 HRS	$e^+ e^-$ 29 GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
1856 ± 36	22	ADAMOVICH	84B EMUL	Photoproduction
1847 ± 7	1	FIORINO	81 EMUL	$\gamma N \rightarrow \bar{D}^0 +$
1863.8 ± 0.5		<sup>1</sup> SCHINDLER	81 MRK2	$e^+ e^-$ 3.77 GeV
1864.7 ± 0.6		<sup>1</sup> TRILLING	81 RVUE	$e^+ e^-$ 3.77 GeV
1863.0 ± 2.5	238	ASTON	80E OMEG	$\gamma p \rightarrow \bar{D}^0$
1860 ± 2	143	<sup>2</sup> AVERY	80 SPEC	$\gamma N \rightarrow D^{*+}$
1869 ± 4	35	<sup>2</sup> AVERY	80 SPEC	$\gamma N \rightarrow D^{*+}$
1854 ± 6	94	<sup>2</sup> ATIYA	79 SPEC	$\gamma N \rightarrow D^0 \bar{D}^0$
1850 ± 15	64	BALTAY	78C HBC	$\nu N \rightarrow K^0 \pi \pi$
1863 ± 3		GOLDHABER	77 MRK1	$D^0, D^+$ recoil spectra
1863.3 ± 0.9		<sup>1</sup> PERUZZI	77 MRK1	$e^+ e^-$ 3.77 GeV
1868 ± 11		PICCOLO	77 MRK1	$e^+ e^-$ 4.03, 4.41 GeV
1865 ± 15	234	GOLDHABER	76 MRK1	$K\pi$ and $K3\pi$

<sup>1</sup>PERUZZI 77 and SCHINDLER 81 errors do not include the 0.13% uncertainty in the absolute SPEAR energy calibration. TRILLING 81 uses the high precision  $J/\psi(1S)$  and  $\psi(2S)$  measurements of ZHOLENTZ 80 to determine this uncertainty and combines the PERUZZI 77 and SCHINDLER 81 results to obtain the value quoted. TRILLING 81 enters the fit in the  $D^\pm$  mass, and PERUZZI 77 and SCHINDLER 81 enter in the  $m_{D^\pm} - m_{D^0}$ , below.

<sup>2</sup>Error does not include possible systematic mass scale shift, estimated to be less than 5 MeV.

## $m_{D^\pm} - m_{D^0}$

The fit includes  $D^\pm$ ,  $D^0$ ,  $D_s^\pm$ ,  $D^{*\pm}$ ,  $D^{*0}$ , and  $D_s^{*\pm}$  mass and mass difference measurements.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
<b>4.79 ± 0.10 OUR FIT</b>	Error includes scale factor of 1.1.		
<b>4.74 ± 0.28 OUR AVERAGE</b>			
4.7 ± 0.3	<sup>3</sup> SCHINDLER	81 MRK2	$e^+ e^-$ 3.77 GeV
5.0 ± 0.8	<sup>3</sup> PERUZZI	77 MRK1	$e^+ e^-$ 3.77 GeV

<sup>3</sup>See the footnote on TRILLING 81 in the  $D^0$  and  $D^\pm$  sections on the mass.

## $D^0$ MEAN LIFE

Measurements with an error  $> 0.05 \times 10^{-12}$  s are omitted from the average, and those with an error  $> 0.1 \times 10^{-12}$  s or that have been superseded by later results have been removed from the Listings.

<u>VALUE (<math>10^{-12}</math> s)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.4126 ± 0.0028 OUR AVERAGE</b>				
0.413 ± 0.003 ± 0.004	35k	AITALA	99E E791	$K^- \pi^+$
0.4085 ± 0.0041 $^{+0.0035}_{-0.0034}$	25k	BONVICINI	99 CLE2	$e^+ e^- \approx \Upsilon(4S)$
0.413 ± 0.004 ± 0.003	16k	FRABETTI	94D E687	$K^- \pi^+$ ,
0.424 ± 0.011 ± 0.007	5118	FRABETTI	91 E687	$K^- \pi^+$ ,
0.417 ± 0.018 ± 0.015	890	ALVAREZ	90 NA14	$K^- \pi^+$ ,
0.388 $^{+0.023}_{-0.021}$	641	<sup>4</sup> BARLAG	90C ACCM	$\pi^-$ Cu 230 GeV
0.48 ± 0.04 ± 0.03	776	ALBRECHT	88I ARG	$e^+ e^-$ 10 GeV
0.422 ± 0.008 ± 0.010	4212	RAAB	88 E691	Photoproduction
0.42 ± 0.05	90	BARLAG	87B ACCM	$K^-$ and $\pi^-$ 200 GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
0.34 $^{+0.06}_{-0.05}$ ± 0.03	58	AMENDOLIA	88 SPEC	Photoproduction
0.46 $^{+0.06}_{-0.05}$	145	AGUILAR-...	87D HYBR	$\pi^- p$ and $pp$
0.50 ± 0.07 ± 0.04	317	CSORNA	87 CLEO	$e^+ e^-$ 10 GeV
0.61 ± 0.09 ± 0.03	50	ABE	86 HYBR	$\gamma p$ 20 GeV
0.47 $^{+0.09}_{-0.08}$ ± 0.05	74	GLADNEY	86 MRK2	$e^+ e^-$ 29 GeV
0.43 $^{+0.07}_{-0.05}$ $^{+0.01}_{-0.02}$	58	USHIDA	86B EMUL	$\nu$ wideband
0.37 $^{+0.10}_{-0.07}$	26	BAILEY	85 SILI	$\pi^-$ Be 200 GeV

<sup>4</sup> BARLAG 90C estimate systematic error to be negligible.

**THERE WILL BE A NOTE ON  $\Delta m$  AND  $\Delta \Gamma$  IN THE 2002 EDITION.**

$$|m_{D_1^0} - m_{D_2^0}|$$

The  $D_1^0$  and  $D_2^0$  are the mass eigenstates of the  $D^0$  meson. To calculate the following limits, we use  $\Delta m = [2r/(1-r)]^{1/2} \hbar/4.126 \times 10^{-13}$  s, where  $r$  is the experimental  $D^0$ - $\bar{D}^0$  mixing ratio.

<u>VALUE (<math>10^{10} \hbar s^{-1}</math>)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt; 7</b>	95	<sup>5</sup> GODANG	00 CLE2	$e^+ e^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<32	90	6,7 AITALA	98 E791	$\pi^-$ nucleus, 500 GeV
<24	90	8 AITALA	96C E791	$\pi^-$ nucleus, 500 GeV
<21	90	7,9 ANJOS	88C E691	Photoproduction

<sup>5</sup>This GODANG 00 limit is inferred from the  $D^0-\bar{D}^0$  mixing ratio  $\Gamma(K^+\pi^-)$  (via  $\bar{D}^0$ )/ $\Gamma(K^-\pi^+)$  given near the end of this  $D^0$  Listings. Decay-time information is used to distinguish DCS decays from  $D^0-\bar{D}^0$  mixing. The limit allows interference between the DCS and mixing ratios, and also allows  $CP$  violation. The strong phase between  $D^0 \rightarrow K^+\pi^-$  and  $\bar{D}^0 \rightarrow K^+\pi^-$  is assumed to be small.

<sup>6</sup>AITALA 98 allows interference between the doubly Cabibbo-suppressed and mixing amplitudes, and also allows  $CP$  violation in this term.

<sup>7</sup>This limit is inferred from the  $D^0-\bar{D}^0$  mixing ratio  $\Gamma(K^+\pi^-$  or  $K^+\pi^-\pi^+\pi^-$  (via  $\bar{D}^0$ ))/ $\Gamma(K^-\pi^+$  or  $K^-\pi^+\pi^+\pi^-$ ) near the end of the  $D^0$  Listings. Decay-time information is used to distinguish doubly Cabibbo-suppressed decays from  $D^0-\bar{D}^0$  mixing.

<sup>8</sup>This limit is inferred from the  $D^0-\bar{D}^0$  mixing ratio  $\Gamma(K^+\ell^-\bar{\nu}_\ell)$  (via  $\bar{D}^0$ )/ $\Gamma(K^-\ell^+\nu_\ell)$  given near the end of the  $D^0$  Listings.

<sup>9</sup>ANJOS 88C assumes no interference between doubly Cabibbo-suppressed and mixing amplitudes. When interference is allowed, the limit degrades by about a factor of two.

### THERE WILL BE A NOTE ON $\Delta m$ AND $\Delta\Gamma$ IN THE 2002 EDITION.

$$(\Gamma_{D_1^0} - \Gamma_{D_2^0})/\Gamma_{D^0} = 2y$$

The  $D_1^0$  and  $D_2^0$  are the mass eigenstates of the  $D^0$  meson. AITALA 99E and LINK 00 use a difference in directly measured decay rates to obtain their limits. The other experiments infer the limits here from limits on mixing, using  $\Delta\Gamma/\Gamma = [8r/(1+r)]^{1/2}$ , where  $r$  is the experimental  $D^0-\bar{D}^0$  mixing ratio. See the footnotes to the entries below.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0684 ± 0.0278 ± 0.0148</b>		10k	10 LINK	00 FOCS	$\gamma$ nucleus

• • • We do not use the following data for averages, fits, limits, etc. • • •

$-0.116 < \Delta\Gamma/\Gamma < 0.020$	95	11	GODANG	00 CLE2	$e^+e^-$
$-0.08 < \Delta\Gamma/\Gamma < 0.12$	90	12	AITALA	99E E791	$K^-\pi^+$ , $K^+K^-$
$ \Delta\Gamma /\Gamma < 0.26$	90	13,14	AITALA	98 E791	$\pi^-$ nucleus, 500 GeV
$ \Delta\Gamma /\Gamma < 0.20$	90	15	AITALA	96C E791	$\pi^-$ nucleus, 500 GeV
$ \Delta\Gamma /\Gamma < 0.17$	90	14,16	ANJOS	88C E691	Photoproduction

<sup>10</sup>LINK 00 measures the lifetime difference between  $D^0 \rightarrow K^-K^+$  ( $CP$  even) decays and  $D^0 \rightarrow K^-\pi^+$  ( $CP$  mixed) decays, or  $y_{CP} = [\Gamma(CP+) - \Gamma(CP-)]/[\Gamma(CP+) + \Gamma(CP-)]$ . We list  $2y_{CP} = \Delta\Gamma/\Gamma$ .

<sup>11</sup>This GODANG 00 limit is inferred from the  $D^0-\bar{D}^0$  mixing ratio  $\Gamma(K^+\pi^-)$  (via  $\bar{D}^0$ )/ $\Gamma(K^-\pi^+)$  given near the end of this  $D^0$  Listings. Decay-time information is used to distinguish DCS decays from  $D^0-\bar{D}^0$  mixing. The limit allows interference between the DCS and mixing ratios, and also allows  $CP$  violation. The phase between  $D^0 \rightarrow K^+\pi^-$  and  $\bar{D}^0 \rightarrow K^+\pi^-$  is assumed to be small. This is a measurement of  $y'$  and is not the same as the  $y_{CP}$  of LINK 00.

<sup>12</sup>AITALA 99E measures  $\Delta\Gamma = 2[\Gamma(D^0 \rightarrow K^+K^-) - \Gamma(D^0 \rightarrow K^-\pi^+)] = +0.04 \pm 0.14 \pm 0.05 \text{ ps}^{-1}$  and thus gets 90%-confidence-level limits  $-0.20 < \Delta\Gamma < +0.28 \text{ ps}^{-1}$ .

- <sup>13</sup> AITALA 98 allows interference between the doubly Cabibbo-suppressed and mixing amplitudes, and also allows *CP* violation in this term.
- <sup>14</sup> This limit is inferred from the  $D^0$ - $\bar{D}^0$  mixing ratio  $\Gamma(K^+\pi^- \text{ or } K^+\pi^-\pi^+\pi^- \text{ (via } \bar{D}^0)) / \Gamma(K^-\pi^+ \text{ or } K^-\pi^+\pi^+\pi^-)$  near the end of the  $D^0$  Listings. Decay-time information is used to distinguish doubly Cabibbo-suppressed decays from  $D^0$ - $\bar{D}^0$  mixing.
- <sup>15</sup> This limit is inferred from the  $D^0$ - $\bar{D}^0$  mixing ratio  $\Gamma(K^+\ell^-\bar{\nu}_\ell \text{ (via } \bar{D}^0)) / \Gamma(K^-\ell^+\nu_\ell)$  given near the end of the  $D^0$  Listings.
- <sup>16</sup> ANJOS 88C assumes no interference between doubly Cabibbo-suppressed and mixing amplitudes. When interference is allowed, the limit degrades by about a factor of two.

## $D^0$ DECAY MODES

$\bar{D}^0$  modes are charge conjugates of the modes below.

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
<b>Inclusive modes</b>		
$\Gamma_1$ $e^+$ anything	( 6.75 ± 0.29 ) %	
$\Gamma_2$ $\mu^+$ anything	( 6.6 ± 0.8 ) %	
$\Gamma_3$ $K^-$ anything	( 53 ± 4 ) %	S=1.3
$\Gamma_4$ $\bar{K}^0$ anything + $K^0$ anything	( 42 ± 5 ) %	
$\Gamma_5$ $K^+$ anything	( 3.4 $^{+0.6}_{-0.4}$ ) %	
$\Gamma_6$ $\eta$ anything	[a] < 13 %	CL=90%
$\Gamma_7$ $\phi$ anything	( 1.7 ± 0.8 ) %	
<b>Semileptonic modes</b>		
$\Gamma_8$ $K^-\ell^+\nu_\ell$	[b] ( 3.47 ± 0.17 ) %	S=1.3
$\Gamma_9$ $K^-e^+\nu_e$	( 3.64 ± 0.18 ) %	
$\Gamma_{10}$ $K^-\mu^+\nu_\mu$	( 3.22 ± 0.17 ) %	
$\Gamma_{11}$ $K^-\pi^0e^+\nu_e$	( 1.6 $^{+1.3}_{-0.5}$ ) %	
$\Gamma_{12}$ $\bar{K}^0\pi^-e^+\nu_e$	( 2.8 $^{+1.7}_{-0.9}$ ) %	
$\Gamma_{13}$ $\bar{K}^*(892)^-e^+\nu_e$ × B( $K^{*-} \rightarrow \bar{K}^0\pi^-$ )	( 1.35 ± 0.22 ) %	
$\Gamma_{14}$ $K^*(892)^-\ell^+\nu_\ell$		
$\Gamma_{15}$ $\bar{K}^*(892)^0\pi^-e^+\nu_e$		
$\Gamma_{16}$ $K^-\pi^+\pi^-\mu^+\nu_\mu$	< 1.2 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{17}$ $(\bar{K}^*(892)\pi)^-\mu^+\nu_\mu$	< 1.4 × 10 <sup>-3</sup>	CL=90%
$\Gamma_{18}$ $\pi^-e^+\nu_e$	( 3.7 ± 0.6 ) × 10 <sup>-3</sup>	
A fraction of the following resonance mode has already appeared above as a submode of a charged-particle mode.		
$\Gamma_{19}$ $K^*(892)^-e^+\nu_e$	( 2.02 ± 0.33 ) %	

### Hadronic modes with a $\bar{K}$ or $\bar{K}K\bar{K}$

$\Gamma_{20}$	$K^- \pi^+$		( 3.83 ± 0.09 ) %	
$\Gamma_{21}$	$\bar{K}^0 \pi^0$		( 2.11 ± 0.21 ) %	S=1.1
$\Gamma_{22}$	$\bar{K}^0 \pi^+ \pi^-$	[c]	( 5.4 ± 0.4 ) %	S=1.2
$\Gamma_{23}$	$\bar{K}^0 \rho^0$		( 1.21 ± 0.17 ) %	
$\Gamma_{24}$	$\bar{K}^0 f_0(980)$ × B( $f_0 \rightarrow \pi^+ \pi^-$ )		( 3.0 ± 0.8 ) × 10 <sup>-3</sup>	
$\Gamma_{25}$	$\bar{K}^0 f_2(1270)$ × B( $f_2 \rightarrow \pi^+ \pi^-$ )		( 2.4 ± 0.9 ) × 10 <sup>-3</sup>	
$\Gamma_{26}$	$\bar{K}^0 f_0(1370)$ × B( $f_0 \rightarrow \pi^+ \pi^-$ )		( 4.3 ± 1.3 ) × 10 <sup>-3</sup>	
$\Gamma_{27}$	$K^*(892)^- \pi^+$ × B( $K^{*-} \rightarrow \bar{K}^0 \pi^-$ )		( 3.4 ± 0.3 ) %	
$\Gamma_{28}$	$K_0^*(1430)^- \pi^+$ × B( $K_0^*(1430)^- \rightarrow \bar{K}^0 \pi^-$ )		( 6.4 ± 1.6 ) × 10 <sup>-3</sup>	
$\Gamma_{29}$	$\bar{K}^0 \pi^+ \pi^-$ nonresonant		( 1.47 ± 0.24 ) %	
$\Gamma_{30}$	$K^- \pi^+ \pi^0$	[c]	( 13.9 ± 0.9 ) %	S=1.3
$\Gamma_{31}$	$K^- \rho^+$		( 10.8 ± 1.0 ) %	
$\Gamma_{32}$	$K^*(892)^- \pi^+$ × B( $K^{*-} \rightarrow K^- \pi^0$ )		( 1.7 ± 0.2 ) %	
$\Gamma_{33}$	$\bar{K}^*(892)^0 \pi^0$ × B( $\bar{K}^{*0} \rightarrow K^- \pi^+$ )		( 2.1 ± 0.3 ) %	
$\Gamma_{34}$	$K^- \pi^+ \pi^0$ nonresonant		( 6.9 ± 2.5 ) × 10 <sup>-3</sup>	
$\Gamma_{35}$	$\bar{K}^0 \pi^0 \pi^0$		—	
$\Gamma_{36}$	$\bar{K}^*(892)^0 \pi^0$ × B( $\bar{K}^{*0} \rightarrow \bar{K}^0 \pi^0$ )		( 1.1 ± 0.2 ) %	
$\Gamma_{37}$	$\bar{K}^0 \pi^0 \pi^0$ nonresonant		( 7.8 ± 2.0 ) × 10 <sup>-3</sup>	
$\Gamma_{38}$	$K^- \pi^+ \pi^+ \pi^-$	[c]	( 7.49 ± 0.31 ) %	
$\Gamma_{39}$	$K^- \pi^+ \rho^0$ total		( 6.3 ± 0.4 ) %	
$\Gamma_{40}$	$K^- \pi^+ \rho^0$ 3-body		( 4.7 ± 2.1 ) × 10 <sup>-3</sup>	
$\Gamma_{41}$	$\bar{K}^*(892)^0 \rho^0$ × B( $\bar{K}^{*0} \rightarrow K^- \pi^+$ )		( 9.8 ± 2.2 ) × 10 <sup>-3</sup>	
$\Gamma_{42}$	$K^- a_1(1260)^+$ × B( $a_1(1260)^+ \rightarrow \pi^+ \pi^+ \pi^-$ )		( 3.6 ± 0.6 ) %	
$\Gamma_{43}$	$\bar{K}^*(892)^0 \pi^+ \pi^-$ total × B( $\bar{K}^{*0} \rightarrow K^- \pi^+$ )		( 1.5 ± 0.4 ) %	
$\Gamma_{44}$	$\bar{K}^*(892)^0 \pi^+ \pi^-$ 3-body × B( $\bar{K}^{*0} \rightarrow K^- \pi^+$ )		( 9.5 ± 2.1 ) × 10 <sup>-3</sup>	
$\Gamma_{45}$	$K_1(1270)^- \pi^+$ × B( $K_1(1270)^- \rightarrow K^- \pi^+ \pi^-$ )	[d]	( 3.6 ± 1.0 ) × 10 <sup>-3</sup>	
$\Gamma_{46}$	$K^- \pi^+ \pi^+ \pi^-$ nonresonant		( 1.74 ± 0.25 ) %	
$\Gamma_{47}$	$\bar{K}^0 \pi^+ \pi^- \pi^0$	[c]	( 10.0 ± 1.2 ) %	

Γ <sub>48</sub>	$\bar{K}^0 \eta \times B(\eta \rightarrow \pi^+ \pi^- \pi^0)$	$(1.6 \pm 0.3) \times 10^{-3}$	
Γ <sub>49</sub>	$\bar{K}^0 \omega \times B(\omega \rightarrow \pi^+ \pi^- \pi^0)$	$(1.9 \pm 0.4) \%$	
Γ <sub>50</sub>	$K^*(892)^- \rho^+$ $\times B(K^{*-} \rightarrow \bar{K}^0 \pi^-)$	$(4.1 \pm 1.6) \%$	
Γ <sub>51</sub>	$\bar{K}^*(892)^0 \rho^0$ $\times B(\bar{K}^{*0} \rightarrow \bar{K}^0 \pi^0)$	$(4.9 \pm 1.1) \times 10^{-3}$	
Γ <sub>52</sub>	$K_1(1270)^- \pi^+$ $\times B(K_1(1270)^- \rightarrow \bar{K}^0 \pi^- \pi^0)$	$(5.1 \pm 1.4) \times 10^{-3}$	[d]
Γ <sub>53</sub>	$\bar{K}^*(892)^0 \pi^+ \pi^-$ 3-body $\times B(\bar{K}^{*0} \rightarrow \bar{K}^0 \pi^0)$	$(4.8 \pm 1.1) \times 10^{-3}$	
Γ <sub>54</sub>	$\bar{K}^0 \pi^+ \pi^- \pi^0$ nonresonant	$(2.1 \pm 2.1) \%$	
Γ <sub>55</sub>	$K^- \pi^+ \pi^0 \pi^0$	—	
Γ <sub>56</sub>	$K^- \pi^+ \pi^+ \pi^- \pi^0$	$(4.0 \pm 0.4) \%$	
Γ <sub>57</sub>	$\bar{K}^*(892)^0 \pi^+ \pi^- \pi^0$ $\times B(\bar{K}^{*0} \rightarrow K^- \pi^+)$	$(1.2 \pm 0.6) \%$	
Γ <sub>58</sub>	$\bar{K}^*(892)^0 \eta$ $\times B(\bar{K}^{*0} \rightarrow K^- \pi^+)$ $\times B(\eta \rightarrow \pi^+ \pi^- \pi^0)$	$(2.9 \pm 0.8) \times 10^{-3}$	
Γ <sub>59</sub>	$K^- \pi^+ \omega \times B(\omega \rightarrow \pi^+ \pi^- \pi^0)$	$(2.7 \pm 0.5) \%$	
Γ <sub>60</sub>	$\bar{K}^*(892)^0 \omega$ $\times B(\bar{K}^{*0} \rightarrow K^- \pi^+)$ $\times B(\omega \rightarrow \pi^+ \pi^- \pi^0)$	$(7 \pm 3) \times 10^{-3}$	
Γ <sub>61</sub>	$\bar{K}^0 \pi^+ \pi^+ \pi^- \pi^-$	$(5.8 \pm 1.6) \times 10^{-3}$	
Γ <sub>62</sub>	$\bar{K}^0 \pi^+ \pi^- \pi^0 \pi^0 (\pi^0)$	—	
Γ <sub>63</sub>	$\bar{K}^0 K^+ K^-$	$(9.4 \pm 1.0) \times 10^{-3}$	
	In the fit as $\frac{1}{2}\Gamma_{75} + \Gamma_{65}$ , where $\frac{1}{2}\Gamma_{75} = \Gamma_{64}$ .		
Γ <sub>64</sub>	$\bar{K}^0 \phi \times B(\phi \rightarrow K^+ K^-)$	$(4.3 \pm 0.5) \times 10^{-3}$	
Γ <sub>65</sub>	$\bar{K}^0 K^+ K^-$ non- $\phi$	$(5.1 \pm 0.8) \times 10^{-3}$	
Γ <sub>66</sub>	$K_S^0 K_S^0 K_S^0$	$(8.3 \pm 1.5) \times 10^{-4}$	
Γ <sub>67</sub>	$K^+ K^- K^- \pi^+$	$(2.1 \pm 0.5) \times 10^{-4}$	
Γ <sub>68</sub>	$K^+ K^- \bar{K}^0 \pi^0$	$(7.2 \pm_{-3.5}^{+4.8}) \times 10^{-3}$	

Fractions of many of the following modes with resonances have already appeared above as submodes of particular charged-particle modes. (Modes for which there are only upper limits and  $\bar{K}^*(892)\rho$  submodes only appear below.)

Γ <sub>69</sub>	$\bar{K}^0 \eta$	$(7.0 \pm 1.0) \times 10^{-3}$	
Γ <sub>70</sub>	$\bar{K}^0 \rho^0$	$(1.21 \pm 0.17) \%$	
Γ <sub>71</sub>	$K^- \rho^+$	$(10.8 \pm 0.9) \%$	S=1.2
Γ <sub>72</sub>	$\bar{K}^0 \omega$	$(2.1 \pm 0.4) \%$	
Γ <sub>73</sub>	$\bar{K}^0 \eta'(958)$	$(1.71 \pm 0.26) \%$	
Γ <sub>74</sub>	$\bar{K}^0 f_0(980)$	$(5.7 \pm 1.6) \times 10^{-3}$	
Γ <sub>75</sub>	$\bar{K}^0 \phi$	$(8.6 \pm 1.0) \times 10^{-3}$	
Γ <sub>76</sub>	$K^- a_1(1260)^+$	$(7.3 \pm 1.1) \%$	

$\Gamma_{77}$	$\bar{K}^0 a_1(1260)^0$	< 1.9	%	CL=90%
$\Gamma_{78}$	$\bar{K}^0 f_2(1270)$	( 4.1 $\pm$ 1.5 )	$\times 10^{-3}$	
$\Gamma_{79}$	$K^- a_2(1320)^+$	< 2	$\times 10^{-3}$	CL=90%
$\Gamma_{80}$	$\bar{K}^0 f_0(1370)$	( 6.9 $\pm$ 2.1 )	$\times 10^{-3}$	
$\Gamma_{81}$	$K^*(892)^- \pi^+$	( 5.0 $\pm$ 0.4 )	%	S=1.2
$\Gamma_{82}$	$\bar{K}^*(892)^0 \pi^0$	( 3.1 $\pm$ 0.4 )	%	
$\Gamma_{83}$	$\bar{K}^*(892)^0 \pi^+ \pi^-$ total	( 2.2 $\pm$ 0.5 )	%	
$\Gamma_{84}$	$\bar{K}^*(892)^0 \pi^+ \pi^-$ 3-body	( 1.42 $\pm$ 0.32 )	%	
$\Gamma_{85}$	$K^- \pi^+ \rho^0$ total	( 6.3 $\pm$ 0.4 )	%	
$\Gamma_{86}$	$K^- \pi^+ \rho^0$ 3-body	( 4.7 $\pm$ 2.1 )	$\times 10^{-3}$	
$\Gamma_{87}$	$\bar{K}^*(892)^0 \rho^0$	( 1.46 $\pm$ 0.32 )	%	
$\Gamma_{88}$	$\bar{K}^*(892)^0 \rho^0$ transverse	( 1.5 $\pm$ 0.5 )	%	
$\Gamma_{89}$	$\bar{K}^*(892)^0 \rho^0$ S-wave	( 2.8 $\pm$ 0.6 )	%	
$\Gamma_{90}$	$\bar{K}^*(892)^0 \rho^0$ S-wave long.	< 3	$\times 10^{-3}$	CL=90%
$\Gamma_{91}$	$\bar{K}^*(892)^0 \rho^0$ P-wave	< 3	$\times 10^{-3}$	CL=90%
$\Gamma_{92}$	$\bar{K}^*(892)^0 \rho^0$ D-wave	( 1.9 $\pm$ 0.6 )	%	
$\Gamma_{93}$	$K^*(892)^- \rho^+$	( 6.1 $\pm$ 2.4 )	%	
$\Gamma_{94}$	$K^*(892)^- \rho^+$ longitudinal	( 2.9 $\pm$ 1.2 )	%	
$\Gamma_{95}$	$K^*(892)^- \rho^+$ transverse	( 3.2 $\pm$ 1.8 )	%	
$\Gamma_{96}$	$K^*(892)^- \rho^+$ P-wave	< 1.5	%	CL=90%
$\Gamma_{97}$	$K^- \pi^+ f_0(980)$	< 1.1	%	CL=90%
$\Gamma_{98}$	$\bar{K}^*(892)^0 f_0(980)$	< 7	$\times 10^{-3}$	CL=90%
$\Gamma_{99}$	$K_1(1270)^- \pi^+$	[d] ( 1.06 $\pm$ 0.29 )	%	
$\Gamma_{100}$	$K_1(1400)^- \pi^+$	< 1.2	%	CL=90%
$\Gamma_{101}$	$\bar{K}_1(1400)^0 \pi^0$	< 3.7	%	CL=90%
$\Gamma_{102}$	$K^*(1410)^- \pi^+$	< 1.2	%	CL=90%
$\Gamma_{103}$	$K_0^*(1430)^- \pi^+$	( 1.04 $\pm$ 0.26 )	%	
$\Gamma_{104}$	$K_2^*(1430)^- \pi^+$	< 8	$\times 10^{-3}$	CL=90%
$\Gamma_{105}$	$\bar{K}_2^*(1430)^0 \pi^0$	< 4	$\times 10^{-3}$	CL=90%
$\Gamma_{106}$	$\bar{K}^*(892)^0 \pi^+ \pi^- \pi^0$	( 1.8 $\pm$ 0.9 )	%	
$\Gamma_{107}$	$\bar{K}^*(892)^0 \eta$	( 1.9 $\pm$ 0.5 )	%	
$\Gamma_{108}$	$K^- \pi^+ \omega$	( 3.0 $\pm$ 0.6 )	%	
$\Gamma_{109}$	$\bar{K}^*(892)^0 \omega$	( 1.1 $\pm$ 0.4 )	%	
$\Gamma_{110}$	$K^- \pi^+ \eta'(958)$	( 7.0 $\pm$ 1.8 )	$\times 10^{-3}$	
$\Gamma_{111}$	$\bar{K}^*(892)^0 \eta'(958)$	< 1.0	$\times 10^{-3}$	CL=90%

### Pionic modes

$\Gamma_{112}$	$\pi^+ \pi^-$	( 1.52 $\pm$ 0.09 )	$\times 10^{-3}$	
$\Gamma_{113}$	$\pi^0 \pi^0$	( 8.4 $\pm$ 2.2 )	$\times 10^{-4}$	
$\Gamma_{114}$	$\pi^+ \pi^- \pi^0$	( 1.1 $\pm$ 0.4 )	%	
$\Gamma_{115}$	$\pi^+ \pi^+ \pi^- \pi^-$	( 7.3 $\pm$ 0.5 )	$\times 10^{-3}$	
$\Gamma_{116}$	$\pi^+ \pi^+ \pi^- \pi^- \pi^0$	( 1.9 $\pm$ 0.4 )	%	
$\Gamma_{117}$	$\pi^+ \pi^+ \pi^+ \pi^- \pi^- \pi^-$	( 4.0 $\pm$ 3.0 )	$\times 10^{-4}$	

### Hadronic modes with a $K\bar{K}$ pair

$\Gamma_{118}$	$K^+ K^-$	$(4.25 \pm 0.16) \times 10^{-3}$	
$\Gamma_{119}$	$K^0 \bar{K}^0$	$(6.5 \pm 1.8) \times 10^{-4}$	S=1.2
$\Gamma_{120}$	$K^0 K^- \pi^+$	$(6.4 \pm 1.0) \times 10^{-3}$	S=1.1
$\Gamma_{121}$	$\bar{K}^*(892)^0 K^0$ $\times B(\bar{K}^{*0} \rightarrow K^- \pi^+)$	$< 1.1 \times 10^{-3}$	CL=90%
$\Gamma_{122}$	$K^*(892)^+ K^-$ $\times B(K^{*+} \rightarrow K^0 \pi^+)$	$(2.3 \pm 0.5) \times 10^{-3}$	
$\Gamma_{123}$	$K^0 K^- \pi^+$ nonresonant	$(2.3 \pm 2.3) \times 10^{-3}$	
$\Gamma_{124}$	$\bar{K}^0 K^+ \pi^-$	$(5.0 \pm 1.0) \times 10^{-3}$	
$\Gamma_{125}$	$K^*(892)^0 \bar{K}^0$ $\times B(K^{*0} \rightarrow K^+ \pi^-)$	$< 5 \times 10^{-4}$	CL=90%
$\Gamma_{126}$	$K^*(892)^- K^+$ $\times B(K^{*-} \rightarrow \bar{K}^0 \pi^-)$	$(1.2 \pm 0.7) \times 10^{-3}$	
$\Gamma_{127}$	$\bar{K}^0 K^+ \pi^-$ nonresonant	$(3.8 \pm_{-1.9}^{+2.3}) \times 10^{-3}$	
$\Gamma_{128}$	$K^+ K^- \pi^0$	$(1.3 \pm 0.4) \times 10^{-3}$	
$\Gamma_{129}$	$K_S^0 K_S^0 \pi^0$	$< 5.9 \times 10^{-4}$	
$\Gamma_{130}$	$K^+ K^- \pi^+ \pi^-$	[e] $(2.50 \pm 0.23) \times 10^{-3}$	
$\Gamma_{131}$	$\phi \pi^+ \pi^- \times B(\phi \rightarrow K^+ K^-)$	$(5.3 \pm 1.4) \times 10^{-4}$	
$\Gamma_{132}$	$\phi \rho^0 \times B(\phi \rightarrow K^+ K^-)$	$(3.0 \pm 1.6) \times 10^{-4}$	
$\Gamma_{133}$	$K^+ K^- \rho^0$ 3-body	$(9.0 \pm 2.3) \times 10^{-4}$	
$\Gamma_{134}$	$K^*(892)^0 K^- \pi^+ + c.c.$ $\times B(K^{*0} \rightarrow K^+ \pi^-)$	[f] $< 5 \times 10^{-4}$	
$\Gamma_{135}$	$K^*(892)^0 \bar{K}^*(892)^0$ $\times B^2(K^{*0} \rightarrow K^+ \pi^-)$	$(6 \pm 2) \times 10^{-4}$	
$\Gamma_{136}$	$K^+ K^- \pi^+ \pi^-$ non- $\phi$	—	
$\Gamma_{137}$	$K^+ K^- \pi^+ \pi^-$ nonresonant	$< 8 \times 10^{-4}$	CL=90%
$\Gamma_{138}$	$K^0 \bar{K}^0 \pi^+ \pi^-$	$(6.8 \pm 2.7) \times 10^{-3}$	
$\Gamma_{139}$	$K^+ K^- \pi^+ \pi^- \pi^0$	$(3.1 \pm 2.0) \times 10^{-3}$	

Fractions of most of the following modes with resonances have already appeared above as submodes of particular charged-particle modes.

$\Gamma_{140}$	$\bar{K}^*(892)^0 K^0$	$< 1.6 \times 10^{-3}$	CL=90%
$\Gamma_{141}$	$K^*(892)^+ K^-$	$(3.5 \pm 0.8) \times 10^{-3}$	
$\Gamma_{142}$	$K^*(892)^0 \bar{K}^0$	$< 8 \times 10^{-4}$	CL=90%
$\Gamma_{143}$	$K^*(892)^- K^+$	$(1.8 \pm 1.0) \times 10^{-3}$	
$\Gamma_{144}$	$\phi \pi^0$	$< 1.4 \times 10^{-3}$	CL=90%
$\Gamma_{145}$	$\phi \eta$	$< 2.8 \times 10^{-3}$	CL=90%
$\Gamma_{146}$	$\phi \omega$	$< 2.1 \times 10^{-3}$	CL=90%



Γ <sub>147</sub>	$\phi\pi^+\pi^-$		$(1.07 \pm 0.28) \times 10^{-3}$	
Γ <sub>148</sub>	$\phi\rho^0$		$(6 \pm 3) \times 10^{-4}$	
Γ <sub>149</sub>	$\phi\pi^+\pi^-$ 3-body		$(7 \pm 5) \times 10^{-4}$	
Γ <sub>150</sub>	$K^*(892)^0 K^- \pi^+ + \text{c.c.}$	[f] < 7	$\times 10^{-4}$	CL=90%
Γ <sub>151</sub>	$K^*(892)^0 K^- \pi^+$			
Γ <sub>152</sub>	$\bar{K}^*(892)^0 K^+ \pi^-$			
Γ <sub>153</sub>	$K^*(892)^0 \bar{K}^*(892)^0$		$(1.4 \pm 0.5) \times 10^{-3}$	

**Radiative modes**

Γ <sub>154</sub>	$\rho^0\gamma$	< 2.4	$\times 10^{-4}$	CL=90%
Γ <sub>155</sub>	$\omega\gamma$	< 2.4	$\times 10^{-4}$	CL=90%
Γ <sub>156</sub>	$\phi\gamma$	< 1.9	$\times 10^{-4}$	CL=90%
Γ <sub>157</sub>	$\bar{K}^*(892)^0\gamma$	< 7.6	$\times 10^{-4}$	CL=90%

**Doubly Cabibbo suppressed (DC) modes,  
 $\Delta C = 2$  forbidden via mixing (C2M) modes,  
 $\Delta C = 1$  weak neutral current (C1) modes, or  
 Lepton Family number (LF) violating modes**

Γ <sub>158</sub>	$K^+ \ell^- \bar{\nu}_\ell$ (via $\bar{D}^0$ )	C2M	< 1.7	$\times 10^{-4}$	CL=90%
Γ <sub>159</sub>	$K^+ \pi^-$	DC	$(1.46 \pm 0.30) \times 10^{-4}$		
Γ <sub>160</sub>	$K^+ \pi^-$ (via $\bar{D}^0$ )	C2M	< 1.6	$\times 10^{-5}$	CL=95%
Γ <sub>161</sub>	$K^+ \pi^- \pi^+ \pi^-$	DC	$(1.9 \pm 2.6) \times 10^{-4}$		
Γ <sub>162</sub>	$K^+ \pi^- \pi^+ \pi^-$ (via $\bar{D}^0$ )	C2M	< 4	$\times 10^{-4}$	CL=90%
Γ <sub>163</sub>	$K^+ \pi^-$ or $K^+ \pi^- \pi^+ \pi^-$ (via $\bar{D}^0$ )		< 1.0	$\times 10^{-3}$	CL=90%
Γ <sub>164</sub>	$\mu^-$ anything (via $\bar{D}^0$ )	C2M	< 4	$\times 10^{-4}$	CL=90%
Γ <sub>165</sub>	$e^+ e^-$	C1	< 6.2	$\times 10^{-6}$	CL=90%
Γ <sub>166</sub>	$\mu^+ \mu^-$	C1	< 4.1	$\times 10^{-6}$	CL=90%
Γ <sub>167</sub>	$\pi^0 e^+ e^-$	C1	< 4.5	$\times 10^{-5}$	CL=90%
Γ <sub>168</sub>	$\pi^0 \mu^+ \mu^-$	C1	< 1.8	$\times 10^{-4}$	CL=90%
Γ <sub>169</sub>	$\eta e^+ e^-$	C1	< 1.1	$\times 10^{-4}$	CL=90%
Γ <sub>170</sub>	$\eta \mu^+ \mu^-$	C1	< 5.3	$\times 10^{-4}$	CL=90%
Γ <sub>171</sub>	$\rho^0 e^+ e^-$	C1	< 1.0	$\times 10^{-4}$	CL=90%
Γ <sub>172</sub>	$\rho^0 \mu^+ \mu^-$	C1	< 2.3	$\times 10^{-4}$	CL=90%
Γ <sub>173</sub>	$\omega e^+ e^-$	C1	< 1.8	$\times 10^{-4}$	CL=90%
Γ <sub>174</sub>	$\omega \mu^+ \mu^-$	C1	< 8.3	$\times 10^{-4}$	CL=90%
Γ <sub>175</sub>	$\phi e^+ e^-$	C1	< 5.2	$\times 10^{-5}$	CL=90%
Γ <sub>176</sub>	$\phi \mu^+ \mu^-$	C1	< 4.1	$\times 10^{-4}$	CL=90%
Γ <sub>177</sub>	$\bar{K}^0 e^+ e^-$		[g] < 1.1	$\times 10^{-4}$	CL=90%
Γ <sub>178</sub>	$\bar{K}^0 \mu^+ \mu^-$		[g] < 2.6	$\times 10^{-4}$	CL=90%
Γ <sub>179</sub>	$\bar{K}^*(892)^0 e^+ e^-$		[g] < 1.4	$\times 10^{-4}$	CL=90%
Γ <sub>180</sub>	$\bar{K}^*(892)^0 \mu^+ \mu^-$		[g] < 1.18	$\times 10^{-3}$	CL=90%

$\Gamma_{181}$	$\pi^+ \pi^- \pi^0 \mu^+ \mu^-$	<i>CI</i>	$< 8.1$	$\times 10^{-4}$	CL=90%
$\Gamma_{182}$	$\mu^\pm e^\mp$	<i>LF</i>	$[h] < 8.1$	$\times 10^{-6}$	CL=90%
$\Gamma_{183}$	$\pi^0 e^\pm \mu^\mp$	<i>LF</i>	$[h] < 8.6$	$\times 10^{-5}$	CL=90%
$\Gamma_{184}$	$\eta e^\pm \mu^\mp$	<i>LF</i>	$[h] < 1.0$	$\times 10^{-4}$	CL=90%
$\Gamma_{185}$	$\rho^0 e^\pm \mu^\mp$	<i>LF</i>	$[h] < 4.9$	$\times 10^{-5}$	CL=90%
$\Gamma_{186}$	$\omega e^\pm \mu^\mp$	<i>LF</i>	$[h] < 1.2$	$\times 10^{-4}$	CL=90%
$\Gamma_{187}$	$\phi e^\pm \mu^\mp$	<i>LF</i>	$[h] < 3.4$	$\times 10^{-5}$	CL=90%
$\Gamma_{188}$	$\bar{K}^0 e^\pm \mu^\mp$	<i>LF</i>	$[h] < 1.0$	$\times 10^{-4}$	CL=90%
$\Gamma_{189}$	$\bar{K}^*(892)^0 e^\pm \mu^\mp$	<i>LF</i>	$[h] < 1.0$	$\times 10^{-4}$	CL=90%

$\Gamma_{190}$  A dummy mode used by the fit.  $(17.2 \pm 3.4) \%$  S=1.1

- [a] This is a weighted average of  $D^\pm$  (44%) and  $D^0$  (56%) branching fractions. See “ $D^+$  and  $D^0 \rightarrow (\eta \text{ anything}) / (\text{total } D^+ \text{ and } D^0)$ ” under “ $D^+$  Branching Ratios” in these Particle Listings.
- [b] This value averages the  $e^+$  and  $\mu^+$  branching fractions, after making a small phase-space adjustment to the  $\mu^+$  fraction to be able to use it as an  $e^+$  fraction; hence our  $\ell^+$  here is really an  $e^+$ .
- [c] The branching fraction for this mode may differ from the sum of the submodes that contribute to it, due to interference effects. See the relevant papers.
- [d] The two experiments measuring this fraction are in serious disagreement. See the Particle Listings.
- [e] The experiments on the division of this charge mode amongst its submodes disagree, and the submode branching fractions here add up to considerably more than the charged-mode fraction.
- [f] However, these upper limits are in serious disagreement with values obtained in another experiment.
- [g] This mode is not a useful test for a  $\Delta C=1$  weak neutral current because both quarks must change flavor in this decay.
- [h] The value is for the sum of the charge states or particle/antiparticle states indicated.

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### CONSTRAINED FIT INFORMATION

An overall fit to 51 branching ratios uses 122 measurements and one constraint to determine 28 parameters. The overall fit has a  $\chi^2 = 64.5$  for 95 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta x_i \delta x_j \rangle / (\delta x_i \delta x_j)$ , in percent, from the fit to the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_9$	6										
$x_{10}$	32	19									
$x_{18}$	1	24	5								
$x_{19}$	1	8	3	2							
$x_{20}$	13	46	42	11	6						
$x_{21}$	1	5	3	1	24	8					
$x_{22}$	1	6	4	2	36	10	66				
$x_{30}$	3	11	9	3	7	23	16	18			
$x_{38}$	5	18	17	4	3	40	4	5	9		
$x_{47}$	1	3	2	1	18	6	33	51	9	4	
$x_{56}$	3	9	8	2	1	19	2	2	4	28	
$x_{65}$	1	3	2	1	16	5	30	46	8	2	
$x_{69}$	1	3	2	1	17	5	58	47	11	2	
$x_{72}$	1	2	2	1	13	4	24	37	6	2	
$x_{75}$	1	4	3	1	21	6	39	60	10	3	
$x_{81}$	1	6	4	1	30	9	56	84	18	4	
$x_{82}$	1	5	4	1	7	10	24	18	43	4	
$x_{84}$	1	3	3	1	0	7	1	1	2	18	
$x_{88}$	1	2	2	0	2	4	3	5	2	9	
$x_{99}$	0	2	1	0	7	3	13	20	4	3	
$x_{107}$	1	3	3	1	2	6	4	4	23	3	
$x_{118}$	8	28	25	7	4	60	5	6	14	24	
$x_{119}$	0	2	1	0	9	3	17	25	4	1	
$x_{120}$	1	4	3	1	14	6	26	39	7	3	
$x_{124}$	1	3	2	1	11	6	20	30	6	2	
$x_{141}$	0	2	1	0	11	3	20	30	5	1	
$x_{190}$	-28	-21	-23	-7	-34	-32	-53	-70	-50	-26	
	$x_2$	$x_9$	$x_{10}$	$x_{18}$	$x_{19}$	$x_{20}$	$x_{21}$	$x_{22}$	$x_{30}$	$x_{38}$	

x56	1									
x65	23	1								
x69	24	1	21							
x72	43	1	17	17						
x75	30	1	7	28	22					
x81	43	2	38	40	31	50				
x82	9	2	8	14	7	11	17			
x84	1	5	0	0	0	0	1	1		
x88	9	3	2	2	4	3	4	1	2	
x99	40	1	9	9	17	12	17	4	1	4
x107	2	1	2	2	2	2	4	10	0	0
x118	3	12	3	3	2	4	6	6	4	2
x119	13	1	11	12	9	15	21	5	0	1
x120	20	1	18	18	14	23	33	7	0	2
x124	15	1	13	14	11	18	25	6	0	2
x141	15	1	14	14	11	18	25	6	0	1
x190	-68	-20	-33	-38	-45	-43	-64	-39	-14	-23
	x47	x56	x65	x69	x72	x75	x81	x82	x84	x88
x107	1									
x118	2	4								
x119	5	1	2							
x120	8	2	4	10						
x124	6	1	3	7	12					
x141	6	1	2	8	12	9				
x190	-34	-25	-20	-18	-30	-24	-23			
	x99	x107	x118	x119	x120	x124	x141			

## $D^0$ BRANCHING RATIOS

See the "Note on  $D$  Mesons" in the  $D^\pm$  Listings.

Some older now obsolete results have been omitted from these Listings.

### Inclusive modes

$\Gamma(e^+ \text{ anything})/\Gamma_{\text{total}}$					$\Gamma_1/\Gamma$
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT	
<b>0.0675 ± 0.0029 OUR AVERAGE</b>					
0.069 ± 0.003 ± 0.005	1670	ALBRECHT	96C ARG	$e^+e^- \approx 10 \text{ GeV}$	
0.0664 ± 0.0018 ± 0.0029	4609	<sup>17</sup> KUBOTA	96B CLE2	$e^+e^- \approx \Upsilon(4S)$	
0.075 ± 0.011 ± 0.004	137	BALTRUSAIT..	.85B MRK3	$e^+e^- 3.77 \text{ GeV}$	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					

0.15 ± 0.05		AGUILAR-...	87E HYBR	$\pi p, pp$	360, 400 GeV
0.055 ± 0.037	12	SCHINDLER	81 MRK2	$e^+ e^-$	3.771 GeV
<sup>17</sup> KUBOTA 96B uses $D^{*+} \rightarrow D^0 \pi^+$ (and charge conjugate) events in which the $D^0$ subsequently decays to $X e^+ \nu_e$ .					

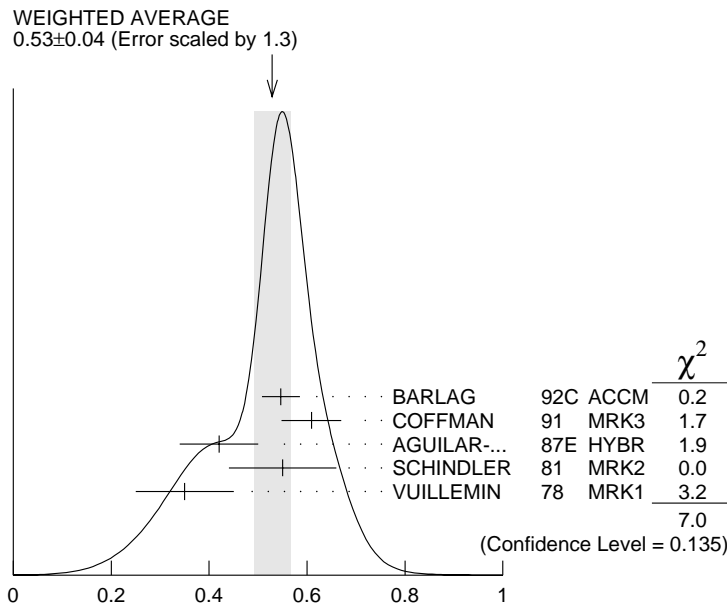
**$\Gamma(\mu^+ \text{ anything})/\Gamma_{\text{total}}$**   **$\Gamma_2/\Gamma$**

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.066 ± 0.008 OUR FIT</b>				
<b>0.060 ± 0.007 ± 0.012</b>	310	ALBRECHT	96C ARG	$e^+ e^- \approx 10$ GeV

**$\Gamma(K^- \text{ anything})/\Gamma_{\text{total}}$**   **$\Gamma_3/\Gamma$**

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.53 ± 0.04 OUR AVERAGE</b>		Error includes scale factor of 1.3. See the ideogram below.		
0.546 <sup>+0.039</sup> <sub>-0.038</sub>		<sup>18</sup> BARLAG	92C ACCM	$\pi^-$ Cu 230 GeV
0.609 ± 0.032 ± 0.052		COFFMAN	91 MRK3	$e^+ e^-$ 3.77 GeV
0.42 ± 0.08		AGUILAR-...	87E HYBR	$\pi p, pp$ 360, 400 GeV
0.55 ± 0.11	121	SCHINDLER	81 MRK2	$e^+ e^-$ 3.771 GeV
0.35 ± 0.10	19	VUILLEMIN	78 MRK1	$e^+ e^-$ 3.772 GeV

<sup>18</sup> BARLAG 92C computes the branching fraction using topological normalization.



$\Gamma(K^- \text{ anything})/\Gamma_{\text{total}}$

$[\Gamma(\bar{K}^0 \text{ anything}) + \Gamma(K^0 \text{ anything})]/\Gamma_{\text{total}} \quad \Gamma_4/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.42 ± 0.05</b>				<b>OUR AVERAGE</b>
0.455 ± 0.050 ± 0.032		COFFMAN	91	MRK3 e <sup>+</sup> e <sup>-</sup> 3.77 GeV
0.29 ± 0.11	13	SCHINDLER	81	MRK2 e <sup>+</sup> e <sup>-</sup> 3.771 GeV
0.57 ± 0.26	6	VUILLEMIN	78	MRK1 e <sup>+</sup> e <sup>-</sup> 3.772 GeV

$\Gamma(K^+ \text{ anything})/\Gamma_{\text{total}} \quad \Gamma_5/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.034<sup>+0.006</sup><sub>-0.004</sub></b>				<b>OUR AVERAGE</b>
0.034 <sup>+0.007</sup> <sub>-0.005</sub>		<sup>19</sup> BARLAG	92C	ACCM π <sup>-</sup> Cu 230 GeV
0.028 ± 0.009 ± 0.004		COFFMAN	91	MRK3 e <sup>+</sup> e <sup>-</sup> 3.77 GeV
0.03 <sup>+0.05</sup> <sub>-0.02</sub>		AGUILAR-...	87E	HYBR πp, pp 360, 400 GeV
0.08 ± 0.03	25	SCHINDLER	81	MRK2 e <sup>+</sup> e <sup>-</sup> 3.771 GeV

<sup>19</sup> BARLAG 92C computes the branching fraction using topological normalization.

$\Gamma(\phi \text{ anything})/\Gamma_{\text{total}} \quad \Gamma_7/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0171<sup>+0.0076</sup><sub>-0.0071</sub> ± 0.0017</b>	9	<sup>20</sup> BAI	00C	BES e <sup>+</sup> e <sup>-</sup> → D <sup>+</sup> D <sup>-*</sup> , D <sup>*+</sup> D <sup>-*</sup>

<sup>20</sup> BAI 00C finds the average (φ anything) branching fraction for the 4.03-GeV mix of D<sup>+</sup> and D<sup>0</sup> mesons to be (1.34 ± 0.52 ± 0.12)%.

————— Semileptonic modes —————

$\Gamma(K^- \ell^+ \nu_\ell)/\Gamma_{\text{total}} \quad \Gamma_8/\Gamma$

We average our K<sup>-</sup>e<sup>+</sup>ν<sub>e</sub> and K<sup>-</sup>μ<sup>+</sup>ν<sub>μ</sub> branching fractions, after multiplying the latter by a phase-space factor of 1.03 to be able to use it with the K<sup>-</sup>e<sup>+</sup>ν<sub>e</sub> fraction. Hence our ℓ<sup>+</sup> here is really an e<sup>+</sup>.

VALUE	DOCUMENT ID	COMMENT
<b>0.0348 ± 0.0017</b>		<b>OUR AVERAGE</b>
0.0364 ± 0.0018	PDG	00 Our Γ(K <sup>-</sup> e <sup>+</sup> ν <sub>e</sub> )/Γ <sub>total</sub>
0.0331 ± 0.0018	PDG	00 1.03 × our Γ(K <sup>-</sup> μ <sup>+</sup> ν <sub>μ</sub> )/Γ <sub>total</sub>

$\Gamma(K^- e^+ \nu_e)/\Gamma_{\text{total}} \quad \Gamma_9/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0364 ± 0.0018</b>				<b>OUR FIT</b>
<b>0.034 ± 0.005 ± 0.004</b>	55	ADLER	89	MRK3 e <sup>+</sup> e <sup>-</sup> 3.77 GeV

$\Gamma(K^- e^+ \nu_e)/\Gamma(K^- \pi^+)$   $\Gamma_9/\Gamma_{20}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.95 ± 0.04</b>				<b>OUR FIT</b>
<b>0.95 ± 0.04</b>				<b>OUR AVERAGE</b>
0.978 ± 0.027 ± 0.044	2510	<sup>21</sup> BEAN	93C	CLE2 e <sup>+</sup> e <sup>-</sup> ≈ γ(4S)
0.90 ± 0.06 ± 0.06	584	<sup>22</sup> CRAWFORD	91B	CLEO e <sup>+</sup> e <sup>-</sup> ≈ 10.5 GeV
0.91 ± 0.07 ± 0.11	250	<sup>23</sup> ANJOS	89F	E691 Photoproduction

- <sup>21</sup> BEAN 93C uses  $K^- \mu^+ \nu_\mu$  as well as  $K^- e^+ \nu_e$  events and makes a small phase-space adjustment to the number of the  $\mu^+$  events to use them as  $e^+$  events. A pole mass of  $2.00 \pm 0.12 \pm 0.18 \text{ GeV}/c^2$  is obtained from the  $q^2$  dependence of the decay rate.
- <sup>22</sup> CRAWFORD 91B uses  $K^- e^+ \nu_e$  and  $K^- \mu^+ \nu_\mu$  candidates to measure a pole mass of  $2.1^{+0.4+0.3}_{-0.2-0.2} \text{ GeV}/c^2$  from the  $q^2$  dependence of the decay rate.
- <sup>23</sup> ANJOS 89F measures a pole mass of  $2.1^{+0.4}_{-0.2} \pm 0.2 \text{ GeV}/c^2$  from the  $q^2$  dependence of the decay rate.

$\Gamma(K^- \mu^+ \nu_\mu)/\Gamma(K^- \pi^+)$   $\Gamma_{10}/\Gamma_{20}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.84 ± 0.04 OUR FIT</b>				
<b>0.84 ± 0.04 OUR AVERAGE</b>				
0.852 ± 0.034 ± 0.028	1897	<sup>24</sup> FRABETTI	95G E687	$\gamma \text{Be } \bar{E}_\gamma = 220 \text{ GeV}$
0.82 ± 0.13 ± 0.13	338	<sup>25</sup> FRABETTI	93I E687	$\gamma \text{Be } \bar{E}_\gamma = 221 \text{ GeV}$
0.79 ± 0.08 ± 0.09	231	<sup>26</sup> CRAWFORD	91B CLEO	$e^+ e^- \approx 10.5 \text{ GeV}$

- <sup>24</sup> FRABETTI 95G extracts the ratio of form factors  $f_-(0)/f_+(0) = -1.3^{+3.6}_{-3.4} \pm 0.6$ , and measures a pole mass of  $1.87^{+0.11+0.07}_{-0.08-0.06} \text{ GeV}/c^2$  from the  $q^2$  dependence of the decay rate.
- <sup>25</sup> FRABETTI 93I measures a pole mass of  $2.1^{+0.7+0.7}_{-0.3-0.3} \text{ GeV}/c^2$  from the  $q^2$  dependence of the decay rate.
- <sup>26</sup> CRAWFORD 91B measures a pole mass of  $2.00 \pm 0.12 \pm 0.18 \text{ GeV}/c^2$  from the  $q^2$  dependence of the decay rate.

$\Gamma(K^- \mu^+ \nu_\mu)/\Gamma(\mu^+ \text{ anything})$   $\Gamma_{10}/\Gamma_2$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.49 ± 0.06 OUR FIT</b>				
<b>0.472 ± 0.051 ± 0.040</b>	232	KODAMA	94 E653	$\pi^-$ emulsion 600 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.32 ± 0.05 ± 0.05	124	KODAMA	91 EMUL	$pA$ 800 GeV

$\Gamma(K^- \pi^0 e^+ \nu_e)/\Gamma_{\text{total}}$   $\Gamma_{11}/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.016^{+0.013}_{-0.005} ± 0.002</b>	4	<sup>27</sup> BAI	91 MRK3	$e^+ e^- \approx 3.77 \text{ GeV}$

- <sup>27</sup> BAI 91 finds that a fraction  $0.79^{+0.15+0.09}_{-0.17-0.03}$  of combined  $D^+$  and  $D^0$  decays to  $\bar{K} \pi e^+ \nu_e$  (24 events) are  $\bar{K}^*(892) e^+ \nu_e$ . BAI 91 uses 56  $K^- e^+ \nu_e$  events to measure a pole mass of  $1.8 \pm 0.3 \pm 0.2 \text{ GeV}/c^2$  from the  $q^2$  dependence of the decay rate.

$\Gamma(\bar{K}^0 \pi^- e^+ \nu_e)/\Gamma_{\text{total}}$   $\Gamma_{12}/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.028^{+0.017}_{-0.008} ± 0.003</b>	6	<sup>28</sup> BAI	91 MRK3	$e^+ e^- \approx 3.77 \text{ GeV}$

- <sup>28</sup> BAI 91 finds that a fraction  $0.79^{+0.15+0.09}_{-0.17-0.03}$  of combined  $D^+$  and  $D^0$  decays to  $\bar{K} \pi e^+ \nu_e$  (24 events) are  $\bar{K}^*(892) e^+ \nu_e$ .

$\Gamma(K^*(892)^- e^+ \nu_e) / \Gamma(K^- e^+ \nu_e)$   $\Gamma_{19} / \Gamma_9$

Unseen decay modes of the  $K^*(892)^-$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.55 ± 0.09 OUR FIT</b>			
<b>0.51 ± 0.18 ± 0.06</b>	CRAWFORD 91B	CLEO	$e^+ e^- \approx 10.5$ GeV

$\Gamma(K^*(892)^- e^+ \nu_e) / \Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{19} / \Gamma_{22}$

Unseen decay modes of the  $\bar{K}^*(892)^-$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.37 ± 0.06 OUR FIT</b>				
<b>0.38 ± 0.06 ± 0.03</b>	152	<sup>29</sup> BEAN	93C	CLE2 $e^+ e^- \approx \Upsilon(4S)$

<sup>29</sup> BEAN 93C uses  $K^{*-} \mu^+ \nu_\mu$  as well as  $K^{*-} e^+ \nu_e$  events and makes a small phase-space adjustment to the number of the  $\mu^+$  events to use them as  $e^+$  events.

$\Gamma(K^*(892)^- \ell^+ \nu_\ell) / \Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{14} / \Gamma_{22}$

This an average of the  $K^*(892)^- e^+ \nu_e$  and  $K^*(892)^- \mu^+ \nu_\mu$  ratios. Unseen decay modes of the  $K^*(892)^-$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.24 ± 0.07 ± 0.06	137	<sup>30</sup> ALEXANDER 90B	CLEO	$e^+ e^-$ 10.5–11 GeV

<sup>30</sup> ALEXANDER 90B cannot exclude extra  $\pi^0$ 's in the final state. See nearby data blocks for more detailed results.

$\Gamma(\bar{K}^*(892)^0 \pi^- e^+ \nu_e) / \Gamma(K^*(892)^- e^+ \nu_e)$   $\Gamma_{15} / \Gamma_{19}$

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.64	90	<sup>31</sup> CRAWFORD 91B	CLEO	$e^+ e^- \approx 10.5$ GeV

<sup>31</sup> The limit on  $(\bar{K}^*(892)\pi)^- \mu^+ \nu_\mu$  below is much stronger.

$\Gamma(K^- \pi^+ \pi^- \mu^+ \nu_\mu) / \Gamma(K^- \mu^+ \nu_\mu)$   $\Gamma_{16} / \Gamma_{10}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.037</b>	90	KODAMA 93B	E653	$\pi^-$ emulsion 600 GeV

$\Gamma((\bar{K}^*(892)\pi)^- \mu^+ \nu_\mu) / \Gamma(K^- \mu^+ \nu_\mu)$   $\Gamma_{17} / \Gamma_{10}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.043</b>	90	<sup>32</sup> KODAMA 93B	E653	$\pi^-$ emulsion 600 GeV

<sup>32</sup> KODAMA 93B searched in  $K^- \pi^+ \pi^- \mu^+ \nu_\mu$ , but the limit includes other  $(\bar{K}^*(892)\pi)^-$  charge states.

$\Gamma(\pi^- e^+ \nu_e) / \Gamma_{\text{total}}$   $\Gamma_{18} / \Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0037 ± 0.0006 OUR FIT</b>				
<b>0.0039<sup>+0.0023</sup><sub>-0.0011</sub> ± 0.0004</b>	7	<sup>33</sup> ADLER 89	MRK3	$e^+ e^-$ 3.77 GeV

<sup>33</sup> This result of ADLER 89 gives  $|\frac{V_{cd}}{V_{cs}} \cdot \frac{f_+^{\pi}(0)}{f_+^K(0)}|^2 = 0.057^{+0.038}_{-0.015} \pm 0.005$ .



$$\Gamma(\pi^- e^+ \nu_e) / \Gamma(K^- e^+ \nu_e)$$

$\Gamma_{18}/\Gamma_9$

VALUE                      EVTS                      DOCUMENT ID                      TECN                      COMMENT

**0.102 ± 0.017 OUR FIT**

**0.101 ± 0.018 OUR AVERAGE**

0.101 ± 0.020 ± 0.003                      91                      <sup>34</sup> FRABETTI                      96B E687                       $\gamma$  Be,  $\bar{E}_\gamma \approx 200$  GeV

0.103 ± 0.039 ± 0.013                      87                      <sup>35</sup> BUTLER                      95 CLE2                      < 0.156 (90% CL)

<sup>34</sup> FRABETTI 96B uses both  $e$  and  $\mu$  events, and makes a small correction to the  $\mu$  events to make them effectively  $e$  events. This result gives  $|\frac{V_{cd}}{V_{cs}} \cdot \frac{f_\pi^+(0)}{f_K^+(0)}|^2 = 0.050 \pm 0.011 \pm 0.002$ .

<sup>35</sup> BUTLER 95 has  $87 \pm 33 \pi^- e^+ \nu_e$  events. The result gives  $|\frac{V_{cd}}{V_{cs}} \cdot \frac{f_\pi^+(0)}{f_K^+(0)}|^2 = 0.052 \pm 0.020 \pm 0.007$ .

————— **Hadronic modes with a  $\bar{K}$  or  $\bar{K}K\bar{K}$**  —————

$$\Gamma(K^- \pi^+) / \Gamma_{\text{total}}$$

$\Gamma_{20}/\Gamma$

We list measurements *before* radiative corrections are made.

VALUE                      EVTS                      DOCUMENT ID                      TECN                      COMMENT

**0.0383 ± 0.0009 OUR FIT**

**0.0385 ± 0.0009 OUR AVERAGE**

0.0382 ± 0.0007 ± 0.0012                      <sup>36</sup> ARTUSO                      98 CLE2                      CLEO average

0.0390 ± 0.0009 ± 0.0012                      5392                      <sup>37</sup> BARATE                      97C ALEP                      From  $Z$  decays

0.045 ± 0.006 ± 0.004                      <sup>38</sup> ALBRECHT                      94 ARG                       $e^+ e^- \approx \Upsilon(4S)$

0.0341 ± 0.0012 ± 0.0028                      1173                      <sup>37</sup> ALBRECHT                      94F ARG                       $e^+ e^- \approx \Upsilon(4S)$

0.0362 ± 0.0034 ± 0.0044                      <sup>37</sup> DECAMP                      91J ALEP                      From  $Z$  decays

0.045 ± 0.008 ± 0.005                      56                      <sup>37</sup> ABACHI                      88 HRS                       $e^+ e^-$  29 GeV

0.042 ± 0.004 ± 0.004                      930                      ADLER                      88C MRK3                       $e^+ e^-$  3.77 GeV

0.041 ± 0.006                      263                      <sup>39</sup> SCHINDLER                      81 MRK2                       $e^+ e^-$  3.771 GeV

0.043 ± 0.010                      130                      <sup>40</sup> PERUZZI                      77 MRK1                       $e^+ e^-$  3.77 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.0381 ± 0.0015 ± 0.0016                      1165                      <sup>41</sup> ARTUSO                      98 CLE2                       $e^+ e^-$  at  $\Upsilon(4S)$

0.0369 ± 0.0011 ± 0.0016                      <sup>42</sup> COAN                      98 CLE2

0.0391 ± 0.0008 ± 0.0017                      4208 <sup>37,43</sup> AKERIB                      93 CLE2                       $e^+ e^- \approx \Upsilon(4S)$

<sup>36</sup> This combines the CLEO results of ARTUSO 98, COAN 98, and AKERIB 93.

<sup>37</sup> ABACHI 88, DECAMP 91J, AKERIB 93, ALBRECHT 94F, and BARATE 97C use  $D^*(2010)^+ \rightarrow D^0 \pi^+$  decays. The  $\pi^+$  is both slow and of low  $p_T$  with respect to the event thrust axis or nearest jet ( $\approx D^{*+}$  direction). The excess number of such  $\pi^+$ 's over background gives the number of  $D^*(2010)^+ \rightarrow D^0 \pi^+$  events, and the fraction with  $D^0 \rightarrow K^- \pi^+$  gives the  $D^0 \rightarrow K^- \pi^+$  branching fraction.

<sup>38</sup> ALBRECHT 94 uses  $D^0$  mesons from  $\bar{B}^0 \rightarrow D^{*+} \ell^- \bar{\nu}_\ell$  decays. This is a different set of events than used by ALBRECHT 94F.

<sup>39</sup> SCHINDLER 81 (MARK-2) measures  $\sigma(e^+ e^- \rightarrow \psi(3770)) \times$  branching fraction to be  $0.24 \pm 0.02$  nb. We use the MARK-3 (ADLER 88C) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6$  nb.

<sup>40</sup> PERUZZI 77 (MARK-1) measures  $\sigma(e^+ e^- \rightarrow \psi(3770)) \times$  branching fraction to be  $0.25 \pm 0.05$  nb. We use the MARK-3 (ADLER 88C) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6$  nb.

<sup>41</sup> ARTUSO 98, following ALBRECHT 94, uses  $D^0$  mesons from  $\bar{B}^0 \rightarrow D^*(2010)^+ X \ell^- \bar{\nu}_\ell$  decays. Our average uses the CLEO average of this value with the values of COAN 98 and AKERIB 93.

<sup>42</sup> COAN 98 assumes that  $\Gamma(B \rightarrow \bar{D}X\ell^+\nu)/\Gamma(B \rightarrow X\ell^+\nu) = 1.0 - 3|V_{ub}/V_{cb}|^2 - 0.010 \pm 0.005$ , the last term accounting for  $\bar{B} \rightarrow D_s^+ K X \ell^- \bar{\nu}$ . COAN 98 is included in the CLEO average in ARTUSO 98.

<sup>43</sup> This AKERIB 93 value does not include radiative corrections; with them, the value is  $0.0395 \pm 0.0008 \pm 0.0017$ . AKERIB 93 is included in the CLEO average in ARTUSO 98.

$\Gamma(\bar{K}^0\pi^0)/\Gamma(K^-\pi^+)$   $\Gamma_{21}/\Gamma_{20}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.55±0.06 OUR FIT</b>	Error includes scale factor of 1.1.			
<b>1.36±0.23±0.22</b>	119	ANJOS	92B E691	$\gamma$ Be 80–240 GeV

$\Gamma(\bar{K}^0\pi^0)/\Gamma(\bar{K}^0\pi^+\pi^-)$   $\Gamma_{21}/\Gamma_{22}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.390±0.031 OUR FIT</b>				
<b>0.378±0.033 OUR AVERAGE</b>				
0.44 ±0.02 ±0.05	1942	PROCARIO	93B CLE2	$e^+e^-$ 10.36–10.7 GeV
0.34 ±0.04 ±0.02	92	<sup>44</sup> ALBRECHT	92P ARG	$e^+e^- \approx 10$ GeV
0.36 ±0.04 ±0.08	104	KINOSHITA	91 CLEO	$e^+e^- \sim 10.7$ GeV

<sup>44</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.

$\Gamma(\bar{K}^0\pi^+\pi^-)/\Gamma_{total}$   $\Gamma_{22}/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.054 ±0.004 OUR FIT</b>	Error includes scale factor of 1.2.			
<b>0.055 ±0.005 OUR AVERAGE</b>				
0.0503±0.0039±0.0049	284	<sup>45</sup> ALBRECHT	94F ARG	$e^+e^- \approx \Upsilon(4S)$
0.064 ±0.005 ±0.010		ADLER	87 MRK3	$e^+e^-$ 3.77 GeV
0.052 ±0.016	32	<sup>46</sup> SCHINDLER	81 MRK2	$e^+e^-$ 3.771 GeV
0.079 ±0.023	28	<sup>47</sup> PERUZZI	77 MRK1	$e^+e^-$ 3.77 GeV

<sup>45</sup> See the footnote on the ALBRECHT 94F measurement of  $\Gamma(K^-\pi^+)/\Gamma_{total}$  for the method used.

<sup>46</sup> SCHINDLER 81 (MARK-2) measures  $\sigma(e^+e^- \rightarrow \psi(3770)) \times$  branching fraction to be  $0.30 \pm 0.08$  nb. We use the MARK-3 (ADLER 88C) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6$  nb.

<sup>47</sup> PERUZZI 77 (MARK-1) measures  $\sigma(e^+e^- \rightarrow \psi(3770)) \times$  branching fraction to be  $0.46 \pm 0.12$  nb. We use the MARK-3 (ADLER 88C) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6$  nb.

$\Gamma(\bar{K}^0\pi^+\pi^-)/\Gamma(K^-\pi^+)$   $\Gamma_{22}/\Gamma_{20}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.42±0.10 OUR FIT</b>	Error includes scale factor of 1.2.			
<b>1.65±0.17 OUR AVERAGE</b>				
1.61±0.10±0.15	856	FRABETTI	94J E687	$\gamma$ Be $\bar{E}_\gamma=220$ GeV
1.7 ±0.8	35	AVERY	80 SPEC	$\gamma N \rightarrow D^{*+}$
2.8 ±1.0	116	PICCOLO	77 MRK1	$e^+e^-$ 4.03, 4.41 GeV

$\Gamma(\bar{K}^0\rho^0)/\Gamma(\bar{K}^0\pi^+\pi^-)$   $\Gamma_{23}/\Gamma_{22}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.223±0.027 OUR AVERAGE</b>	Error includes scale factor of 1.2.		
0.350±0.028±0.067	FRABETTI	94G E687	$\gamma$ Be, $\bar{E}_\gamma \approx 220$ GeV
0.227±0.032±0.009	ALBRECHT	93D ARG	$e^+e^- \approx 10$ GeV
0.215±0.051±0.037	ANJOS	93 E691	$\gamma$ Be 90–260 GeV
0.20 ±0.06 ±0.03	FRABETTI	92B E687	$\gamma$ Be $\bar{E}_\gamma=221$ GeV
0.12 ±0.01 ±0.07	ADLER	87 MRK3	$e^+e^-$ 3.77 GeV

$\Gamma(\overline{K}^0 f_0(980))/\Gamma(\overline{K}^0 \pi^+ \pi^-)$   $\Gamma_{74}/\Gamma_{22}$

Unseen decay modes of the  $f_0(980)$  are included.

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.105 ± 0.029 OUR AVERAGE</b>			
0.131 ± 0.031 ± 0.034	FRABETTI	94G E687	$\gamma$ Be, $\overline{E}_\gamma \approx 220$ GeV
0.088 ± 0.035 ± 0.012	ALBRECHT	93D ARG	$e^+ e^- \approx 10$ GeV

$\Gamma(\overline{K}^0 f_2(1270))/\Gamma(\overline{K}^0 \pi^+ \pi^-)$   $\Gamma_{78}/\Gamma_{22}$

Unseen decay modes of the  $f_2(1270)$  are included.

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.076 ± 0.028 OUR AVERAGE</b>			
0.065 ± 0.025 ± 0.030	FRABETTI	94G E687	$\gamma$ Be, $\overline{E}_\gamma \approx 220$ GeV
0.088 ± 0.037 ± 0.014	ALBRECHT	93D ARG	$e^+ e^- \approx 10$ GeV

$\Gamma(\overline{K}^0 f_0(1370))/\Gamma(\overline{K}^0 \pi^+ \pi^-)$   $\Gamma_{80}/\Gamma_{22}$

Unseen decay modes of the  $f_0(1370)$  are included.

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.13 ± 0.04 OUR AVERAGE</b>			
0.123 ± 0.035 ± 0.049	FRABETTI	94G E687	$\gamma$ Be, $\overline{E}_\gamma \approx 220$ GeV
0.131 ± 0.045 ± 0.021	ALBRECHT	93D ARG	$e^+ e^- \approx 10$ GeV

$\Gamma(K^*(892)^- \pi^+)/\Gamma(\overline{K}^0 \pi^+ \pi^-)$   $\Gamma_{81}/\Gamma_{22}$

Unseen decay modes of the  $K^*(892)^-$  are included.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.93 ± 0.04 OUR FIT</b>				Error includes scale factor of 1.1.
<b>0.96 ± 0.04 OUR AVERAGE</b>				
0.938 ± 0.054 ± 0.038		FRABETTI	94G E687	$\gamma$ Be, $\overline{E}_\gamma \approx 220$ GeV
1.08 ± 0.063 ± 0.045		ALBRECHT	93D ARG	$e^+ e^- \approx 10$ GeV
0.720 ± 0.145 ± 0.185		ANJOS	93 E691	$\gamma$ Be 90–260 GeV
0.96 ± 0.12 ± 0.075		FRABETTI	92B E687	$\gamma$ Be $\overline{E}_\gamma = 221$ GeV
0.84 ± 0.06 ± 0.08		ADLER	87 MRK3	$e^+ e^- 3.77$ GeV
1.05 $\begin{smallmatrix} +0.23 \\ -0.26 \end{smallmatrix}$ $\begin{smallmatrix} +0.07 \\ -0.09 \end{smallmatrix}$	25	SCHINDLER	81 MRK2	$e^+ e^- 3.771$ GeV

$\Gamma(K_0^*(1430)^- \pi^+)/\Gamma(\overline{K}^0 \pi^+ \pi^-)$   $\Gamma_{103}/\Gamma_{22}$

Unseen decay modes of the  $\overline{K}_0^*(1430)^-$  are included.

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.19 ± 0.05 OUR AVERAGE</b>			
0.176 ± 0.044 ± 0.047	FRABETTI	94G E687	$\gamma$ Be, $\overline{E}_\gamma \approx 220$ GeV
0.208 ± 0.055 ± 0.034	ALBRECHT	93D ARG	$e^+ e^- \approx 10$ GeV

$\Gamma(K_2^*(1430)^- \pi^+)/\Gamma(\overline{K}^0 \pi^+ \pi^-)$   $\Gamma_{104}/\Gamma_{22}$

Unseen decay modes of the  $\overline{K}_2^*(1430)^-$  are included.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt;0.15</b>	90	ALBRECHT	93D ARG	$e^+ e^- \approx 10$ GeV

$\Gamma(\bar{K}^0 \pi^+ \pi^- \text{ nonresonant}) / \Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{29} / \Gamma_{22}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.27 ± 0.04 OUR AVERAGE</b>			
0.263 ± 0.024 ± 0.041	ANJOS	93 E691	$\gamma$ Be 90–260 GeV
0.26 ± 0.08 ± 0.05	FRABETTI	92B E687	$\gamma$ Be $\bar{E}_\gamma = 221$ GeV
0.33 ± 0.05 ± 0.10	ADLER	87 MRK3	$e^+ e^-$ 3.77 GeV

$\Gamma(K^- \pi^+ \pi^0) / \Gamma_{\text{total}}$   $\Gamma_{30} / \Gamma$

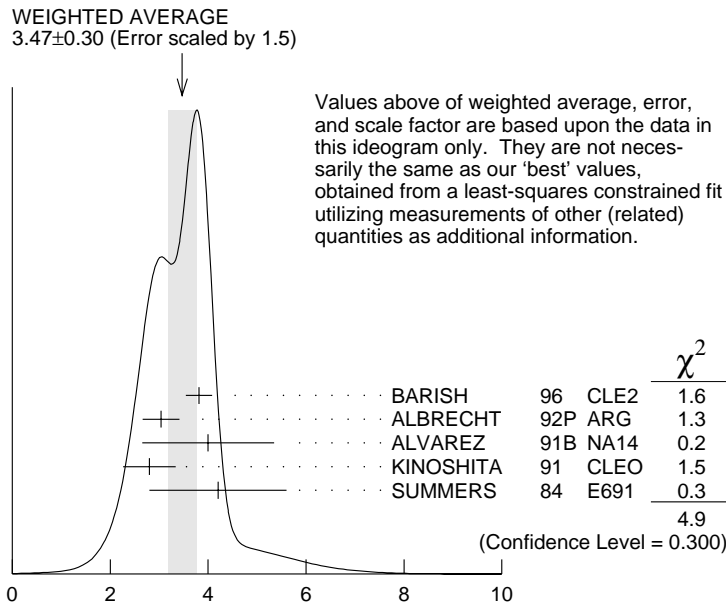
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.139 ± 0.009 OUR FIT</b>				Error includes scale factor of 1.3.
<b>0.131 ± 0.016 OUR AVERAGE</b>				
0.133 ± 0.012 ± 0.013	931	ADLER	88C MRK3	$e^+ e^-$ 3.77 GeV
0.117 ± 0.043	37	<sup>48</sup> SCHINDLER	81 MRK2	$e^+ e^-$ 3.771 GeV

<sup>48</sup>SCHINDLER 81 (MARK-2) measures  $\sigma(e^+ e^- \rightarrow \psi(3770)) \times$  branching fraction to be  $0.68 \pm 0.23$  nb. We use the MARK-3 (ADLER 88C) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6$  nb.

$\Gamma(K^- \pi^+ \pi^0) / \Gamma(K^- \pi^+)$   $\Gamma_{30} / \Gamma_{20}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>3.63 ± 0.23 OUR FIT</b>				Error includes scale factor of 1.4.
<b>3.47 ± 0.30 OUR AVERAGE</b>				Error includes scale factor of 1.5. See the ideogram below.
3.81 ± 0.07 ± 0.26	10k	BARISH	96 CLE2	$e^+ e^- \approx \Upsilon(4S)$
3.04 ± 0.16 ± 0.34	931	<sup>49</sup> ALBRECHT	92P ARG	$e^+ e^- \approx 10$ GeV
4.0 ± 0.9 ± 1.0	69	ALVAREZ	91B NA14	Photoproduction
2.8 ± 0.14 ± 0.52	1050	KINOSHITA	91 CLEO	$e^+ e^- \sim 10.7$ GeV
4.2 ± 1.4	41	SUMMERS	84 E691	Photoproduction

<sup>49</sup>This value is calculated from numbers in Table 1 of ALBRECHT 92P.



$$\Gamma(K^- \pi^+ \pi^0) / \Gamma(K^- \pi^+)$$

$\Gamma(K^- \rho^+)/\Gamma(K^- \pi^+ \pi^0)$   $\Gamma_{31}/\Gamma_{30}$

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.78 ± 0.05 OUR AVERAGE</b>				
0.765 ± 0.041 ± 0.054		FRABETTI	94G E687	$\gamma$ Be, $\bar{E}_\gamma \approx 220$ GeV
0.647 ± 0.039 ± 0.150		ANJOS	93 E691	$\gamma$ Be 90–260 GeV
0.81 ± 0.03 ± 0.06		ADLER	87 MRK3	$e^+ e^-$ 3.77 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.31 <sup>+0.20</sup> <sub>-0.14</sub>	13	SUMMERS	84 E691	Photoproduction
0.85 <sup>+0.11</sup> <sub>-0.15</sub> <sup>+0.09</sup> <sub>-0.10</sub>	31	SCHINDLER	81 MRK2	$e^+ e^-$ 3.771 GeV

$\Gamma(K^*(892)^- \pi^+)/\Gamma(K^- \pi^+ \pi^0)$   $\Gamma_{81}/\Gamma_{30}$

Unseen decay modes of the  $K^*(892)^-$  are included.

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.36 ± 0.04 OUR FIT</b> Error includes scale factor of 1.3.			
<b>0.28 ± 0.04 OUR AVERAGE</b>			
0.444 ± 0.084 ± 0.147	FRABETTI	94G E687	$\gamma$ Be, $\bar{E}_\gamma \approx 220$ GeV
0.252 ± 0.033 ± 0.035	ANJOS	93 E691	$\gamma$ Be 90–260 GeV
0.36 ± 0.06 ± 0.09	ADLER	87 MRK3	$e^+ e^-$ 3.77 GeV

$\Gamma(\bar{K}^*(892)^0 \pi^0)/\Gamma(K^- \pi^+ \pi^0)$   $\Gamma_{82}/\Gamma_{30}$

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.227 ± 0.027 OUR FIT</b>			
<b>0.221 ± 0.029 OUR AVERAGE</b>			
0.248 ± 0.047 ± 0.023	FRABETTI	94G E687	$\gamma$ Be, $\bar{E}_\gamma \approx 220$ GeV
0.213 ± 0.027 ± 0.035	ANJOS	93 E691	$\gamma$ Be 90–260 GeV
0.20 ± 0.03 ± 0.05	ADLER	87 MRK3	$e^+ e^-$ 3.77 GeV

$\Gamma(K^- \pi^+ \pi^0 \text{ nonresonant})/\Gamma(K^- \pi^+ \pi^0)$   $\Gamma_{34}/\Gamma_{30}$

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.049 ± 0.018 OUR AVERAGE</b> Error includes scale factor of 1.1.				
0.101 ± 0.033 ± 0.040		FRABETTI	94G E687	$\gamma$ Be, $\bar{E}_\gamma \approx 220$ GeV
0.036 ± 0.004 ± 0.018		ANJOS	93 E691	$\gamma$ Be 90–260 GeV
0.09 ± 0.02 ± 0.04		ADLER	87 MRK3	$e^+ e^-$ 3.77 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.51 ± 0.22	21	SUMMERS	84 E691	Photoproduction

$\Gamma(\bar{K}^*(892)^0 \pi^0)/\Gamma(\bar{K}^0 \pi^0)$   $\Gamma_{82}/\Gamma_{21}$

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.49 ± 0.23 OUR FIT</b> Error includes scale factor of 1.1.				
<b>1.65 <sup>+0.39</sup><sub>-0.31</sub> ± 0.20</b>	122	PROCARIO	93B CLE2	$\bar{K}^0 \pi^0 \pi^0$ Dalitz plot

$\Gamma(\bar{K}_2^*(1430)^0 \pi^0)/\Gamma(\bar{K}^*(892)^0 \pi^0)$   $\Gamma_{105}/\Gamma_{82}$

Unseen decay modes of the  $\bar{K}_2^*(1430)^0$  and  $\bar{K}^*(892)^0$  are included.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>&lt; 0.12</b>	90	PROCARIO	93B CLE2	$\bar{K}^0 \pi^0 \pi^0$ Dalitz plot

$\Gamma(\overline{K}^0 \pi^0 \pi^0 \text{ nonresonant}) / \Gamma(\overline{K}^0 \pi^0)$

$\Gamma_{37} / \Gamma_{21}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.37 ± 0.08 ± 0.04</b>	76	PROCARIO	93B CLE2	$\overline{K}^0 \pi^0 \pi^0$ Dalitz plot

$\Gamma(K^- \pi^+ \pi^+ \pi^-) / \Gamma_{\text{total}}$

$\Gamma_{38} / \Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.0749 ± 0.0031 OUR FIT**

**0.075 ± 0.006 OUR AVERAGE** Error includes scale factor of 1.3. See the ideogram below.

0.079 ± 0.015 ± 0.009		<sup>50</sup> ALBRECHT	94 ARG	$e^+ e^- \approx \gamma(4S)$
0.0680 ± 0.0027 ± 0.0057	1430	<sup>51</sup> ALBRECHT	94F ARG	$e^+ e^- \approx \gamma(4S)$
0.091 ± 0.008 ± 0.008	992	ADLER	88C MRK3	$e^+ e^-$ 3.77 GeV
0.117 ± 0.025	185	<sup>52</sup> SCHINDLER	81 MRK2	$e^+ e^-$ 3.771 GeV
0.062 ± 0.019	44	<sup>53</sup> PERUZZI	77 MRK1	$e^+ e^-$ 3.77 GeV

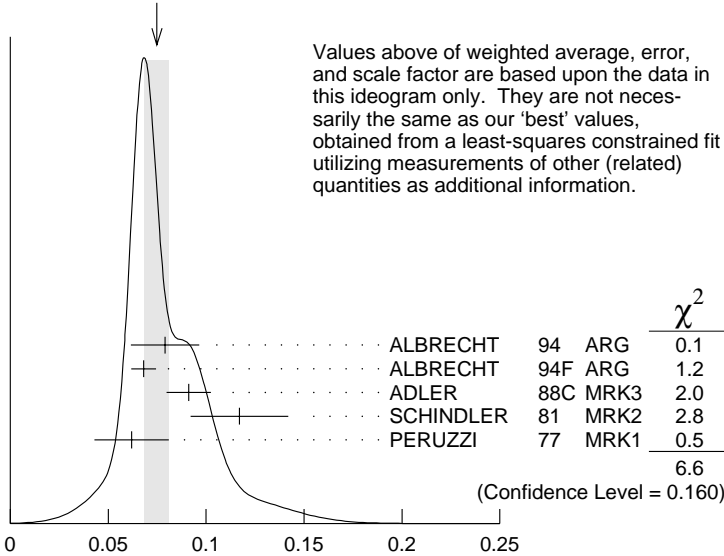
<sup>50</sup> ALBRECHT 94 uses  $D^0$  mesons from  $\overline{B}^0 \rightarrow D^{*+} \ell^- \overline{\nu}_\ell$  decays. This is a different set of events than used by ALBRECHT 94F.

<sup>51</sup> See the footnote on the ALBRECHT 94F measurement of  $\Gamma(K^- \pi^+) / \Gamma_{\text{total}}$  for the method used.

<sup>52</sup> SCHINDLER 81 (MARK-2) measures  $\sigma(e^+ e^- \rightarrow \psi(3770)) \times$  branching fraction to be  $0.68 \pm 0.11$  nb. We use the MARK-3 (ADLER 88C) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6$  nb.

<sup>53</sup> PERUZZI 77 (MARK-1) measures  $\sigma(e^+ e^- \rightarrow \psi(3770)) \times$  branching fraction to be  $0.36 \pm 0.10$  nb. We use the MARK-3 (ADLER 88C) value of  $\sigma = 5.8 \pm 0.5 \pm 0.6$  nb.

WEIGHTED AVERAGE  
0.075 ± 0.006 (Error scaled by 1.3)



$\Gamma(K^- \pi^+ \pi^+ \pi^-) / \Gamma_{\text{total}}$

$$\Gamma(K^- \pi^+ \pi^+ \pi^-) / \Gamma(K^- \pi^+) \quad \Gamma_{38} / \Gamma_{20}$$

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>1.96 ± 0.07 OUR FIT</b>				
<b>1.97 ± 0.09 OUR AVERAGE</b>				
1.94 ± 0.07 <sup>+0.09</sup> <sub>-0.11</sub>		JUN	00 SELX	Σ <sup>-</sup> nucleus, 600 GeV
1.7 ± 0.2 ± 0.2	1745	ANJOS	92C E691	γBe 90–260 GeV
1.90 ± 0.25 ± 0.20	337	ALVAREZ	91B NA14	Photoproduction
2.12 ± 0.16 ± 0.09		BORTOLETTO88	CLEO	e <sup>+</sup> e <sup>-</sup> 10.55 GeV
2.0 ± 0.9	48	BAILEY	86 ACCM	π <sup>-</sup> Be fixed target
2.17 ± 0.28 ± 0.23		ALBRECHT	85F ARG	e <sup>+</sup> e <sup>-</sup> 10 GeV
2.0 ± 1.0	10	BAILEY	83B SPEC	π <sup>-</sup> Be → D <sup>0</sup>
2.2 ± 0.8	214	PICCOLO	77 MRK1	e <sup>+</sup> e <sup>-</sup> 4.03, 4.41 GeV

$$\Gamma(K^- \pi^+ \rho^0 \text{ total}) / \Gamma(K^- \pi^+ \pi^+ \pi^-) \quad \Gamma_{39} / \Gamma_{38}$$

This includes  $K^- a_1(1260)^+$ ,  $\bar{K}^*(892)^0 \rho^0$ , etc. The next entry gives the specifically 3-body fraction. We rely on the MARK III and E691 full amplitude analyses of the  $K^- \pi^+ \pi^+ \pi^-$  channel for values of the resonant substructure.

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.835 ± 0.035 OUR AVERAGE</b>			
0.80 ± 0.03 ± 0.05	ANJOS	92C E691	γBe 90–260 GeV
0.855 ± 0.032 ± 0.030	COFFMAN	92B MRK3	e <sup>+</sup> e <sup>-</sup> 3.77 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.98 ± 0.12 ± 0.10	ALVAREZ	91B NA14	Photoproduction

$$\Gamma(K^- \pi^+ \rho^0 \text{ 3-body}) / \Gamma(K^- \pi^+ \pi^+ \pi^-) \quad \Gamma_{40} / \Gamma_{38}$$

We rely on the MARK III and E691 full amplitude analyses of the  $K^- \pi^+ \pi^+ \pi^-$  channel for values of the resonant substructure.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.063 ± 0.028 OUR AVERAGE</b>				
0.05 ± 0.03 ± 0.02		ANJOS	92C E691	γBe 90–260 GeV
0.084 ± 0.022 ± 0.04		COFFMAN	92B MRK3	e <sup>+</sup> e <sup>-</sup> 3.77 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.77 ± 0.06 ± 0.06	54	ALVAREZ	91B NA14	Photoproduction
0.85 <sup>+0.11</sup> <sub>-0.22</sub>	180	PICCOLO	77 MRK1	e <sup>+</sup> e <sup>-</sup> 4.03, 4.41 GeV

<sup>54</sup>This value is for  $\rho^0 (K^- \pi^+)$ -nonresonant. ALVAREZ 91B cannot determine what fraction of this is  $K^- a_1(1260)^+$ .

$$\Gamma(\bar{K}^*(892)^0 \rho^0) / \Gamma(K^- \pi^+ \pi^+ \pi^-) \quad \Gamma_{87} / \Gamma_{38}$$

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included. We rely on the MARK III and E691 full amplitude analyses of the  $K^- \pi^+ \pi^+ \pi^-$  channel for values of the resonant substructure.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.195 ± 0.03 ± 0.03</b>		ANJOS	92C E691	γBe 90–260 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.34 ± 0.09 ± 0.09		ALVAREZ	91B NA14	Photoproduction
0.75 ± 0.3	5	BAILEY	83B SPEC	πBe → D <sup>0</sup>
0.15 <sup>+0.16</sup> <sub>-0.15</sub>	20	PICCOLO	77 MRK1	e <sup>+</sup> e <sup>-</sup> 4.03, 4.41 GeV

$\Gamma(\bar{K}^*(892)^0 \rho^0 \text{ transverse})/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{88}/\Gamma_{38}$

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.20 ± 0.07 OUR FIT</b>			
<b>0.213 ± 0.024 ± 0.075</b>	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV

$\Gamma(\bar{K}^*(892)^0 \rho^0 \text{ S-wave})/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{89}/\Gamma_{38}$

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.375 ± 0.045 ± 0.06</b>	ANJOS	92C E691	$\gamma$ Be 90–260 GeV

$\Gamma(\bar{K}^*(892)^0 \rho^0 \text{ S-wave long.})/\Gamma_{\text{total}}$   $\Gamma_{90}/\Gamma$

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.003</b>	90	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV

$\Gamma(\bar{K}^*(892)^0 \rho^0 \text{ P-wave})/\Gamma_{\text{total}}$   $\Gamma_{91}/\Gamma$

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.003</b>	90	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.009	90	ANJOS	92C E691	$\gamma$ Be 90–260 GeV

$\Gamma(\bar{K}^*(892)^0 \rho^0 \text{ D-wave})/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{92}/\Gamma_{38}$

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.255 ± 0.045 ± 0.06</b>	ANJOS	92C E691	$\gamma$ Be 90–260 GeV

$\Gamma(K^- \pi^+ f_0(980))/\Gamma_{\text{total}}$   $\Gamma_{97}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.011</b>	90	ANJOS	92C E691	$\gamma$ Be 90–260 GeV

$\Gamma(\bar{K}^*(892)^0 f_0(980))/\Gamma_{\text{total}}$   $\Gamma_{98}/\Gamma$

Unseen decay modes of the  $\bar{K}^*(892)^0$  and  $f_0(980)$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.007</b>	90	ANJOS	92C E691	$\gamma$ Be 90–260 GeV

$\Gamma(K^- a_1(1260)^+)/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{76}/\Gamma_{38}$

Unseen decay modes of the  $a_1(1260)^+$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.97 ± 0.14 OUR AVERAGE</b>			
0.94 ± 0.13 ± 0.20	ANJOS	92C E691	$\gamma$ Be 90–260 GeV
0.984 ± 0.048 ± 0.16	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV

$\Gamma(K^- a_2(1320)^+)/\Gamma_{\text{total}}$   $\Gamma_{79}/\Gamma$

Unseen decay modes of the  $a_2(1320)^+$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.002</b>	90	ANJOS	92C E691	$\gamma$ Be 90–260 GeV
• • • We do not use the following data for averages, fits, limits, etc. • • •				
<0.006	90	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV



$\Gamma(K_1(1270)^- \pi^+)/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{99}/\Gamma_{38}$

Unseen decay modes of the  $K_1(1270)^-$  are included. The MARK3 and E691 experiments disagree considerably here.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
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**0.14 ± 0.04 OUR FIT**

**0.194 ± 0.056 ± 0.088**      COFFMAN      92B MRK3       $e^+ e^-$  3.77 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.013      90      ANJOS      92C E691       $\gamma$ Be 90–260 GeV

$\Gamma(K_1(1400)^- \pi^+)/\Gamma_{total}$   $\Gamma_{100}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
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**<0.012**      90      COFFMAN      92B MRK3       $e^+ e^-$  3.77 GeV

$\Gamma(K^*(1410)^- \pi^+)/\Gamma_{total}$   $\Gamma_{102}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
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**<0.012**      90      COFFMAN      92B MRK3       $e^+ e^-$  3.77 GeV

$\Gamma(\bar{K}^*(892)^0 \pi^+ \pi^- total)/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{83}/\Gamma_{38}$

This includes  $\bar{K}^*(892)^0 \rho^0$ , etc. The next entry gives the specifically 3-body fraction.

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
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**0.30 ± 0.06 ± 0.03**      ANJOS      92C E691       $\gamma$ Be 90–260 GeV

$\Gamma(\bar{K}^*(892)^0 \pi^+ \pi^- 3\text{-body})/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{84}/\Gamma_{38}$

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
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**0.19 ± 0.04 OUR FIT**

**0.18 ± 0.04 OUR AVERAGE**

0.165 ± 0.03 ± 0.045      ANJOS      92C E691       $\gamma$ Be 90–260 GeV

0.210 ± 0.027 ± 0.06      COFFMAN      92B MRK3       $e^+ e^-$  3.77 GeV

$\Gamma(K^- \pi^+ \pi^+ \pi^- \text{nonresonant})/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{46}/\Gamma_{38}$

VALUE	DOCUMENT ID	TECN	COMMENT
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**0.233 ± 0.032 OUR AVERAGE**

0.23 ± 0.02 ± 0.03      ANJOS      92C E691       $\gamma$ Be 90–260 GeV

0.242 ± 0.025 ± 0.06      COFFMAN      92B MRK3       $e^+ e^-$  3.77 GeV

$\Gamma(\bar{K}^0 \pi^+ \pi^- \pi^0)/\Gamma_{total}$   $\Gamma_{47}/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.100 ± 0.012 OUR FIT**

**0.103 ± 0.022 ± 0.025**      140      COFFMAN      92B MRK3       $e^+ e^-$  3.77 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.134 <sup>+0.032</sup>/<sub>-0.033</sub>      <sup>55</sup>BARLAG      92C ACCM       $\pi^-$  Cu 230 GeV

<sup>55</sup>BARLAG 92C computes the branching fraction using topological normalization.

$\Gamma(\bar{K}^0 \pi^+ \pi^- \pi^0) / \Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{47} / \Gamma_{22}$

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**1.84 ± 0.20 OUR FIT**

**1.86 ± 0.23 OUR AVERAGE**

1.80 ± 0.20 ± 0.21	190	<sup>56</sup> ALBRECHT	92P ARG	e <sup>+</sup> e <sup>-</sup> ≈ 10 GeV
2.8 ± 0.8 ± 0.8	46	ANJOS	92C E691	γ Be 90–260 GeV
1.85 ± 0.26 ± 0.30	158	KINOSHITA	91 CLEO	e <sup>+</sup> e <sup>-</sup> ~ 10.7 GeV

<sup>56</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.

$\Gamma(\bar{K}^0 \eta) / \Gamma(K^- \pi^+)$   $\Gamma_{69} / \Gamma_{20}$

Unseen decay modes of the η are included.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 0.64$	90	ALBRECHT	89D ARG	e <sup>+</sup> e <sup>-</sup> 10 GeV
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$\Gamma(\bar{K}^0 \eta) / \Gamma(\bar{K}^0 \pi^0)$   $\Gamma_{69} / \Gamma_{21}$

Unseen decay modes of the η are included.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**0.33 ± 0.04 OUR FIT**

0.32 ± 0.04 ± 0.03	225	PROCARIO	93B CLE2	η → γγ
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$\Gamma(\bar{K}^0 \eta) / \Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{69} / \Gamma_{22}$

Unseen decay modes of the η are included.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**0.130 ± 0.017 OUR FIT**

0.14 ± 0.02 ± 0.02	80	PROCARIO	93B CLE2	η → π <sup>+</sup> π <sup>-</sup> π <sup>0</sup>
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$\Gamma(\bar{K}^0 \omega) / \Gamma(K^- \pi^+)$   $\Gamma_{72} / \Gamma_{20}$

Unseen decay modes of the ω are included.

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**0.54 ± 0.10 OUR FIT**

1.00 ± 0.36 ± 0.20	ALBRECHT	89D ARG	e <sup>+</sup> e <sup>-</sup> 10 GeV
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$\Gamma(\bar{K}^0 \omega) / \Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{72} / \Gamma_{22}$

Unseen decay modes of the ω are included.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**0.38 ± 0.07 OUR FIT**

**0.33 ± 0.09 OUR AVERAGE** Error includes scale factor of 1.1.

0.29 ± 0.08 ± 0.05	16	<sup>57</sup> ALBRECHT	92P ARG	e <sup>+</sup> e <sup>-</sup> ≈ 10 GeV
0.54 ± 0.14 ± 0.16	40	KINOSHITA	91 CLEO	e <sup>+</sup> e <sup>-</sup> ~ 10.7 GeV

<sup>57</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.

$\Gamma(\bar{K}^0 \omega) / \Gamma(\bar{K}^0 \pi^+ \pi^- \pi^0)$   $\Gamma_{72} / \Gamma_{47}$

Unseen decay modes of the ω are included.

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**0.21 ± 0.04 OUR FIT**

0.220 ± 0.048 ± 0.0116	COFFMAN	92B MRK3	e <sup>+</sup> e <sup>-</sup> 3.77 GeV
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$\Gamma(\bar{K}^0 \eta'(958))/\Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{73}/\Gamma_{22}$

Unseen decay modes of the  $\eta'(958)$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.32 ± 0.04 OUR AVERAGE</b>				
0.31 ± 0.02 ± 0.04	594	PROCARIO	93B CLE2	$\eta' \rightarrow \eta \pi^+ \pi^-, \rho^0 \gamma$
0.37 ± 0.13 ± 0.06	18	<sup>58</sup> ALBRECHT	92P ARG	$e^+ e^- \approx 10 \text{ GeV}$

<sup>58</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.

$\Gamma(K^*(892)^- \rho^+)/\Gamma(\bar{K}^0 \pi^+ \pi^- \pi^0)$   $\Gamma_{93}/\Gamma_{47}$

Unseen decay modes of the  $K^*(892)^-$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.606 ± 0.188 ± 0.126</b>	COFFMAN	92B MRK3	$e^+ e^- 3.77 \text{ GeV}$

$\Gamma(K^*(892)^- \rho^+ \text{longitudinal})/\Gamma(\bar{K}^0 \pi^+ \pi^- \pi^0)$   $\Gamma_{94}/\Gamma_{47}$

Unseen decay modes of the  $K^*(892)^-$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.290 ± 0.111</b>	COFFMAN	92B MRK3	$e^+ e^- 3.77 \text{ GeV}$

$\Gamma(K^*(892)^- \rho^+ \text{transverse})/\Gamma(\bar{K}^0 \pi^+ \pi^- \pi^0)$   $\Gamma_{95}/\Gamma_{47}$

Unseen decay modes of the  $K^*(892)^-$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.317 ± 0.180</b>	COFFMAN	92B MRK3	$e^+ e^- 3.77 \text{ GeV}$

$\Gamma(K^*(892)^- \rho^+ P\text{-wave})/\Gamma_{\text{total}}$   $\Gamma_{96}/\Gamma$

Unseen decay modes of the  $K^*(892)^-$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.015</b>	90	<sup>59</sup> COFFMAN	92B MRK3	$e^+ e^- 3.77 \text{ GeV}$

<sup>59</sup> Obtained using other  $\bar{K}^*(892) \rho P\text{-wave}$  limits and isospin relations.

$\Gamma(\bar{K}^*(892)^0 \rho^0 \text{transverse})/\Gamma(\bar{K}^0 \pi^+ \pi^- \pi^0)$   $\Gamma_{88}/\Gamma_{47}$

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.15 ± 0.06 OUR FIT</b>			
<b>0.126 ± 0.111</b>	COFFMAN	92B MRK3	$e^+ e^- 3.77 \text{ GeV}$

$\Gamma(\bar{K}^0 a_1(1260)^0)/\Gamma_{\text{total}}$   $\Gamma_{77}/\Gamma$

Unseen decay modes of the  $a_1(1260)^0$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.019</b>	90	COFFMAN	92B MRK3	$e^+ e^- 3.77 \text{ GeV}$

$\Gamma(K_1(1270)^- \pi^+)/\Gamma(\bar{K}^0 \pi^+ \pi^- \pi^0)$   $\Gamma_{99}/\Gamma_{47}$

Unseen decay modes of the  $K_1(1270)^-$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.106 ± 0.028 OUR FIT</b>			
<b>0.10 ± 0.03</b>	COFFMAN	92B MRK3	$e^+ e^- 3.77 \text{ GeV}$

$\Gamma(\bar{K}_1(1400)^0 \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{101}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.037</b>	90	COFFMAN	92B MRK3	$e^+ e^- 3.77 \text{ GeV}$

$\Gamma(\bar{K}^*(892)^0 \pi^+ \pi^- \text{3-body}) / \Gamma(\bar{K}^0 \pi^+ \pi^- \pi^0)$   $\Gamma_{84}/\Gamma_{47}$

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.14 ± 0.04 OUR FIT</b>			Error includes scale factor of 1.1.
<b>0.191 ± 0.105</b>	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV

$\Gamma(\bar{K}^0 \pi^+ \pi^- \pi^0 \text{nonresonant}) / \Gamma(\bar{K}^0 \pi^+ \pi^- \pi^0)$   $\Gamma_{54}/\Gamma_{47}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.210 ± 0.147 ± 0.150</b>	COFFMAN	92B MRK3	$e^+ e^-$ 3.77 GeV

$\Gamma(K^- \pi^+ \pi^0 \pi^0) / \Gamma_{\text{total}}$   $\Gamma_{55}/\Gamma$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.177 ± 0.029		<sup>60</sup> BARLAG	92C ACCM	$\pi^-$ Cu 230 GeV
0.149 ± 0.037 ± 0.030	24	<sup>61</sup> ADLER	88C MRK3	$e^+ e^-$ 3.77 GeV
0.209 <sup>+0.074</sup> <sub>-0.043</sub> ± 0.012	9	<sup>60</sup> AGUILAR-...	87F HYBR	$\pi p, p p$ 360, 400 GeV

<sup>60</sup> AGUILAR-BENITEZ 87F and BARLAG 92C compute the branching fraction using topological normalization. They do not distinguish the presence of a third  $\pi^0$ , and thus are not included in the average.

<sup>61</sup> ADLER 88C uses an absolute normalization method finding this decay channel opposite a detected  $\bar{D}^0 \rightarrow K^+ \pi^-$  in pure  $D\bar{D}$  events.

$\Gamma(K^- \pi^+ \pi^+ \pi^- \pi^0) / \Gamma(K^- \pi^+)$   $\Gamma_{56}/\Gamma_{20}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.05 ± 0.10 OUR FIT**

<b>0.98 ± 0.11 ± 0.11</b>	225	<sup>62</sup> ALBRECHT	92P ARG	$e^+ e^- \approx 10$ GeV
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<sup>62</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.

$\Gamma(K^- \pi^+ \pi^+ \pi^- \pi^0) / \Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{56}/\Gamma_{38}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.54 ± 0.05 OUR FIT**

**0.56 ± 0.07 OUR AVERAGE**

0.55 ± 0.07 <sup>+0.12</sup> <sub>-0.09</sub>	167	KINOSHITA	91 CLEO	$e^+ e^- \sim 10.7$ GeV
0.57 ± 0.06 ± 0.05	180	ANJOS	90D E691	Photoproduction

$\Gamma(\bar{K}^*(892)^0 \pi^+ \pi^- \pi^0) / \Gamma(K^- \pi^+ \pi^+ \pi^- \pi^0)$   $\Gamma_{106}/\Gamma_{56}$

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
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<b>0.45 ± 0.15 ± 0.15</b>	ANJOS	90D E691	Photoproduction
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$\Gamma(\bar{K}^*(892)^0 \eta) / \Gamma(K^- \pi^+)$   $\Gamma_{107}/\Gamma_{20}$

Unseen decay modes of the  $\bar{K}^*(892)^0$  and  $\eta$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.49 ± 0.12 OUR FIT**

<b>0.58 ± 0.19<sup>+0.24</sup></b> <b>-0.28</b>	46	KINOSHITA	91 CLEO	$e^+ e^- \sim 10.7$ GeV
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$\Gamma(\bar{K}^*(892)^0 \eta) / \Gamma(K^- \pi^+ \pi^0)$   $\Gamma_{107} / \Gamma_{30}$

Unseen decay modes of the  $\bar{K}^*(892)^0$  and  $\eta$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.134 ± 0.034 OUR FIT</b>				
<b>0.13 ± 0.02 ± 0.03</b>	214	PROCARIO	93B CLE2	$\bar{K}^{*0} \eta \rightarrow K^- \pi^+ / \gamma \gamma$

$\Gamma(K^- \pi^+ \omega) / \Gamma(K^- \pi^+)$   $\Gamma_{108} / \Gamma_{20}$

Unseen decay modes of the  $\omega$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.78 ± 0.12 ± 0.10</b>	99	<sup>63</sup> ALBRECHT	92P ARG	$e^+ e^- \approx 10$ GeV

<sup>63</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.

$\Gamma(\bar{K}^*(892)^0 \omega) / \Gamma(K^- \pi^+)$   $\Gamma_{109} / \Gamma_{20}$

Unseen decay modes of the  $\bar{K}^*(892)^0$  and  $\omega$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.28 ± 0.11 ± 0.04</b>	17	<sup>64</sup> ALBRECHT	92P ARG	$e^+ e^- \approx 10$ GeV

<sup>64</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.

$\Gamma(\bar{K}^*(892)^0 \omega) / \Gamma(K^- \pi^+ \pi^+ \pi^- \pi^0)$   $\Gamma_{109} / \Gamma_{56}$

Unseen decay modes of the  $\bar{K}^*(892)^0$  and  $\omega$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>• • •</b>				We do not use the following data for averages, fits, limits, etc. • • •

<0.44	90	<sup>65</sup> ANJOS	90D E691	Photoproduction
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<sup>65</sup> Recovered from the published limit,  $\Gamma(\bar{K}^*(892)^0 \omega) / \Gamma_{\text{total}}$ , in order to make our normalization consistent.

$\Gamma(K^- \pi^+ \eta'(958)) / \Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{110} / \Gamma_{38}$

Unseen decay modes of the  $\eta'(958)$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.093 ± 0.014 ± 0.019</b>	286	PROCARIO	93B CLE2	$\eta' \rightarrow \eta \pi^+ \pi^-, \rho^0 \gamma$

$\Gamma(\bar{K}^*(892)^0 \eta'(958)) / \Gamma(K^- \pi^+ \eta'(958))$   $\Gamma_{111} / \Gamma_{110}$

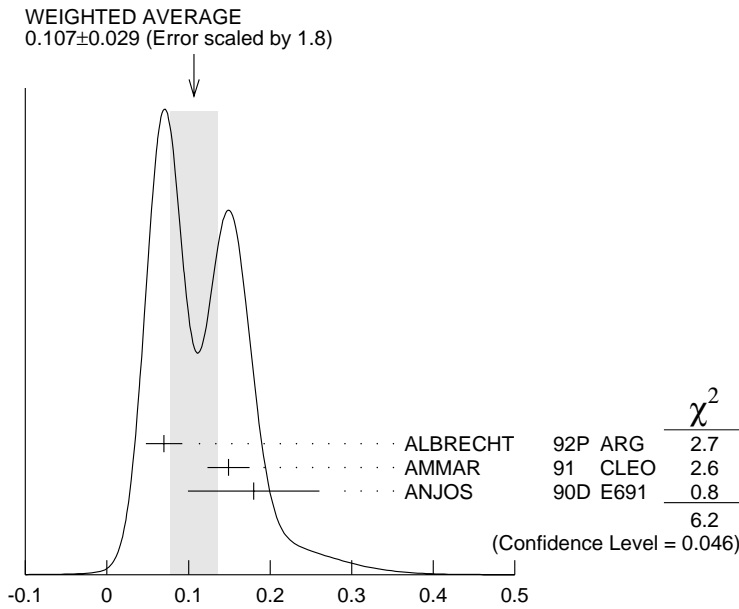
Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.15</b>	90	PROCARIO	93B CLE2	

$\Gamma(\bar{K}^0 \pi^+ \pi^+ \pi^- \pi^-) / \Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{61} / \Gamma_{22}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.107 ± 0.029 OUR AVERAGE</b>				Error includes scale factor of 1.8. See the ideogram below.
0.07 ± 0.02 ± 0.01	11	<sup>66</sup> ALBRECHT	92P ARG	$e^+ e^- \approx 10$ GeV
0.149 ± 0.026	56	AMMAR	91 CLEO	$e^+ e^- \approx 10.5$ GeV
0.18 ± 0.07 ± 0.04	6	ANJOS	90D E691	Photoproduction

<sup>66</sup> This value is calculated from numbers in Table 1 of ALBRECHT 92P.



$$\Gamma(\bar{K}^0 \pi^+ \pi^+ \pi^- \pi^-) / \Gamma(\bar{K}^0 \pi^+ \pi^-)$$

$$\Gamma(\bar{K}^0 \pi^+ \pi^- \pi^0 \pi^0 (\pi^0)) / \Gamma_{\text{total}} \quad \Gamma_{62} / \Gamma$$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.106 <sup>+0.073</sup> <sub>-0.029</sub> ± 0.006	4	<sup>67</sup> AGUILAR-...	87F HYBR	$\pi p, p p$ 360, 400 GeV
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<sup>67</sup> AGUILAR-BENITEZ 87F computes the branching fraction using topological normalization, and does not distinguish the presence of a third  $\pi^0$ .

$$\Gamma(\bar{K}^0 K^+ K^-) / \Gamma(\bar{K}^0 \pi^+ \pi^-) \quad \Gamma_{63} / \Gamma_{22} = (\Gamma_{65} + \frac{1}{2} \Gamma_{75}) / \Gamma_{22}$$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.172 ± 0.014 OUR FIT**

**0.178 ± 0.019 OUR AVERAGE**

0.20 ± 0.05 ± 0.04	47	FRABETTI	92B E687	$\gamma \text{Be } \bar{E}_\gamma = 221 \text{ GeV}$
0.170 ± 0.022	136	AMMAR	91 CLEO	$e^+ e^- \approx 10.5 \text{ GeV}$
0.24 ± 0.08		BEBEK	86 CLEO	$e^+ e^-$ near $\Upsilon(4S)$
0.185 ± 0.055	52	ALBRECHT	85B ARG	$e^+ e^-$ 10 GeV

$$\Gamma(\bar{K}^0 \phi) / \Gamma(\bar{K}^0 \pi^+ \pi^-) \quad \Gamma_{75} / \Gamma_{22}$$

Unseen decay modes of the  $\phi$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.158 ± 0.016 OUR FIT**

**0.156 ± 0.017 OUR AVERAGE**

0.13 ± 0.06 ± 0.02	13	FRABETTI	92B E687	$\gamma \text{Be } \bar{E}_\gamma = 221 \text{ GeV}$
0.163 ± 0.023	63	AMMAR	91 CLEO	$e^+ e^- \approx 10.5 \text{ GeV}$
0.155 ± 0.033	56	ALBRECHT	87E ARG	$e^+ e^-$ 10 GeV
0.14 ± 0.05	29	BEBEK	86 CLEO	$e^+ e^-$ near $\Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

0.186±0.052                      26                      ALBRECHT                      85B ARG                      See ALBRECHT 87E

**$\Gamma(\bar{K}^0 K^+ K^- \text{non-}\phi)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$**   **$\Gamma_{65}/\Gamma_{22}$**

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.093±0.014 OUR FIT</b>				
<b>0.088±0.019 OUR AVERAGE</b>				
0.11 ±0.04 ±0.03	20	FRABETTI	92B E687	$\gamma$ Be $\bar{E}_\gamma = 221$ GeV
0.084±0.020		ALBRECHT	87E ARG	$e^+ e^- 10$ GeV

**$\Gamma(K_S^0 \bar{K}_S^0 K_S^0)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$**   **$\Gamma_{66}/\Gamma_{22}$**

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.0154±0.0025 OUR AVERAGE</b>				
0.0139±0.0019±0.0024	61	ASNER	96B CLE2	$e^+ e^- \approx \gamma(4S)$
0.035 ±0.012 ±0.006	10	FRABETTI	94J E687	$\gamma$ Be $\bar{E}_\gamma = 220$ GeV
0.016 ±0.005	22	AMMAR	91 CLEO	$e^+ e^- \approx 10.5$ GeV
0.017 ±0.007 ±0.005	5	ALBRECHT	90C ARG	$e^+ e^- \approx 10$ GeV

**$\Gamma(K^+ K^- K^- \pi^+)/\Gamma(K^- \pi^+ \pi^+ \pi^-)$**   **$\Gamma_{67}/\Gamma_{38}$**

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.0028±0.0007±0.0001</b>	20	FRABETTI	95C E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV

**$\Gamma(K^+ K^- \bar{K}^0 \pi^0)/\Gamma_{\text{total}}$**   **$\Gamma_{68}/\Gamma$**

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.0072<sup>+0.0048</sup><sub>-0.0035</sub></b>	<sup>68</sup> BARLAG	92C ACCM	$\pi^-$ Cu 230 GeV

<sup>68</sup>BARLAG 92C computes the branching fraction using topological normalization.

**Pionic modes**

**$\Gamma(\pi^+ \pi^-)/\Gamma(K^- \pi^+)$**   **$\Gamma_{112}/\Gamma_{20}$**

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.0397±0.0021 OUR AVERAGE</b>				
0.040 ±0.002 ±0.003	2043	AITALA	98C E791	$\pi^-$ nucleus, 500 GeV
0.043 ±0.007 ±0.003	177	FRABETTI	94C E687	$\gamma$ Be $\bar{E}_\gamma = 220$ GeV
0.0348±0.0030±0.0023	227	SELEN	93 CLE2	$e^+ e^- \approx \gamma(4S)$
0.048 ±0.013 ±0.008	51	ADAMOVICH	92 OMEG	$\pi^- 340$ GeV
0.055 ±0.008 ±0.005	120	ANJOS	91D E691	Photoproduction
0.040 ±0.007 ±0.006	57	ALBRECHT	90C ARG	$e^+ e^- \approx 10$ GeV
0.050 ±0.007 ±0.005	110	ALEXANDER	90 CLEO	$e^+ e^- 10.5-11$ GeV
0.033 ±0.010 ±0.006	39	BALTRUSAIT.	85E MRK3	$e^+ e^- 3.77$ GeV
0.033 ±0.015		ABRAMS	79D MRK2	$e^+ e^- 3.77$ GeV

**$\Gamma(\pi^0 \pi^0)/\Gamma(K^- \pi^+)$**   **$\Gamma_{113}/\Gamma_{20}$**

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.022±0.004±0.004</b>	40	SELEN	93 CLE2	$e^+ e^- \approx \gamma(4S)$

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{114}/\Gamma$

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.011 ± 0.004 ± 0.002</b>	10	<sup>69</sup> BALTRUSAIT..85E	MRK3	$e^+e^-$ 3.77 GeV
0.0390 <sup>+0.0100</sup> <sub>-0.0095</sub>		<sup>70</sup> BARLAG	92C ACCM	$\pi^-$ Cu 230 GeV

<sup>69</sup> All the BALTRUSAITIS 85E events are consistent with  $\rho^0\pi^0$ .

<sup>70</sup> BARLAG 92C computes the branching fraction using topological normalization. Possible contamination by extra  $\pi^0$ 's may partly explain the unexpectedly large value.

$\Gamma(\pi^+\pi^+\pi^-\pi^-)/\Gamma(K^-\pi^+\pi^+\pi^-)$   $\Gamma_{115}/\Gamma_{38}$

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.098 ± 0.006 OUR AVERAGE</b>				
0.095 ± 0.007 ± 0.002	814	FRABETTI	95C E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV
0.115 ± 0.023 ± 0.016	64	ADAMOVICH	92 OMEG	$\pi^-$ 340 GeV
0.108 ± 0.024 ± 0.008	79	FRABETTI	92 E687	$\gamma$ Be
0.102 ± 0.013	345	<sup>71</sup> AMMAR	91 CLEO	$e^+e^- \approx 10.5$ GeV
0.096 ± 0.018 ± 0.007	66	ANJOS	91 E691	$\gamma$ Be 80–240 GeV

<sup>71</sup> AMMAR 91 finds  $1.25 \pm 0.25 \pm 0.25$   $\rho^0$ 's per  $\pi^+\pi^+\pi^-\pi^-$  decay, but can't untangle the resonant substructure ( $\rho^0\rho^0$ ,  $a_1^\pm\pi^\mp$ ,  $\rho^0\pi^+\pi^-$ ).

$\Gamma(\pi^+\pi^+\pi^-\pi^-\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{116}/\Gamma$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.0192<sup>+0.0041</sup><sub>-0.0038</sub></b>	<sup>72</sup> BARLAG	92C ACCM	$\pi^-$ Cu 230 GeV

<sup>72</sup> BARLAG 92C computes the branching fraction using topological normalization.

$\Gamma(\pi^+\pi^+\pi^+\pi^-\pi^-\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{117}/\Gamma$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.0004 ± 0.0003</b>	<sup>73</sup> BARLAG	92C ACCM	$\pi^-$ Cu 230 GeV

<sup>73</sup> BARLAG 92C computes the branching fraction using topological normalization.

————— **Hadronic modes with a  $K\bar{K}$  pair** —————

$\Gamma(K^+K^-)/\Gamma(K^-\pi^+)$   $\Gamma_{118}/\Gamma_{20}$

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.1109 ± 0.0033 OUR FIT</b>				
<b>0.1109 ± 0.0033 OUR AVERAGE</b>				
0.109 ± 0.003 ± 0.003	3317	AITALA	98C E791	$\pi^-$ nucleus, 500 GeV
0.116 ± 0.007 ± 0.007	1102	ASNER	96B CLE2	$e^+e^- \approx \gamma(4S)$
0.109 ± 0.007 ± 0.009	581	FRABETTI	94C E687	$\gamma$ Be $\bar{E}_\gamma = 220$ GeV
0.107 ± 0.029 ± 0.015	103	ADAMOVICH	92 OMEG	$\pi^-$ 340 GeV
0.138 ± 0.027 ± 0.010	155	FRABETTI	92 E687	$\gamma$ Be
0.16 ± 0.05	34	ALVAREZ	91B NA14	Photoproduction
0.107 ± 0.010 ± 0.009	193	ANJOS	91D E691	Photoproduction
0.10 ± 0.02 ± 0.01	131	ALBRECHT	90C ARG	$e^+e^- \approx 10$ GeV
0.117 ± 0.010 ± 0.007	249	ALEXANDER	90 CLEO	$e^+e^-$ 10.5–11 GeV
0.122 ± 0.018 ± 0.012	118	BALTRUSAIT..85E	MRK3	$e^+e^-$ 3.77 GeV
0.113 ± 0.030		ABRAMS	79D MRK2	$e^+e^-$ 3.77 GeV



$\Gamma(K^+ K^-)/\Gamma(\pi^+ \pi^-)$

$\Gamma_{118}/\Gamma_{112}$

The unused results here are redundant with  $\Gamma(K^+ K^-)/\Gamma(K^- \pi^+)$  and  $\Gamma(\pi^+ \pi^-)/\Gamma(K^- \pi^+)$  measurements by the same experiments.

VALUE	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$2.75 \pm 0.15 \pm 0.16$	AITALA	98C E791	$\pi^-$ nucleus, 500 GeV
$2.53 \pm 0.46 \pm 0.19$	FRABETTI	94C E687	$\gamma$ Be $\bar{E}_\gamma = 220$ GeV
$2.23 \pm 0.81 \pm 0.46$	ADAMOVICH	92 OMEG	$\pi^-$ 340 GeV
$1.95 \pm 0.34 \pm 0.22$	ANJOS	91D E691	Photoproduction
$2.5 \pm 0.7$	ALBRECHT	90C ARG	$e^+ e^- \approx 10$ GeV
$2.35 \pm 0.37 \pm 0.28$	ALEXANDER	90 CLEO	$e^+ e^-$ 10.5–11 GeV

$\Gamma(K^0 \bar{K}^0)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$

$\Gamma_{119}/\Gamma_{22}$

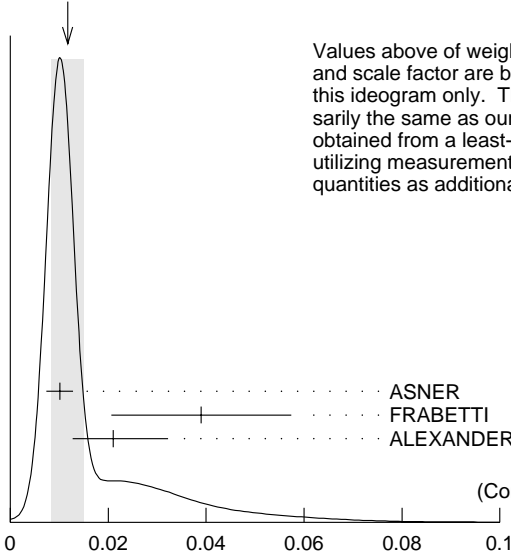
VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.0120 ± 0.0033 OUR FIT** Error includes scale factor of 1.3.

**0.0117 ± 0.0033 OUR AVERAGE** Error includes scale factor of 1.3. See the ideogram below.

$0.0101 \pm 0.0022 \pm 0.0016$	26	ASNER	96B CLE2	$e^+ e^- \approx \gamma(4S)$
$0.039 \pm 0.013 \pm 0.013$	20	FRABETTI	94J E687	$\gamma$ Be $\bar{E}_\gamma = 220$ GeV
$0.021 \begin{smallmatrix} +0.011 \\ -0.008 \end{smallmatrix} \pm 0.002$	5	ALEXANDER	90 CLEO	$e^+ e^-$ 10.5–11 GeV

WEIGHTED AVERAGE  
0.0117 ± 0.0033 (Error scaled by 1.3)



Values above of weighted average, error, and scale factor are based upon the data in this ideogram only. They are not necessarily the same as our 'best' values, obtained from a least-squares constrained fit utilizing measurements of other (related) quantities as additional information.

$\Gamma(K^0 \bar{K}^0)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$

$\Gamma(K^0 \bar{K}^0)/\Gamma(K^+ K^-)$   $\Gamma_{119}/\Gamma_{118}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.15±0.04 OUR FIT</b>				Error includes scale factor of 1.2.
<b>0.24±0.16</b>	4	<sup>74</sup> CUMALAT	88	SPEC $nN$ 0–800 GeV

<sup>74</sup> Includes a correction communicated to us by the authors of CUMALAT 88.

$\Gamma(K^0 K^- \pi^+)/\Gamma(K^- \pi^+)$   $\Gamma_{120}/\Gamma_{20}$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.168±0.026 OUR FIT</b>			Error includes scale factor of 1.1.
<b>0.16 ±0.06</b>	<sup>75</sup> ANJOS	91	E691 $\gamma$ Be 80–240 GeV

<sup>75</sup> The factor 100 at the top of column 2 of Table I of ANJOS 91 should be omitted.

$\Gamma(K^0 K^- \pi^+)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{120}/\Gamma_{22}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.118±0.018 OUR FIT</b>				Error includes scale factor of 1.1.
<b>0.119±0.021 OUR AVERAGE</b>				Error includes scale factor of 1.3.
0.108±0.019	61	AMMAR	91	CLEO $e^+ e^- \approx 10.5$ GeV
0.16 ±0.03 ±0.02	39	ALBRECHT	90C	ARG $e^+ e^- \approx 10$ GeV

$\Gamma(\bar{K}^*(892)^0 K^0)/\Gamma(K^- \pi^+)$   $\Gamma_{140}/\Gamma_{20}$

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
• • •			We do not use the following data for averages, fits, limits, etc. • • •
0.00 <sup>+0.03</sup> <sub>-0.00</sub>	<sup>76</sup> ANJOS	91	E691 $\gamma$ Be 80–240 GeV

<sup>76</sup> The factor 100 at the top of column 2 of Table I of ANJOS 91 should be omitted.

$\Gamma(\bar{K}^*(892)^0 K^0)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{140}/\Gamma_{22}$

Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.029</b>	90	AMMAR	91	CLEO $e^+ e^- \approx 10.5$ GeV
• • •				We do not use the following data for averages, fits, limits, etc. • • •
<0.03	90	ALBRECHT	90C	ARG $e^+ e^- \approx 10$ GeV

$\Gamma(K^*(892)^+ K^-)/\Gamma(K^- \pi^+)$   $\Gamma_{141}/\Gamma_{20}$

Unseen decay modes of the  $K^*(892)^+$  are included.

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.090±0.020 OUR FIT</b>			
<b>0.16<sup>+0.08</sup><sub>-0.06</sub></b>	<sup>77</sup> ANJOS	91	E691 $\gamma$ Be 80–240 GeV

<sup>77</sup> The factor 100 at the top of column 2 of Table I of ANJOS 91 should be omitted.

$\Gamma(K^*(892)^+ K^-)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{141}/\Gamma_{22}$

Unseen decay modes of the  $K^*(892)^+$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.064±0.014 OUR FIT</b>				Error includes scale factor of 1.1.
<b>0.058±0.014 OUR AVERAGE</b>				
0.064±0.018	23	AMMAR	91	CLEO $e^+ e^- \approx 10.5$ GeV
0.05 ±0.02 ±0.01	15	ALBRECHT	90C	ARG $e^+ e^- \approx 10$ GeV

$\Gamma(K^0 K^- \pi^+ \text{ nonresonant})/\Gamma(K^- \pi^+)$   $\Gamma_{123}/\Gamma_{20}$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.06±0.06</b>	<sup>78</sup> ANJOS	91	E691 $\gamma$ Be 80–240 GeV

<sup>78</sup> The factor 100 at the top of column 2 of Table I of ANJOS 91 should be omitted.

$\Gamma(\bar{K}^0 K^+ \pi^-)/\Gamma(K^- \pi^+)$   $\Gamma_{124}/\Gamma_{20}$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.129±0.025 OUR FIT</b>			
<b>0.10 ±0.05</b>	<sup>79</sup> ANJOS	91	E691 $\gamma$ Be 80–240 GeV

<sup>79</sup> The factor 100 at the top of column 2 of Table I of ANJOS 91 should be omitted.

$\Gamma(\bar{K}^0 K^+ \pi^-)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{124}/\Gamma_{22}$

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>0.091±0.018 OUR FIT</b>				
<b>0.098±0.020</b>	55	AMMAR	91	CLEO $e^+ e^- \approx 10.5$ GeV

$\Gamma(K^*(892)^0 \bar{K}^0)/\Gamma(K^- \pi^+)$   $\Gamma_{142}/\Gamma_{20}$

Unseen decay modes of the  $K^*(892)^0$  are included.

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.00^{+0.04}_{-0.00}$	<sup>80</sup> ANJOS	91	E691 $\gamma$ Be 80–240 GeV
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<sup>80</sup> The factor 100 at the top of column 2 of Table I of ANJOS 91 should be omitted.

$\Gamma(K^*(892)^0 \bar{K}^0)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{142}/\Gamma_{22}$

Unseen decay modes of the  $K^*(892)^0$  are included.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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<b>&lt;0.015</b>	90	AMMAR	91	CLEO $e^+ e^- \approx 10.5$ GeV
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$\Gamma(K^*(892)^- K^+)/\Gamma(K^- \pi^+)$   $\Gamma_{143}/\Gamma_{20}$

Unseen decay modes of the  $K^*(892)^-$  are included.

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.00^{+0.03}_{-0.00}$	<sup>81</sup> ANJOS	91	E691 $\gamma$ Be 80–240 GeV
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<sup>81</sup> The factor 100 at the top of column 2 of Table I of ANJOS 91 should be omitted.

$\Gamma(K^*(892)^- K^+)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{143}/\Gamma_{22}$

Unseen decay modes of the  $K^*(892)^-$  are included.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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<b>0.034±0.019</b>	12	AMMAR	91	CLEO $e^+ e^- \approx 10.5$ GeV
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$\Gamma(\bar{K}^0 K^+ \pi^- \text{ nonresonant})/\Gamma(K^- \pi^+)$   $\Gamma_{127}/\Gamma_{20}$

<u>VALUE</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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$0.10^{+0.06}_{-0.05}$	<sup>82</sup> ANJOS	91	E691 $\gamma$ Be 80–240 GeV
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<sup>82</sup> The factor 100 at the top of column 2 of Table I of ANJOS 91 should be omitted.

$\Gamma(K^+ K^- \pi^0)/\Gamma(K^- \pi^+ \pi^0)$					$\Gamma_{128}/\Gamma_{30}$
<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>0.0095 ± 0.0026</b>	151	ASNER	96B CLE2	$e^+ e^- \approx \gamma(4S)$	

$\Gamma(K_S^0 K_S^0 \pi^0)/\Gamma_{total}$					$\Gamma_{129}/\Gamma$
<u>VALUE</u>		<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>&lt;0.00059</b>		ASNER	96B CLE2	$e^+ e^- \approx \gamma(4S)$	

$\Gamma(\phi \pi^0)/\Gamma_{total}$					$\Gamma_{144}/\Gamma$
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>&lt;0.0014</b>	90	ALBRECHT	94i ARG	$e^+ e^- \approx 10$ GeV	

$\Gamma(\phi \eta)/\Gamma_{total}$					$\Gamma_{145}/\Gamma$
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>&lt;0.0028</b>	90	ALBRECHT	94i ARG	$e^+ e^- \approx 10$ GeV	

$\Gamma(\phi \omega)/\Gamma_{total}$					$\Gamma_{146}/\Gamma$
<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>&lt;0.0021</b>	90	ALBRECHT	94i ARG	$e^+ e^- \approx 10$ GeV	

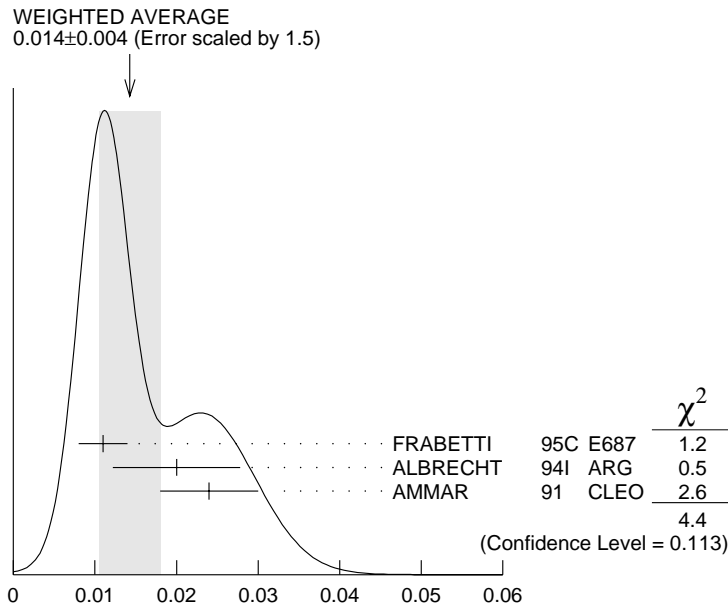
$\Gamma(K^+ K^- \pi^+ \pi^-)/\Gamma(K^- \pi^+ \pi^+ \pi^-)$					$\Gamma_{130}/\Gamma_{38}$
<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>0.0334 ± 0.0028 OUR AVERAGE</b>					
0.0313 ± 0.0037 ± 0.0036	136	AITALA	98D E791	$\pi^-$ nucleus, 500 GeV	
0.035 ± 0.004 ± 0.002	244	FRABETTI	95C E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV	
0.041 ± 0.007 ± 0.005	114	ALBRECHT	94i ARG	$e^+ e^- \approx 10$ GeV	
0.0314 ± 0.010	89	AMMAR	91 CLEO	$e^+ e^- \approx 10.5$ GeV	
0.028 <sup>+0.008</sup> / <sub>-0.007</sub>		ANJOS	91 E691	$\gamma$ Be 80–240 GeV	

$\Gamma(\phi \pi^+ \pi^-)/\Gamma(K^- \pi^+ \pi^+ \pi^-)$					$\Gamma_{147}/\Gamma_{38}$
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Unseen decay modes of the  $\phi$  are included.

<u>VALUE</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>	
<b>0.014 ± 0.004 OUR AVERAGE</b>		Error includes scale factor of 1.5. See the ideogram below.			
0.011 ± 0.003		FRABETTI	95C E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV	
0.020 ± 0.006 ± 0.005	28	ALBRECHT	94i ARG	$e^+ e^- \approx 10$ GeV	
0.024 ± 0.006	34	<sup>83</sup> AMMAR	91 CLEO	$e^+ e^- \approx 10.5$ GeV	
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
0.0076 <sup>+0.0066</sup> / <sub>-0.0049</sub>	3	ANJOS	91 E691	$\gamma$ Be 80–240 GeV	

<sup>83</sup>AMMAR 91 measures  $\phi \rho^0$ , but notes that  $\phi \rho^0$  dominates  $\phi \pi^+ \pi^-$ . We put the measurement here to keep from having more  $\phi \rho^0$  than  $\phi \pi^+ \pi^-$ .



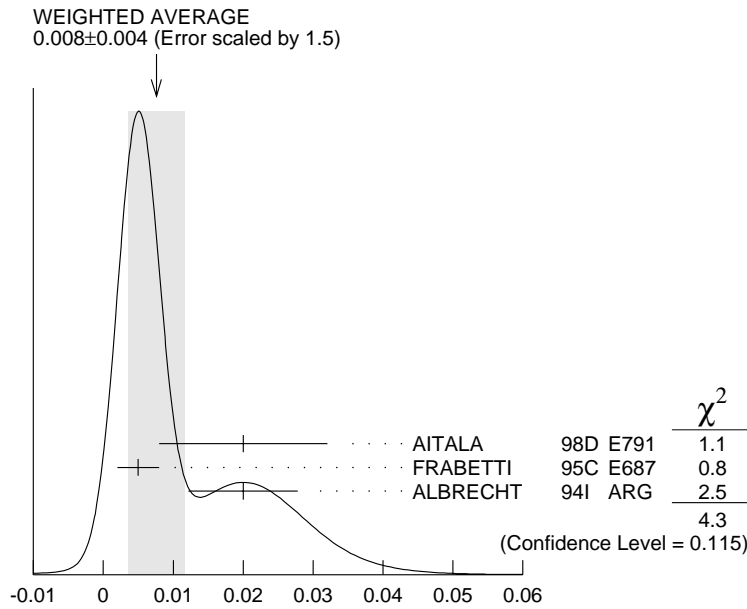
$$\Gamma(\phi \pi^+ \pi^-) / \Gamma(K^- \pi^+ \pi^+ \pi^-)$$

$$\Gamma(\phi \rho^0) / \Gamma(K^- \pi^+ \pi^+ \pi^-)$$

$$\Gamma_{148} / \Gamma_{38}$$

Unseen decay modes of the  $\phi$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>0.008 \pm 0.004</math> OUR AVERAGE</b>		Error includes scale factor of 1.5. See the ideogram below.		
$0.02 \pm 0.009 \pm 0.008$		AITALA	98D E791	$\pi^-$ nucleus, 500 GeV
$0.005 \pm 0.003$		FRABETTI	95C E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV
$0.020 \pm 0.006 \pm 0.005$	28	ALBRECHT	94I ARG	$e^+ e^- \approx 10$ GeV



$$\Gamma(\phi\rho^0)/\Gamma(K^-\pi^+\pi^+\pi^-)$$

$$\Gamma(\phi\pi^+\pi^-\text{3-body})/\Gamma(K^-\pi^+\pi^+\pi^-) \quad \Gamma_{149}/\Gamma_{38}$$

Unseen decay modes of the  $\phi$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.009±0.004±0.005</b>		AITALA	98D E791	$\pi^-$ nucleus, 500 GeV
••• We do not use the following data for averages, fits, limits, etc. •••				
<0.006	90	FRABETTI	95C E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV

$$\Gamma(K^+K^-\rho^0\text{3-body})/\Gamma(K^-\pi^+\pi^+\pi^-) \quad \Gamma_{133}/\Gamma_{38}$$

VALUE	DOCUMENT ID	TECN	COMMENT
<b>0.012 ±0.003</b>	FRABETTI	95C E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV

$$\Gamma(K^*(892)^0K^-\pi^+ + \text{c.c.})/\Gamma(K^-\pi^+\pi^+\pi^-) \quad \Gamma_{150}/\Gamma_{38}$$

Unseen decay modes of the  $K^*(892)^0$  are included.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.01</b>	90	<sup>84</sup> AITALA	98D E791	$\pi^-$ nucleus, 500 GeV
••• We do not use the following data for averages, fits, limits, etc. •••				
<0.017	90	<sup>84</sup> FRABETTI	95C E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV
0.010 <sup>+0.016</sup> <sub>-0.010</sub>		ANJOS	91 E691	$\gamma$ Be 80–240 GeV

<sup>84</sup> These upper limits are in conflict with values in the next two data blocks.

$\Gamma(K^*(892)^0 K^- \pi^+)/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{151}/\Gamma_{38}$

The  $K^{*0} K^- \pi^+$  and  $\bar{K}^{*0} K^+ \pi^-$  modes are distinguished by the charge of the pion in  $D^*(2010)^\pm \rightarrow D^0 \pi^\pm$  decays. Unseen decay modes of the  $K^*(892)^0$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.043 \pm 0.014 \pm 0.009$	55	<sup>85</sup> ALBRECHT	94I ARG	$e^+ e^- \approx 10$ GeV
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<sup>85</sup> This ALBRECHT 94I value is in conflict with upper limits given above.

$\Gamma(\bar{K}^*(892)^0 K^+ \pi^-)/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{152}/\Gamma_{38}$

The  $K^{*0} K^- \pi^+$  and  $\bar{K}^{*0} K^+ \pi^-$  modes are distinguished by the charge of the pion in  $D^*(2010)^\pm \rightarrow D^0 \pi^\pm$  decays. Unseen decay modes of the  $\bar{K}^*(892)^0$  are included.

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.023 \pm 0.013 \pm 0.009$	30	<sup>86</sup> ALBRECHT	94I ARG	$e^+ e^- \approx 10$ GeV
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<sup>86</sup> This ALBRECHT 94I value is in conflict with upper limits given above.

$\Gamma(K^*(892)^0 \bar{K}^*(892)^0)/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{153}/\Gamma_{38}$

Unseen decay modes of the  $K^*(892)^0$  and  $\bar{K}^*(892)^0$  are included.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.018 ± 0.007 OUR AVERAGE** Error includes scale factor of 1.2.

$0.016 \pm 0.006$			FRABETTI	95C E687	$\gamma$ Be, $\bar{E}_\gamma \approx 200$ GeV
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$0.036^{+0.020}_{-0.016}$	11		ANJOS	91 E691	$\gamma$ Be 80–240 GeV
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.02	90		AITALA	98D E791	$\pi^-$ nucleus, 500 GeV
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<0.033	90	<sup>87</sup>	AMMAR	91 CLEO	$e^+ e^- \approx 10.5$ GeV
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<sup>87</sup> A corrected value (G. Moneti, private communication).

$\Gamma(K^+ K^- \pi^+ \pi^- \text{ non-}\phi)/\Gamma_{\text{total}}$   $\Gamma_{136}/\Gamma$

VALUE	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.0017 \pm 0.0005$	<sup>88</sup> BARLAG	92C ACCM	$\pi^-$ Cu 230 GeV
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<sup>88</sup> BARLAG 92C computes the branching fraction using topological normalization.

$\Gamma(K^+ K^- \pi^+ \pi^- \text{ nonresonant})/\Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{137}/\Gamma_{38}$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
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**<0.011** 90 FRABETTI 95C E687  $\gamma$  Be,  $\bar{E}_\gamma \approx 200$  GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$0.001^{+0.011}_{-0.001}$		ANJOS	91 E691	$\gamma$ Be 80–240 GeV
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$\Gamma(K^0 \bar{K}^0 \pi^+ \pi^-)/\Gamma(\bar{K}^0 \pi^+ \pi^-)$   $\Gamma_{138}/\Gamma_{22}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
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**0.126 ± 0.038 ± 0.030** 25 ALBRECHT 94I ARG  $e^+ e^- \approx 10$  GeV

$\Gamma(K^+ K^- \pi^+ \pi^- \pi^0)/\Gamma_{\text{total}}$   $\Gamma_{139}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.0031 ± 0.0020</b>		<sup>89</sup> BARLAG	92C ACCM	$\pi^-$ Cu 230 GeV

<sup>89</sup> BARLAG 92C computes the branching fraction using topological normalization.

———— Radiative modes ————

$\Gamma(\rho^0 \gamma)/\Gamma_{\text{total}}$   $\Gamma_{154}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN
<b>&lt; 2.4 × 10<sup>-4</sup></b>	90	ASNER	98 CLE2

$\Gamma(\omega \gamma)/\Gamma_{\text{total}}$   $\Gamma_{155}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN
<b>&lt; 2.4 × 10<sup>-4</sup></b>	90	ASNER	98 CLE2

$\Gamma(\phi \gamma)/\Gamma_{\text{total}}$   $\Gamma_{156}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN
<b>&lt; 1.9 × 10<sup>-4</sup></b>	90	ASNER	98 CLE2

$\Gamma(\bar{K}^*(892)^0 \gamma)/\Gamma_{\text{total}}$   $\Gamma_{157}/\Gamma$

VALUE	CL%	DOCUMENT ID	TECN
<b>&lt; 7.6 × 10<sup>-4</sup></b>	90	ASNER	98 CLE2

———— Rare or forbidden modes ————

$\Gamma(K^+ \ell^- \bar{\nu}_\ell \text{ (via } \bar{D}^0))/\Gamma(K^- \ell^+ \nu_\ell)$   $\Gamma_{158}/\Gamma_8$

This is a  $D^0$ - $\bar{D}^0$  mixing limit without the complications of possible doubly-Cabibbo-suppressed decays that occur when using hadronic modes. For the limits on  $|m_{D_1^0} - m_{D_2^0}|$  and  $(\Gamma_{D_1^0} - \Gamma_{D_2^0})/\Gamma_{D^0}$  that come from the best mixing limit, see near the beginning of these  $D^0$  Listings.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt; 0.005</b>	90	<sup>90</sup> AITALA	96C E791	$\pi^-$ nucleus, 500 GeV

<sup>90</sup> AITALA 96C uses  $D^{*+} \rightarrow D^0 \pi^+$  (and charge conjugate) decays to identify the charm at production and  $D^0 \rightarrow K^- \ell^+ \nu_\ell$  (and charge conjugate) decays to identify the charm at decay.

$\Gamma(K^+ \pi^-)/\Gamma(K^- \pi^+)$   $\Gamma_{159}/\Gamma_{20}$

The  $D^0 \rightarrow K^+ \pi^-$  decay can occur directly by doubly Cabibbo-suppressed (DCS) decay, or indirectly by  $D^0 \rightarrow \bar{D}^0$  mixing followed by  $\bar{D}^0 \rightarrow K^+ \pi^-$  decay. The experiments here use the charge of the pion in  $D^*(2010)^\pm \rightarrow (D^0 \text{ or } \bar{D}^0) \pi^\pm$  decay to tell whether a  $D^0$  or a  $\bar{D}^0$  was born. Some of the experiments can use the decay-time information to disentangle the two modes. Here, we list the DCS branching ratio; in the next data block we give the limits on the mixing ratio.

Some early limits have been omitted from this Listing; see our 1998 (EPJ **C3** 1) edition.



VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.0038 ± 0.0008</b>					<b>OUR AVERAGE</b> Error includes scale factor of 1.1.
0.00332 <sup>+0.00063</sup> <sub>-0.00065</sub>	± 0.00040	45	<sup>91</sup> GODANG	00 CLE2	$e^+ e^-$
0.0068 <sup>+0.0034</sup> <sub>-0.0033</sub>	± 0.0007		<sup>92</sup> AITALA	98 E791	$\pi^-$ nucleus, 500 GeV
0.0184 ± 0.0059	± 0.0034	19	<sup>93</sup> BARATE	98W ALEP	$e^+ e^-$ at $Z^0$
0.0077 ± 0.0025	± 0.0025	19	<sup>94</sup> CINABRO	94 CLE2	$e^+ e^- \approx \Upsilon(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.011	90		<sup>94</sup> AMMAR	91 CLEO	$e^+ e^- \approx 10.5$ GeV
<0.015	90	1 ± 6	<sup>95</sup> ANJOS	88C E691	Photoproduc- tion
<0.014	90		<sup>94</sup> ALBRECHT	87K ARG	$e^+ e^-$ 10 GeV

<sup>91</sup> This GODANG 00 result assumes no  $D^0$ - $\bar{D}^0$  mixing; the DCS ratio becomes  $0.0048 \pm 0.0012 \pm 0.0004$  when mixing is allowed.

<sup>92</sup> This AITALA 98 result assumes no  $D^0$ - $\bar{D}^0$  mixing; the DCS ratio becomes  $0.0090^{+0.0120}_{-0.0109} \pm 0.0044$  when mixing is allowed.

<sup>93</sup> BARATE 98W gets  $0.0177^{+0.0060}_{-0.0056} \pm 0.0031$  for the DCS ratio when mixing is allowed, assuming no interference between the DCS and mixing amplitudes.

<sup>94</sup> CINABRO 94, AMMAR 91, and ALBRECHT 87K cannot distinguish between doubly Cabibbo-suppressed decay and  $D^0$ - $\bar{D}^0$  mixing.

<sup>95</sup> ANJOS 88C allows mixing but assumes no interference between the DCS and mixing amplitudes. When interference is allowed, the limit degrades to 0.049.

### $\Gamma(K^+ \pi^- \text{ (via } \bar{D}^0)) / \Gamma(K^- \pi^+)$

$\Gamma_{160} / \Gamma_{20}$

This is a  $D^0$ - $\bar{D}^0$  mixing limit. The experiments here (1) use the charge of the pion in  $D^*(2010)^\pm \rightarrow (D^0 \text{ or } \bar{D}^0) \pi^\pm$  decay to tell whether a  $D^0$  or a  $\bar{D}^0$  was born; and (2) use the decay-time distribution to disentangle doubly Cabibbo-suppressed decay and mixing. For the limits on  $|m_{D_1^0} - m_{D_2^0}|$  and  $(\Gamma_{D_1^0} - \Gamma_{D_2^0}) / \Gamma_{D_0}$  that come from the best mixing limit, see near the beginning of these  $D^0$  Listings.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.00041</b>	95		<sup>96</sup> GODANG	00 CLE2	$e^+ e^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.0092	95		<sup>97</sup> BARATE	98W ALEP	$e^+ e^-$ at $Z^0$
<0.005	90	1 ± 4	<sup>98</sup> ANJOS	88C E691	Photoproduction

<sup>96</sup> This GODANG 00 result assumes that the strong phase between  $D^0 \rightarrow K^+ \pi^-$  and  $\bar{D}^0 \rightarrow K^+ \pi^-$  is small, and limits only  $D^0 \rightarrow \bar{D}^0$  transitions via off-shell intermediate states. The limit on transitions via on-shell intermediate states is 0.0017.

<sup>97</sup> This BARATE 98W result assumes no interference between the DCS and mixing amplitudes. When interference is allowed, the limit degrades to 0.036 (95%CL).

<sup>98</sup> This ANJOS 88C result assumes no interference between the DCS and mixing amplitudes. When interference is allowed, the limit degrades to 0.019. Combined with results on  $K^\pm \pi^\mp \pi^+ \pi^-$ , the limit is, assuming no interference, 0.0037.

$\Gamma(K^+ \pi^- \pi^+ \pi^-) / \Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{161} / \Gamma_{38}$

The  $D^0 \rightarrow K^+ \pi^- \pi^+ \pi^-$  decay can occur directly by doubly Cabibbo-suppressed (DCS) decay, or indirectly by  $D^0 \rightarrow \bar{D}^0$  mixing followed by  $\bar{D}^0 \rightarrow K^+ \pi^- \pi^+ \pi^-$  decay. The experiments here use the charge of the pion in  $D^*(2010)^\pm \rightarrow (D^0 \text{ or } \bar{D}^0) \pi^\pm$  decay to tell whether a  $D^0$  or a  $\bar{D}^0$  was born. Some of the experiments can use the decay-time information to disentangle the two modes. Here, we list the DCS branching ratio; in the next data block we give the limits on the mixing ratio.

Some early limits have been omitted from this Listing; see our 1998 (EPJ **C3** 1) edition.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$0.0025^{+0.0036}_{-0.0034} \pm 0.0003$			<sup>99</sup> AITALA	98 E791	$\pi^-$ nucleus, 500 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

<0.018	90		<sup>100</sup> AMMAR	91 CLEO	$e^+ e^- \approx 10.5$ GeV
<0.018	90	$5 \pm 12$	<sup>101</sup> ANJOS	88C E691	Photoproduction

<sup>99</sup> AITALA 98 uses the charge of the pion in  $D^{*\pm} \rightarrow (D^0 \text{ or } \bar{D}^0) \pi^\pm$  to tell whether a  $D^0$  or a  $\bar{D}^0$  was born. This result assumes no  $D^0$ - $\bar{D}^0$  mixing; it becomes  $-0.0020^{+0.0117}_{-0.0106} \pm 0.0035$  when mixing is allowed and decay-time information is used to distinguish doubly Cabibbo-suppressed decays from mixing.

<sup>100</sup> AMMAR 91 cannot distinguish between doubly Cabibbo-suppressed decay and  $D^0$ - $\bar{D}^0$  mixing.

<sup>101</sup> ANJOS 88C uses decay-time information to distinguish doubly Cabibbo-suppressed (DCS) decays from  $D^0$ - $\bar{D}^0$  mixing. However, the result assumes no interference between the DCS and mixing amplitudes. When interference is allowed, the limit degrades to 0.033.

$\Gamma(K^+ \pi^- \pi^+ \pi^- \text{ (via } \bar{D}^0)) / \Gamma(K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{162} / \Gamma_{38}$

This is a  $D^0$ - $\bar{D}^0$  mixing limit. The experiments here (1) use the charge of the pion in  $D^*(2010)^\pm \rightarrow (D^0 \text{ or } \bar{D}^0) \pi^\pm$  decay to tell whether a  $D^0$  or a  $\bar{D}^0$  was born; and (2) use the decay-time distribution to disentangle doubly Cabibbo-suppressed decay and mixing. For the limits on  $|m_{D_1^0} - m_{D_2^0}|$  and  $(\Gamma_{D_1^0} - \Gamma_{D_2^0}) / \Gamma_{D^0}$  that come from the best mixing limit, see near the beginning of these  $D^0$  Listings.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<0.005	90	$0 \pm 4$	<sup>102</sup> ANJOS	88C E691	Photoproduction

<sup>102</sup> ANJOS 88C uses decay-time information to distinguish doubly Cabibbo-suppressed (DCS) decays from  $D^0$ - $\bar{D}^0$  mixing. However, the result assumes no interference between the DCS and mixing amplitudes. When interference is allowed, the limit degrades to 0.007. Combined with results on  $K^\pm \pi^\mp$ , the limit is, assuming no interference, 0.0037.

$\Gamma(K^+ \pi^- \text{ or } K^+ \pi^- \pi^+ \pi^- \text{ (via } \bar{D}^0)) / \Gamma(K^- \pi^+ \text{ or } K^- \pi^+ \pi^+ \pi^-)$   $\Gamma_{163} / \Gamma_0$

This is a  $D^0$ - $\bar{D}^0$  mixing limit. For the limits on  $|m_{D_1^0} - m_{D_2^0}|$  and  $(\Gamma_{D_1^0} - \Gamma_{D_2^0}) / \Gamma_{D^0}$  that come from the best mixing limit, see near the beginning of these  $D^0$  Listings.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<0.0085	90	<sup>103</sup> AITALA	98 E791	$\pi^-$ nucleus, 500 GeV

<sup>103</sup> AITALA 98 uses decay-time information to distinguish doubly Cabibbo-suppressed decays from  $D^0$ - $\bar{D}^0$  mixing. The fit allows interference between the two amplitudes, and also allows  $CP$  violation in this term. The central value obtained is  $0.0039^{+0.0036}_{-0.0032} \pm 0.0016$ . When interference is disallowed, the result becomes  $0.0021 \pm 0.0009 \pm 0.0002$ .

### $\Gamma(\mu^- \text{ anything (via } \bar{D}^0))/\Gamma(\mu^+ \text{ anything})$

$\Gamma_{164}/\Gamma_2$

This is a  $D^0-\bar{D}^0$  mixing limit. See the somewhat better limits above.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>&lt;0.0056</b>	90	LOUIS	86 SPEC	$\pi^-$ W 225 GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
<0.012	90	BENVENUTI	85 CNTR	$\mu$ C, 200 GeV
<0.044	90	BODEK	82 SPEC	$\pi^-$ , pFe $\rightarrow$ $D^0$

### $\Gamma(e^+ e^-)/\Gamma_{\text{total}}$

$\Gamma_{165}/\Gamma$

A test for the  $\Delta C = 1$  weak neutral current. Allowed by first-order weak interaction combined with electromagnetic interaction.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;6.2 <math>\times 10^{-6}</math></b>	90		AITALA	99G E791	$\pi^-$ N 500 GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
<8.19 $\times 10^{-6}$	90		PRIPSTEIN	00 E789	p nucleus, 800 GeV
<1.3 $\times 10^{-5}$	90	0	FREYBERGER	96 CLE2	$e^+ e^- \approx \gamma(4S)$
<1.3 $\times 10^{-4}$	90		ADLER	88 MRK3	$e^+ e^-$ 3.77 GeV
<1.7 $\times 10^{-4}$	90	7	ALBRECHT	88G ARG	$e^+ e^-$ 10 GeV
<2.2 $\times 10^{-4}$	90	8	HAAS	88 CLEO	$e^+ e^-$ 10 GeV

### $\Gamma(\mu^+ \mu^-)/\Gamma_{\text{total}}$

$\Gamma_{166}/\Gamma$

A test for the  $\Delta C = 1$  weak neutral current. Allowed by first-order weak interaction combined with electromagnetic interaction.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.1 <math>\times 10^{-6}</math></b>	90		ADAMOVICH	97 BEAT	$\pi^-$ Cu, W 350 GeV
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
<1.56 $\times 10^{-5}$	90		PRIPSTEIN	00 E789	p nucleus, 800 GeV
<5.2 $\times 10^{-6}$	90		AITALA	99G E791	$\pi^-$ N 500 GeV
<4.2 $\times 10^{-6}$	90		ALEXOPOU...	96 E771	p Si, 800 GeV
<3.4 $\times 10^{-5}$	90	1	FREYBERGER	96 CLE2	$e^+ e^- \approx \gamma(4S)$
<7.6 $\times 10^{-6}$	90	0	ADAMOVICH	95 BEAT	See ADAMOVICH 97
<4.4 $\times 10^{-5}$	90	0	KODAMA	95 E653	$\pi^-$ emulsion 600 GeV
<3.1 $\times 10^{-5}$	90	104	MISHRA	94 E789	$-4.1 \pm 4.8$ events
<7.0 $\times 10^{-5}$	90	3	ALBRECHT	88G ARG	$e^+ e^-$ 10 GeV
<1.1 $\times 10^{-5}$	90		LOUIS	86 SPEC	$\pi^-$ W 225 GeV
<3.4 $\times 10^{-4}$	90		AUBERT	85 EMC	Deep inelast. $\mu^-$ N

<sup>104</sup> Here MISHRA 94 uses "the statistical approach advocated by the PDG." For an alternate approach, giving a limit of  $9 \times 10^{-6}$  at 90% confidence level, see the paper.

### $\Gamma(\pi^0 e^+ e^-)/\Gamma_{\text{total}}$

$\Gamma_{167}/\Gamma$

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>&lt;4.5 <math>\times 10^{-5}</math></b>	90	0	FREYBERGER	96 CLE2	$e^+ e^- \approx \gamma(4S)$

**$\Gamma(\pi^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$**   **$\Gamma_{168}/\Gamma$**

A test for the  $\Delta C=1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.8 \times 10^{-4}</math></b>	90	2	KODAMA	95 E653	$\pi^-$ emulsion 600 GeV
••• We do not use the following data for averages, fits, limits, etc. •••					
$<5.4 \times 10^{-4}$	90	3	FREYBERGER	96 CLE2	$e^+ e^- \approx \gamma(4S)$

**$\Gamma(\eta e^+ e^-)/\Gamma_{\text{total}}$**   **$\Gamma_{169}/\Gamma$**

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.1 \times 10^{-4}</math></b>	90	0	FREYBERGER	96 CLE2	$e^+ e^- \approx \gamma(4S)$

**$\Gamma(\eta \mu^+ \mu^-)/\Gamma_{\text{total}}$**   **$\Gamma_{170}/\Gamma$**

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;5.3 \times 10^{-4}</math></b>	90	0	FREYBERGER	96 CLE2	$e^+ e^- \approx \gamma(4S)$

**$\Gamma(\rho^0 e^+ e^-)/\Gamma_{\text{total}}$**   **$\Gamma_{171}/\Gamma$**

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.0 \times 10^{-4}</math></b>	90	2	<sup>105</sup> FREYBERGER	96 CLE2	$e^+ e^- \approx \gamma(4S)$

••• We do not use the following data for averages, fits, limits, etc. •••

$<4.5 \times 10^{-4}$  90 2 HAAS 88 CLEO  $e^+ e^-$  10 GeV

<sup>105</sup> This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 1.8 \times 10^{-4}$  using a photon pole amplitude model.

**$\Gamma(\rho^0 \mu^+ \mu^-)/\Gamma_{\text{total}}$**   **$\Gamma_{172}/\Gamma$**

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;2.3 \times 10^{-4}</math></b>	90	0	KODAMA	95 E653	$\pi^-$ emulsion 600 GeV

••• We do not use the following data for averages, fits, limits, etc. •••

$<4.9 \times 10^{-4}$  90 1 <sup>106</sup> FREYBERGER 96 CLE2  $e^+ e^- \approx \gamma(4S)$

$<8.1 \times 10^{-4}$  90 5 HAAS 88 CLEO  $e^+ e^-$  10 GeV

<sup>106</sup> This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 4.5 \times 10^{-4}$  using a photon pole amplitude model.

**$\Gamma(\omega e^+ e^-)/\Gamma_{\text{total}}$**   **$\Gamma_{173}/\Gamma$**

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.8 \times 10^{-4}</math></b>	90	1	<sup>107</sup> FREYBERGER	96 CLE2	$e^+ e^- \approx \gamma(4S)$

<sup>107</sup> This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 2.7 \times 10^{-4}$  using a photon pole amplitude model.

$\Gamma(\omega\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{174}/\Gamma$

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<8.3 \times 10^{-4}$	90	0	108 FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$

<sup>108</sup>This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 6.5 \times 10^{-4}$  using a photon pole amplitude model.

$\Gamma(\phi e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{175}/\Gamma$

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<5.2 \times 10^{-5}$	90	2	109 FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$

<sup>109</sup>This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 7.6 \times 10^{-5}$  using a photon pole amplitude model.

$\Gamma(\phi\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{176}/\Gamma$

A test for the  $\Delta C = 1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<4.1 \times 10^{-4}$	90	0	110 FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$

<sup>110</sup>This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 2.4 \times 10^{-4}$  using a photon pole amplitude model.

$\Gamma(\bar{K}^0 e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{177}/\Gamma$

Allowed by first-order weak interaction combined with electromagnetic interaction.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<1.1 \times 10^{-4}$	90	0	FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<1.7 \times 10^{-3}$	90		ADLER	89C MRK3	$e^+e^-$ 3.77 GeV
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$\Gamma(\bar{K}^0\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{178}/\Gamma$

Allowed by first-order weak interaction combined with electromagnetic interaction.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<2.6 \times 10^{-4}$	90	2	KODAMA	95 E653	$\pi^-$ emulsion 600 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$<6.7 \times 10^{-4}$	90	1	FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$
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$\Gamma(\bar{K}^*(892)^0 e^+e^-)/\Gamma_{\text{total}}$   $\Gamma_{179}/\Gamma$

Allowed by first-order weak interaction combined with electromagnetic interaction.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<1.4 \times 10^{-4}$	90	1	111 FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$

<sup>111</sup>This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 2.0 \times 10^{-4}$  using a photon pole amplitude model.

$\Gamma(\bar{K}^*(892)^0\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{180}/\Gamma$

Allowed by first-order weak interaction combined with electromagnetic interaction.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$<1.18 \times 10^{-3}$	90	1	112 FREYBERGER 96	CLE2	$e^+e^- \approx \gamma(4S)$

<sup>112</sup>This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 1.0 \times 10^{-3}$  using a photon pole amplitude model.

$\Gamma(\pi^+\pi^-\pi^0\mu^+\mu^-)/\Gamma_{\text{total}}$   $\Gamma_{181}/\Gamma$

A test for the  $\Delta C=1$  weak neutral current. Allowed by higher-order electroweak interactions.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$< 8.1 \times 10^{-4}$	90	1	KODAMA	95 E653	$\pi^-$ emulsion 600 GeV

$\Gamma(\mu^\pm e^\mp)/\Gamma_{\text{total}}$   $\Gamma_{182}/\Gamma$

A test of lepton family number conservation.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$< 8.1 \times 10^{-6}$	90		AITALA	99G E791	$\pi^- N$ 500 GeV

• • • We do not use the following data for averages, fits, limits, etc. • • •

$< 1.72 \times 10^{-5}$	90		PRIPSTEIN	00 E789	$p$ nucleus, 800 GeV
$< 1.9 \times 10^{-5}$	90	2	<sup>113</sup> FREYBERGER	96 CLE2	$e^+e^- \approx \gamma(4S)$
$< 1.0 \times 10^{-4}$	90	4	ALBRECHT	88G ARG	$e^+e^-$ 10 GeV
$< 2.7 \times 10^{-4}$	90	9	HAAS	88 CLEO	$e^+e^-$ 10 GeV
$< 1.2 \times 10^{-4}$	90		BECKER	87C MRK3	$e^+e^-$ 3.77 GeV
$< 9 \times 10^{-4}$	90		PALKA	87 SILI	200 GeV $\pi p$
$< 21 \times 10^{-4}$	90	0	<sup>114</sup> RILES	87 MRK2	$e^+e^-$ 29 GeV

<sup>113</sup>This is the corrected result given in the erratum to FREYBERGER 96.

<sup>114</sup>RILES 87 assumes  $B(D \rightarrow K\pi) = 3.0\%$  and has production model dependency.

$\Gamma(\pi^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{183}/\Gamma$

A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$< 8.6 \times 10^{-5}$	90	2	FREYBERGER	96 CLE2	$e^+e^- \approx \gamma(4S)$

$\Gamma(\eta e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{184}/\Gamma$

A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$< 1.0 \times 10^{-4}$	90	0	FREYBERGER	96 CLE2	$e^+e^- \approx \gamma(4S)$

$\Gamma(\rho^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{185}/\Gamma$

A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$< 4.9 \times 10^{-5}$	90	0	<sup>115</sup> FREYBERGER	96 CLE2	$e^+e^- \approx \gamma(4S)$

<sup>115</sup>This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 5.0 \times 10^{-5}$  using a photon pole amplitude model.

$\Gamma(\omega e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{186}/\Gamma$

A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
$< 1.2 \times 10^{-4}$	90	0	<sup>116</sup> FREYBERGER	96 CLE2	$e^+e^- \approx \gamma(4S)$

<sup>116</sup>This FREYBERGER 96 limit is obtained using a phase-space model. The same limit is obtained using a photon pole amplitude model.

**$\Gamma(\phi e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{187}/\Gamma$**

A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;3.4 \times 10^{-5}</math></b>	90	0	117 FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

117 This FREYBERGER 96 limit is obtained using a phase-space model. The limit changes to  $< 3.3 \times 10^{-5}$  using a photon pole amplitude model.

**$\Gamma(\bar{K}^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{188}/\Gamma$**

A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.0 \times 10^{-4}</math></b>	90	0	FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

**$\Gamma(\bar{K}^*(892)^0 e^\pm \mu^\mp)/\Gamma_{\text{total}}$   $\Gamma_{189}/\Gamma$**

A test of lepton family number conservation. The value is for the sum of the two charge states.

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>&lt;1.0 \times 10^{-4}</math></b>	90	0	118 FREYBERGER 96	CLE2	$e^+ e^- \approx \gamma(4S)$

118 This FREYBERGER 96 limit is obtained using a phase-space model. The same limit is obtained using a photon pole amplitude model.

**$D^0$  CP-VIOLATING DECAY-RATE ASYMMETRIES**

**$A_{CP}(K^+ K^-)$  in  $D^0, \bar{D}^0 \rightarrow K^+ K^-$**

This is the difference between  $D^0$  and  $\bar{D}^0$  partial widths for these modes divided by the sum of the widths. The  $D^0$  and  $\bar{D}^0$  are distinguished by the charge of the parent  $D^*$ :  $D^{*+} \rightarrow D^0 \pi^+$  and  $D^{*-} \rightarrow D^0 \pi^-$ .

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.009 ± 0.021 OUR AVERAGE</b>					
-0.001 ± 0.022 ± 0.015		3330	119 LINK	00B FOCS	
-0.010 ± 0.049 ± 0.012		609	119 AITALA	98C E791	-0.093 $< A_{CP} < +0.073$ (90% CL)
+0.080 ± 0.061			BARTELT	95 CLE2	-0.022 $< A_{CP} < +0.18$ (90%CL)
+0.024 ± 0.084			119 FRABETTI	94I E687	-0.11 $< A_{CP} < +0.16$ (90% CL)

119 FRABETTI 94I, AITALA 98C, and LINK 00B measure  $N(D^0 \rightarrow K^+ K^-)/N(D^0 \rightarrow K^- \pi^+)$ , the ratio of numbers of events observed, and similarly for the  $\bar{D}^0$ .

**$A_{CP}(\pi^+ \pi^-)$  in  $D^0, \bar{D}^0 \rightarrow \pi^+ \pi^-$**

This is the difference between  $D^0$  and  $\bar{D}^0$  partial widths for these modes divided by the sum of the widths. The  $D^0$  and  $\bar{D}^0$  are distinguished by the charge of the parent  $D^*$ :  $D^{*+} \rightarrow D^0 \pi^+$  and  $D^{*-} \rightarrow D^0 \pi^-$ .

VALUE	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.03 ± 0.04 OUR AVERAGE</b>					
+0.048 ± 0.039 ± 0.025		1177	120 LINK	00B FOCS	
-0.049 ± 0.078 ± 0.030		343	120 AITALA	98C E791	-0.186 $< A_{CP} < +0.088$ (90% CL)

120 AITALA 98C and LINK 00B measure  $N(D^0 \rightarrow \pi^+ \pi^-)/N(D^0 \rightarrow K^- \pi^+)$ , the ratio of numbers of events observed, and similarly for the  $\bar{D}^0$ .

### $A_{CP}(K_S^0 \phi)$ in $D^0, \bar{D}^0 \rightarrow K_S^0 \phi$

This is the difference between  $D^0$  and  $\bar{D}^0$  partial widths for these modes divided by the sum of the widths. The  $D^0$  and  $\bar{D}^0$  are distinguished by the charge of the parent  $D^*$ :  $D^{*+} \rightarrow D^0 \pi^+$  and  $D^{*-} \rightarrow \bar{D}^0 \pi^-$ .

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.028 \pm 0.094$	BARTELT	95 CLE2	$-0.182 < A_{CP} < +0.126$ (90%CL)

### $A_{CP}(K_S^0 \pi^0)$ in $D^0, \bar{D}^0 \rightarrow K_S^0 \pi^0$

This is the difference between  $D^0$  and  $\bar{D}^0$  partial widths for these modes divided by the sum of the widths. The  $D^0$  and  $\bar{D}^0$  are distinguished by the charge of the parent  $D^*$ :  $D^{*+} \rightarrow D^0 \pi^+$  and  $D^{*-} \rightarrow \bar{D}^0 \pi^-$ .

VALUE	DOCUMENT ID	TECN	COMMENT
$-0.018 \pm 0.030$	BARTELT	95 CLE2	$-0.067 < A_{CP} < +0.031$ (90%CL)

### $A_{CP}(K^\pm \pi^\mp)$ in $D^0 \rightarrow K^+ \pi^-, \bar{D}^0 \rightarrow K^- \pi^+$

This is the difference between  $D^0$  and  $\bar{D}^0$  partial widths for these modes divided by the sum of the widths. The  $D^0$  and  $\bar{D}^0$  are distinguished by the charge of the parent  $D^*$ :  $D^{*+} \rightarrow D^0 \pi^+$  and  $D^{*-} \rightarrow \bar{D}^0 \pi^-$ .

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
$+0.02^{+0.19}_{-0.20} \pm 0.01$	45 121	GODANG	00 CLE2	$-0.43 < A_{CP} < +0.34$ (95%CL)

<sup>121</sup>This GODANG 00 result assumes no  $D^0$ - $\bar{D}^0$  mixing; it becomes  $-0.01^{+0.16}_{-0.17} \pm 0.01$  when mixing is allowed.

## $D^0$ PRODUCTION CROSS SECTION AT $\psi(3770)$

A compilation of the cross sections for the direct production of  $D^0$  mesons at or near the  $\psi(3770)$  peak in  $e^+ e^-$  production.

VALUE (nanobarns)	DOCUMENT ID	TECN	COMMENT
$5.8 \pm 0.5 \pm 0.6$	122 ADLER	88C MRK3	$e^+ e^-$ 3.768 GeV
$7.3 \pm 1.3$	123 PARTRIDGE	84 CBAL	$e^+ e^-$ 3.771 GeV
$8.00 \pm 0.95 \pm 1.21$	124 SCHINDLER	80 MRK2	$e^+ e^-$ 3.771 GeV
$11.5 \pm 2.5$	125 PERUZZI	77 MRK1	$e^+ e^-$ 3.774 GeV

<sup>122</sup>This measurement compares events with one detected  $D$  to those with two detected  $D$  mesons, to determine the the absolute cross section. ADLER 88C find the ratio of cross sections (neutral to charged) to be  $1.36 \pm 0.23 \pm 0.14$ .

<sup>123</sup>This measurement comes from a scan of the  $\psi(3770)$  resonance and a fit to the cross section. PARTRIDGE 84 measures  $6.4 \pm 1.15$  nb for the cross section. We take the phase space division of neutral and charged  $D$  mesons in  $\psi(3770)$  decay to be 1.33, and we assume that the  $\psi(3770)$  is an isosinglet to evaluate the cross sections. The noncharm decays (e.g. radiative) of the  $\psi(3770)$  are included in this measurement and may amount to a few percent correction.

<sup>124</sup>This measurement comes from a scan of the  $\psi(3770)$  resonance and a fit to the cross section. SCHINDLER 80 assume the phase space division of neutral and charged  $D$  mesons in  $\psi(3770)$  decay to be 1.33, and that the  $\psi(3770)$  is an isosinglet. The noncharm decays (e.g. radiative) of the  $\psi(3770)$  are included in this measurement and may amount to a few percent correction.

<sup>125</sup>This measurement comes from a scan of the  $\psi(3770)$  resonance and a fit to the cross section. The phase space division of neutral and charged  $D$  mesons in  $\psi(3770)$  decay



is taken to be 1.33, and  $\psi(3770)$  is assumed to be an isosinglet. The noncharm decays (e.g. radiative) of the  $\psi(3770)$  are included in this measurement and may amount to a few percent correction. We exclude this measurement from the average because of uncertainties in the contamination from  $\tau$  lepton pairs. Also see RAPIDIS 77.

## $D^0$ REFERENCES

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JUN	00	PRL 84 1857	S.Y. Jun <i>et al.</i>	(FNAL SELEX Collab.)
LINK	00	PL B485 62	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
LINK	00B	PL B491 232	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
Also	00D	PL B495 443 (errata)	J.M. Link <i>et al.</i>	(FNAL FOCUS Collab.)
PDG	00	EPJ C15 1	D.E. Groom <i>et al.</i>	
PRIPSTEIN	00	PR D61 032005	D. Pripstein <i>et al.</i>	(FNAL E789 Collab.)
AITALA	99E	PRL 83 32	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
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BONVICINI	99	PRL 82 4586	G. Bonvicini <i>et al.</i>	(CLEO Collab.)
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AITALA	98C	PL B421 405	E.M. Aitala <i>et al.</i>	(FNAL E791 Collab.)
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ALBRECHT	96C	PL B374 249	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
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KUBOTA	96B	PR D54 2994	Y. Kubota <i>et al.</i>	(CLEO Collab.)
ADAMOVICH	95	PL B353 563	M.I. Adamovich <i>et al.</i>	(CERN BEATRICE Collab.)
BARTELT	95	PR D52 4860	J.E. Bartelt <i>et al.</i>	(CLEO Collab.)
BUTLER	95	PR D52 2656	F. Butler <i>et al.</i>	(CLEO Collab.)
FRABETTI	95C	PL B354 486	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRABETTI	95G	PL B364 127	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
KODAMA	95	PL B345 85	K. Kodama <i>et al.</i>	(FNAL E653 Collab.)
ALBRECHT	94	PL B324 249	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	94F	PL B340 125	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	94I	ZPHY C64 375	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
CINABRO	94	PRL 72 1406	D. Cinabro <i>et al.</i>	(CLEO Collab.)
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FRABETTI	94G	PL B331 217	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
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ALBRECHT	93D	PL B308 435	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
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FRABETTI	93I	PL B315 203	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
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ALBRECHT	92P	ZPHY C56 7	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ANJOS	92B	PR D46 R1	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
ANJOS	92C	PR D46 1941	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
BARLAG	92C	ZPHY C55 383	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
Also	90D	ZPHY C48 29	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
COFFMAN	92B	PR D45 2196	D.M. Coffman <i>et al.</i>	(Mark III Collab.)
Also	90	PRL 64 2615	J. Adler <i>et al.</i>	(Mark III Collab.)

FRABETTI	92	PL B281 167	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
FRABETTI	92B	PL B286 195	P.L. Frabetti <i>et al.</i>	(FNAL E687 Collab.)
ALVAREZ	91B	ZPHY C50 11	M.P. Alvarez <i>et al.</i>	(CERN NA14/2 Collab.)
AMMAR	91	PR D44 3383	R. Ammar <i>et al.</i>	(CLEO Collab.)
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COFFMAN	91	PL B263 135	D.M. Coffman <i>et al.</i>	(Mark III Collab.)
CRAWFORD	91B	PR D44 3394	G. Crawford <i>et al.</i>	(CLEO Collab.)
DECAMP	91J	PL B266 218	D. Decamp <i>et al.</i>	(ALEPH Collab.)
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KINOSHITA	91	PR D43 2836	K. Kinoshita <i>et al.</i>	(CLEO Collab.)
KODAMA	91	PRL 66 1819	K. Kodama <i>et al.</i>	(FNAL E653 Collab.)
ALBRECHT	90C	ZPHY C46 9	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALEXANDER	90	PRL 65 1184	J. Alexander <i>et al.</i>	(CLEO Collab.)
ALEXANDER	90B	PRL 65 1531	J. Alexander <i>et al.</i>	(CLEO Collab.)
ALVAREZ	90	ZPHY C47 539	M.P. Alvarez <i>et al.</i>	(CERN NA14/2 Collab.)
ANJOS	90D	PR D42 2414	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
BARLAG	90C	ZPHY C46 563	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
ADLER	89	PRL 62 1821	J. Adler <i>et al.</i>	(Mark III Collab.)
ADLER	89C	PR D40 906	J. Adler <i>et al.</i>	(Mark III Collab.)
ALBRECHT	89D	ZPHY C43 181	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ANJOS	89F	PRL 62 1587	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
ABACHI	88	PL B205 411	S. Abachi <i>et al.</i>	(HRS Collab.)
ADLER	88	PR D37 2023	J. Adler <i>et al.</i>	(Mark III Collab.)
ADLER	88C	PRL 60 89	J. Adler <i>et al.</i>	(Mark III Collab.)
ALBRECHT	88G	PL B209 380	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	88I	PL B210 267	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
AMENDOLIA	88	EPL 5 407	S.R. Amendolia <i>et al.</i>	(NA1 Collab.)
ANJOS	88C	PRL 60 1239	J.C. Anjos <i>et al.</i>	(FNAL E691 Collab.)
BORTOLETTO	88	PR D37 1719	D. Bortoletto <i>et al.</i>	(CLEO Collab.)
Also	89D	PR D39 1471 erratum	D. Bortoletto <i>et al.</i>	(CLEO Collab.)
CUMALAT	88	PL B210 253	J.P. Cumalat <i>et al.</i>	(E-400 Collab.)
HAAS	88	PRL 60 1614	P. Haas <i>et al.</i>	(CLEO Collab.)
RAAB	88	PR D37 2391	J.R. Raab <i>et al.</i>	(FNAL E691 Collab.)
ADAMOVICH	87	EPL 4 887	M.I. Adamovich <i>et al.</i>	
ADLER	87	PL B196 107	J. Adler <i>et al.</i>	(Mark III Collab.)
AGUILAR-...	87D	PL B193 140	M. Aguilar-Benitez <i>et al.</i>	(LEBC-EHS Collab.)
Also	88B	ZPHY C40 321	M. Aguilar-Benitez <i>et al.</i>	(LEBC-EHS Collab.)
AGUILAR-...	87E	ZPHY C36 551	M. Aguilar-Benitez <i>et al.</i>	(LEBC-EHS Collab.)
Also	88B	ZPHY C40 321	M. Aguilar-Benitez <i>et al.</i>	(LEBC-EHS Collab.)
AGUILAR-...	87F	ZPHY C36 559	M. Aguilar-Benitez <i>et al.</i>	(LEBC-EHS Collab.)
Also	88	ZPHY C38 520 erratum	M. Aguilar-Benitez <i>et al.</i>	(LEBC-EHS Collab.)
ALBRECHT	87E	ZPHY C33 359	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	87K	PL B199 447	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
BARLAG	87B	ZPHY C37 17	S. Barlag <i>et al.</i>	(ACCMOR Collab.)
BECKER	87C	PL B193 147	J.J. Becker <i>et al.</i>	(Mark III Collab.)
Also	87D	PL B198 590 erratum	J.J. Becker <i>et al.</i>	(Mark III Collab.)
CSORNA	87	PL B191 318	S.E. Csorna <i>et al.</i>	(CLEO Collab.)
PALKA	87	PL B189 238	H. Palka <i>et al.</i>	(ACCMOR Collab.)
RILES	87	PR D35 2914	K. Riles <i>et al.</i>	(Mark II Collab.)
ABE	86	PR D33 1	K. Abe <i>et al.</i>	
BAILEY	86	ZPHY C30 51	R. Bailey <i>et al.</i>	(ACCMOR Collab.)
BEBEK	86	PRL 56 1893	C. Bebek <i>et al.</i>	(CLEO Collab.)
GLADNEY	86	PR D34 2601	L. Gladney <i>et al.</i>	(Mark II Collab.)
LOUIS	86	PRL 56 1027	W.C. Louis <i>et al.</i>	(PRIN, CHIC, ISU)
USHIDA	86B	PRL 56 1771	N. Ushida <i>et al.</i>	(AICH, FNAL, KOBE, SEOU+)
ALBRECHT	85B	PL 158B 525	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
ALBRECHT	85F	PL 150B 235	H. Albrecht <i>et al.</i>	(ARGUS Collab.)
AUBERT	85	PL 155B 461	J.J. Aubert <i>et al.</i>	(EMC Collab.)
BAILEY	85	ZPHY C28 357	R. Bailey <i>et al.</i>	(ABCCMR Collab.)
BALTRUSAIT-...	85B	PRL 54 1976	R.M. Baltrusaitis <i>et al.</i>	(Mark III Collab.)
BALTRUSAIT-...	85E	PRL 55 150	R.M. Baltrusaitis <i>et al.</i>	(Mark III Collab.)
BENVENUTI	85	PL 158B 531	A.C. Benvenuti <i>et al.</i>	(BCDMS Collab.)
ADAMOVICH	84B	PL 140B 123	M.I. Adamovich <i>et al.</i>	(CERN WA58 Collab.)
DERRICK	84	PRL 53 1971	M. Derrick <i>et al.</i>	(HRS Collab.)
PARTRIDGE	84	Thesis CALT-68-1150	R.A. Partridge	(Crystal Ball Collab.)
SUMMERS	84	PRL 52 410	D.J. Summers <i>et al.</i>	(UCSB, CARL, COLO+)

BAILEY	83B	PL 132B 237	R. Bailey <i>et al.</i>	(ACCMOR Collab.)
BODEK	82	PL 113B 82	A. Bodek <i>et al.</i>	(ROCH, CIT, CHIC, FNAL+)
FIORINO	81	LNC 30 166	A. Fiorino <i>et al.</i>	
SCHINDLER	81	PR D24 78	R.H. Schindler <i>et al.</i>	(Mark II Collab.)
TRILLING	81	PRPL 75 57	G.H. Trilling	(LBL, UCB) J
ASTON	80E	PL 94B 113	D. Aston <i>et al.</i>	(BONN, CERN, EPOL, GLAS+)
AVERY	80	PRL 44 1309	P. Avery <i>et al.</i>	(ILL, FNAL, COLU)
SCHINDLER	80	PR D21 2716	R.H. Schindler <i>et al.</i>	(Mark II Collab.)
ZHOLENTZ	80	PL 96B 214	A.A. Zholents <i>et al.</i>	(NOVO)
Also	81	SJNP 34 814	A.A. Zholents <i>et al.</i>	(NOVO)
		Translated from YAF 34	1471.	
ABRAMS	79D	PRL 43 481	G.S. Abrams <i>et al.</i>	(Mark II Collab.)
ATIYA	79	PRL 43 414	M.S. Atiya <i>et al.</i>	(COLU, ILL, FNAL)
BALTAY	78C	PRL 41 73	C. Baltay <i>et al.</i>	(COLU, BNL)
VUILLEMIN	78	PRL 41 1149	V. Vuillemin <i>et al.</i>	(Mark I Collab.)
GOLDHABER	77	PL 69B 503	G. Goldhaber <i>et al.</i>	(Mark I Collab.)
PERUZZI	77	PRL 39 1301	I. Peruzzi <i>et al.</i>	(Mark I Collab.)
PICCOLO	77	PL 70B 260	M. Piccolo <i>et al.</i>	(Mark I Collab.)
RAPIDIS	77	PRL 39 526	P.A. Rapidis <i>et al.</i>	(Mark I Collab.)
GOLDHABER	76	PRL 37 255	G. Goldhaber <i>et al.</i>	(Mark I Collab.)

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